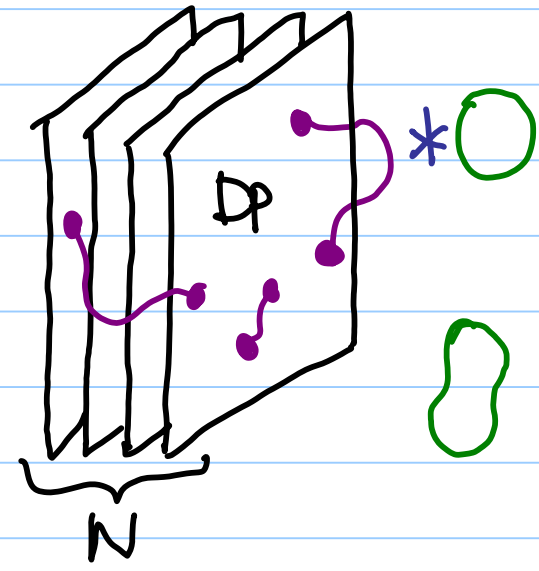


AdS/CFT: a holographic duality

Decoupling limit :-
 $El_s \rightarrow 0$ such that
 $(g_{YM}^2 N) E^{p-3}$ is fixed;
 $\equiv \lambda(E) = (2\pi)^{p-2} (g_s N) (l_s E)^{p-3}$.



Turns off interactions between open & closed strings

Taking the limit simplifies life for open & closed strings:

OPEN	CLOSED
Large- N gauge theory. Use 't Hooft scaling \Rightarrow	Lose asymptotically flat part of SUGRA spacetime; get metric
$\left(\lambda = g_{YM}^2 N ; \frac{1}{N} \right)$	$ds^2 = H^{1/2} (-dt^2 - K + d\vec{x}_{\parallel}^2) + H^{1/2} \left(\frac{dr^2}{k} + r^2 d\Omega_{p-1}^2 \right)$
Dimensionless 't Hooft coupling is finite; $\lambda E^{p-3} = (2\pi)^{p-2} (g_s N) (l_s E)^{p-3}$	with $H = c_p g_s N \left(\frac{l_s}{r} \right)^{7-p}$; $k = \left[- \left(\frac{r_H}{r} \right)^{7-p} \right]$
SU(N) SYM theory	Double perturbation series in $(\alpha' \mathcal{R}, g_s e^{\Phi})$

IMSY: phase diagrams

(2)

- For D3, geometry of BPS stack of N D-branes is

$$ds^2 = + \frac{r^2}{R^2} (-dt^2 + d\vec{x}_3^2) + R^2 \left(\frac{dr^2}{r^2} + d\Omega_{S^5}^2 \right) \quad R^4 = 4\pi g_s l_s^4$$

{Poincaré patch of AdS_5 } \times $\{S^5\}$.

AdS/CFT: same system, exactly equivalent degrees of freedom.
DUALITY

- Different situation ensues for D_p , $p \neq 3$. First, SYM theory either super-renormalizable or non-renormalizable.
 For $p > 3$, string theory unambiguously completes theory in UV
 For $p < 3$, theory strongly coupled in IR (rather than UV)

- SYM not valid $\forall E = U \equiv r/l_s^2$;

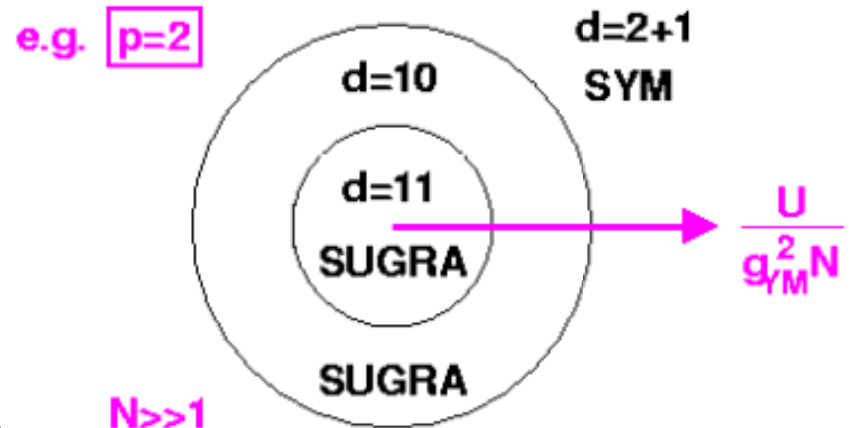
SUGRA complementary.

$e\Phi \rightarrow$ large near $r=0$

$\alpha' R \rightarrow$ large as $r \rightarrow \infty$

(monotonicity ex decoupling limit)

Resolution of null singularity 😊



Gauge Theory Duals of Gravity Phenomena

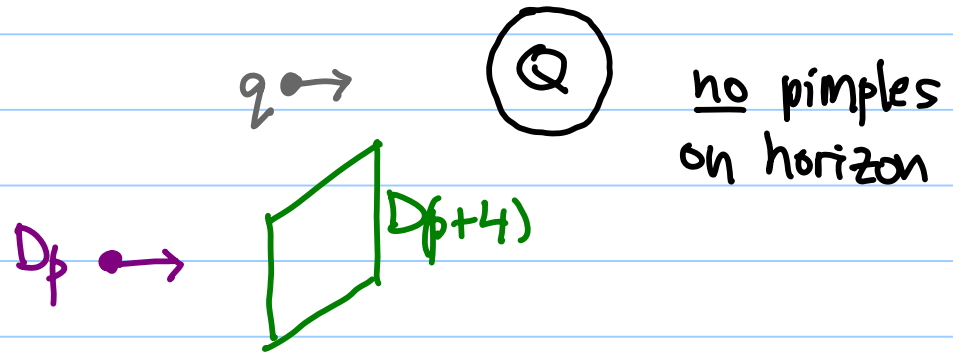
③

- Phenomena where gravity & gauge theory inform each other.

e.g. ① No-hair/CMW duality

Take $D_p \rightarrow D(p+4)$:

- Find boldness for $p=0,1$ (power-law; log) but localization possible for $p \geq 2$.



e.g. ② Gregory-Laflamme transition of black brane wrapped on S^1 .

- Numerical relativity informs (but not yet settles) endpoint debate.
- See phase transition via eigenvalue distribution in large- N field theory in $d=H+1$ relevant to D-branes.

e.g. ③ - Many large- N field theories have (!) a Hagedorn transition even at zero coupling!

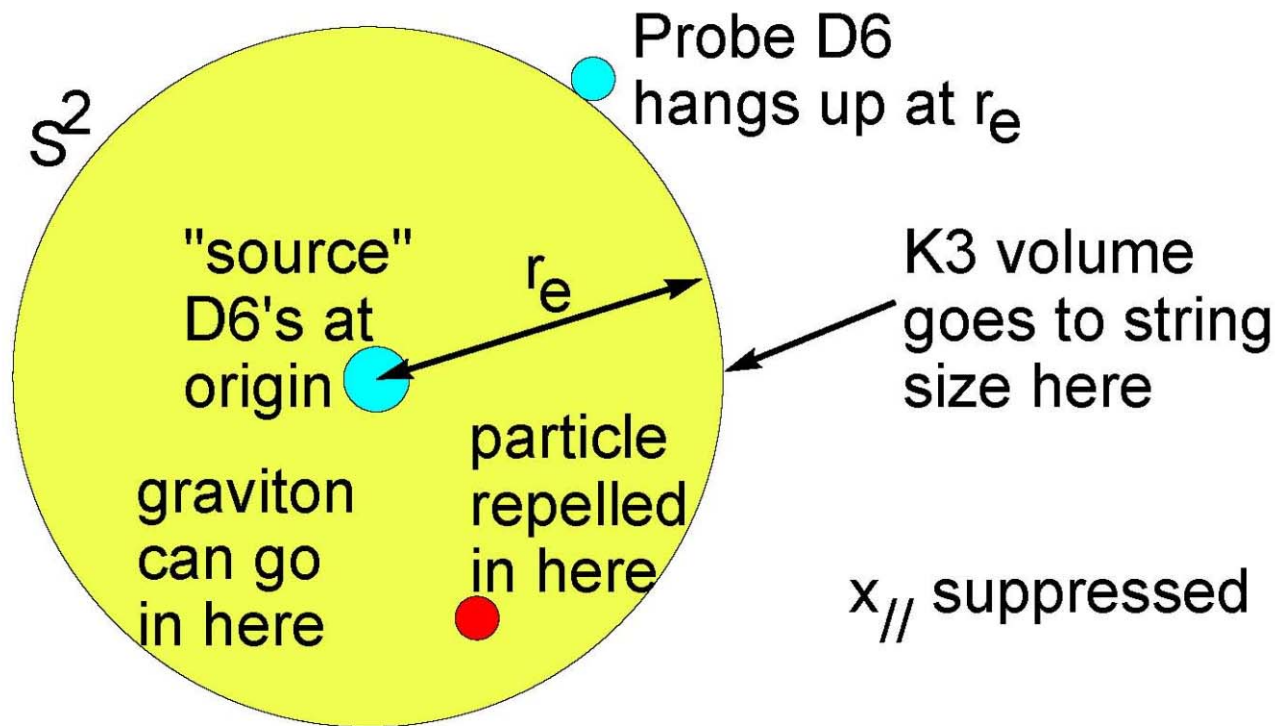
e.g. ④ - In-principle resolution of BH information problem.

- Maldacena proposal: eternal BH = $CFT_1 \otimes CFT_2$.

e.g. trying for pure $\mathcal{N}=2$ duality

(4)

- Various people tried but failed to take decoupling limit for brane systems with 8 supercharges ($\mathcal{N}=2$ in $d=4$) but no matter/hypermultiplets. Not all SUGRA geometries have weakly coupled dual in region where $\alpha' R \gtrsim 1$, $g_{\text{eff}} \gtrsim 1$!
- Two dual realizations: $D(p+4)$ on $K3$; fractional Dp .
- Enhancement mechanism: resolution of timelike naked singularity ("repulson")



- People also turned to deforming AdS/CFT (lots of great work!)

Spacetime Singularity Resolution

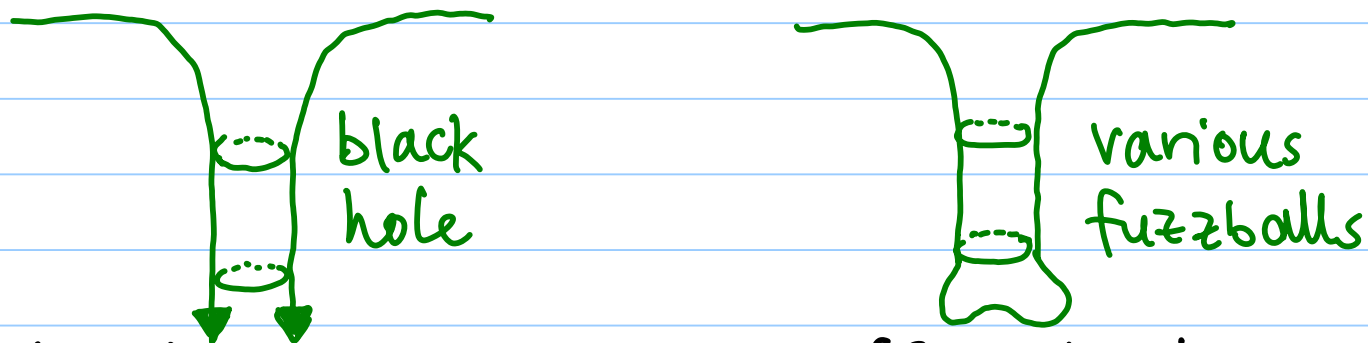
⑤

- Two ways classical spacetimes break down: curvature singularities and regions of bad chronology (i.e. with CTCs). String theory resolves these in one of two ways:
 - rules them out ab initio: some spacetimes do not have a quantum continuation - figments of classical imagination!;
 - resolves them by fuzzing out pointlike singularity or preventing even strongly curved regions altogether.
- Classical GR singularity theorems in form (note that SUGRA matter fields do not violate WEC) but only classically, and do not apply (as originally derived) with brane action sources $S_{Dp} = -T_p \int d^{D-p}x e^{-\Phi} \sqrt{-\det(2\pi F_{\alpha\beta} + P(g_{\alpha\beta} + B_{\alpha\beta}))} - T_p \int_{\mathcal{M}} e^{2\pi F + B} \wedge \bigoplus_n C_n$.
- Are sources needed, or are fluxes enough - for AF brane spacetimes?
- String theory problems always have precisely specified, unique field content (at least, in $d=10$!) \Rightarrow "S-branes" are for us string theory S_{Dp} -branes which are understood in CFT.

Mathur's fuzzball program

⑥

- Central proposal: SBH arises because of coarse-graining over microstates — and microstates are geometries with no horizon or singularity or entropy!
- So far, limited class of examples, but intriguing to most string theorists. Doubts regarding generalizability: potential stumbling block is insistence that microstates are valid SUGRA spacetimes. * Toronto group: pert. non-ext.
- All examples so far BPS. Horizon at end of ∞ throat; fuzzballs are effectively capped throats [2-charges; 3.]



- Infalling observer sees different physics than for BH but several aspects (e.g. delay time) work out great.

Unstable D-branes

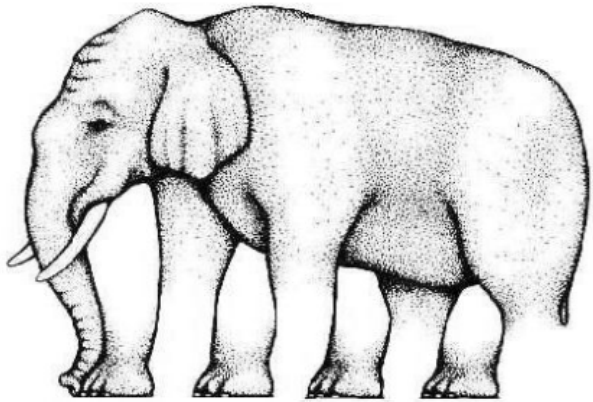
⑦

- Obvious reason why finding these spacetimes is both important -and hard: (Several bogus claims/papers in this area exist...)
 - time dependence
 - brane has codimension & \perp directions have r, Ω \Rightarrow at best, trying for solution of PDE's in (t, r)
- Spacetimes with right asymptotics seem to be (generically) singular. Looked at possible resolution via source action, but still singular, apparently.
- Sen (et al.) have CFT description (t -dependent) for decay of unstable D_p -branes. Strong indications arising that open string [field] theory has necessary degrees of freedom to describe closed string physics - collective modes...
- If want string theory connection for your "S-brane" spacetime, need
 - precise string theoretic massless fields
 - quantum continuation of spacetime

Take-home messages; & Many Frontiers

②

- Solid, exciting accounting for BPS BH microstates, beyond leading order. Exact agreement of entropy and greybody factors for near-BPS $d=5,4$ cases. Qualitative agreement even for far-from-BPS cases - as in Guijosa's work & presentation here. New calculations for BPS black rings also exciting successes.
- Gravity/gauge correspondences = explicit realizations of holography. Closed string theory (SUGRA with corrections) and SYM QFT. Usefully inform each other. Other stringy examples abound, e.g. BPS black hole entropy \leftrightarrow topological string theory \leftrightarrow matrix models



\leftarrow Nonperturbative
string theory

String theorists \rightarrow
(us 😊)



▶ Let's seek more joint triumphs ♥♥