

As in classical mechanics, in FM

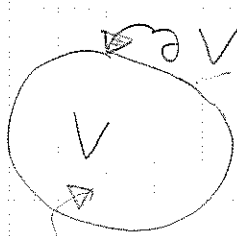
" $\vec{F} = m\vec{a}$ " can be cast as a consequence of momentum (non-) conservation

(i.e. $\vec{p} = m\vec{v}$, $\dot{\vec{p}} = m\vec{a} = 0$ - conserved \vec{p} if no reason for it to change, i.e. a force acts)

In FM, there are, in the nonrelativistic limit, several important conservation laws - mass, momentum, energy, angular momentum.

(mass alone is not ^{necessarily} conserved in relativistic fluids)

mass conservation



- fixed volume in fluid w/ mass density $\rho(\vec{x}, t)$ & velocity $\vec{v}(\vec{x}, t)$

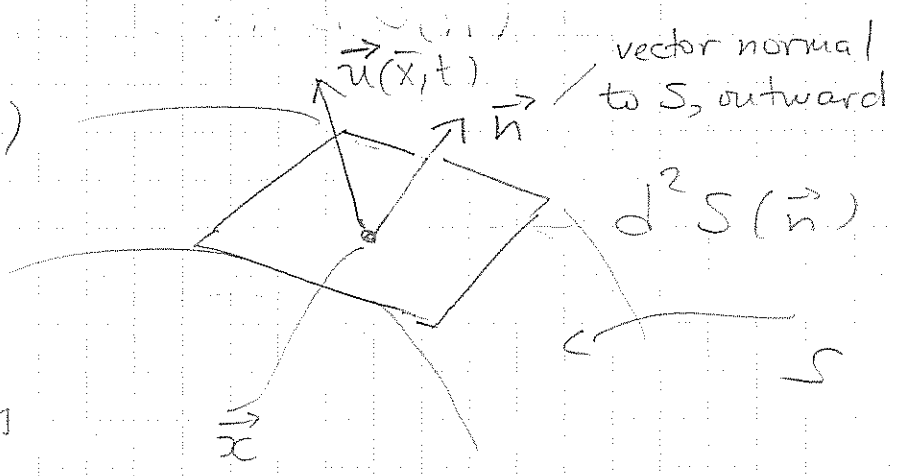
total mass inside V can change only due to fluid coming into V or going out of V through its boundary ∂V

(replace "mass" by "charge" \Rightarrow get charge conservation in E & M)

total mass inside is $\int \rho(\vec{x}, t) d^3x$

now $\rho(\vec{x}, t) \vec{u}(\vec{x}, t) \cdot \vec{n} d^2S(\vec{n})$

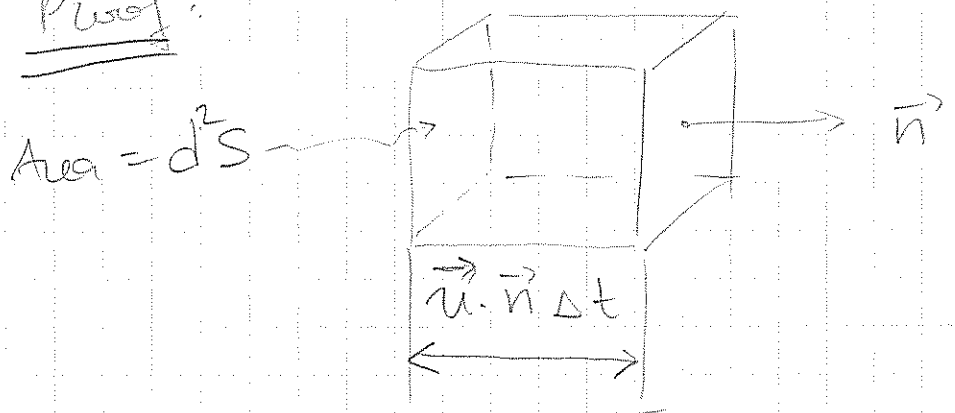
is the mass which flows through $d^2S(\vec{n})$ in unit time:



$\vec{u}(\vec{x}, t) \cdot \vec{n} =$
= component of velocity \perp to $d^2S(\vec{n})$

$\rho \vec{u} \cdot \vec{n} =$ (mass going through $d^2S(\vec{n})$ in unit time)
= flux of mass:

Proof:



all particles inside $d^2S(\vec{n})$ in time Δt

So total mass inside  is $\rho \vec{u} \cdot \vec{n} \cdot \Delta t$

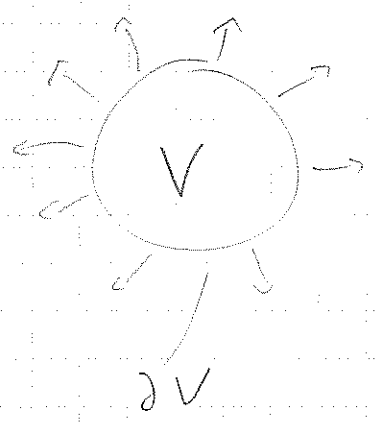
and in Δt it will go through $d^2S(\vec{n}) \Rightarrow$

\Rightarrow flux of mass is $\rho \vec{u} \cdot \vec{n}$

□

So now we have learned that,
given an arbitrary closed volume V
(fixed) of fluid:

$$\frac{d}{dt} \int_V \rho(\vec{x}, t) d^3x = - \oint_{\partial V} \rho(\vec{x}, t) \vec{u}(\vec{x}, t) \cdot \vec{n} d^2S(\vec{n})$$



mass in V changes
since fluid
leaves V (as in pic.)

hence "-" sign:
 $\vec{u} \cdot \vec{n} > 0 \iff$ flux leaves
 $\vec{u} \cdot \vec{n} < 0 \iff$ flux comes in

now some math: $\vec{n} d^2S(\vec{n}) \equiv d^2\vec{S}$

$$\oint_{\partial V} \vec{A} \cdot d^2\vec{S} = \int_V \vec{\nabla} \cdot \vec{A} dV \quad (\text{where } dV \equiv d^3x)$$

(Proof by integration by parts: $\text{rhs} = \int_a^b \frac{d}{dx} f(x) dx \iff \int dV$
 $\text{lhs} = f(b) - f(a) \iff \oint_{\partial V} f dS$)

$$\vec{\nabla} = \left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right) \equiv \left\{ \frac{\partial}{\partial x^i} \right\}_{i=1,2,3} = \left\{ \nabla^i \right\}$$

In the future it will be very useful to write using indices

$$\vec{n} d^2 S(\vec{n}) = d\vec{S} = \underbrace{\{d^2 S^i, i=1,2,3\}}$$

i-th comp. surface element
(-orient. of)

$$\oint_{S=\partial V} \vec{A} \cdot d\vec{S} = \oint_{S=\partial V} A^i d^2 \sigma^i = \int_V \nabla_i A^i dV$$

Σ over *i* understood

it may happen that *A* carries another index -

$$\text{then } \oint_{S=\partial V} A^i_j d^2 \sigma^i = \int_V \nabla^i A^i_j dV$$

index *j*
"goes for the ride"

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$$\oint_{\partial V=S} p \vec{u} d^2 \vec{S} = \oint_{\partial V=S} p u^i d^2 S^i = \int_V \nabla^i (p u^i) dV$$

$\underbrace{p u^i}_{\equiv A^i}$

hence

$$\frac{d}{dt} \int p(\vec{x}, t) d^3x = - \oint p(\vec{x}, t) \vec{u}(\vec{x}, t) \cdot d^2\vec{S}$$

Vis fixed $\int_V \left(\frac{d}{dt} \text{ goes here!} \right) \delta V = S$

$$\int_V \frac{d}{dt} (p(\vec{x}, t)) d^3x = - \int_V \vec{\nabla} \cdot (p(\vec{x}, t) \vec{u}(\vec{x}, t)) dV$$

$$= - \int_V \nabla^i (p(\vec{x}, t) u^i(\vec{x}, t)) dV$$

$$\int_V dV \left[\frac{d}{dt} (p(\vec{x}, t)) + \nabla^i (p(\vec{x}, t) u^i(\vec{x}, t)) \right] = 0$$

V is arbitrary : $\int_V dV f(\vec{x}) = 0, \forall V \Rightarrow f = 0$

hence

$$\frac{d}{dt} \rho + \nabla^i (\rho u^i) = 0$$

$\rho u^i \leftrightarrow$ flux of fluid

continuity equation \equiv mass conservation

(E&M: $\rho \leftrightarrow$ charge density
 $\rho \vec{u} \leftrightarrow$ current density)

All conservation laws have this kind of form (differential expression)

an incompressible fluid is when

$\rho = \text{const}$ (doesn't change w/ \vec{x} , t , that is)

continuity eqn. sez:

$$\underbrace{\frac{d}{dt} \rho + \nabla^i \rho u^i}_0 + \underbrace{\rho \nabla^i u^i}_0 = 0$$

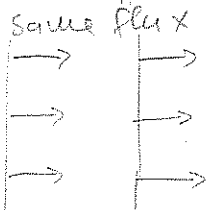
ρ - indep of \vec{x}, t

since $\rho \neq 0$

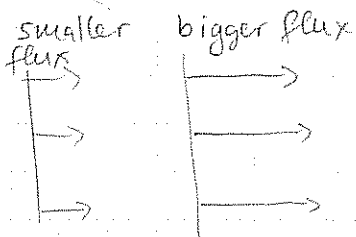
$$\nabla \cdot \vec{u}(\vec{x}, t) = 0$$

- for an incompressible fluid
- as we'll see, many equations simplify a lot for such fluids (e.g. water \approx incompressible)
- (- air - not so much -)

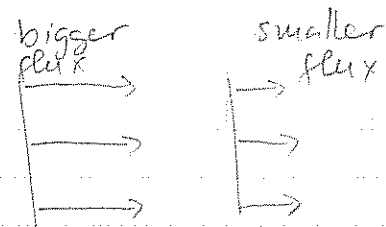
Ex: in 1-d flow $u = u(x, t)$



vs.



vs.



$$\partial_x u_x = 0$$

no mass pile-up
density = const

$$\partial_x u_x > 0$$

need source of
mass

$$\partial_x u_x < 0$$

need sink of
mass