

Noble Lectures: Toronto, April 2010

Snowball Earth and Snowball Earths

Part I: Meet the Proterozoic

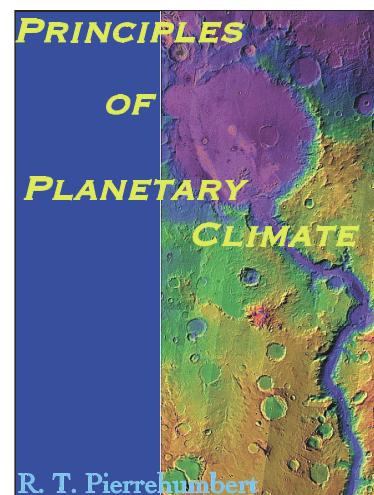
Raymond T. Pierrehumbert
The University of Chicago

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Neoproterozoic Climate review article

Ann. Rev. Earth Plan. Sci. (preprint July 2010)

(and ...)



JUNE, 2010

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The Snowball State

- Consider a planet that is all or mostly ocean
- If the planet freezes over completely, the high albedo can keep it frozen.
- This is the Snowball state. It takes a great deal of warming to break out.
- There is a multitude of geological evidence that this happened on Earth twice around 650 million years ago (the Neoproterozoic), and once around 2.5 billion years ago (the beginning of the Paleoproterozoic)
- Evidence for glaciers near equator is unequivocal. Evidence for totally frozen ocean is more indirect
- This is a habitability crisis for planets in general.

My collaborators on Snowball work

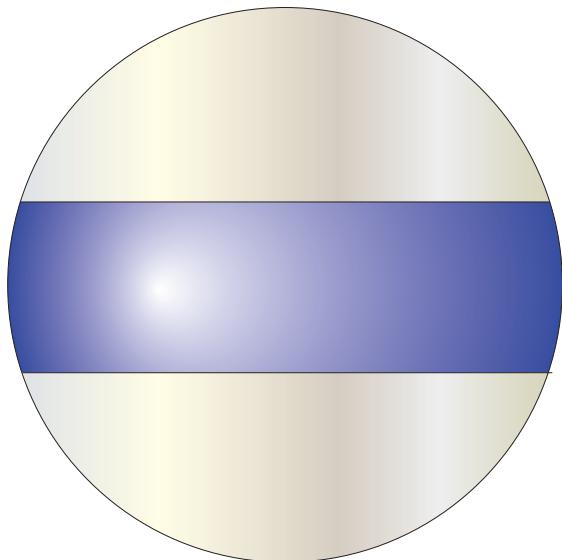
- Chris Poulsen
- Jason Goodman
- Yannick Donnadieu, Guillaume Le Hir, Yves Godderis
- Jonathan Mitchell, Xavier Levine, Rodrigo Caballero
- Dorian Abbot, Aiko Voight
- Paul Hoffman and Dan Schrag

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The main questions

- Under what circumstances does a planet enter a global glaciation?
- Under what circumstances can you get out?
- What is the climate like while you're there?
- What is the climate like after you get out?

Additional questions: Tropical waterbelt worlds?



- Stable states with low-latitude ice margins?
- Are waterbelt states cold enough to have land ice sheets?

Geological evidence for a hard snowball

- Glacial deposits (diamictites, dropstones) at low latitudes
- Cap carbonates
- Reappearance of Banded Iron Formations (global ocean anoxia marker)
- Isotopic evidence of very high CO_2 after deglaciation
- Iridium anomaly (long accumulation on global ice cover)

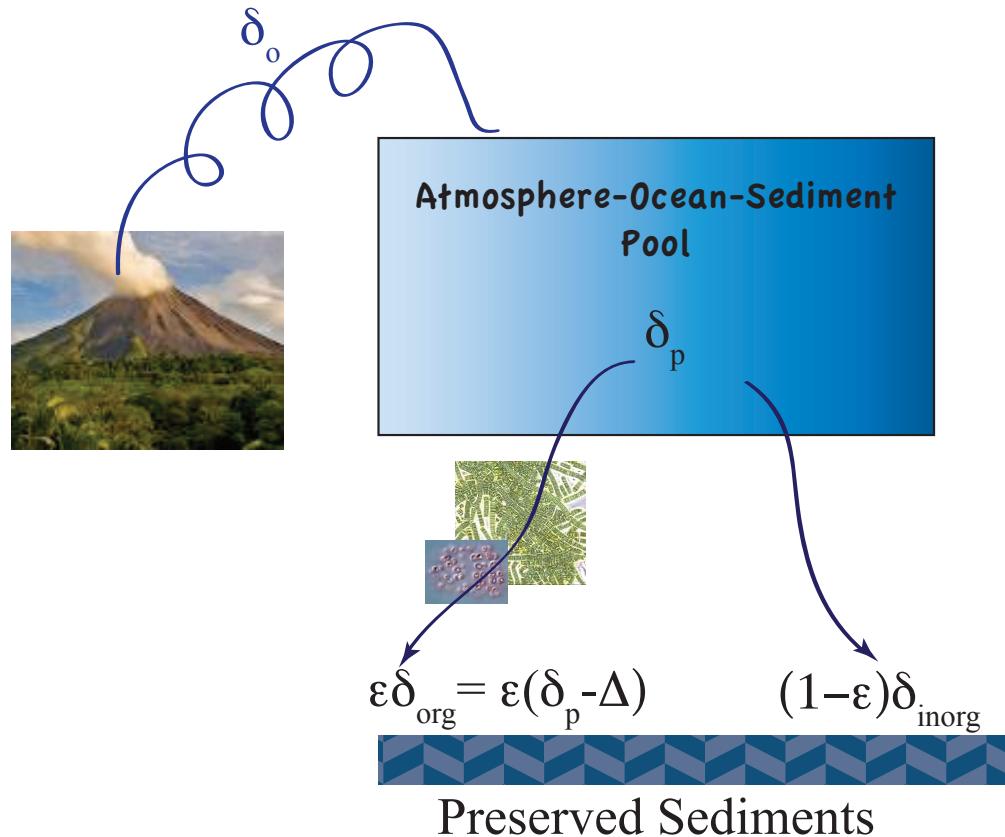
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Dramatis Personae

- Carbon cycle (CO_2 , CH_4 , organic carbon burial)
- Oxygen and Sulfur (SO_4^{--} and O_2 ; pyrite, H_2S)
- Glaciation

$\delta^{13}C$ Primer

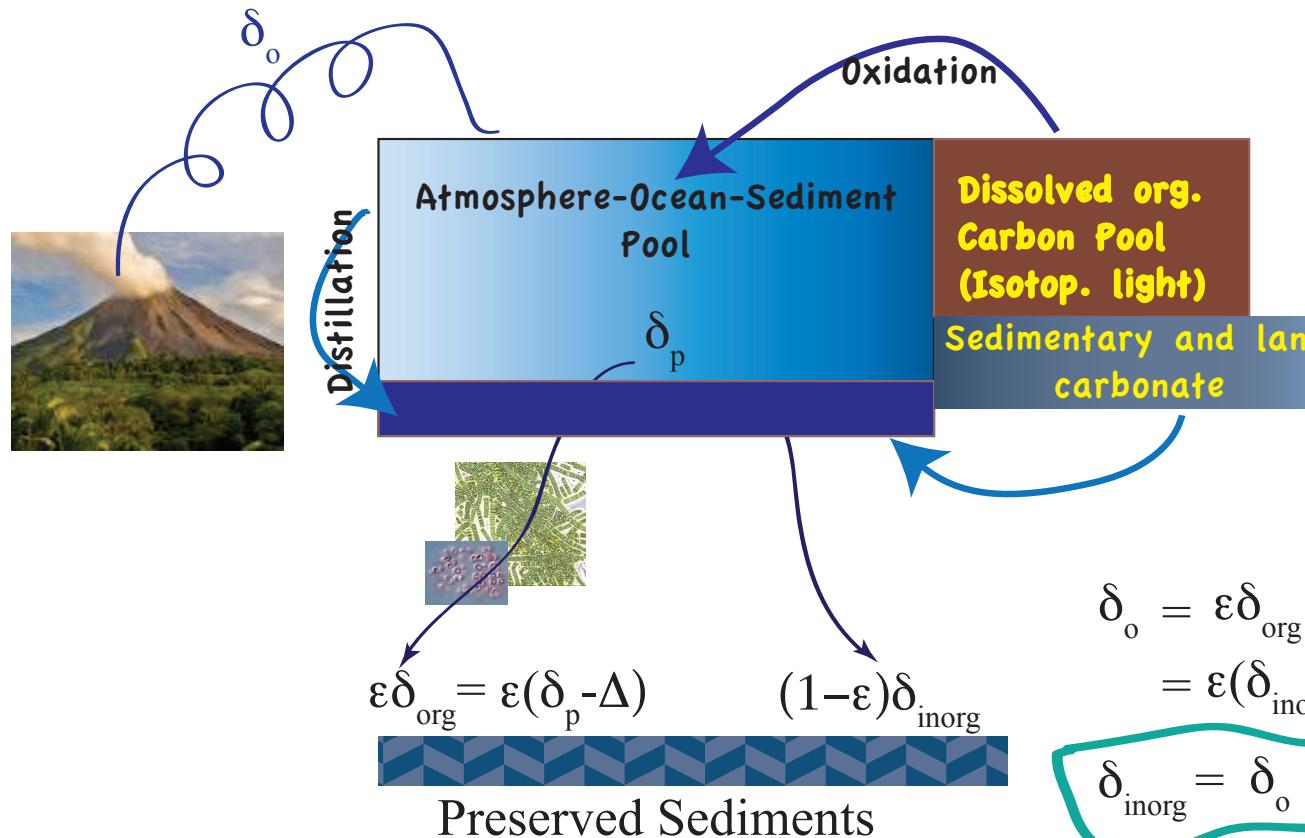
- Stable isotopes ^{12}C and ^{13}C
- ^{12}C is dominant
- Photosynthetic/chemosynthetic life prefers ^{12}C (light carbon)
- Organic matter $\delta^{13}C$ shifted negative relative to pool it is made from



$$\begin{aligned}\delta_o &= \varepsilon \delta_{org} + (1-\varepsilon) \delta_{inorg} \\ &= \varepsilon (\delta_{inorg} - \Delta) + (1-\varepsilon) \delta_{inorg}\end{aligned}$$

$$\delta_{inorg} = \delta_o + \varepsilon \Delta$$

$$\delta_o \approx -6\text{\textperthousand}, \Delta \approx 25\text{\textperthousand}$$

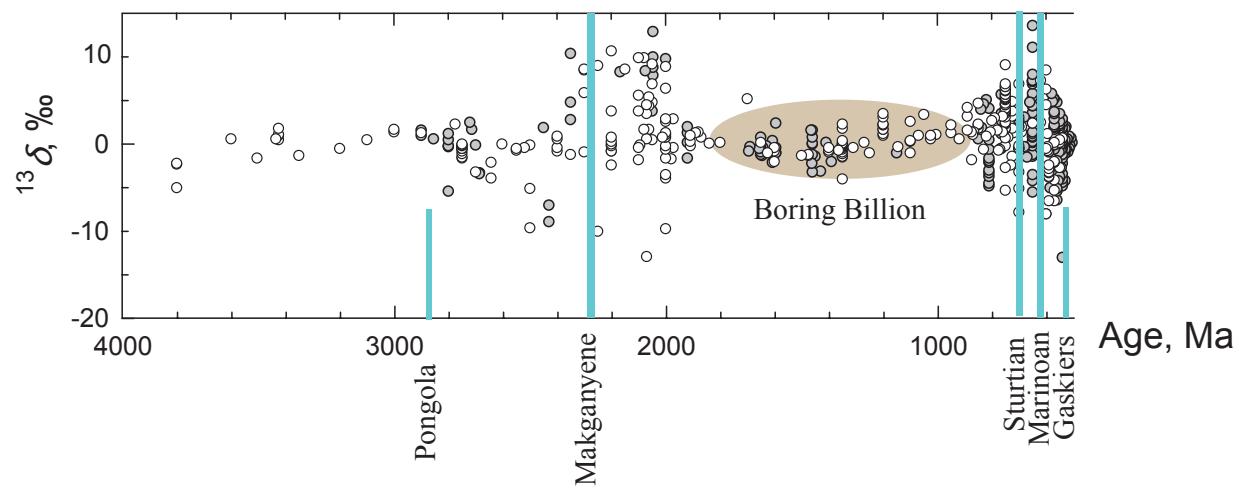


$$\begin{aligned}\delta_o &= \epsilon\delta_{org} + (1-\epsilon)\delta_{inorg} \\ &= \epsilon(\delta_{inorg} - \Delta) + (1-\epsilon)\delta_{inorg}\end{aligned}$$

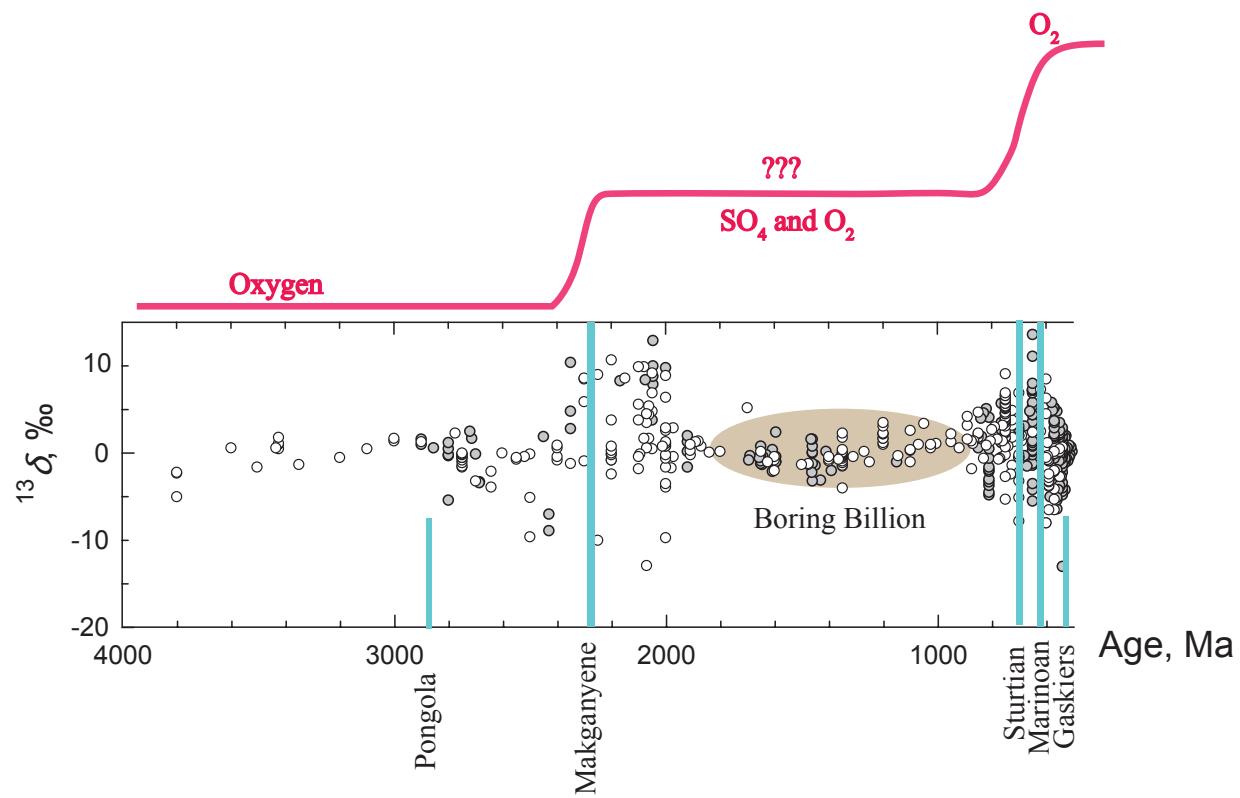
$$\delta_{inorg} = \delta_o + \epsilon\Delta$$

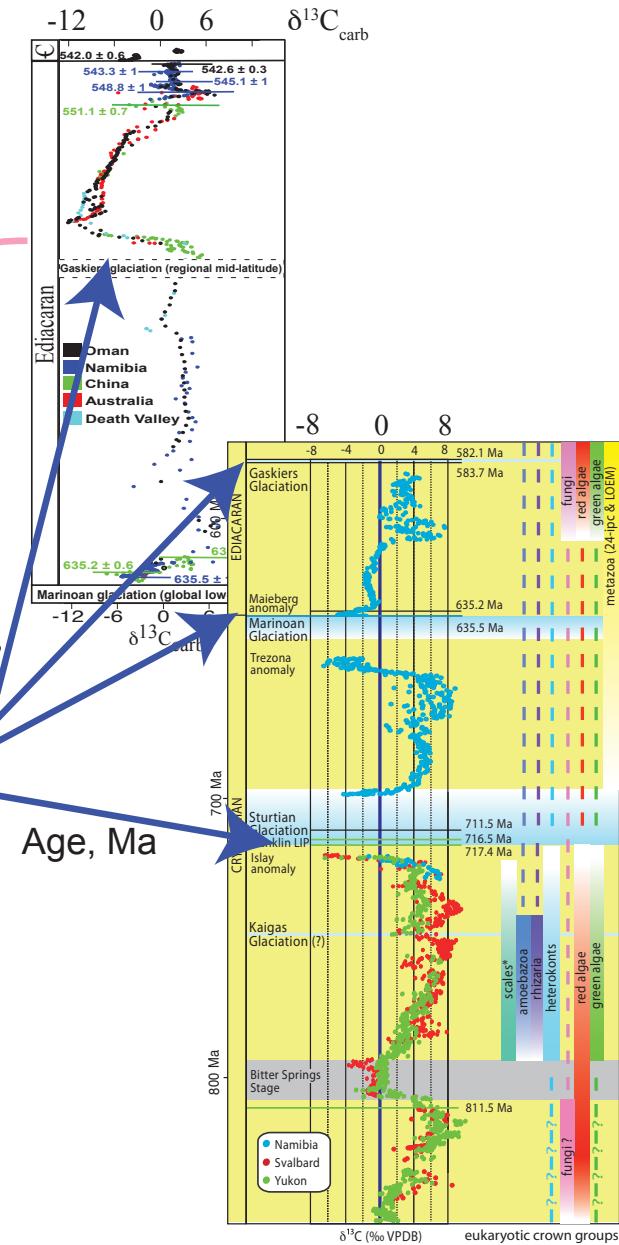
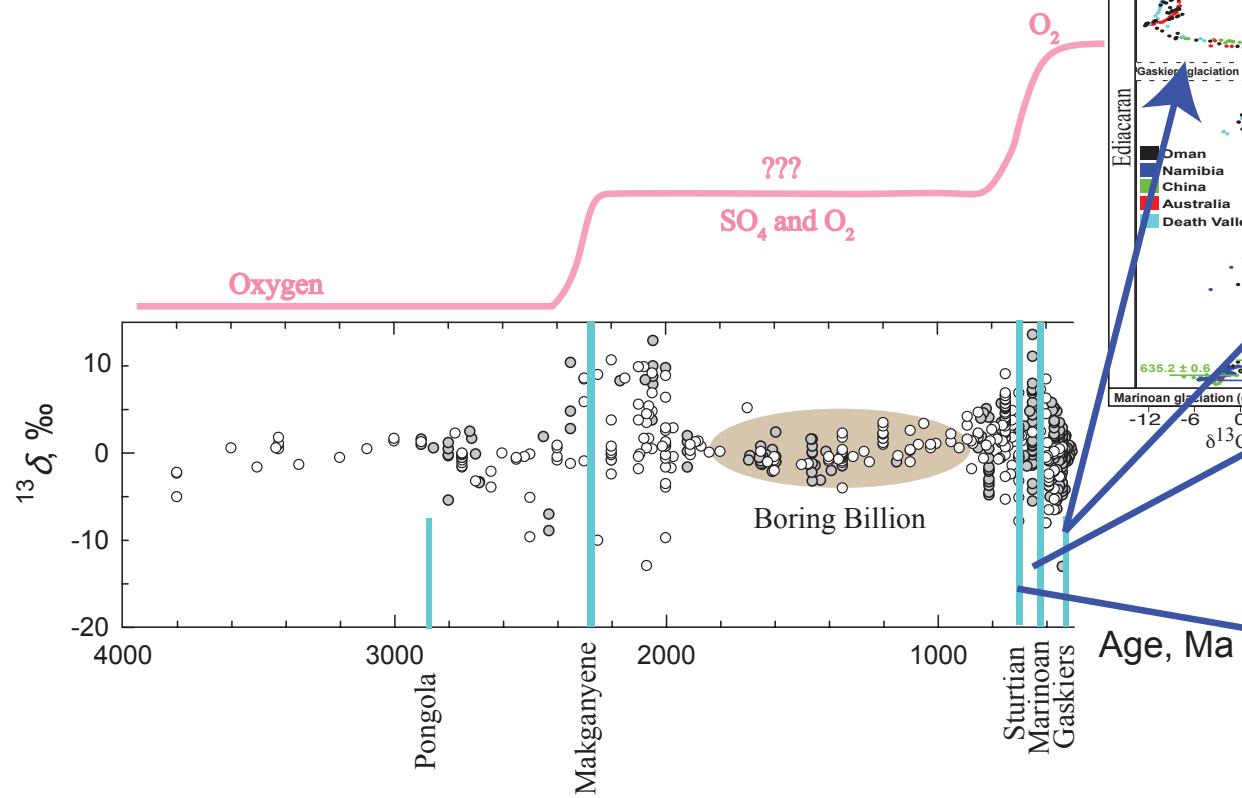
$$\delta_o \approx -6\text{\textperthousand}, \Delta \approx 25\text{\textperthousand}$$

$\delta^{13}\text{C}_{\text{carb}}$



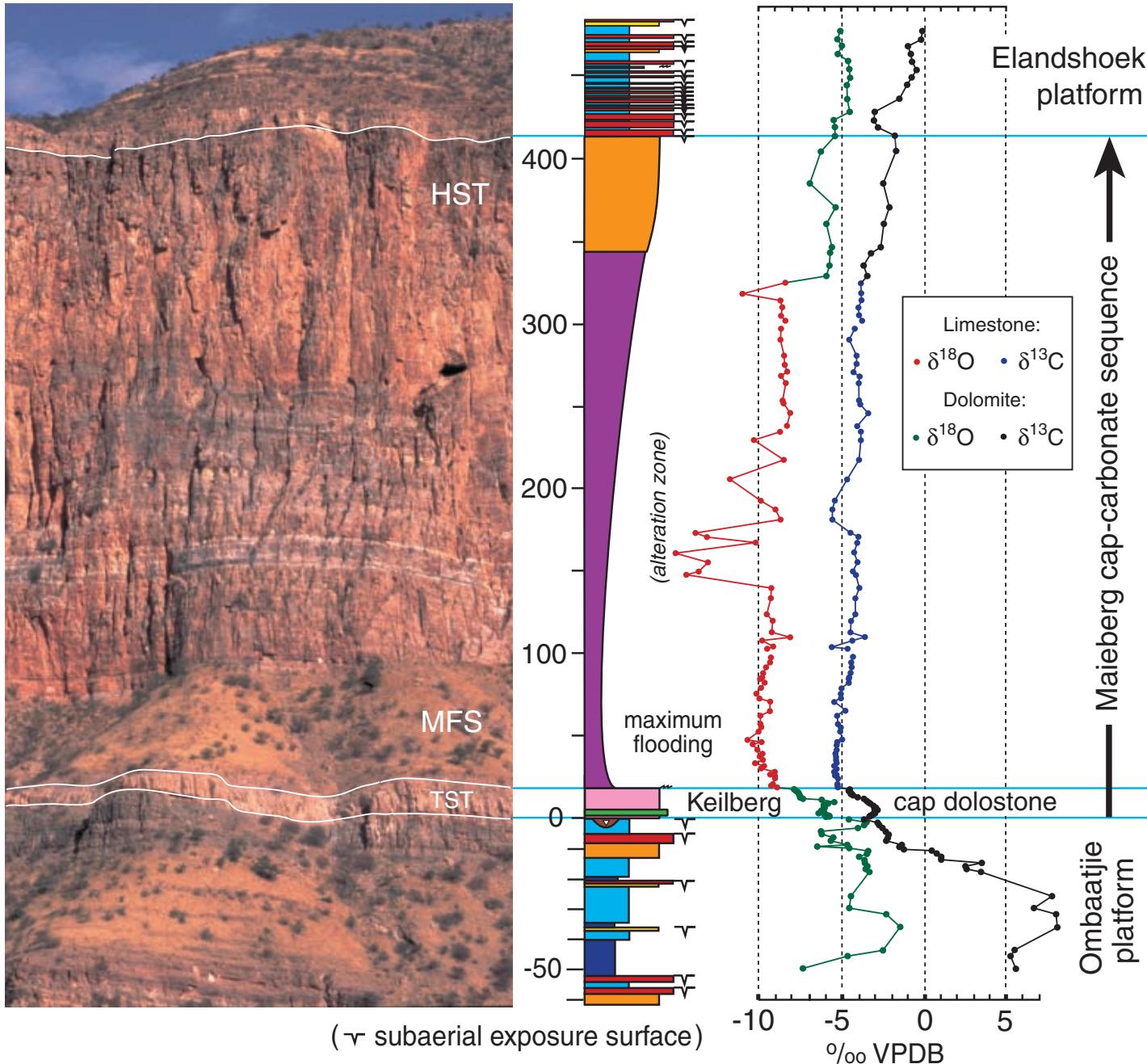
$\delta^{13}\text{C}_{\text{carb}}$





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Cap Carbonates



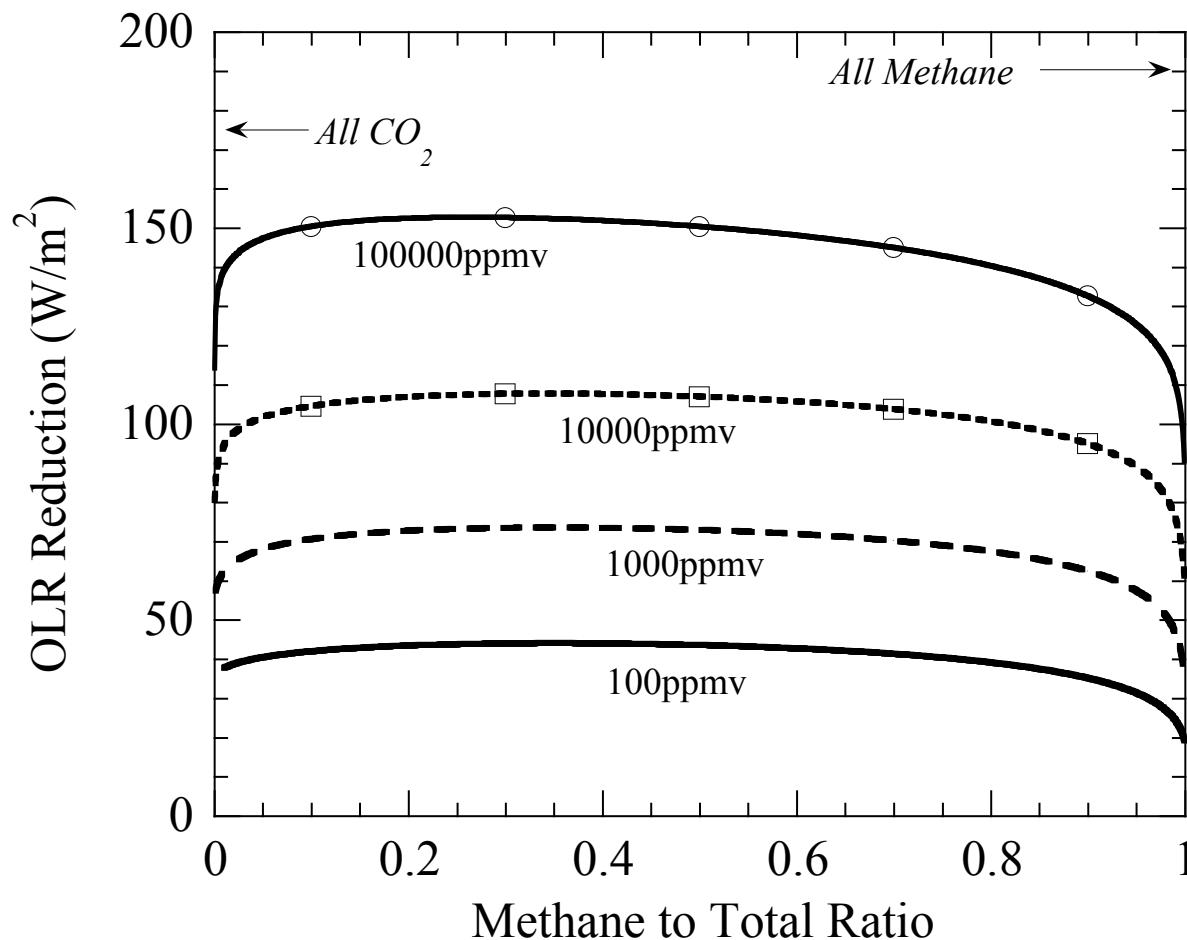
The cap carbonate story

- Features indicate rapid deposition (e.g. giant wave ripples, aragonite fans)
- Interpreted as postglacial, formed when carbonate from land dissolves under hot acid rain in the postglacial hothouse, and washes into ocean.
- Appears to require buildup to very high CO_2 followed by rapid deglaciation – bifurcation between wildly different states.
- i.e. strongly suggestive of a Hard Snowball state.

Carbon Cycle and Climate: DOC pool, CO_2 and oxygen

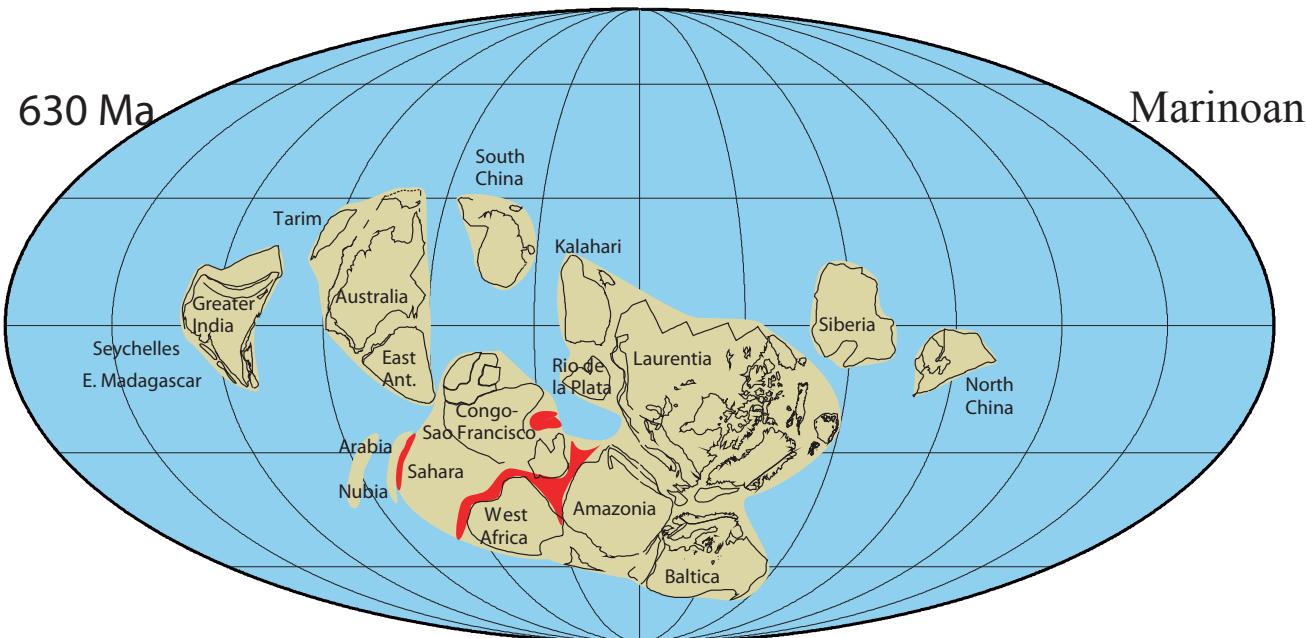
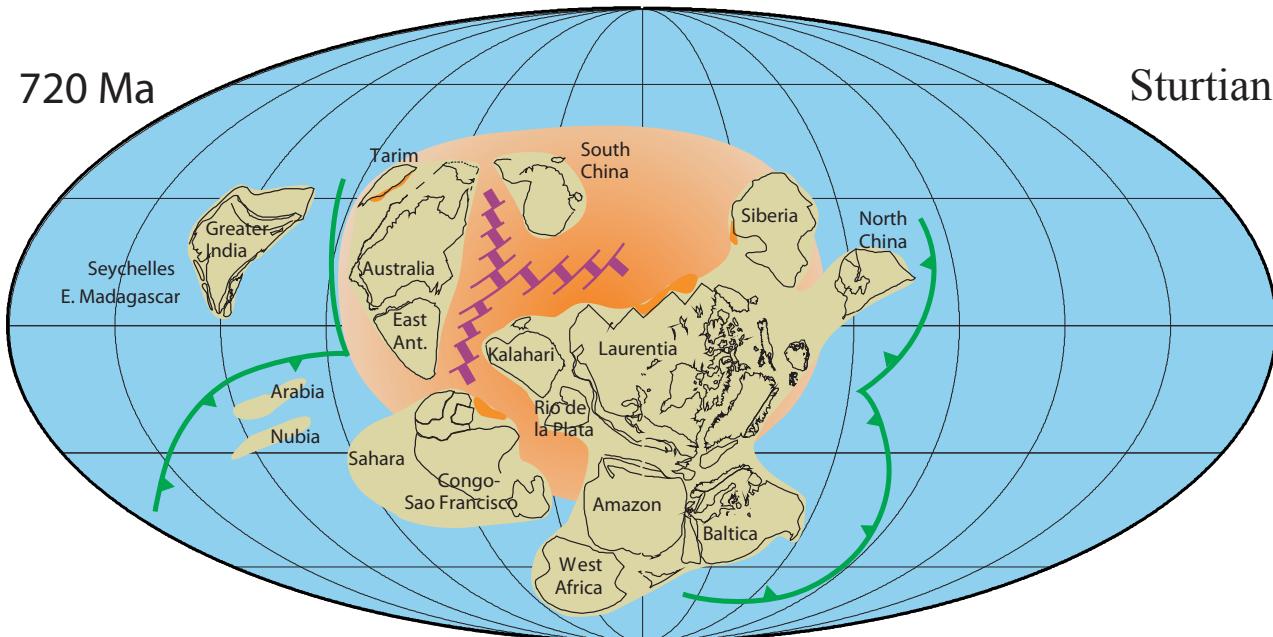
- Oxygenic photosynthesis and organic carbon burial (perhaps in DOC pool) convert a greenhouse gas (CO_2) into a non-greenhouse gas O_2 .
- Oxygen may accumulate in the form of SO_4^{--} or oxidized iron instead of free O_2 .
- Sulfate reducing bacteria can use SO_4^{--} to oxidize DOC pool
- DOC provides a pool of light carbon.
- Oxidation of DOC into CO_2 produces warming and negative isotopic excursion

Carbon Cycle and Climate: CO_2 vs CH_4



Carbon Cycle and Climate: CO_2 vs CH_4

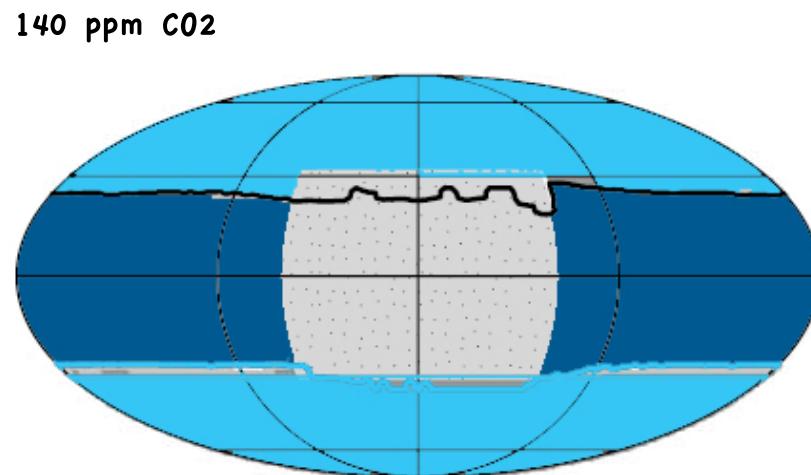
- Hence a methane-dominated greenhouse doesn't cause cooling when methane oxidizes to CO_2 in the air
- However CO_2 is more soluble (via carbonate equilibrium) than methane, so this allows some reduction of greenhouse effect
- Note that methane is isotopically light, so if methane carbon winds up in carbonate sediment, it drives a negative ^{13}C excursion



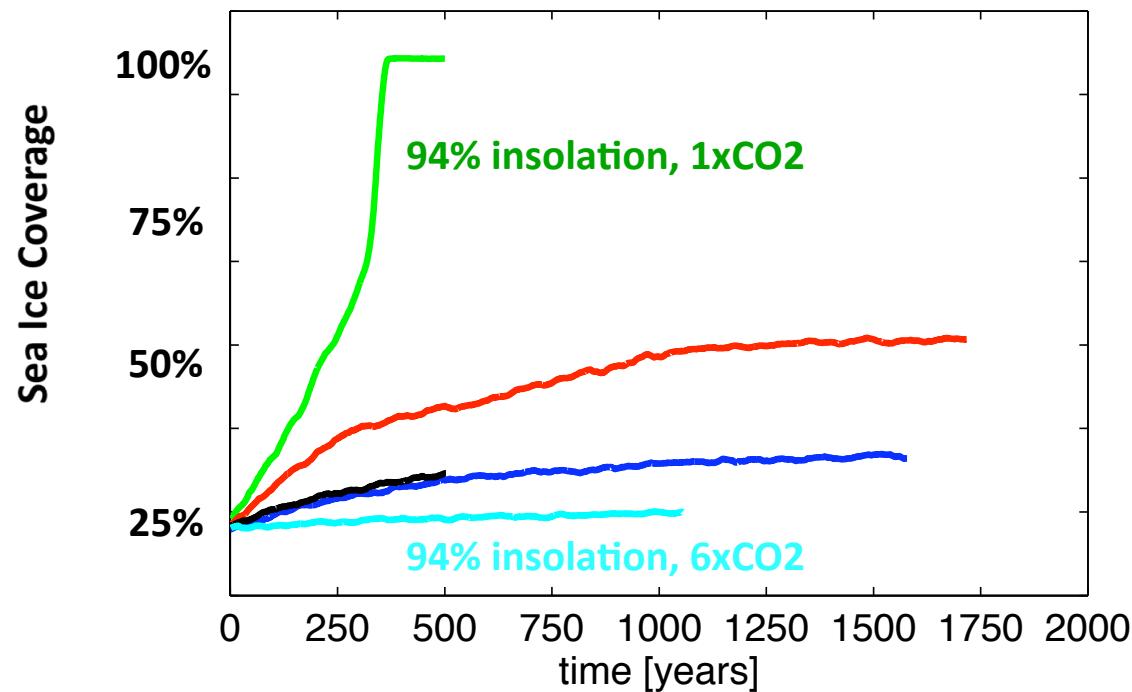
Jumping into the deep end: Ocean dynamics and initiation

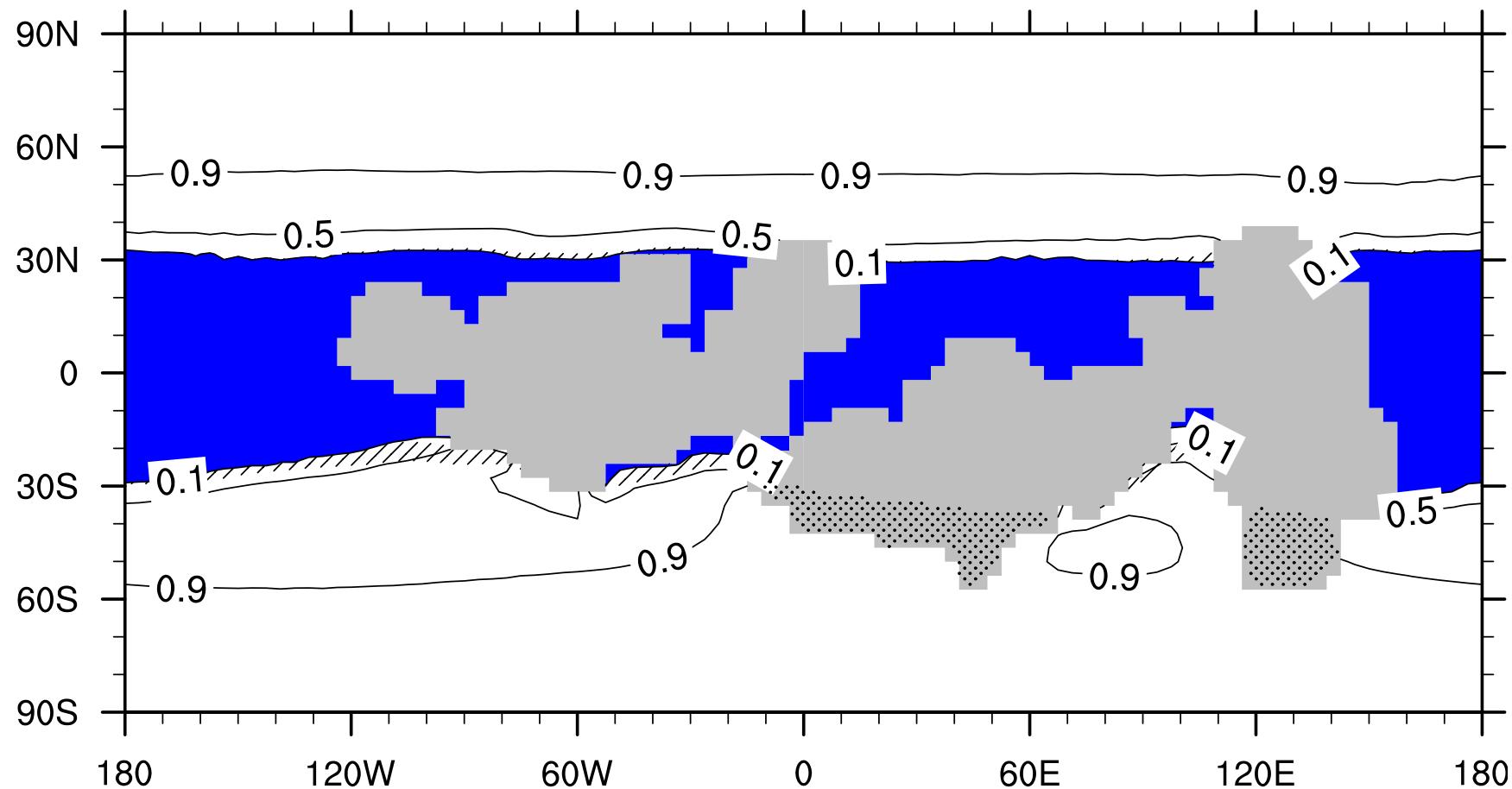
Poulsen and Pierrehumbert *GRL* 2001, ...,

Poulsen and Jacob *Paleocean*. 2004



The MPI coupled global climate model initiates a Snowball at reasonable Neoproterozoic conditions.

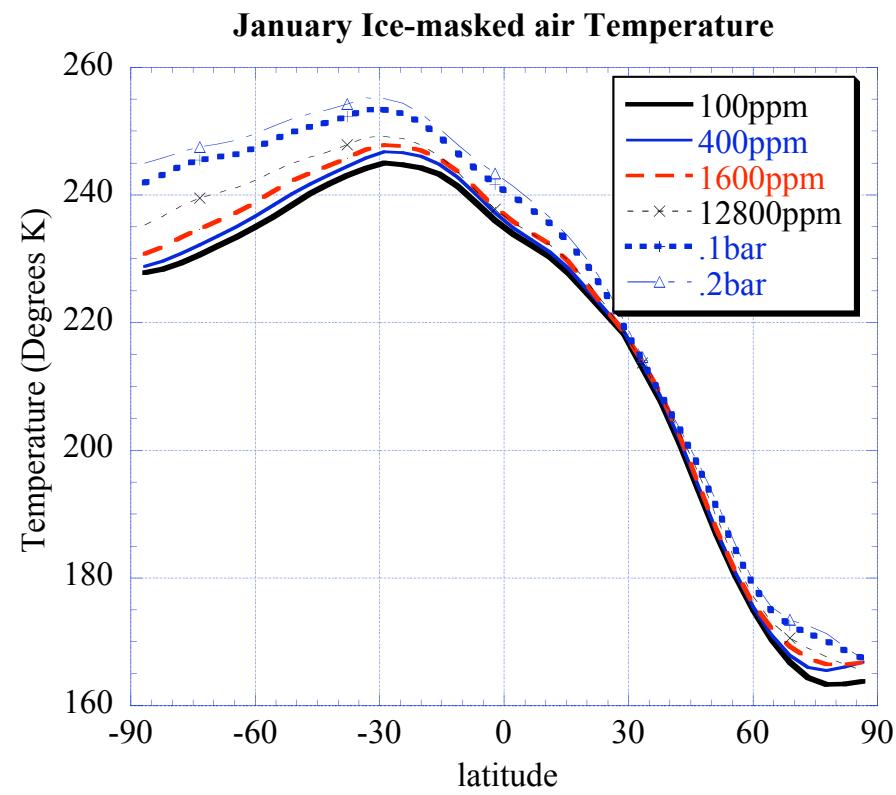




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Jumping into the deep end: Deglaciation

Pierrehumbert, *Nature* 2004, *J. Geophys Res* 2005.

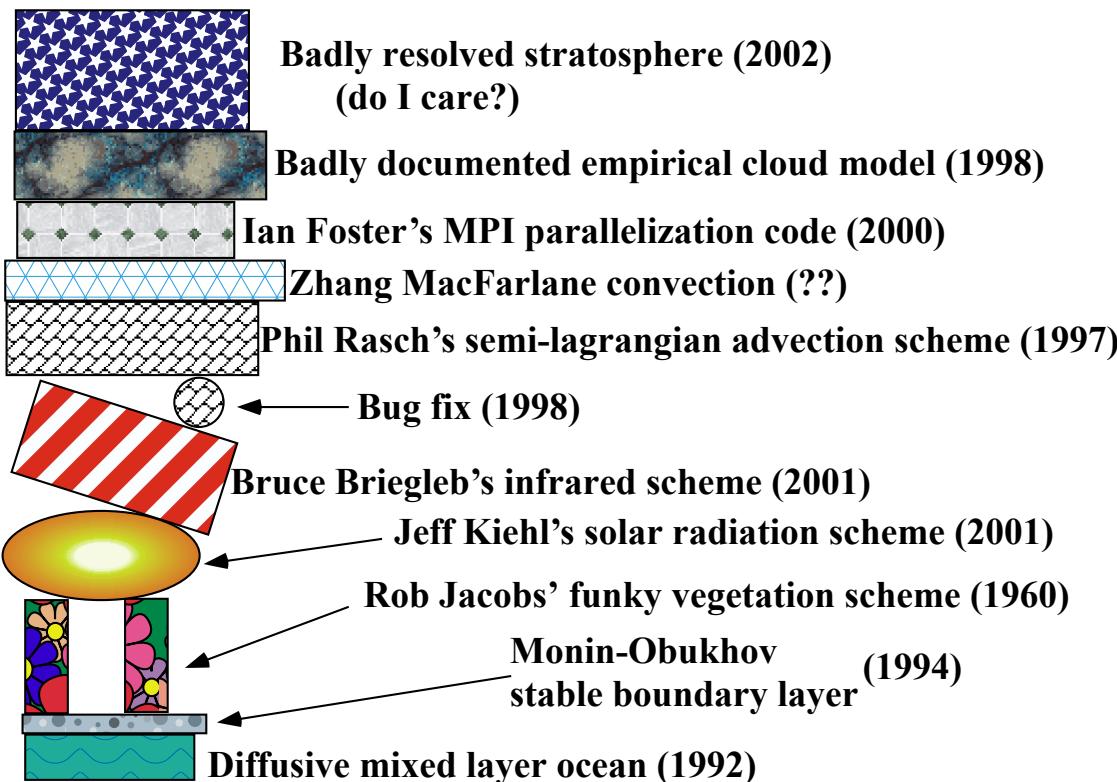


Advantages of using a GCM

- Includes physics you might not have thought was important, or known about
- Collective behavior between parts you didn't think would interact
- Very hard to systematically represent synoptic eddy transports in EBM's

(But there are also some disadvantages ...)

Stratigraphic analysis of an Atmospheric GCM



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Snowball Earth and Snowball Earths Part II: Simulations, theory,mathematics

Raymond T. Pierrehumbert
The University of Chicago

The Snowball provides a rich lode of idealized climate dynamics problems

- Snowball state:
 - Dry dynamics, no water vapor feedback
 - Clouds are radiatively active, but quasi-passive tracers
 - No ocean dynamics or thermal inertia
- We'll start with Earthlike configurations, then generalize to exoplanets with exotic orbits and spin states

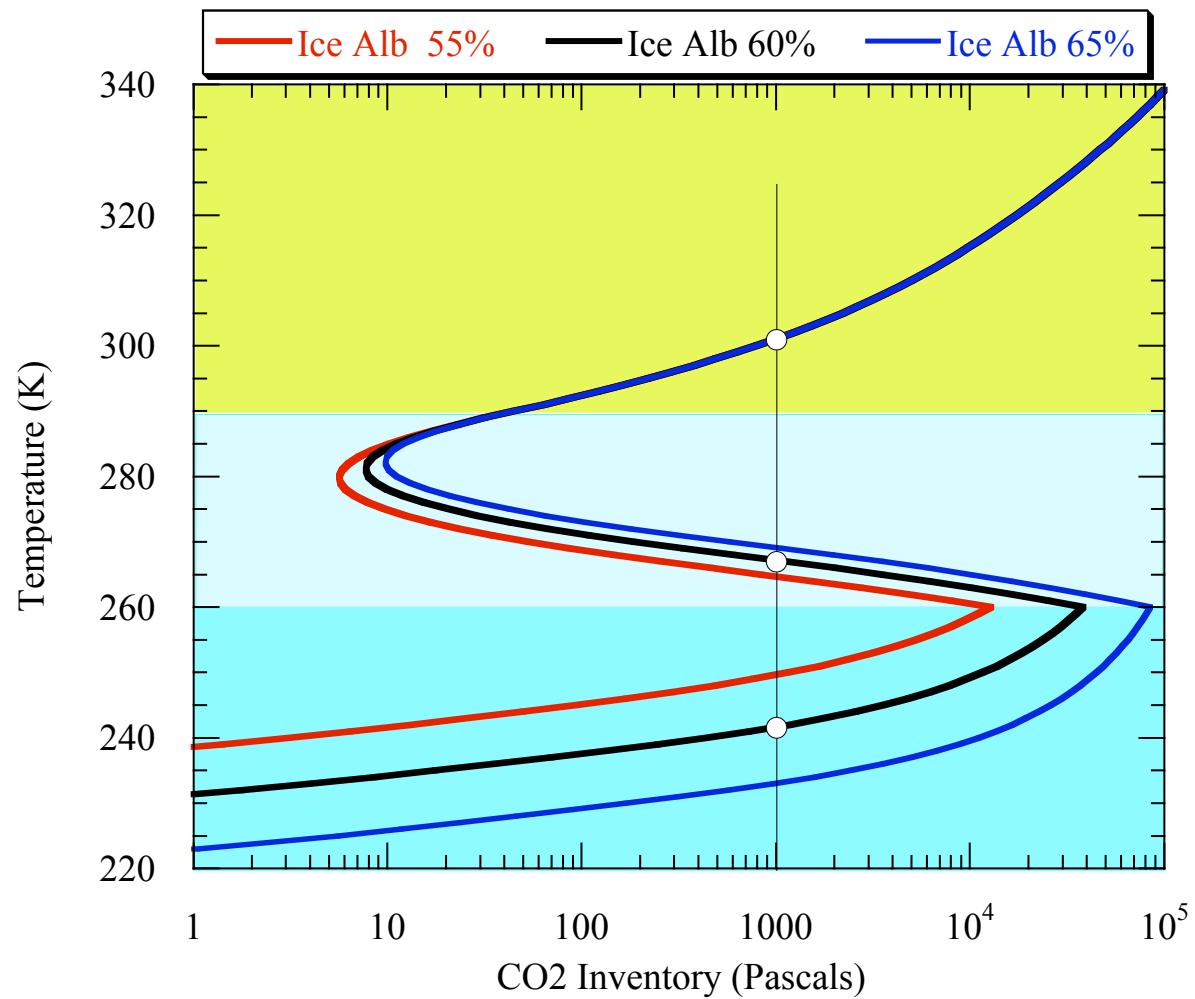
Our theoretical toolkit for understanding what's going on Snowball simulations

- Energy balance models (various flavors)
- Radiative-convective column models
- Surface energy balance models
- Thermal diffusion in ice with nonlinear upper flux B.C.
- Hadley cell theory (various flavors)
- WTG models
- Synoptic eddy transport theory ???
- 1D "shallow ice" models (Sea Glaciers)
- Precipitation-Temperature relation models
- Carbon cycle box models (Geochemical weathering)

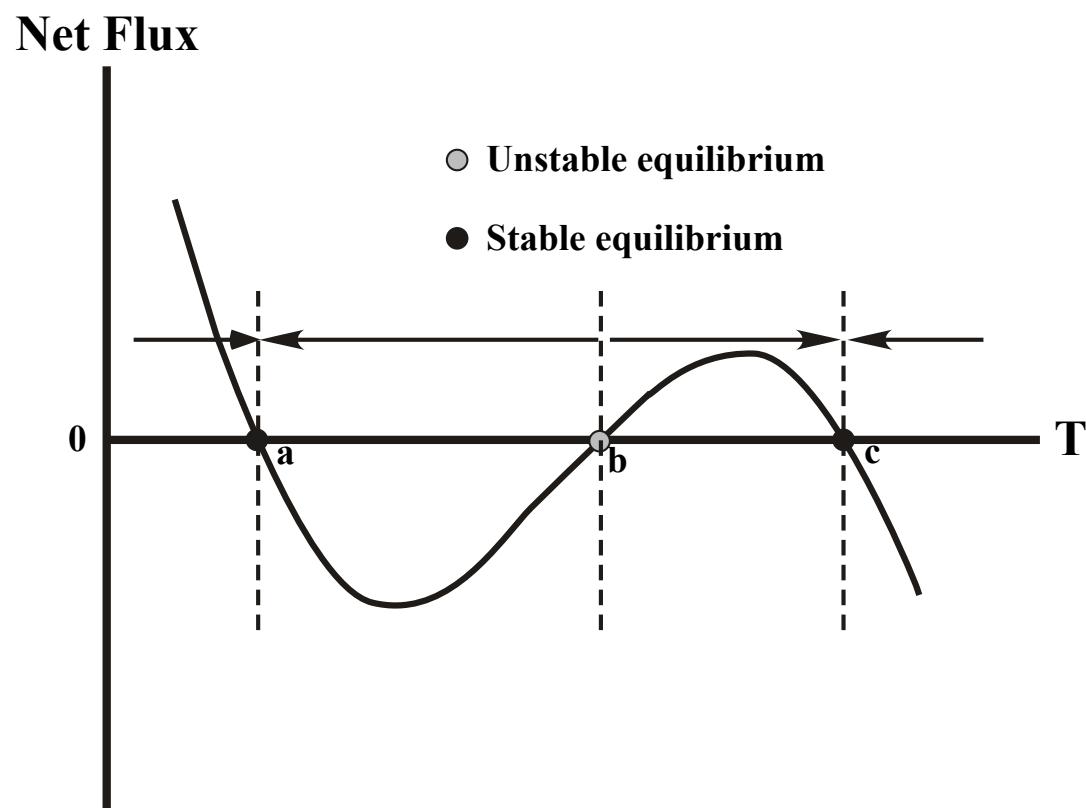
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Zero D Steady Snowball Model

$$\frac{1}{4}(1 - \alpha(T))L_{\odot} = OLR(T, CO_2)$$



Attractor basins



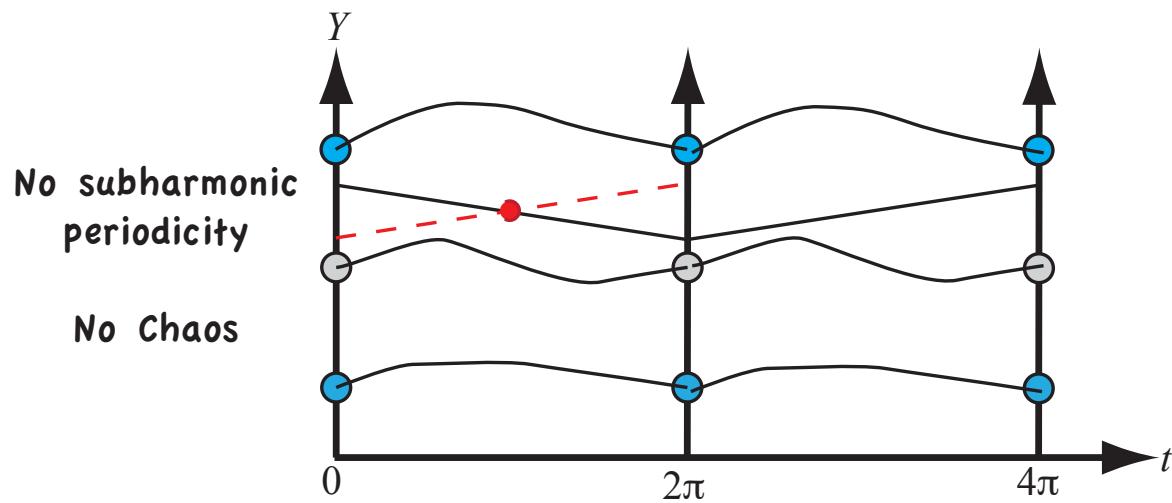
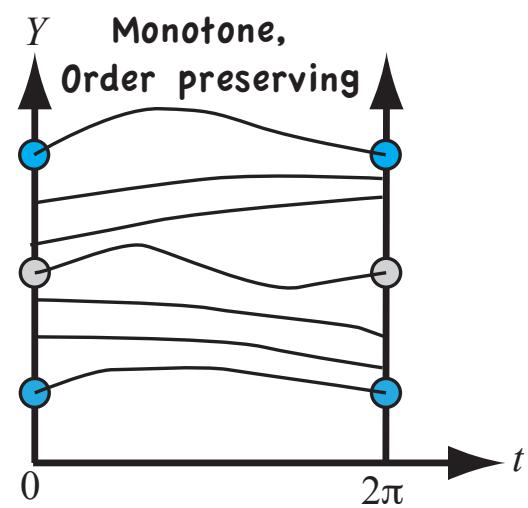
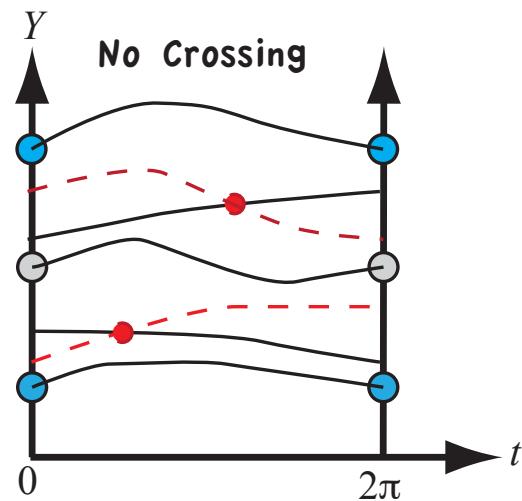
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Zero-D nonlinear with seasonal forcing

$$\frac{dY}{dt} = F(Y, t)$$

$$F(Y, t + 2\pi) = F(Y, t)$$

$F < 0$ for large Y , $F > 0$ for $Y \rightarrow 0$



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1-D nonlinear PDE with seasonal forcing

$$\partial_t E(T) = \partial_y (1 - y^2) D \partial_y T + (1 - \alpha(h_i)) S(y, t) - OLR(T)$$

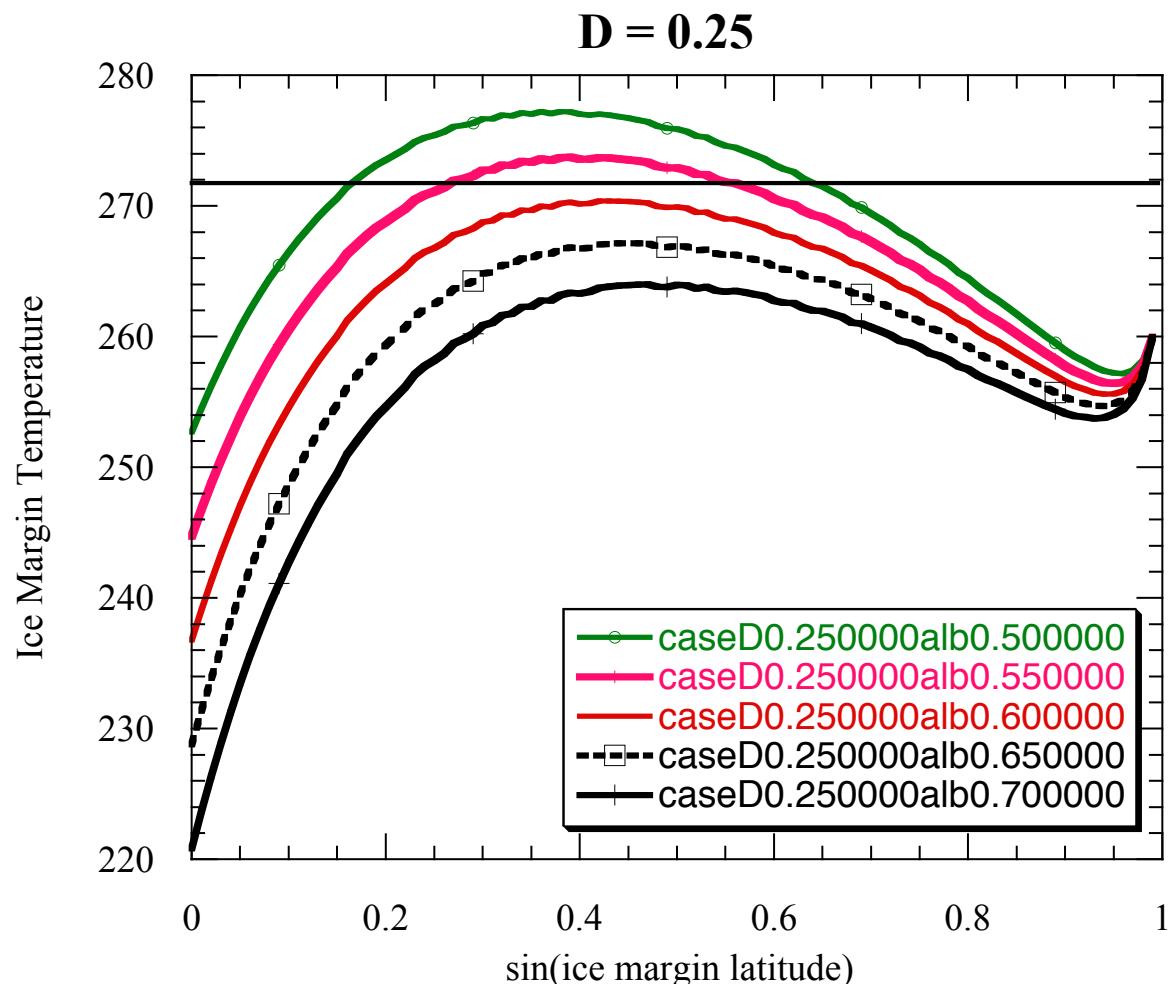
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Steady-State becomes an ODE

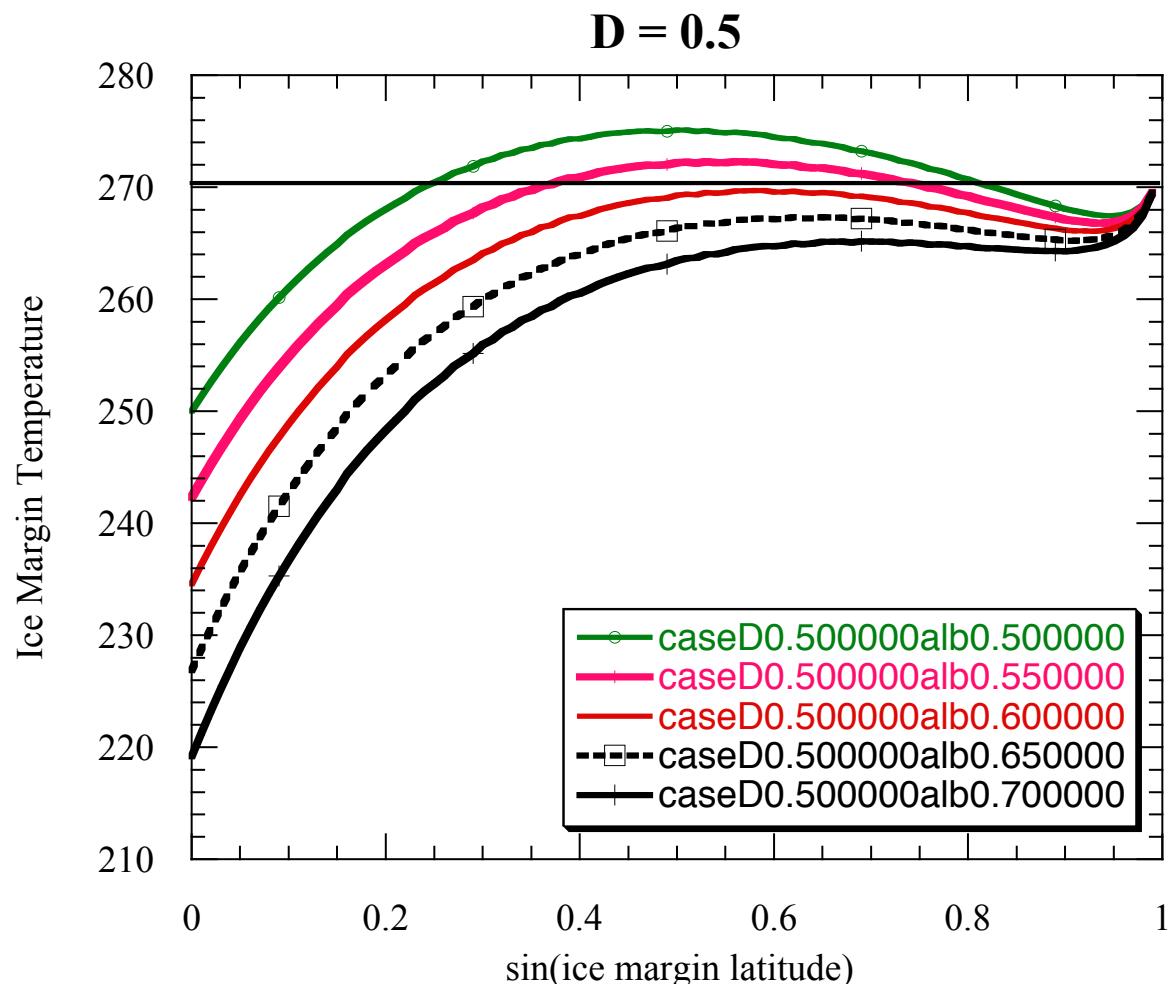
$$0 = \partial_y(1 - y^2)D\partial_y T + (1 - \alpha(h_i))S(y, t) - OLR(T)$$

Plot ice margin T vs ice margin position

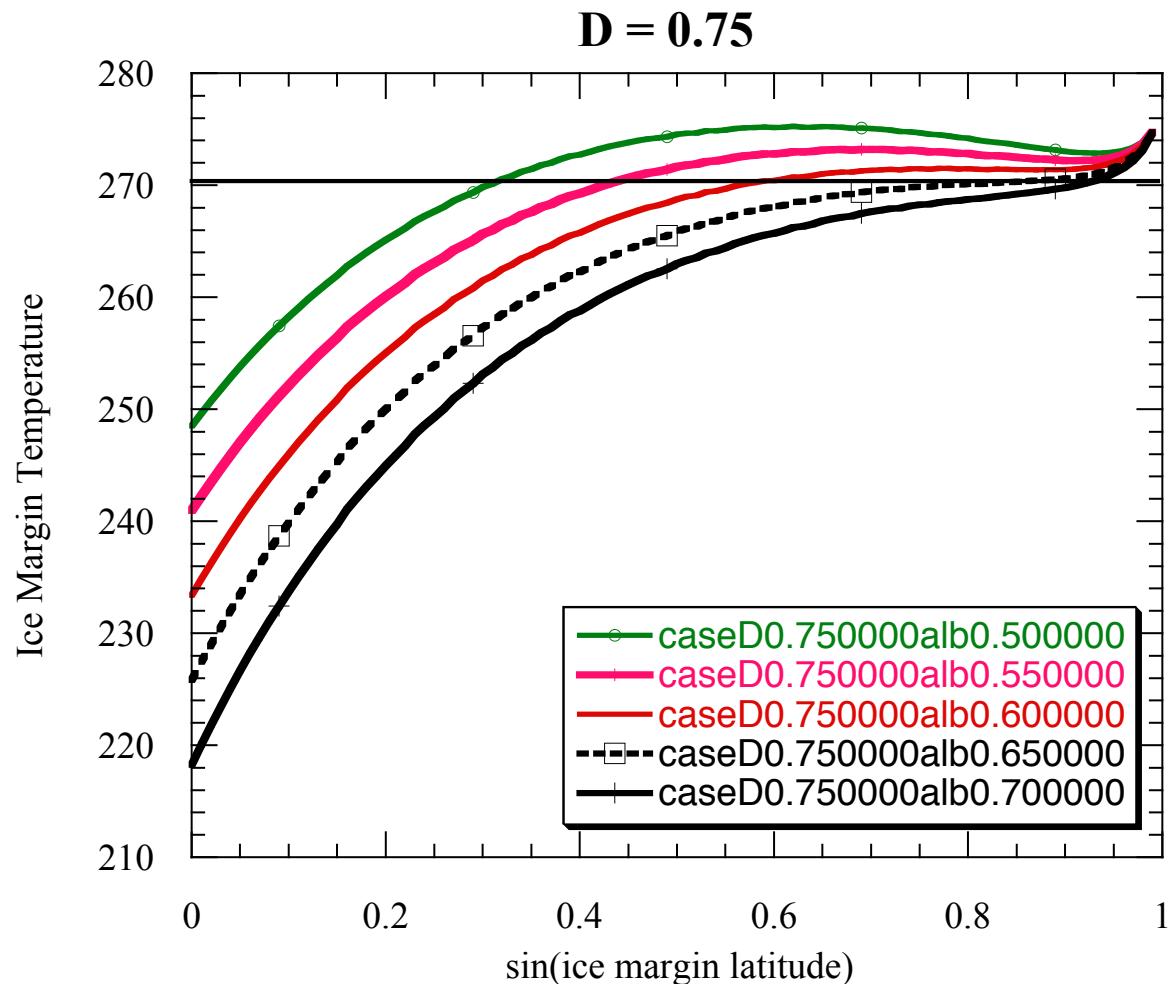
Steady State



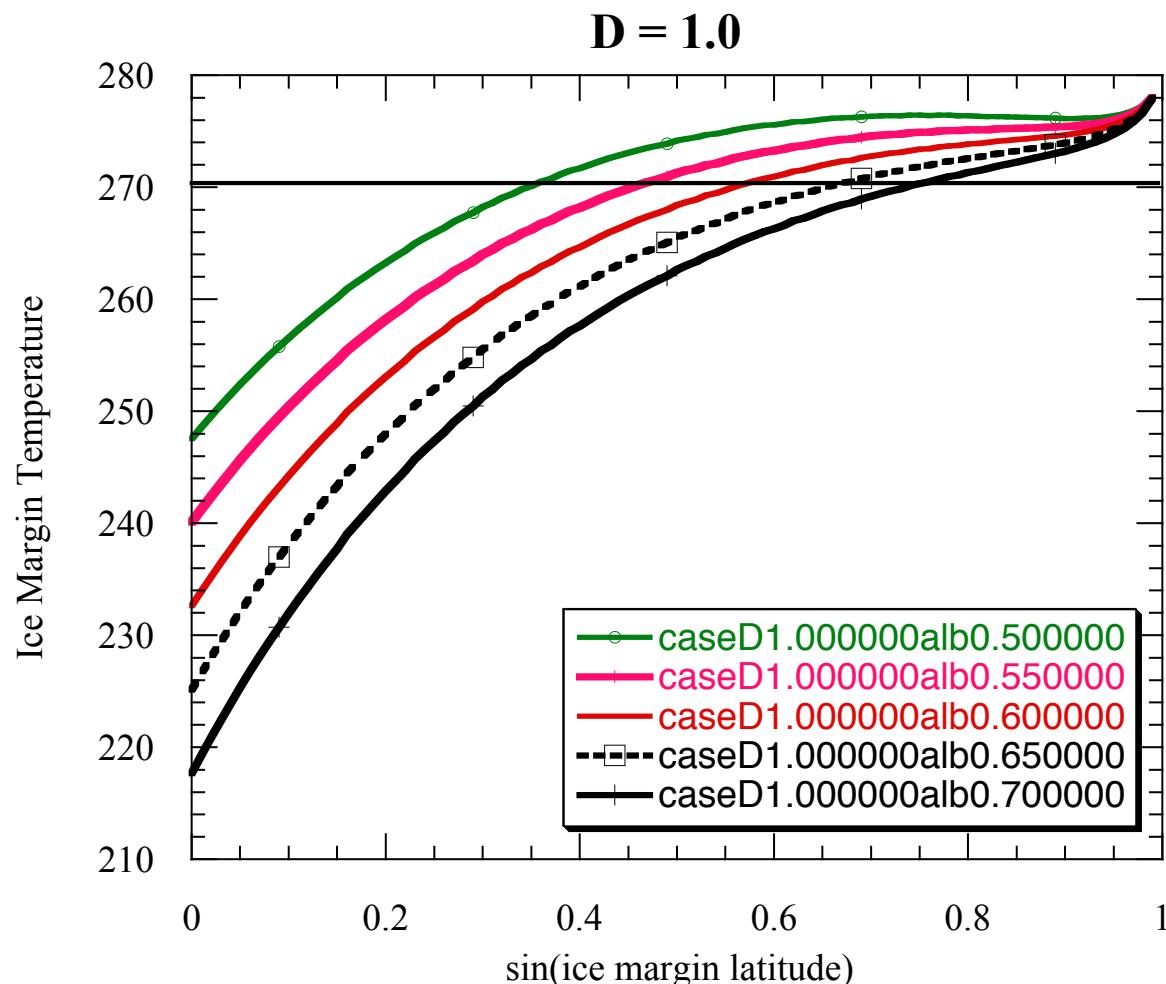
Steady State



Steady State

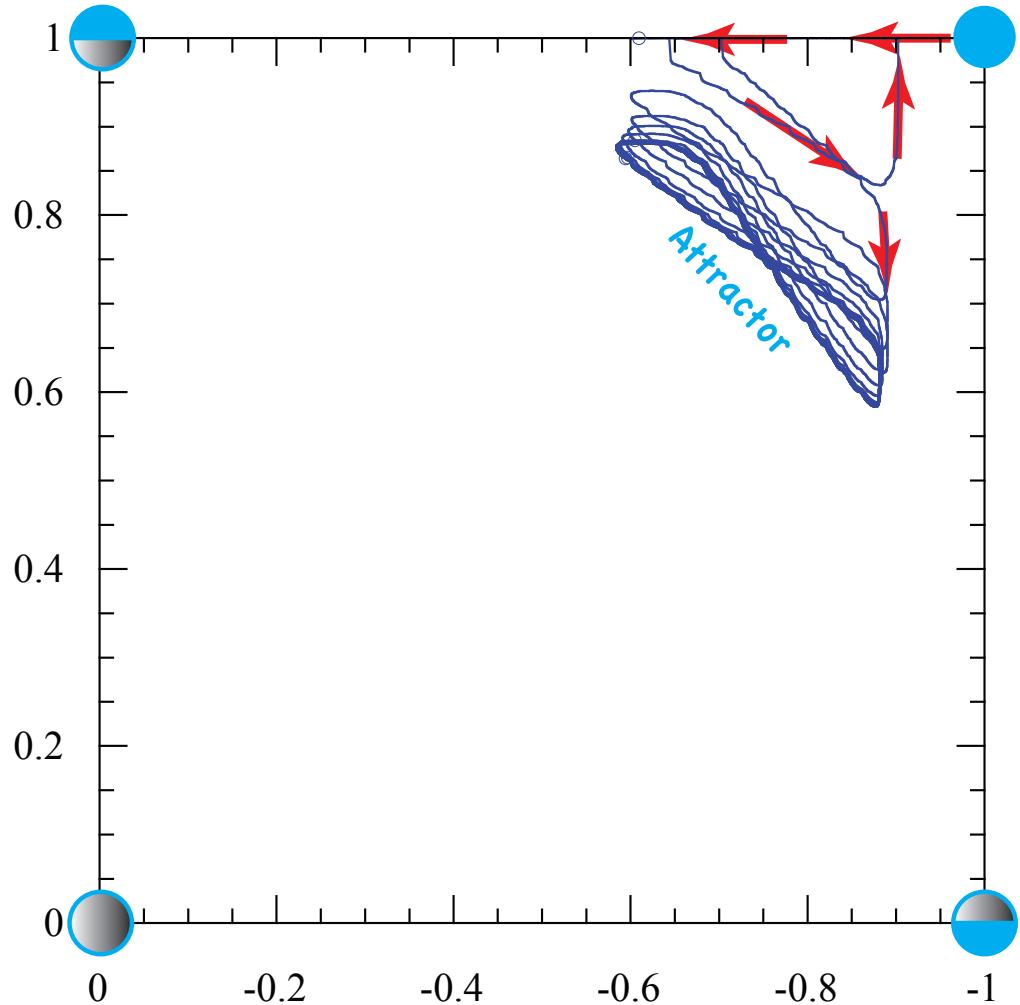


Steady State



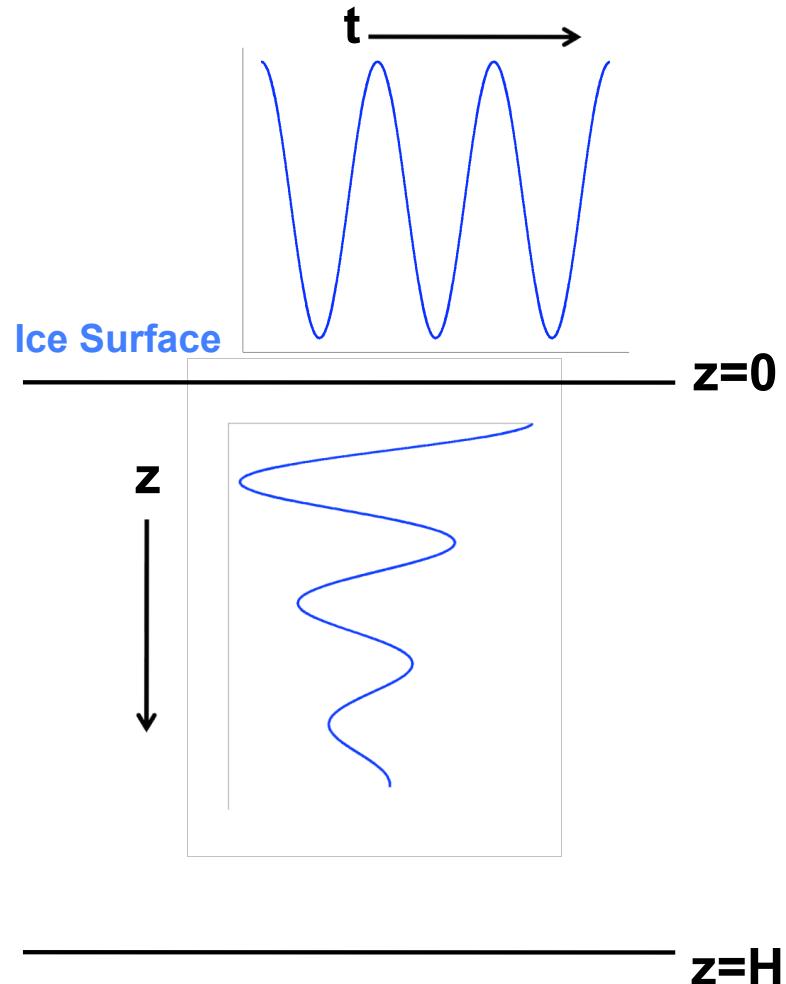
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Attractor for 1D PDE Seasonal Snowball EBM

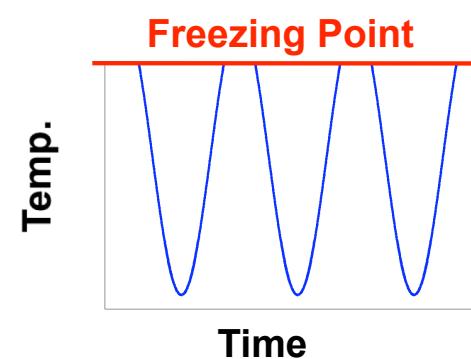


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An exaggerated surface temperature diurnal cycle can lead to false deglaciations.

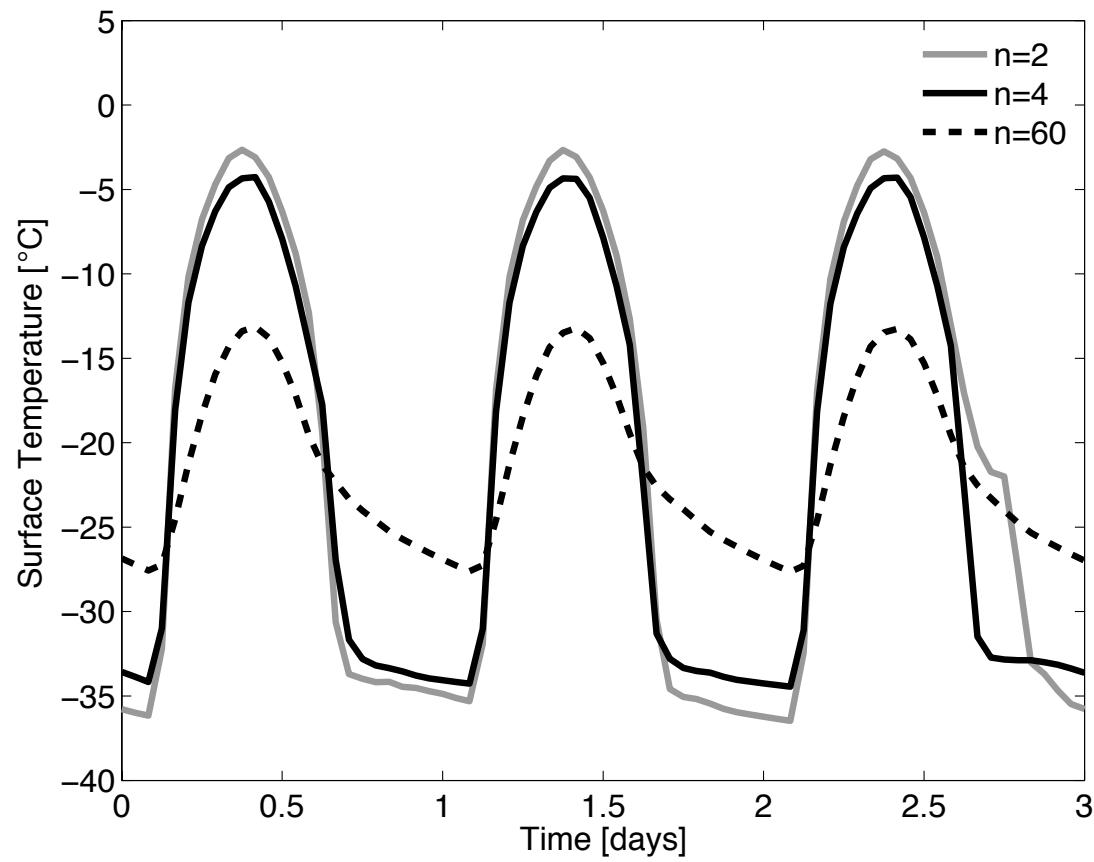


We care because melting occurs at noon

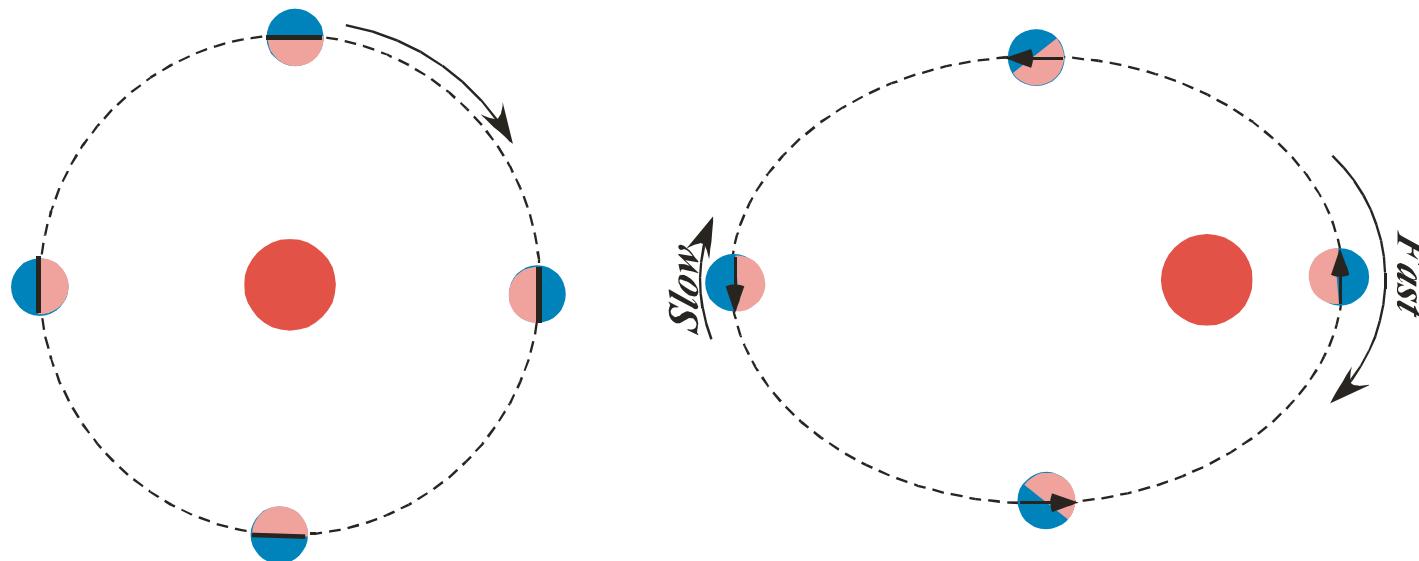


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Drastic reduction in diurnal cycle occurs when you increase sea ice resolution below a single column atmosphere as well.



Review: Tide-locked and quasi-synchronous states

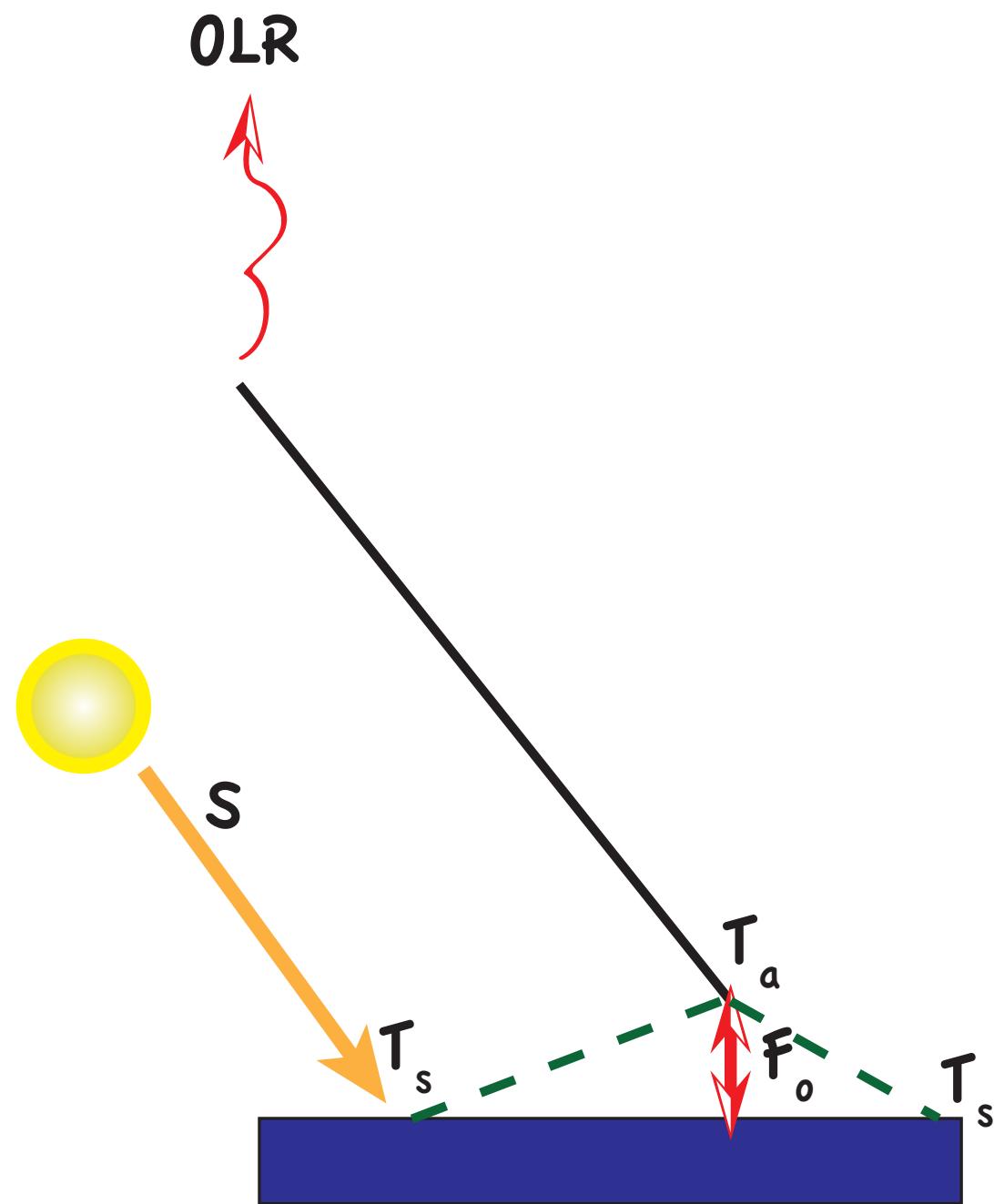


Review: The basic dynamical regime

- Orbital/diurnal periods of 20-100 Earth days
- Coriolis effects at short end comparable to Titan
- Low Coriolis force implies efficient atmospheric heat transport. Nearly isothermal free atmosphere.
- Can use WTG approximation globally
- Applies to 3:2 spin-orbit resonance case as well

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For WTG Snowball, surface budget is what makes it interesting

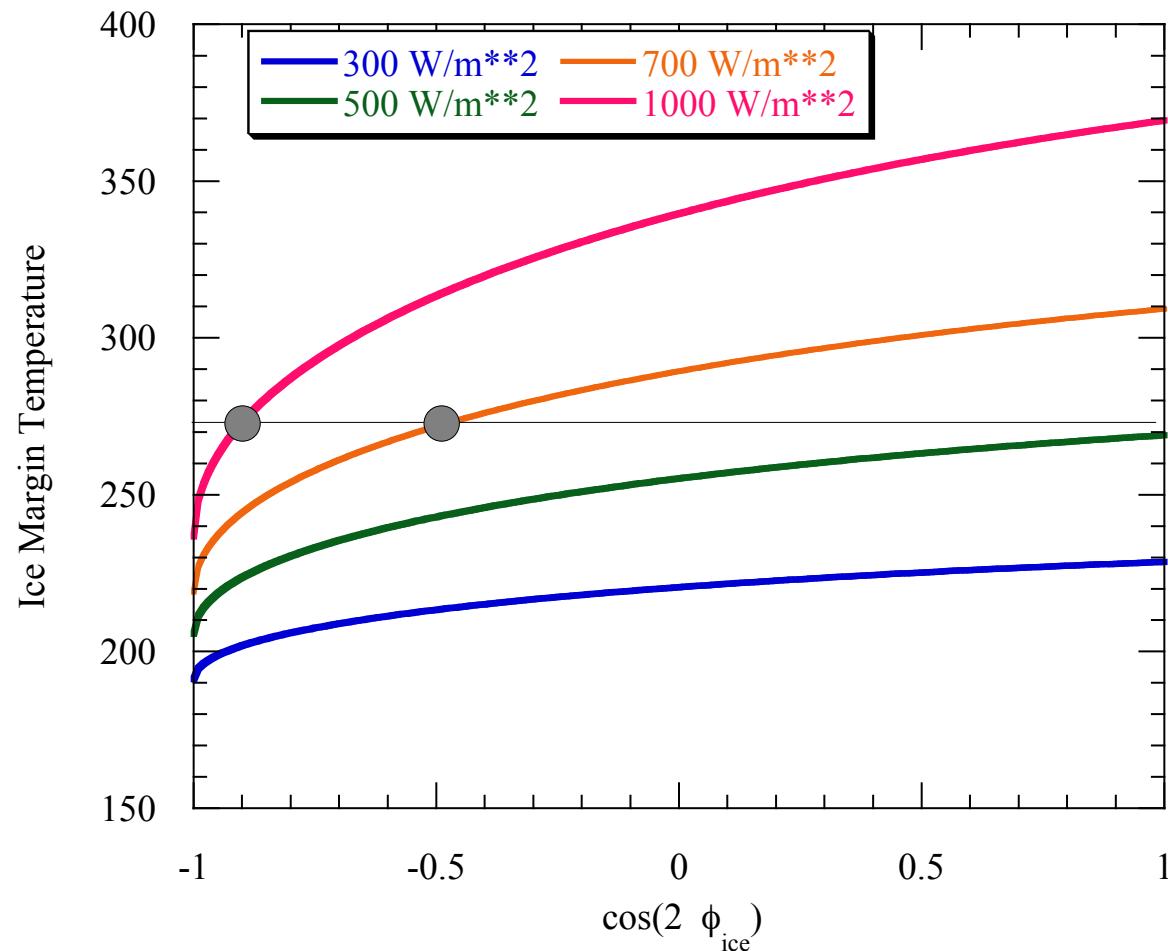


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$$OLR(T_a) = L_{\odot} f(\phi_i)$$

$$F_o(T_a) + a_s \cdot (T_s(\phi) - T_a) = L_{\odot} \cos \phi$$

Equilibrium results for a tide-locked planet



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Planet, Planet shining bright

In the Kepler's wobbly sight

Are you like Venus dry and hot?

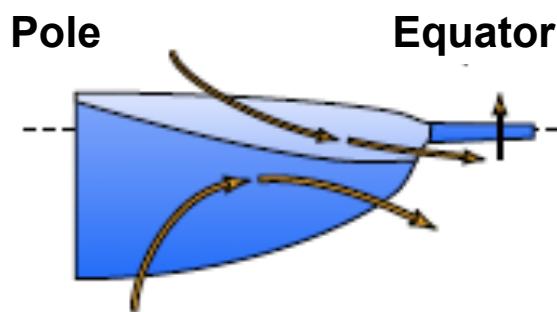
Or icy Snowball that is not?

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The hydrological cycle drives ice flow.

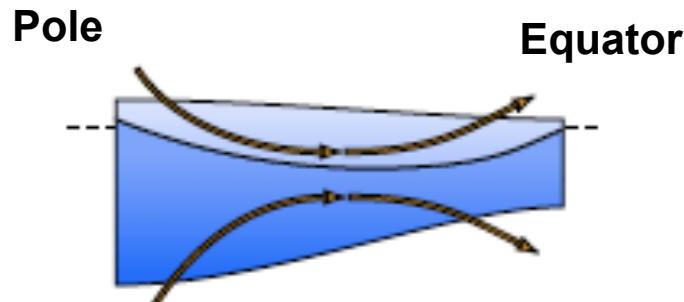
Sea Ice Elevator

[Hoffman and Schrag, 1999]



Ice flow from higher latitudes

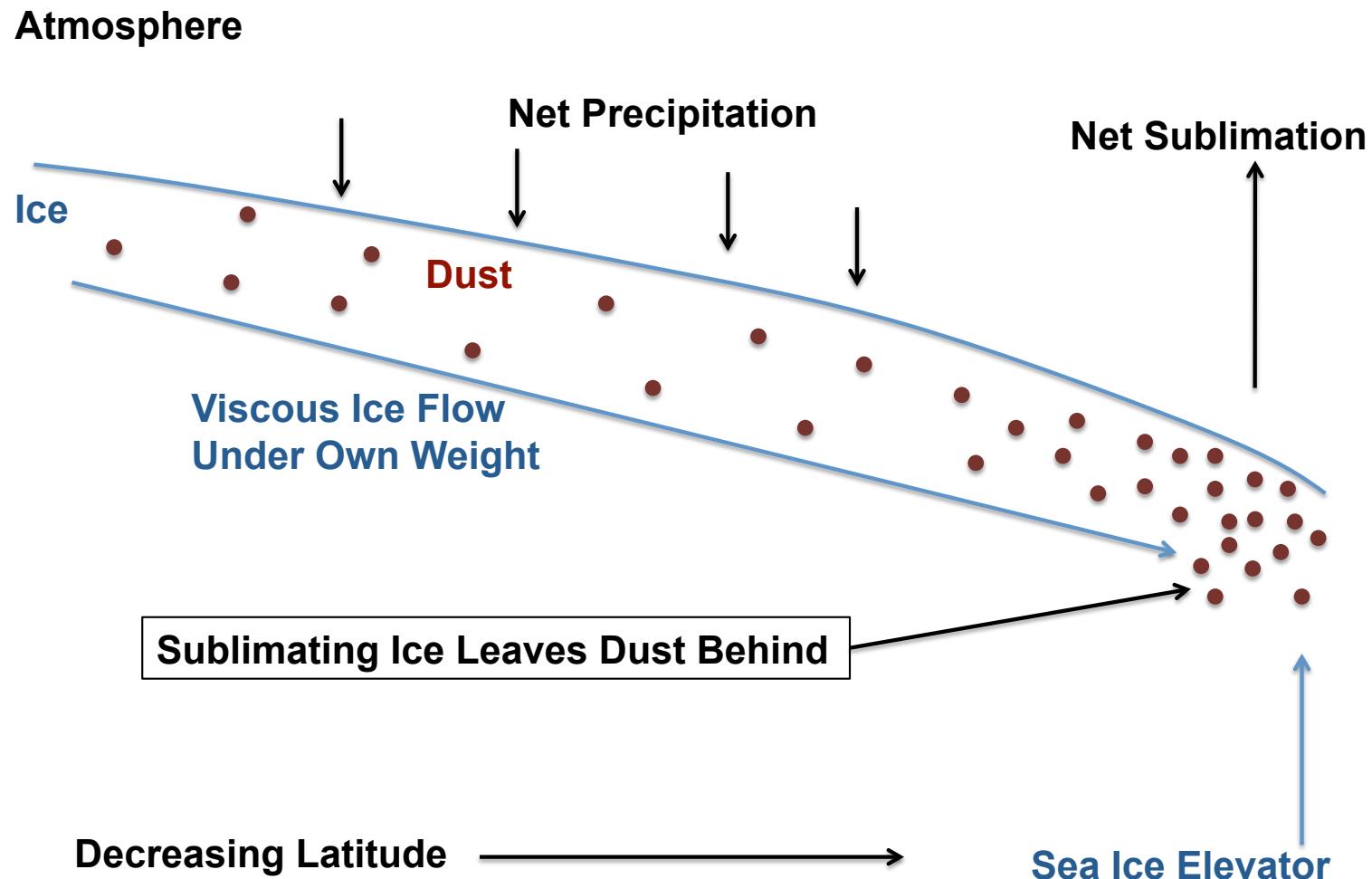
[Goodman and Pierrehumbert, 2003]



Figures from Goodman (2006)

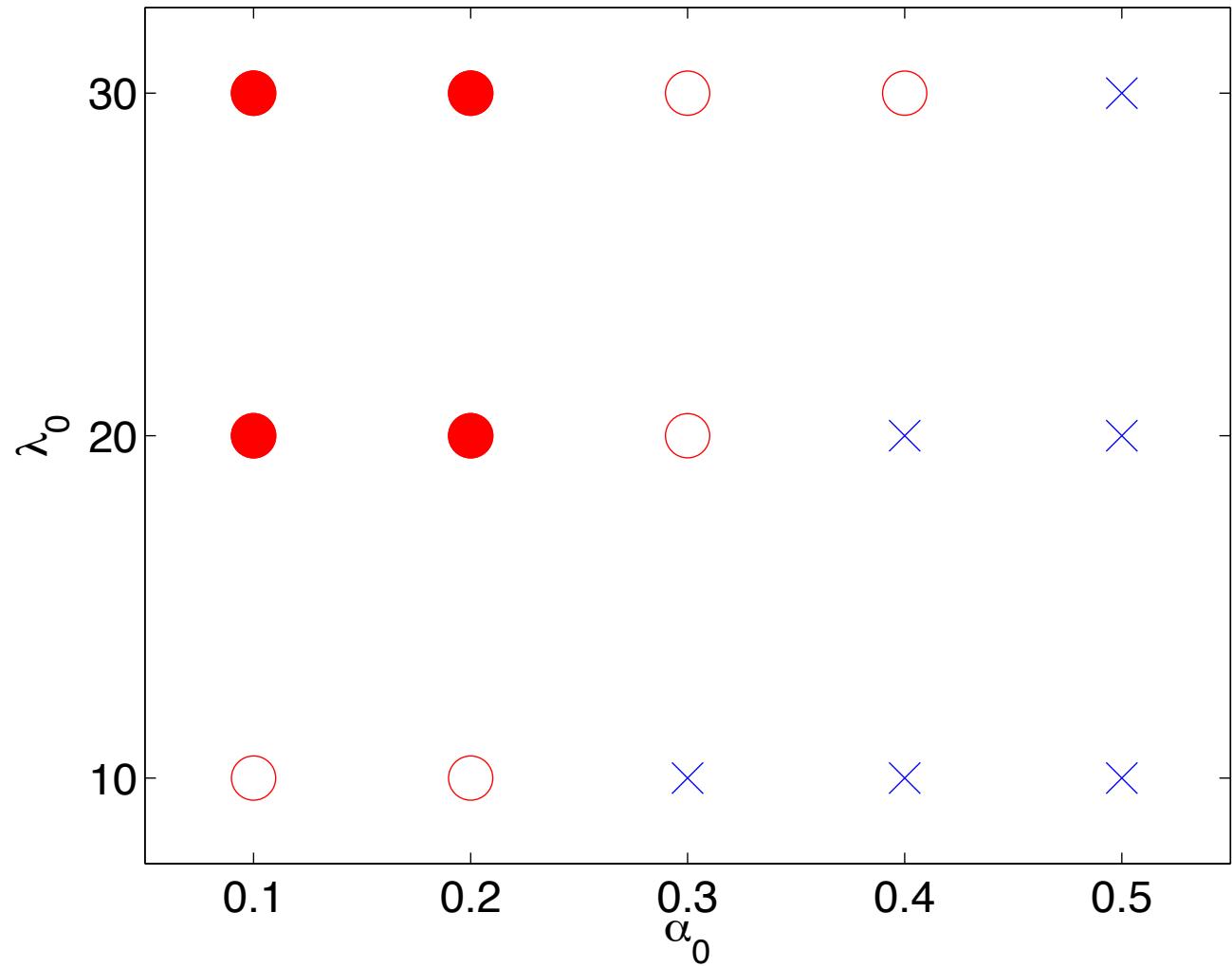
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Ice flow should trap dust in tropics.



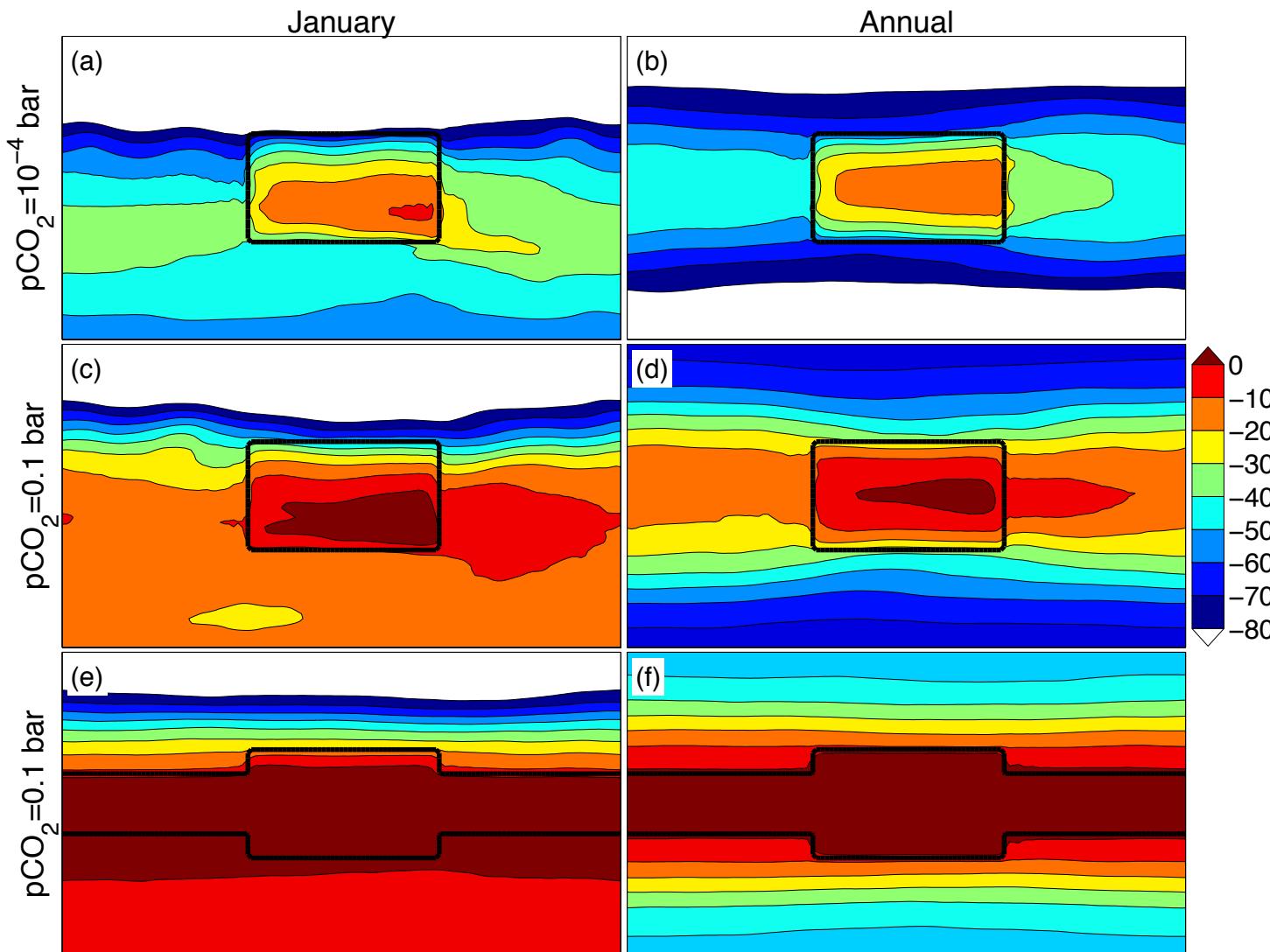
ABBOT AND PIERREHUMBERT: DUSTY SNOWBALL

X - 33



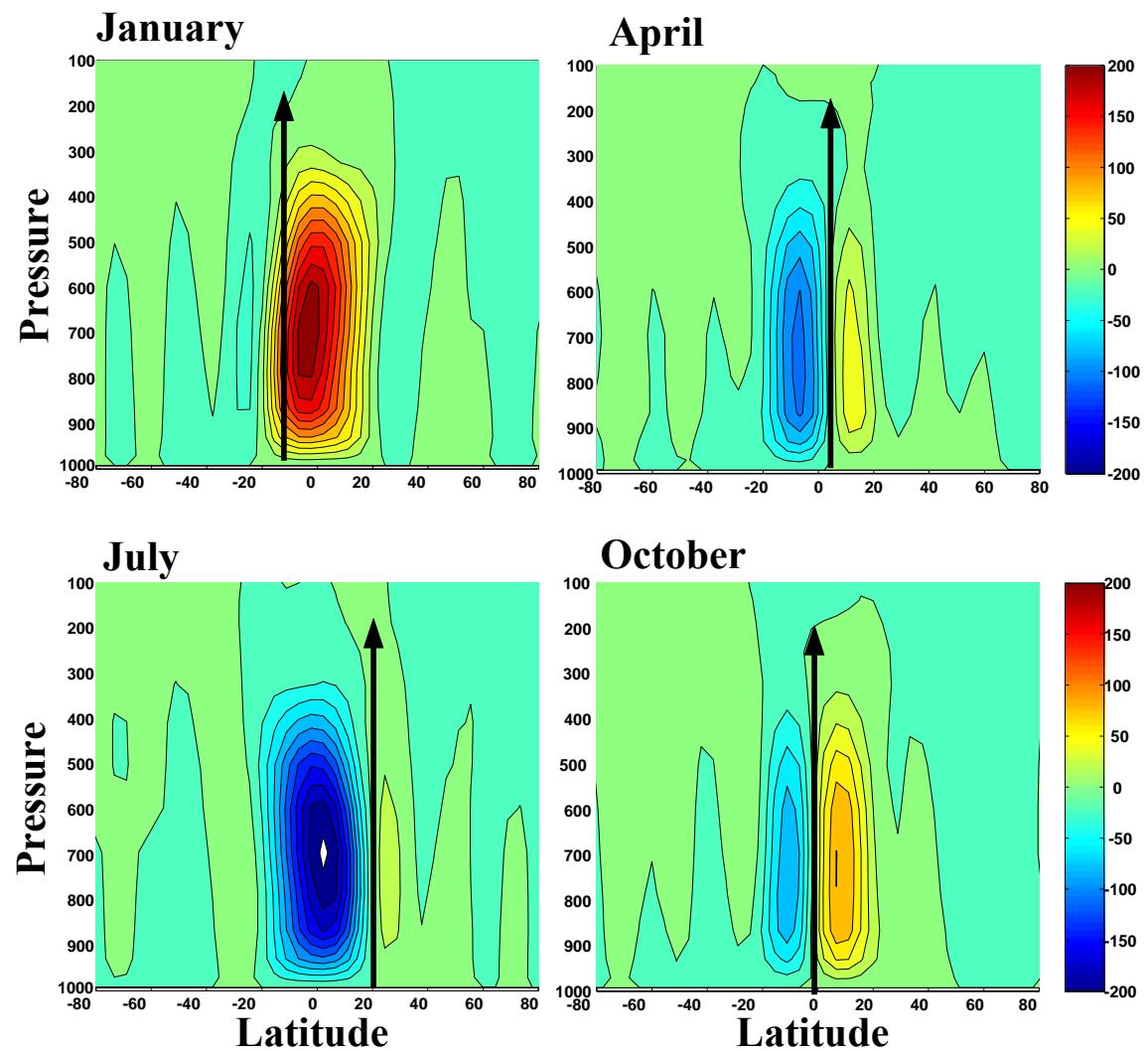
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ABBOT AND PIERREHUMBERT: DUSTY SNOWBALL



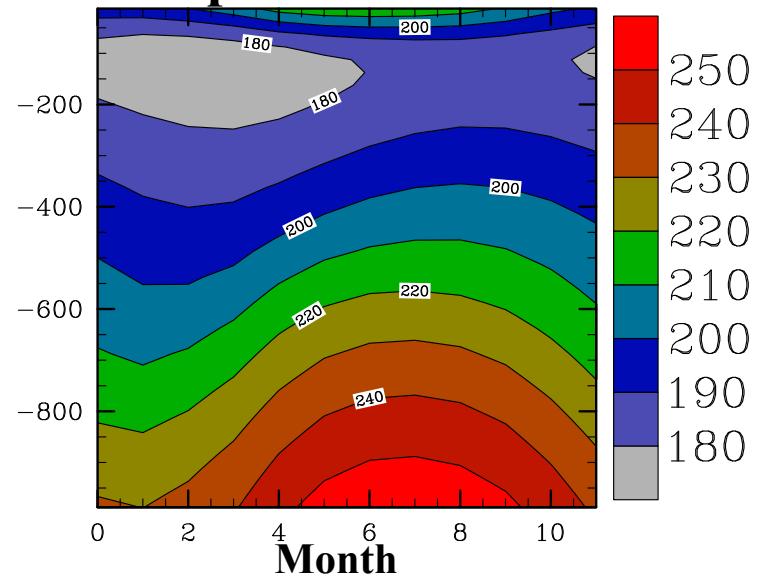
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Snowball Hadley Cell

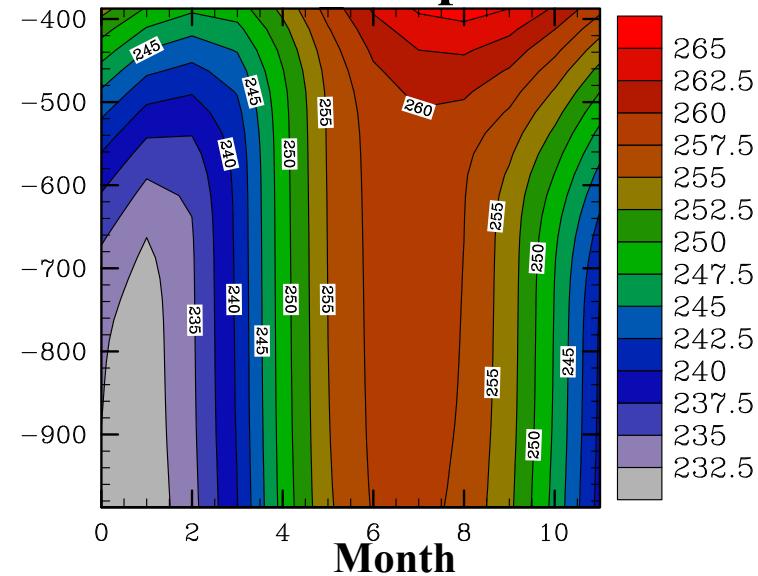


RC: Weak Winter lapse rate doesn't need dynamics

Temperature

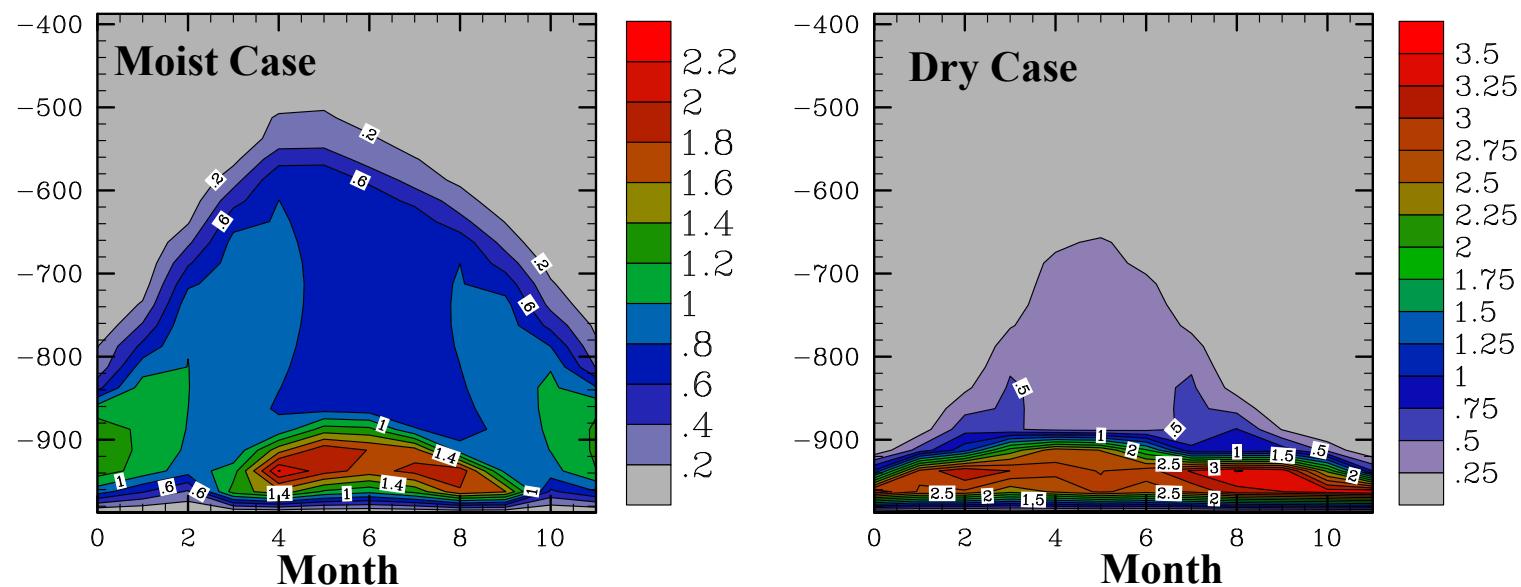


Potential Temp

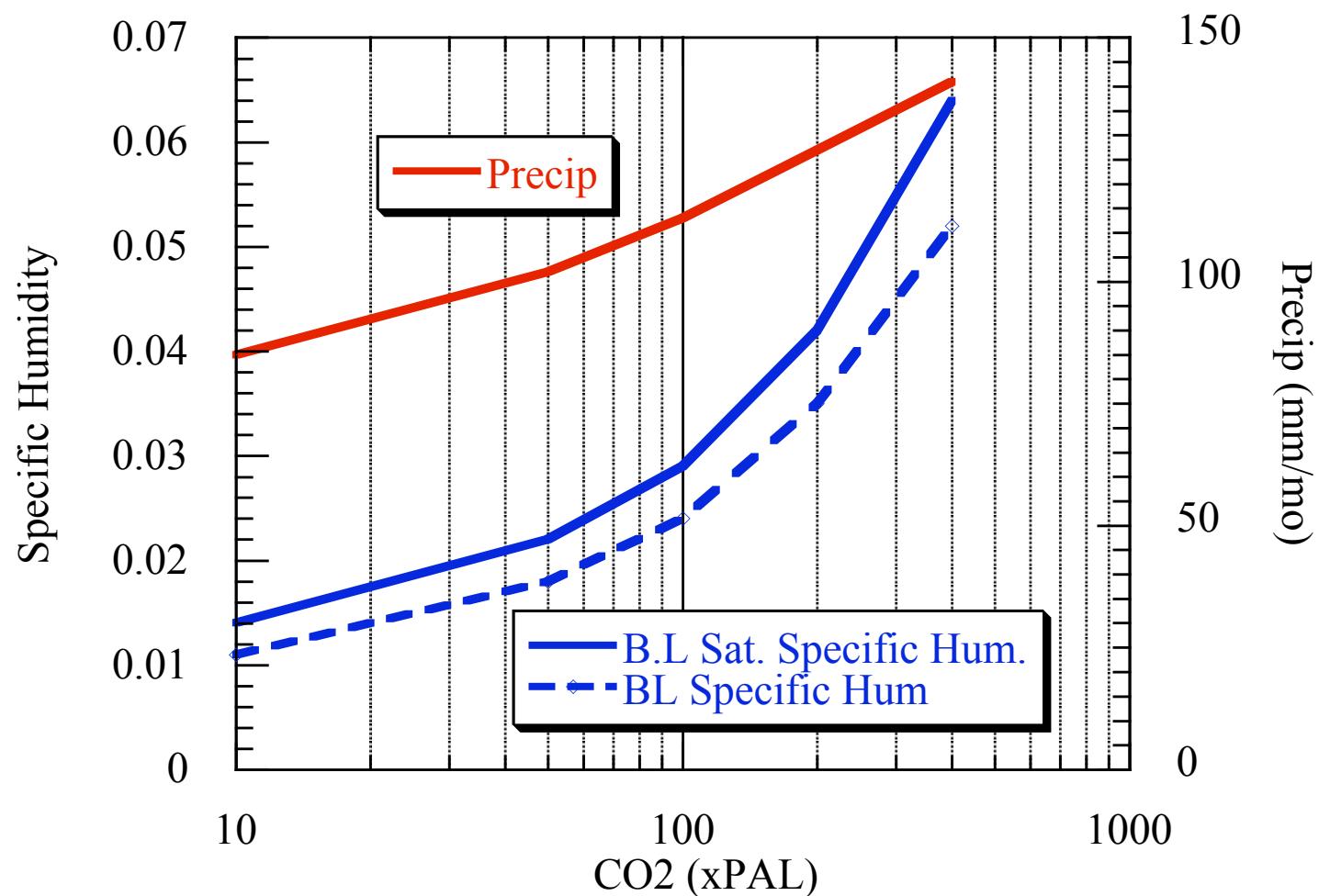


Convection is shallow and weak

Convective heating, K/da



Precipitation and temperature in the postglacial hothouse



Conclusions and prospects

- Critical question: Why Snowball glaciations at beginning and end of Proterozoic, but not any glaciation for a billion years in between? Similarly for ^{13}C fluctuations
- Is the Shuram anomaly real? Is it caused by the same things as the Shuram and Marinoan anomalies? If so, why no glaciation?
- With ocean heat transport, stable low-latitude sea ice margins are quite robust, though you can freeze over at low enough CO_2
- Very unlikely that there are waterbelt states cold enough to allow land glaciation, but need to do calculation with coupled ice sheets.
- Hard to reconcile waterbelt (nee slushball) states with geochemical signatures. Waterbelt (incomplete Snowball) state is implausible.
- It would be promising to explore EBM's based on MSE diffusion

- Albedo! Albedo! Albedo!
- Fate of surface transient melt water

- Dust moraine provides viable means for deglaciating Snowball
- Need to interactively couple sea-glacier and dust cycle to EBM. 3D sea glaciers
- Small-particle ice clouds could also help deglaciation, but cold convective cloud problem is untouched

- WTG model of the tide-locked waterworld reproduces the "swimming pool" state
- Extension to optically thick CO_2 atmospheres (easy)
- Extension to quasisynchronous eccentric orbit and 3:2 spin-orbit cases

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Conclusions and prospects

In order to use a GCM safely and productively, it is essential to have a very good theoretical toolkit