

low R

So let's now study the Stokes problem

- a body moving w/ $\vec{U} = \text{const}$ thru a viscous incompressible fluid
- take a sphere now (explicit soltn exists)
- low R \equiv "Stokes' equations" - neglect inertia & stationary

Stokes eqns

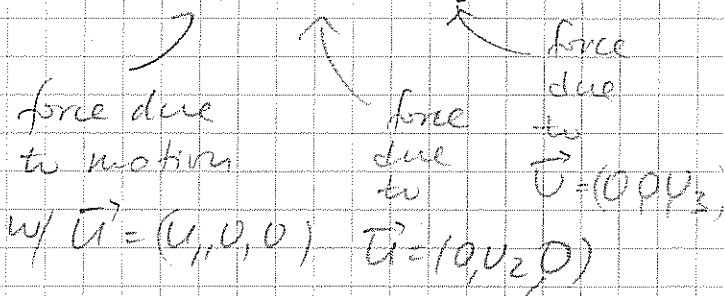
$$\begin{cases} \mu^2 \Delta \vec{u} = \vec{\nabla} p \\ \vec{\nabla} \cdot \vec{u} = 0 \end{cases}$$

assume no body force for now

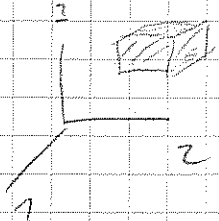
W.B. - Stokes eqns are linear

- any boundary motion or body force affects soltns linearly

- take a solid body moving w/ $\vec{U} = (U_1, U_2, U_3)$ then force on it is $\vec{F} = \vec{F}_1 + \vec{F}_2 + \vec{F}_3$



- take it a cube oriented along



by symmetry same α

then $\vec{F}_1 = (\alpha U_1, 0, 0)$, $\vec{F}_2 = (0, \alpha U_2, 0)$, $\vec{F}_3 = (0, 0, \alpha U_3)$

$$\Rightarrow \vec{F} = \vec{F}_1 + \vec{F}_2 + \vec{F}_3 = \alpha \vec{U}$$

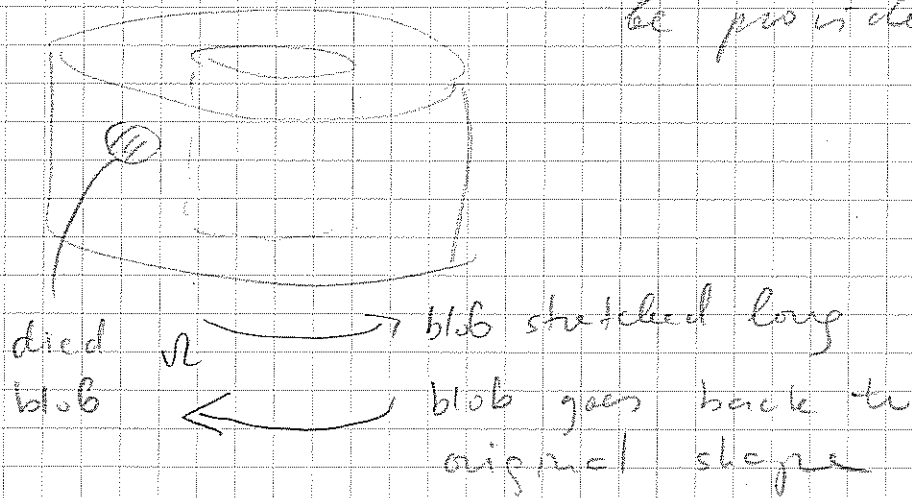
Moral: at low R , the drag force on a cube is independent on its orientation.

(at higher R , not true anymore)

- also because of linearity, if sign of velocity @ boundaries is reversed - the sign of solution is reversed

- "reversibility" of very viscous (low- R) flows

- great videos (see links to be provided)



- Finally, as opposed to $N-S$. In general, solutions of Stokes' eqns: (a) exist (b) are unique when velocity field is specified on ∂V of some volume V , if $R < \pi\sqrt{3}$ (!)

there are various ways to write Stokes' eqns:

take $\vec{\nabla} \cdot (-) \Rightarrow \Delta p = 0$

take $\vec{\nabla} \times (-) \Rightarrow \nabla^2 \vec{\omega} = 0$

$(\vec{\omega} = \vec{\nabla} \times \vec{v})$

\Rightarrow pressure field & vorticity field are both harmonic (their Laplacian vanishes)

INTERLUDE, NEEDED:

Especially simple are "2d" incompressible flows!

Ex. 1: a "genuine" 2d flow $\vec{u} = (u_x(x,y,t), u_y(x,y,t), 0)$

$\vec{\nabla} \cdot \vec{u} = 0 \Rightarrow \partial_x u_x + \partial_y u_y = 0$

* Clearly, if $u_x(x,y,t) = \partial_y \psi(x,y,t)$

‡ $u_y(x,y,t) = -\partial_x \psi(x,y,t)$

continuity eqn is satisfied, as

$\partial_x u_x + \partial_y u_y = \partial_x \partial_y \psi - \partial_y \partial_x \psi = 0$

* Conversely, every solution of $\partial_x u_x + \partial_y u_y = 0$ can be locally written as

$u_x = \partial_y \psi$

$u_y = -\partial_x \psi$

(Poincaré lemma)

\Leftarrow (there can be global obstructions)

Given u_x, u_y, ψ can be found as follows \longrightarrow

$$\left. \begin{aligned} \frac{\partial}{\partial y} \psi(x, y) &= u_x(x, y) \\ \frac{\partial}{\partial x} \psi(x, y) &= -u_y(x, y) \end{aligned} \right\} \begin{array}{l} \text{(set of 2)} \\ \text{a linear diff. eqns} \end{array}$$

From 1st $\psi(x, y) = \int_{y_0}^y u_x(x, y') dy' + (\text{y-indep.})$

2nd $\psi(x, y) = -\int_{x_0}^x u_y(x', y) dx' + (\text{x-indep.})$

& these can be combined into

$$\psi(x, y) - \psi(x_0, y_0) = \int_{(x_0, y_0)}^{(x, y)} (u_x dy - u_y dx) \quad \text{contour}$$

Proof:

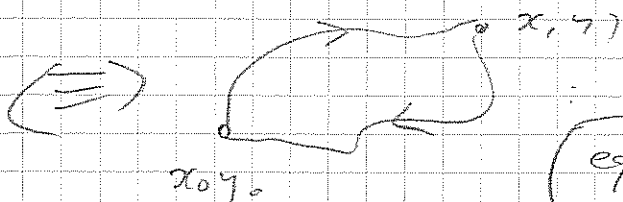
Clearly taking $\frac{\partial \psi}{\partial x} = \frac{1}{\Delta x} \left(\int_{(x_0, y_0)}^{(x+\Delta x, y)} - \int_{(x_0, y_0)}^{(x, y)} (u_x dy - u_y dx) \right) \Big|_{\Delta x \rightarrow 0}$

$$= \frac{1}{\Delta x} \int_{(x, y)}^{(x+\Delta x, y)} (-u_y dx) = -\frac{\Delta x u_y}{\Delta x} = -u_y$$

(same goes for $\frac{\partial \psi}{\partial y}$)

↑ since y does not change

$\psi(x, y)$ is independent of the contour taken \Leftrightarrow



integral over closed contour $\equiv 0$.

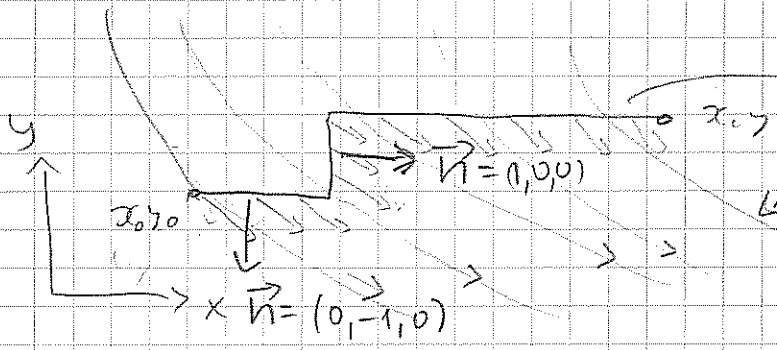
(equivalent to $\nabla \cdot \vec{u} = 0$ inside)
(so there must be only fluid inside)

the reason this introduction of ψ is convenient is that

- * it solves 2d $\nabla \cdot \vec{u} = 0$
- * we get further eqn depending on one scalar ψ , not a vector

look @ ψ 's def. nice more to discrete meaning:

$$\psi(x, y) - \psi(x_0, y_0) = \int_{(x_0, y_0)}^{(x, y)} (u_x dy + (u_y - 1) dx)$$



take this contour
 an ex. of a velocity field
 (\vec{u} tangent to lines @ every point)

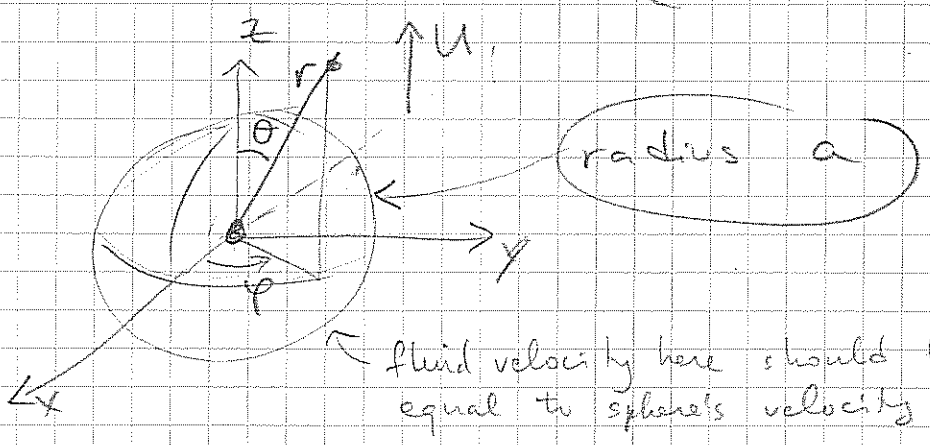
normal vector to "surface" \Rightarrow above $\int = \underbrace{\vec{u} \cdot \vec{n}}_{\text{exactly the flux}} dl$

so $\psi(x, y) - \psi(x_0, y_0) \equiv$
 \equiv (flux of volume of incompressible fluid thru
 \forall line joining x, y w/ x_0, y_0)

$\psi(x, y) \equiv$ "stream fun"

$\Psi(x, y)$ greatly used in 2d
ideal incompressible flows

? about our Stokes problem?



- * fix coordinate system w/ spherical particle
 - * so condition on velocity of fluid @ surface is that it equals U
 - * we'll take $\vec{U} \parallel \hat{z}$
- at instantaneous position!

What to expect?

- linearity sez $\vec{F} = \text{const } \vec{U}$
- drag = 0 if $\mu = 0$
- so $\vec{F} = \text{const } \mu \vec{U}$
- should depend on a as well -

$$[M] = [\rho v] = \frac{\text{kg}}{\text{m}^3} \frac{\text{m}^2}{\text{s}}$$

$$[U] = \frac{\text{m}}{\text{s}}$$

$$[\mu U] = \frac{\text{kg}}{\text{s}^2} \rightarrow \text{to get force, need } \mu$$

$$[a \mu \vec{U}] = \frac{\text{kg} \cdot \text{m}}{\text{s}^2} \quad \checkmark$$

so $\vec{F} = (\text{const}) \mu a \vec{U}$; we are after dimensionless #.

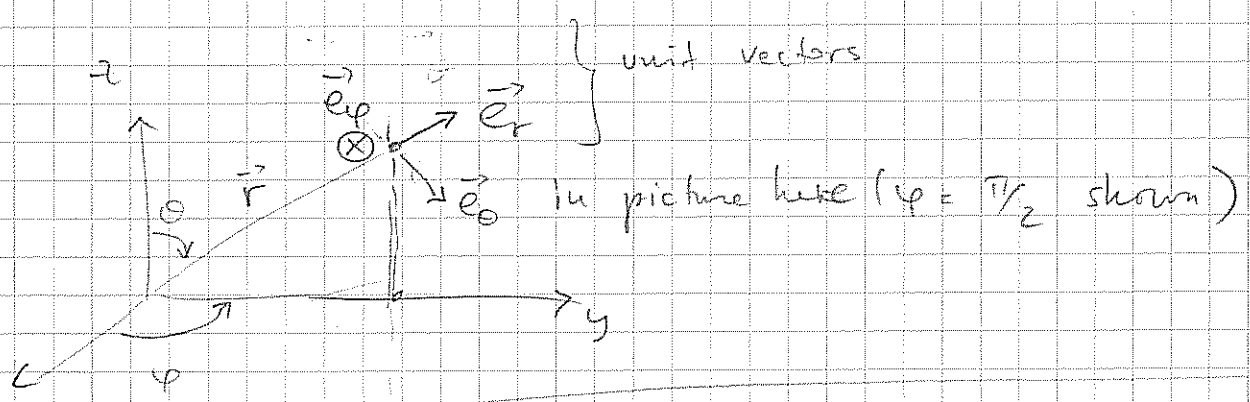
\uparrow dimensionless

Answer is $\vec{F} = \underbrace{6\pi\mu a}_{\downarrow} \vec{U}$

- not just for the fun of solving diff eqns!
- low-R drag force useful to measuring μ for various substances! (so 6π matters, really!)

So let's get it!

- let's use $\vec{x} = (r, \theta, \varphi)$ coordinates
- b.c. are φ independent, so expect soltn to only depend on r & θ . \rightarrow "2 dimensionality"
- $\rightarrow \vec{u} = \vec{e}_r u_r + \vec{e}_\theta u_\theta, u_\varphi = 0$



Gradient in spherical coordinates: $\vec{\nabla} F = \vec{e}_r \frac{\partial F}{\partial r} + \vec{e}_\theta \frac{1}{r} \frac{\partial F}{\partial \theta} + \vec{e}_\varphi \frac{1}{r \sin \theta} \frac{\partial F}{\partial \varphi}$

Divergence in spherical coordinates: $\vec{\nabla} \cdot \vec{F} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 F_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta F_\theta) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \varphi} (\sin \theta F_\varphi)$

Laplacian - u: $\Delta F = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \frac{\partial F}{\partial r}) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta \frac{\partial F}{\partial \theta}) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 F}{\partial \varphi^2}$

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curl: $\nabla \times \vec{F} = \vec{e}_r \left(\frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta F_\varphi) - \frac{\partial F_\theta}{\partial \varphi} \right)$
 $+ \vec{e}_\theta \left(\frac{1}{\sin \theta} \frac{\partial F_r}{\partial \varphi} - \frac{\partial}{\partial r} (r F_\varphi) \right) \frac{1}{r}$
 $+ \vec{e}_\varphi \left(\frac{\partial}{\partial r} (r F_\theta) - \frac{\partial F_r}{\partial \theta} \right) \frac{1}{r}$

will use as we need

→ 1st: solve $\nabla \cdot \vec{u} = 0$ by symmetry $u_\varphi = 0, u_r, u_\theta(r, \theta)$ only.

$\vec{u} = \vec{e}_r u_r + \vec{e}_\theta u_\theta \Rightarrow$ so

$\nabla \cdot \vec{u} = 0 \Rightarrow 0 = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 u_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta u_\theta)$
 \Downarrow x (by $r^2 \neq \sin \theta$)

$0 = \frac{\partial}{\partial r} (r^2 \sin \theta u_r) + \frac{\partial}{\partial \theta} (r \sin \theta u_\theta)$

(recall $0 = \frac{\partial}{\partial x} u_x + \frac{\partial}{\partial y} u_y \Rightarrow u_x = \partial_y \psi$
 $u_y = -\partial_x \psi$)

so $r^2 \sin \theta u_r = \partial_\theta \psi(r, \theta)$

$r \sin \theta u_\theta = -\partial_r \psi(r, \theta)$

hence taking $u_r = \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \psi$

$u_\theta = -\frac{1}{r \sin \theta} \frac{\partial}{\partial r} \psi$ solves axisymm. $\nabla \cdot \vec{u} = 0$

Stokes' eqns $\rightarrow \mu \Delta \vec{u} = \vec{\nabla} p$ imply $(\vec{\nabla} \times \dots)$

$$\Delta(\vec{\nabla} \times \vec{u}) = 0$$

$\vec{\nabla} \times \vec{u}$, for \vec{u} given by ψ , is = ?

$$\vec{\omega} = \vec{\nabla} \times \vec{u} = (\text{for } \vec{u} = u_r, u_\theta, \varphi\text{-indep.})$$

$$= \vec{e}_\varphi \frac{1}{r} \left(\frac{\partial}{\partial r} (r u_\theta) - \frac{\partial}{\partial \theta} (u_r) \right)$$

- vorticity in this axisymmetric flow is $\sim \vec{e}_\varphi$ only

$$\text{so } \omega_\varphi = \frac{1}{r} \frac{\partial}{\partial r} (r u_\theta) - \frac{1}{r} \frac{\partial}{\partial \theta} u_r = \text{sub } u_r, u_\theta \text{ from } \psi$$

$$= \frac{1}{r} \frac{\partial}{\partial r} \left(\frac{-1}{\sin \theta} \frac{\partial \psi}{\partial r} \right) - \frac{1}{r} \frac{\partial}{\partial \theta} \left(\frac{1}{r^2 \sin \theta} \frac{\partial \psi}{\partial \theta} \right)$$

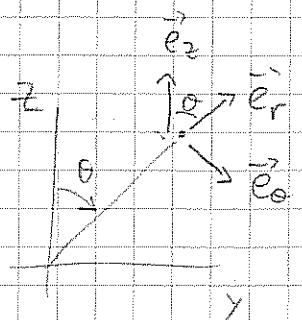
$$= -\frac{1}{r \sin \theta} \frac{\partial^2 \psi}{\partial r^2} - \frac{1}{r^3} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin \theta} \frac{\partial \psi}{\partial \theta} \right)$$

and we're told $\Delta \omega_\varphi = 0$ \Leftarrow (Stokes' eqn.)

+ b.c. @ $r=a$ are $\vec{u}|_{r=a} = \vec{e}_z U$

$$\vec{u}|_{r=a} \cdot \vec{e}_r \equiv U_r|_{r=a} = U \vec{e}_z \cdot \vec{e}_r = U \cos \theta$$

$$\vec{u}|_{r=a} \cdot \vec{e}_\theta \equiv U_\theta|_{r=a} = U \vec{e}_z \cdot \vec{e}_\theta = U \cos(\theta + \frac{\pi}{2}) = -U \sin \theta$$



$$\text{so } U_r|_{r=a} = U \cos \theta = \frac{1}{a^2 \sin \theta} \frac{\partial \psi}{\partial \theta} \Big|_{r=a}$$

$$U_\theta|_{r=a} = -U \sin \theta = -\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r} \Big|_{r=a}$$

So we have

$$\begin{cases} \frac{\partial \psi}{\partial \theta} \Big|_{r=a} = \frac{1}{2} U a^2 \sin 2\theta \\ \frac{\partial \psi}{\partial r} \Big|_{r=a} = U a \sin^2 \theta \end{cases}$$

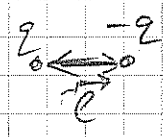
+ Now we must solve $\Delta \psi = 0 \iff (\Delta \vec{w} = 0)$
 subject to above b.c. for ψ @ $r=a$.

+ away from sphere, as $r \rightarrow \infty$, it must be that
 $\vec{u} / \text{fluid} \rightarrow 0$, so we're looking for a "Jyng" soln @ ∞

+ recall simple facts: "Gauss' law" $\Delta \psi = 4\pi \rho$
 \uparrow \uparrow
 el. static potential charge density.

• $\rho = Q \delta^{(3)}(\vec{x})$
 \uparrow
 point charge @ $\vec{x}=0$

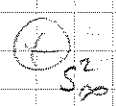
field is $\psi \sim \frac{1}{r}$




$\vec{d} = e\vec{l}$ dipole

field is $\psi \sim \frac{\vec{d} \cdot \vec{r}}{r^3}$

← these solve Laplace = 0 away from $\vec{r}=0$

charge $\psi \sim \frac{1}{r}$ so $\int \psi \sim Q$


dipole $\psi \sim \frac{1}{r^2}$ so $\int = 0$ (no charge inside)


So/lt of $\Delta(\text{anything}) = 0 \Leftrightarrow$ expansion in multipoles

- monopole $\sim 1/r$

- dipole $\sim \frac{1}{r^2}$

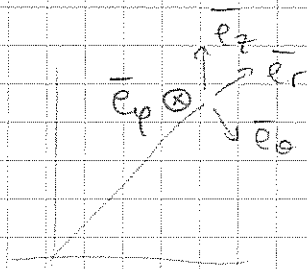
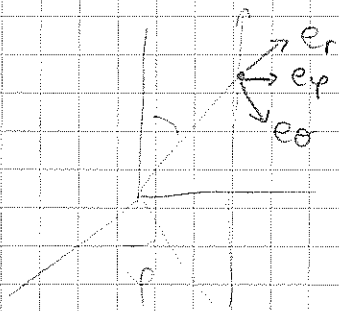
- quadrupole $\sim \frac{1}{r^3}$

Here $\Delta \vec{w} = 0 \Rightarrow \vec{w}$ is a vector

dipole term: $\vec{w} = \text{const} \frac{\vec{U} \times \vec{r}}{r^3}$ is leading (no monopole possible, as no scalar)

Of course, $\vec{U} \times \vec{r}$, w/ $\vec{U} \parallel \hat{z}$

can only have nonzero φ .



$\vec{e}_z \times \vec{e}_r = \vec{e}_\phi \sin \theta$ & $\vec{e}_z \times \vec{e}_\phi = -\vec{e}_r \sin \theta$
 $\vec{e}_z \times \vec{e}_\theta = -\vec{e}_\phi \cos \theta$

$\int_0 \vec{U} \times \vec{r} = U \vec{e}_z \times (r \vec{e}_r) = U \vec{e}_\phi \sin \theta r$

So $w_\varphi = \frac{U \sin \theta r}{r^3} \times \text{const.} \Leftarrow$ determined r by solving $\Delta \vec{w} = 0$ & taking dipole term Ansatz (consistent w/ $\vec{w} \rightarrow 0$ as $r \rightarrow \infty$)

on the other hand

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$$\omega_{\varphi} = -\frac{1}{r \sin \theta} \frac{\partial^2 \varphi}{\partial r^2} - \frac{1}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin \theta} \frac{\partial \varphi}{\partial \theta} \right) \quad \text{p. (126)}$$

by def of $\vec{\omega} = \nabla \times \vec{u}$ & soln of $\nabla \cdot \vec{u} = 0$.

So, equate the two:

$$-\frac{1}{r \sin \theta} \frac{\partial^2 \varphi}{\partial r^2} - \frac{1}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin \theta} \frac{\partial \varphi}{\partial \theta} \right) = C \frac{U \sin \theta}{r^2}$$

or

$$\frac{\partial^2 \varphi}{\partial r^2} + \frac{\sin \theta}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin \theta} \frac{\partial \varphi}{\partial \theta} \right) = -C \frac{U \sin^2 \theta}{r}$$

$$+ \left. \frac{\partial \varphi}{\partial r} \right|_{r=a} = U a \sin^2 \theta$$

$$\left. \frac{\partial \varphi}{\partial \theta} \right|_{r=a} = \frac{1}{2} U a^2 \sin 2\theta = \frac{1}{2} U a^2 \frac{\partial}{\partial \theta} (\sin^2 \theta)$$

storing

@ B.C. : take $\varphi \sim \sin^2 \theta$.

hence take

$$\varphi = f(r) U \sin^2 \theta \quad \text{as a try.}$$

b.c. says $f'(r) \rightarrow a'$ as $r \rightarrow a$

$$\dagger \quad f(r) \rightarrow \frac{a^2}{2} \quad \text{as } r \rightarrow a.$$

w/ $\Psi = f(r) U \sin^2 \theta$, $f(a) = \frac{a^2}{2}$
 $f'(a) = a$

eqn for f is

$$- f''(r) U + \sin^2 \theta \frac{f(r) U}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin \theta} \frac{\partial \sin^2 \theta}{\partial \theta} \right) = -C \frac{U \sin^2 \theta}{r}$$

$$\frac{\partial}{\partial \theta} \frac{2 \sin \theta \cos \theta}{\sin^2 \theta} = -2 \cot \theta$$

$$f''(r) U \sin^2 \theta - \frac{2f(r) U \sin^2 \theta}{r^2} = -C \frac{U \sin^2 \theta}{r}$$

so $f'' - \frac{2f}{r^2} = -\frac{C}{r}$

General solution:

$$f = \beta r^\alpha$$

$$\alpha(\alpha-1)\beta r^{\alpha-2} = 2\beta r^{\alpha-2} = -\frac{C}{r}$$

$$(\alpha(\alpha-1) - 2) \beta r^{\alpha-2} = -\frac{C}{r}$$

$\Rightarrow \alpha = 1, \beta = \frac{1}{2}$ solves inhomogeneous eqn.

$\nabla \alpha(\alpha-1) = 2$ solves homogeneous
 $\alpha = 2$
 $\alpha = -1$

Generally:

$$f = \frac{1}{2} C \frac{1}{r} + B r^2 + D \frac{1}{r}$$

Now as $r \rightarrow \infty$, $f/r^2 \rightarrow 0$ ($\frac{1}{r} \frac{\partial f}{\partial r} \rightarrow 0$)

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Velocity field following from $\psi = U f(r) \sin^2 \theta$ is

(DIY) $\vec{u} = \vec{U} \frac{1}{r} \frac{\partial f}{\partial r} + \vec{r} \frac{\vec{r} \cdot \vec{U}}{r^2} \left(\frac{2f}{r^2} - \frac{1}{r} \frac{\partial f}{\partial r} \right)$

($\vec{U} = U \vec{e}_z$)

hence $B=0$ & $f = \frac{1}{2} C r + \frac{D}{r}$

$$f' = \frac{1}{2} C - \frac{D}{r^2}$$

$$f(a) = \frac{a^2}{2} = \frac{C}{2} a + \frac{D}{a} \quad \Rightarrow \quad C = \frac{3}{2} a$$

$$f'(a) = a = \frac{C}{2} - \frac{D}{a^2} \quad \Rightarrow \quad D = -\frac{a^3}{4}$$

check: $\frac{3}{4} a^2 - \frac{1}{4} a^2 = \frac{a^2}{2}$

$$\frac{3}{4} a + \frac{a^3}{4 a^2} = a \quad \checkmark$$

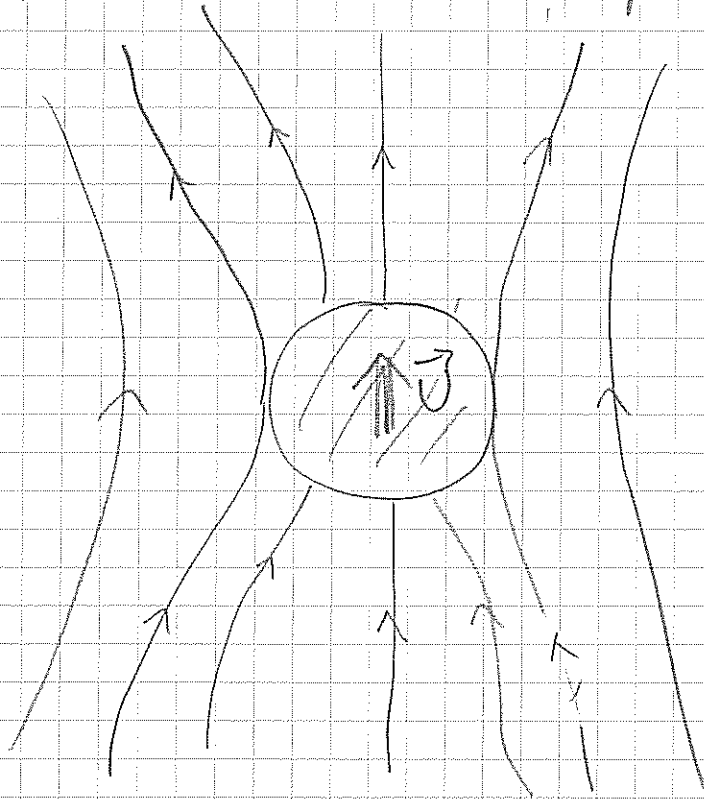
$$\therefore \psi = U r^2 \sin^2 \theta \left(\frac{3}{4} \frac{a}{r} - \frac{a^3}{4 r^3} \right)$$

$\psi(r, \theta)$ - stream function

recall $u_r = \frac{1}{r^2 \sin \theta} \partial_\theta \psi$

$u_\theta = -\frac{1}{r \sin \theta} \partial_r \psi$

Sketch - DIY - use Maple, Mathematica, whatever



= this is much more symmetric than I plotted
(axisymmetric, really)

= velocity falls off like $\frac{1}{r}$ as $r \rightarrow \infty$

$$\vec{u} \approx \vec{U} \frac{3a \sin^2 \theta \cos \theta}{4r}$$

so $\mu \Delta u \sim \frac{\mu a}{r^3} U$ far away

while $\rho (\vec{u} \cdot \nabla) \vec{u} \approx \rho \cdot \vec{U} \cdot \nabla \vec{u} \sim \rho \frac{1}{r^2} a U^2$
 ↑ relative to sphere velocity

depending on how small R becomes down at sufficient large r .

$$\frac{\text{inertia}}{\text{viscosity}} \sim \frac{\rho a U^2}{r^2} / \frac{\mu a}{r^3} \sim \frac{r U}{\nu} \sim \frac{r}{a} \left(\frac{Ua}{\nu} \right) \sim \left(\frac{r}{a} \right) R$$

- not so good for arrays!

Finally - the force!

Need \leftarrow normal to sphere

$$\sigma_{ij} \Big|_{r=a} n_j = \left(-p \delta_{ij} + \mu \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \right) \Big|_{r=a} n_j = (*)$$

DIY: use velocity field $\vec{u} = \vec{U} \frac{1}{r} f' + \vec{r} \frac{\vec{r} \cdot \vec{U}}{r^2} \left(\frac{2f}{r^2} - \frac{1}{r} f' \right)$
to show that

$$(*) = \left(-p n_i + \mu n_i \vec{U} \cdot \vec{n} \left(-\frac{f''}{r} + \frac{6f'}{r^2} - \frac{10f}{r^3} \right) + \mu U_i \left(\frac{f''}{r} - \frac{2f'}{r^2} + \frac{2f}{r^3} \right) \right) \Big|_{r=a} = (**)$$

to find $p \rightarrow$ need to use $\vec{\nabla} p = \mu \Delta \vec{u}$ or, just a pressure @ ∞

but smarter, since $\vec{\omega} = \vec{\nabla} \times \vec{u}$, note that

$$\vec{\nabla} \times \vec{\omega} = \vec{\nabla} \times (\vec{\nabla} \times \vec{u}) = -(\vec{\nabla} \cdot \vec{\nabla}) \vec{u} = -\Delta \vec{u}$$

$\Rightarrow \vec{\nabla} p = \mu \vec{\nabla} \times \vec{\omega}$; also recall $\Delta p = 0$, too:

but $\vec{\omega} = C \frac{\vec{U} \times \vec{r}}{r^3}$, hence $p = \mu C \frac{\vec{u} \cdot \vec{x}}{r^3} + p_0$

same coefft fixed by $\vec{\nabla} p = \mu \vec{\nabla} \times \vec{\omega}$.

→ then put p & f & C into $(*)$

to find

$$\sigma_{ij} n_j \Big|_{r=a} = -p_0 n_i - \frac{3\mu U_i}{2a}$$

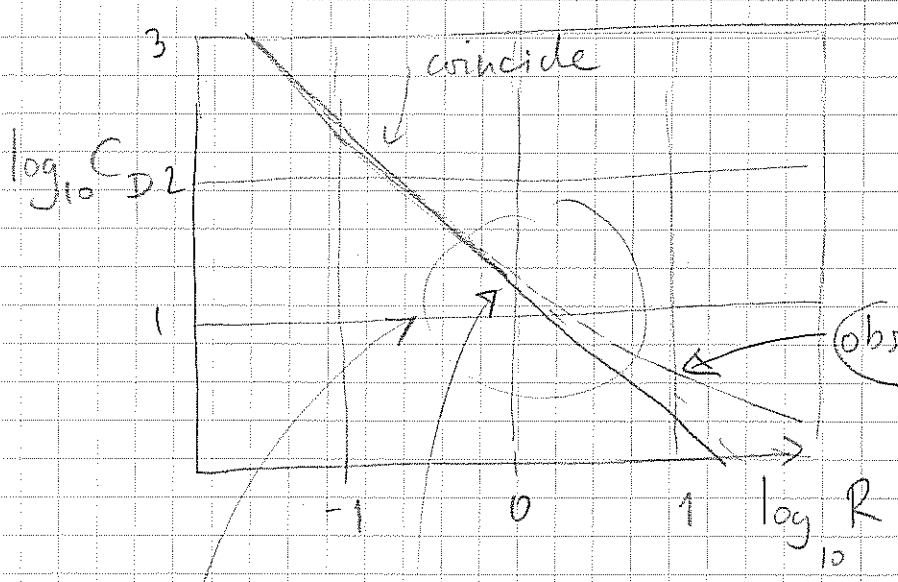
this does not contribute
after $a^2 \int dp d\theta \sin\theta n^i = 0$

but
since $U_i = U \delta_{i3}$

we have $\underbrace{a^2 \int dp d\theta \sin\theta}_{4\pi a^2} \frac{3\mu U_i}{2a} = 6\pi a \mu U_i$

So force on sphere is

$$\vec{F} = -\vec{U} 6\pi a \mu$$



$$C_D = \frac{|\vec{F}|}{\left(\frac{1}{2} \rho U^2\right) (\pi a^2)} = \frac{24}{R}$$

drag
x-section

$$R = \frac{2aU}{\nu}$$

Stokes' law
(an 2nd order improvement)

(2 is historical, here)