

Origin of Mass of the Higgs Boson

Christopher T. Hill
Fermilab

University of Toronto, April 28, 2015

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i.e., what is the origin of the Higgs Boson mass itself?

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We now know that a fundamental Higgs Boson exists and explains the masses of quarks, leptons, W and Z

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i.e., what is the origin of the Higgs Boson mass itself?

This is either very sobering, or it presents theoretical opportunities

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If true, it's a serious challenge to
our understanding of naturalness

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What is the custodial symmetry?

The world of masslessness
features a symmetry:

Scale Invariance

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Scale Invariance is (almost)
always broken by quantum
effects

Feynman Loops $\propto \hbar$

Scale Symmetry in QCD
is broken by quantum loops
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The Origin of the Nucleon Mass
(aka, most of the visible mass in
The Universe)

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$$\frac{dg}{d \ln \mu} = \beta(g)$$


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$$\beta(g) = \beta_0 g^3$$

where


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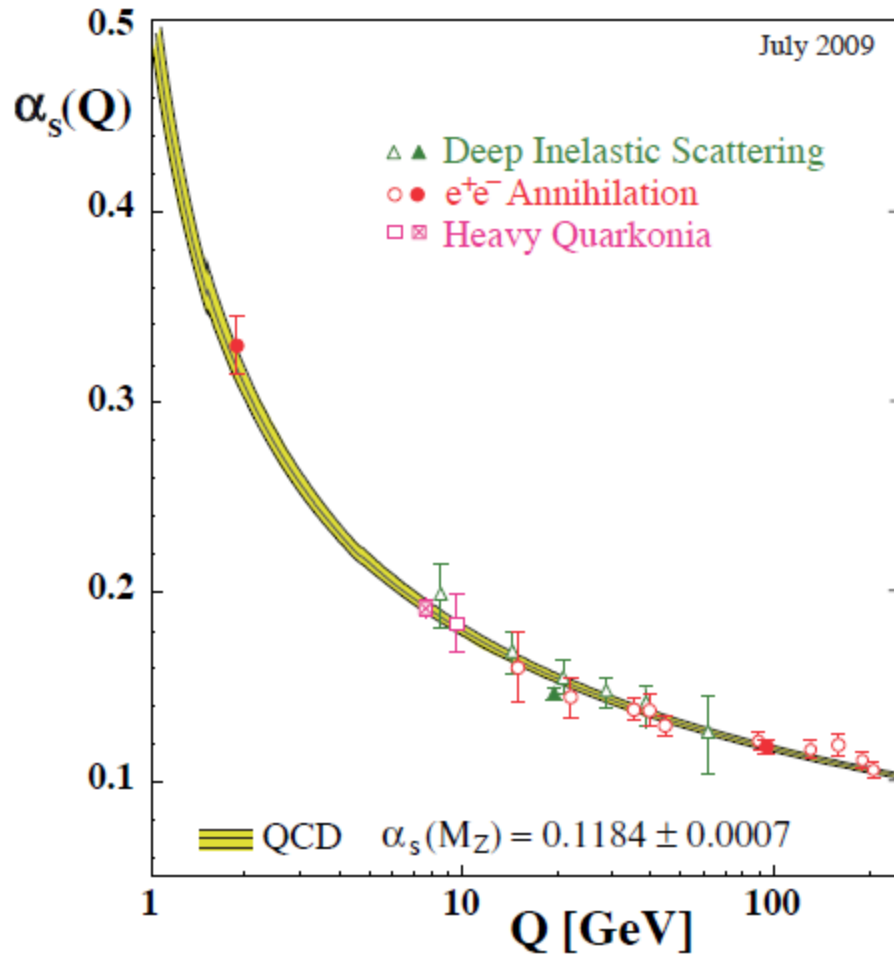
“running
coupling constant”

$$\alpha_s(k^2) \equiv \frac{g_s^2(k^2)}{4\pi} \approx \frac{1}{|\beta_0| \ln(k^2/\Lambda^2)}$$



$\Lambda = 200 \text{ MeV}$

S. Burby and C. Maxwell
arXiv:hep-ph/0011203



QCD running
coupling constant

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A Puzzle: Murray Gell-Mann lecture ca 1975

Noether current of
Scale symmetry

$$S_\mu = x^\nu T_{\mu\nu}$$

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Stress Tensor $T_{\mu\nu} = \text{Tr}(G_{\mu\rho}G^\rho_\nu) - \frac{1}{4}g_{\mu\nu} \text{Tr}(G_{\rho\sigma}G^{\rho\sigma})$

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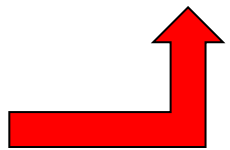
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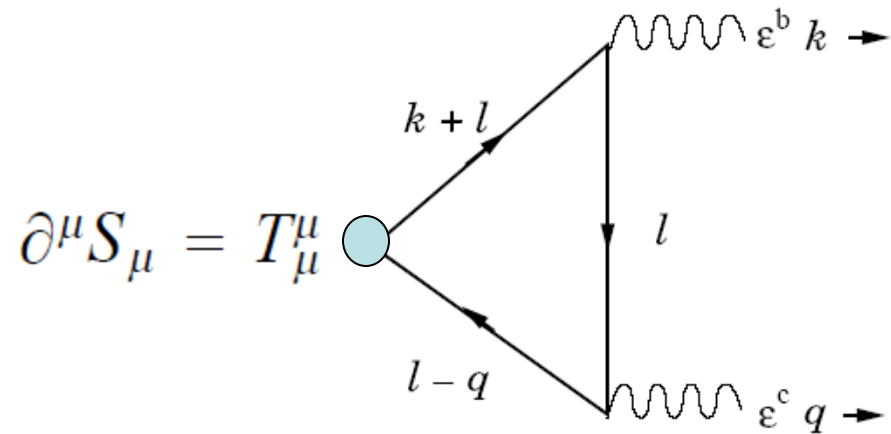
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QCD is scale invariant!!!!????



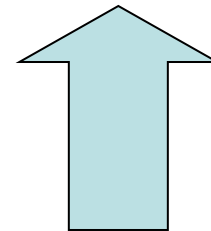
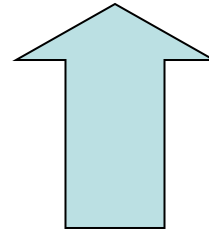
Resolution: The Scale Anomaly



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$$\partial_{\mu} S^{\mu} = \frac{\beta(g)}{g} \text{Tr} G_{\mu\nu} G^{\mu\nu} = \mathcal{O}(\hbar)$$

Origin of Mass in QCD
= Quantum Mechanics



See Murraypalooza talk:

Conjecture on the physical implications of the scale anomaly.

[Christopher T. Hill . hep-th/0510177](#)

't Hooft Naturalness:

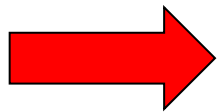
“Small ratios of physical parameters are controlled by symmetries. In the limit that a ratio goes to zero, there is enhanced symmetry (“custodial symmetry”).”

$$\frac{\Lambda}{M} = \exp\left(-\frac{8\pi^2}{|b_0|g^2(M)}\right) \quad b_0 \propto \hbar.$$

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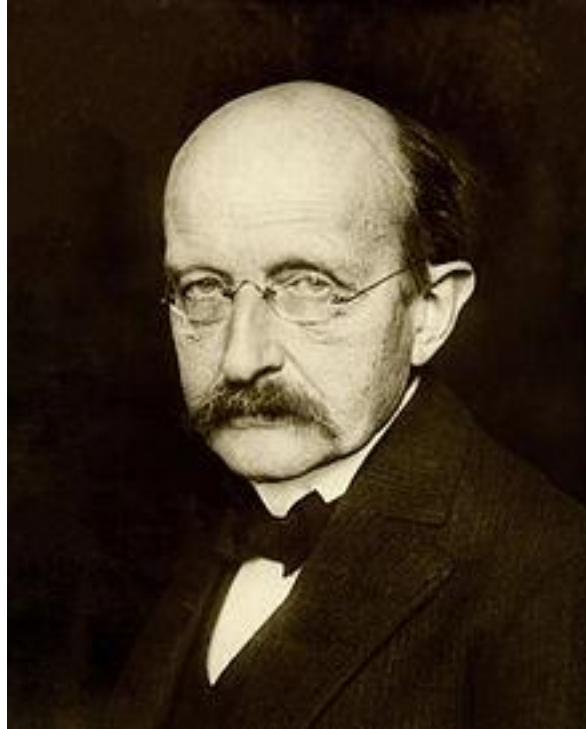
\hbar



0

Classical Scale Invariance
is the "Custodial Symmetry" of Λ_{QCD}

Conjecture:



Max Planck

All mass is a
quantum phenomenon.

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- (3) Extended Technicolor
- (4) Multiscale Technicolor
- (5) Walking Extended Technicolor
- (6) Topcolor Assisted Technicolor
- (7) Top Seesaw
- (8) Supersymmetric Walking Extended Technicolor
- (9) Strong dynamics from extra-dimensions
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Mass extinction of theories on July 4th 2012



Susy is still alive.



Susy is still alive?

If so, where is it?

How much fine tuning should
we tolerate?

Weak Scale SUSY was seriously challenged
before the LHC turned on (e.g. EDM's)

MSSM now copes with severe direct limits;
Some nMSSM models survive

If SUSY is the custodial symmetry
we may see it in LHC RUN-II

Why EDM's are so powerful:

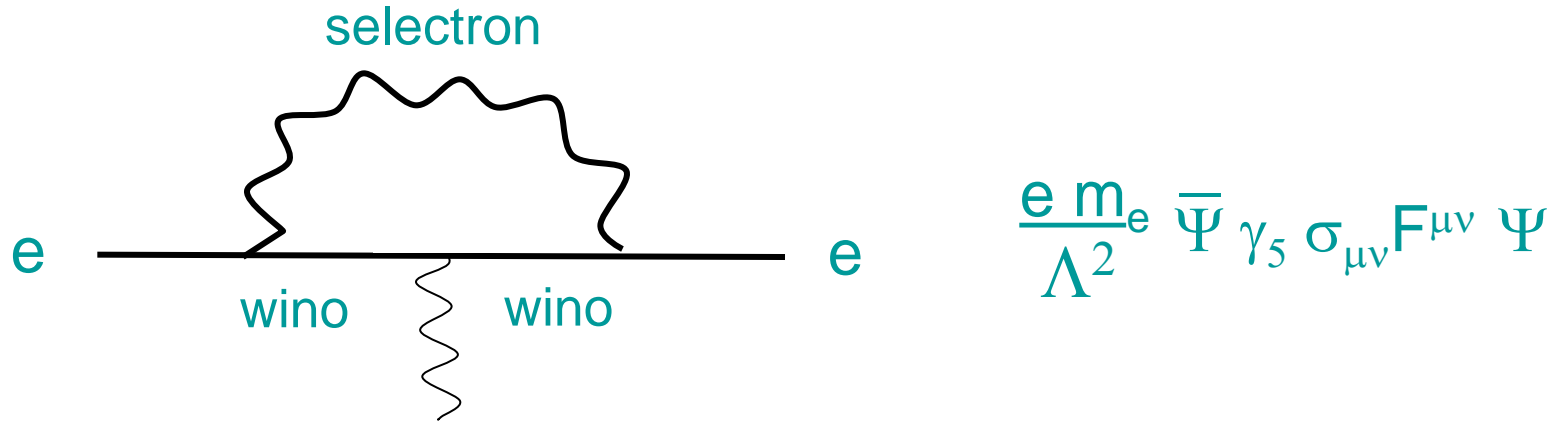
$$\frac{e m_e}{\Lambda^2} \bar{\Psi} \gamma_5 \sigma_{\mu\nu} F^{\mu\nu} \Psi$$

$$d_e = \frac{e m_e}{\Lambda^2} = 0.2 \times 10^{-16} \text{ (e-cm)} \times \frac{(m_e/\text{MeV})}{(\Lambda/\text{GeV})^2}$$

Current limit: $d_e < 10^{-27} \text{ e-cm}$

$$\Lambda > 1.4 \times 10^5 \text{ GeV}$$

Are EDM's telling us something about SUSY?:



$$1/(\Lambda)^2 = (\alpha \sin(\gamma) / 4\pi \sin^2 \theta) (1/M_{\text{selectron}})^2$$

$$M_{\text{selectron}} > 6.8 \times 10^3 \text{ GeV} (\sin(\gamma))^{1/2}$$

Future limit: $d_e < 10^{-29} \text{ e-cm} \text{ -- } 10^{-32} \text{ e-cm} ?$

Can a perturbatively
light Higgs Boson mass
come from quantum mechanics?

Bardeen: Classical Scale Invariance
could be the custodial symmetry
of a fundamental, perturbatively
light Higgs Boson in $SU(3) \times SU(2) \times U(1)$.

The only manifestations of
Classical Scale Invariance breaking by
quantum loops are $d = 4$ scale anomalies.

On naturalness in the standard model.

William A. Bardeen

FERMILAB-CONF-95-391-T, Aug 1995. 5pp.

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quantum loops are $d = 4$ scale anomalies.
There is no meaning to the "quadratic divergence"
as a source of scale breaking.

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There are possible additive effects from higher mass scales: $\delta m_H^2 = \alpha^p M_{GUT}^2 + \alpha^q M_{Planck}^2$.

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But the existence of the low mass Higgs may be telling us that such effects are absent!

Something seems to be missing in our understanding of scale symmetry and fine-tuning.

This is profoundly important in sculpting our view of the physical world!

Treat this as a
phenomenological question:

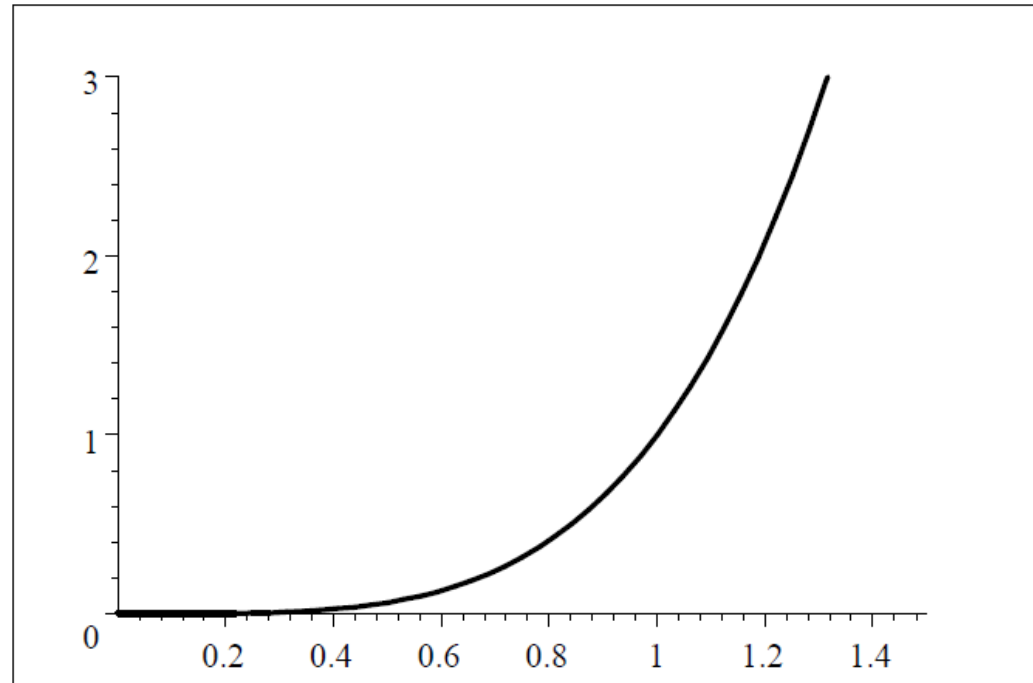
Is the Higgs Potential Generated by
Infra-red (scale-breaking)
Quantum Loop Effects?

i.e., is the Higgs potential a
Coleman-Weinberg Potential?

Start with the Classically Scale Invariant Higgs Potential

$$\frac{\lambda}{2} |H|^4$$

$$\langle H \rangle = v$$

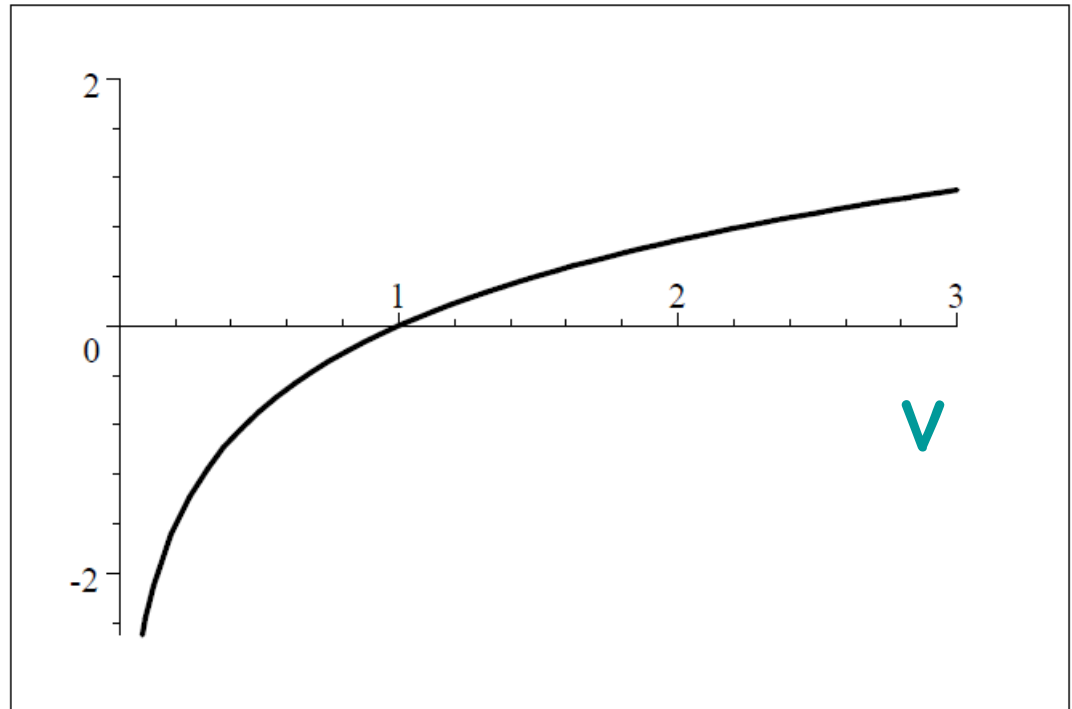


V

Scale Invariance \rightarrow Quartic Potential \rightarrow No VEV

Quantum loops generate a logarithmic "running" of the quartic coupling

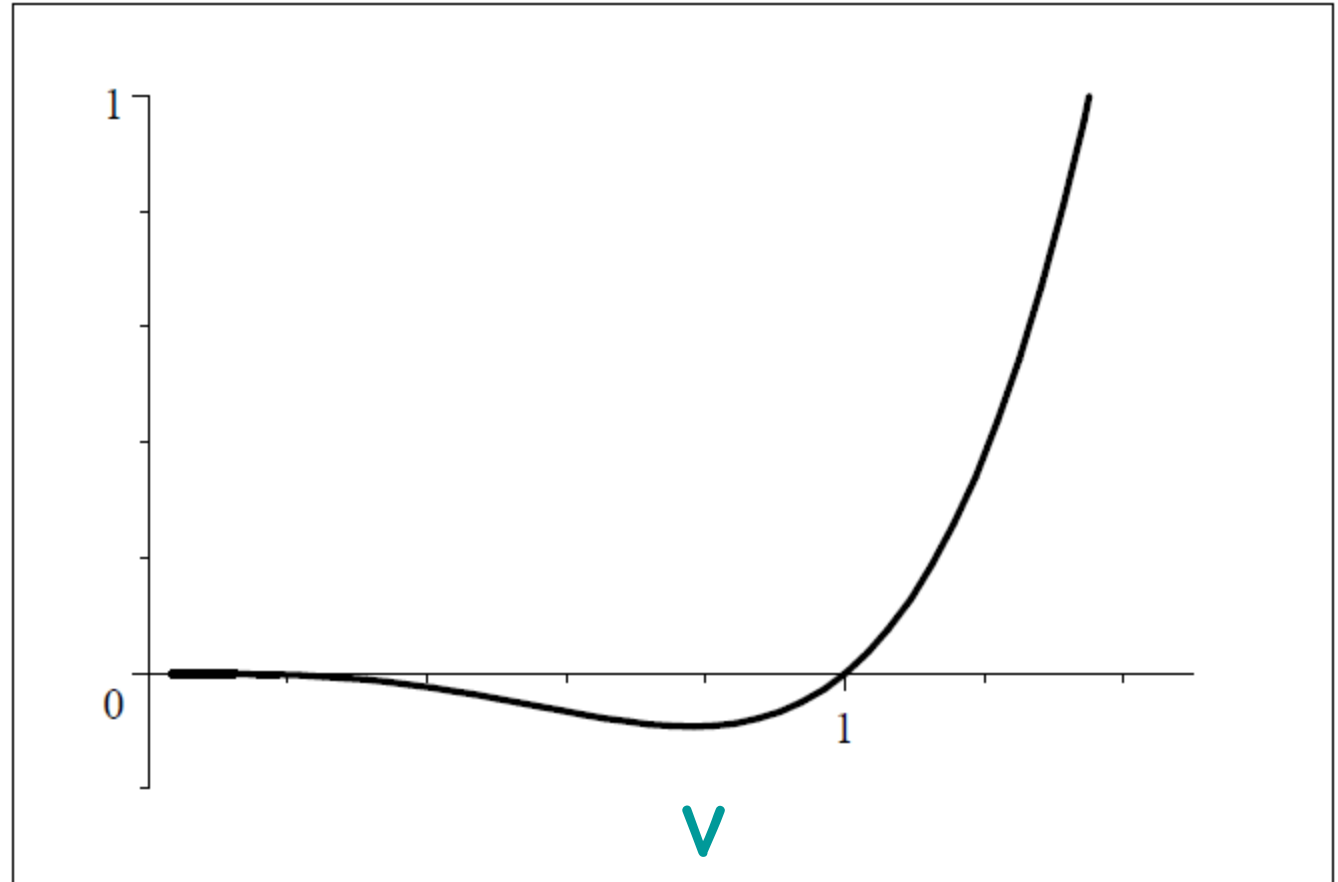
$$\lambda(v) \propto \hbar \beta \log(v/M)$$



Nature chooses a particular trajectory determined by dimensionless cc's.

Result: "Coleman-Weinberg Potential"

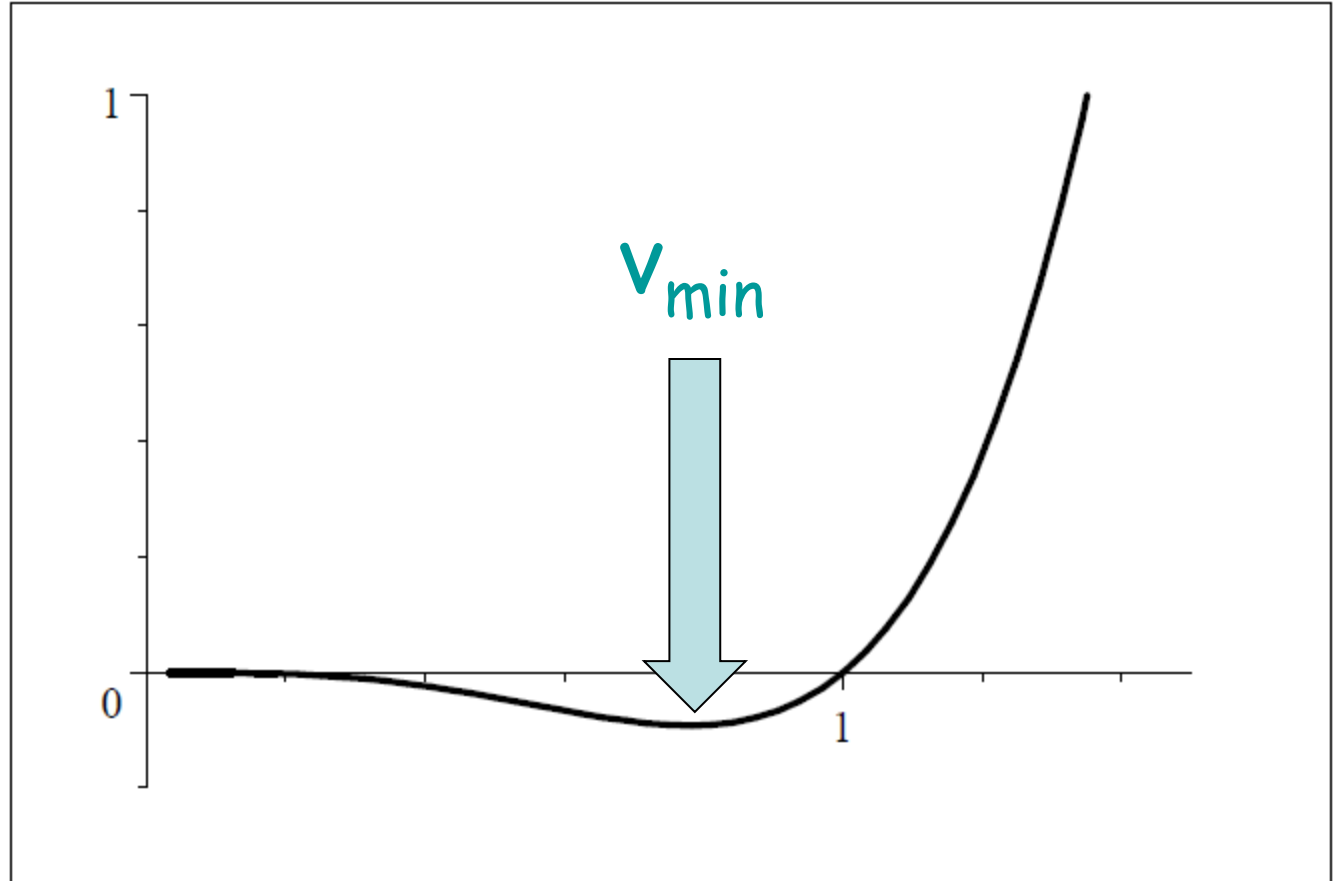
$$\frac{\tilde{\lambda}(v)}{2} \times v^4$$



Potential Minimum arises from running of λ
i.e. Quantum Mechanics

Result: "Coleman-Weinberg Potential

$$\frac{\tilde{\lambda}(v)}{2} \times v^4$$



Require: $\beta > 0$ $\lambda < 0$ for a minimum, positive curvature

Boson Mass is determined by curvature at minimum

An Improved Coleman-Weinberg Potential

$$S = \int d^4x \mathcal{L} = \int d^4x \left(\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right)$$

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Improved Stress tensor:
Callan, Coleman, Jackiw

$$\tilde{T}_{\mu\nu} = T_{\mu\nu} + Q_{\mu\nu}$$

$$= \frac{2}{3} \partial_\mu \phi \partial_\nu \phi - \frac{1}{6} \eta_{\mu\nu} \partial_\rho \phi \partial^\rho \phi - \frac{1}{3} \phi \partial_\mu \partial_\nu \phi + \frac{1}{3} \eta_{\mu\nu} \phi \partial^2 \phi + \eta_{\mu\nu} V(\phi)$$

Trace of improved stress tensor:

$$\tilde{T}^\mu_\mu = \phi \partial^2 \phi + 4V(\phi) = -\phi \frac{\delta}{\delta \phi} V(\phi) + 4V(\phi)$$

Traceless for a classical scale invariant theory:

$$V(\phi) = \frac{\lambda}{4} \phi^4, \quad \longrightarrow \quad \tilde{T}^\mu_\mu = 0 \quad \text{Conserved scale current}$$

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Running coupling constant: $\frac{\delta}{\delta \phi} \lambda(\phi) = \beta(\lambda)$

$$\longrightarrow \quad \tilde{T}_\mu^\mu = -\frac{\beta(\lambda)}{\lambda} V(\phi)$$

Trace Anomaly associated with running coupling

Improved Coleman-Weinberg Potential is the solution to the equations:

$$\tilde{T}_\mu^\mu = -\frac{\beta(\lambda)}{\lambda}V(\phi) \Rightarrow$$

$$\phi \frac{\delta}{\delta \phi} V(\phi) - 4V(\phi) = \frac{\beta}{\lambda} V(\phi)$$

$$\frac{d\lambda(\mu)}{d \ln \mu} = \beta(\lambda)$$

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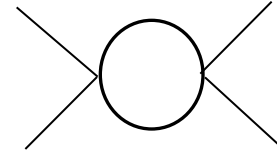
$$\phi \frac{\delta}{\delta \phi} V(\phi) - 4V(\phi) = \frac{\beta}{\lambda} V(\phi)$$

$$\frac{d\lambda(\mu)}{d \ln \mu} = \beta(\lambda)$$

The solution is: $V(\phi) = \frac{1}{2} \lambda(\phi) \phi^4$

Example: ϕ^4 Field theory

$$\frac{d\lambda}{d\ln(\phi)} = \beta(\lambda) = \frac{9\lambda^2}{32\pi^2}$$

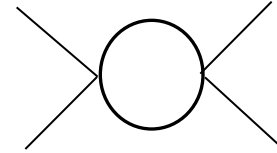


$$V_{RG} = \frac{\lambda}{4}\phi^4 + \hbar\frac{9\lambda^2}{32\pi^2}\phi^4\ln(\phi/M) = \hbar\frac{m_h^4}{32\pi^2}(\phi/v)^4\ln(\phi/\tilde{M})$$

agrees with CW original result \log (path Integral)

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Example: Scalar Electrodynamics

$$V(\phi) = \frac{\lambda_0}{2} |\phi|^4 + \frac{1}{16\pi^2} (5\lambda^2 - 6\lambda e^2 + 6e^4) |\phi|^4 \ln\left(\frac{|\phi|}{M}\right)$$

$$\phi_c^2 = 2|\phi|^2 \quad \text{and} \quad \frac{\lambda_{CW}}{4!} \phi_c^4 = \frac{\lambda_0}{2} |\phi|^4$$

$$V(\phi'_c) = \frac{\lambda_{CW}}{4!} \phi_c'^4 + \left(\frac{5\lambda_{CW}^2}{1152\pi^2} + \frac{3e^4}{64\pi^2} \right) \phi_c'^4 \ln\left(\frac{\phi_c'^2}{M'^2}\right)$$

(C.6)

agrees with CW original result BUT with canonical normalization

The Renormalization Group generates the Coleman Weinberg potential

Expand about minimum:

$$V_{CW}(h) = \frac{1}{2}\lambda(v + h/\sqrt{2})(v + h/\sqrt{2})^4$$

$$\left. \frac{dV}{dh} \right|_{h=0} = \sqrt{2}v^3 \left(\lambda + \frac{1}{4}\beta \right) = 0$$

$$\beta_1(\lambda_i(v)) = -4\lambda_1(v) \quad \text{at minimum}$$

$$\frac{d^2V}{dh^2} = m_h^2 = \left(3\lambda + \frac{7}{4}\beta \right) v^2$$

$$m_h^2 = -4\lambda v^2 = \beta v^2 > 0$$

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We label all relevant coupling constants that enter in any order of the loop diagrams for the running of λ (*e.g.*, g_{top} , g_2 , g_{QCD} , etc.) as λ_i . We denote the scalar quartic (Higgs) coupling as $\lambda \equiv \lambda_1$ with β -function $\beta_1(\lambda_i)$. Each λ_i has its own β_i :

$$\frac{d\lambda_i}{d \ln(\mu)} = \beta_i(\lambda_j)$$

$$v\lambda_1'(v) = \beta_1$$

$$v^2\lambda_1''(v) = \beta_j \frac{\partial \beta_1}{\partial \lambda_j} - \beta_1$$

$$v^3\lambda_1'''(v) = \beta_i\beta_j \frac{\partial^2 \beta_1}{\partial \lambda_i \partial \lambda_j} + \beta_j \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta_1}{\partial \lambda_i} - 3\beta_j \frac{\partial \beta_i}{\partial \lambda_i} + 2\beta_1$$

$$v^4\lambda_1''''(v) = \beta_i\beta_j\beta_k \frac{\partial^3 \beta_1}{\partial \lambda_i \partial \lambda_j \partial \lambda_k} + \beta_k \frac{\partial \beta_j}{\partial \lambda_k} \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta_1}{\partial \lambda_i} + \beta_k\beta_j \frac{\partial^2 \beta_i}{\partial \lambda_j \partial \lambda_k} \frac{\partial \beta_1}{\partial \lambda_i} + 3\beta_k\beta_j \frac{\partial \beta_i}{\partial \lambda_k} \frac{\partial^2 \beta_1}{\partial \lambda_i \partial \lambda_j} - 6\beta_i\beta_j \frac{\partial^2 \beta_1}{\partial \lambda_i \partial \lambda_j} - 6\beta_k \frac{\partial \beta_j}{\partial \lambda_k} \frac{\partial \beta_1}{\partial \lambda_j} + 11\beta_i \frac{\partial \beta_1}{\partial \lambda_i} - 6\beta_1 \quad (21)$$

The Coleman-Weinberg Potential is completely determined by β -functions:

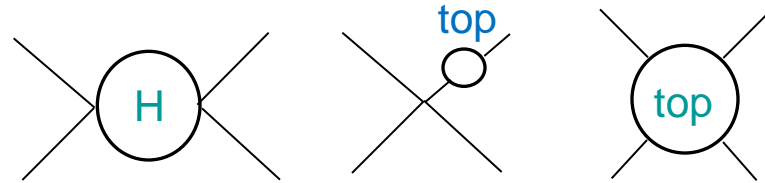
$$\begin{aligned}
 V_{CW}(h) = & -\frac{1}{8}\beta_1 v^4 + \frac{1}{2}v^2 h^2 \left(\beta_1 + \frac{1}{4}\beta_j \frac{\partial \beta_1}{\partial \lambda_j} \right) \\
 & + \frac{5}{6\sqrt{2}} v h^3 \left(\beta_1 + \frac{9}{20}\beta_i \frac{\partial \beta_1}{\partial \lambda_i} + \frac{1}{20}\beta_j \beta_i \frac{\partial^2 \beta_1}{\partial \lambda_j \partial \lambda_i} \right. \\
 & \quad \left. + \frac{1}{20}\beta_j \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta_1}{\partial \lambda_i} \right) \\
 & + \frac{11}{48} h^4 \left(\beta_1 + \frac{35}{44}\beta_i \frac{\partial \beta_1}{\partial \lambda_i} + \frac{5}{22}\beta_j \beta_i \frac{d^2 \beta_1}{\partial \lambda_j \partial \lambda_i} \right. \\
 & \quad + \frac{5}{22}\beta_j \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta_1}{\partial \lambda_i} + \frac{1}{44}\beta_k \beta_j \beta_i \frac{d^3 \beta_1}{\partial \lambda_k \partial \lambda_j \partial \lambda_i} \\
 & \quad + \frac{1}{44}\beta_k \frac{\partial \beta_j}{\partial \lambda_k} \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta_1}{\partial \lambda_i} + \frac{1}{44}\beta_j \beta_i \frac{d^2 \beta_i}{\partial \lambda_j \partial \lambda_i} \frac{\partial \beta_1}{\partial \lambda_i} \\
 & \quad \left. + \frac{3}{44}\beta_j \beta_k \frac{\partial \beta_i}{\partial \lambda_k} \frac{d^2 \beta_1}{\partial \lambda_j \partial \lambda_i} \right) \\
 & + \frac{h^5}{40\sqrt{2}v} \left(\beta + \frac{25}{12}\beta_i \frac{d\beta}{d\lambda_i} + \frac{35}{24}\beta_j \beta_i \frac{d^2 \beta}{d\lambda_j d\lambda_i} \right. \\
 & \quad + \frac{35}{24}\beta_j \frac{d\beta_i}{d\lambda_j} \frac{d\beta}{d\lambda_i} + \frac{5}{12}\beta_k \beta_j \beta_i \frac{d^3 \beta}{d\lambda_k d\lambda_j d\lambda_i} \\
 & \quad + \frac{5}{12}\beta_k \frac{d\beta_j}{d\lambda_k} \frac{d\beta_i}{d\lambda_j} \frac{d\beta}{d\lambda_i} + \frac{5}{12}\beta_j \beta_i \frac{d^2 \beta_i}{d\lambda_j d\lambda_i} \frac{d\beta}{d\lambda_i} \\
 & \quad + \frac{5}{4}\beta_j \beta_k \frac{d\beta_i}{d\lambda_k} \frac{d^2 \beta}{d\lambda_j d\lambda_i} + \frac{1}{24}\beta_i \beta_j \beta_k \beta_\ell \frac{\partial^4 \beta}{\partial \lambda_i \partial \lambda_j \partial \lambda_k \partial \lambda_\ell} \\
 & \quad + \frac{1}{24}\beta_\ell \frac{\partial \beta_k}{\partial \lambda_\ell} \frac{\partial \beta_j}{\partial \lambda_k} \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta}{\partial \lambda_i} + \frac{1}{4}\beta_i \beta_j \beta_k \frac{\partial \beta_\ell}{\partial \lambda_k} \frac{\partial^3 \beta}{\partial \lambda_i \partial \lambda_j \partial \lambda_\ell} \\
 & \quad + \frac{1}{8}\beta_\ell \beta_k \frac{\partial \beta_j}{\partial \lambda_k} \frac{\partial^2 \beta_i}{\partial \lambda_j \partial \lambda_\ell} \frac{\partial \beta}{\partial \lambda_i} + \frac{1}{6}\beta_i \beta_\ell \frac{\partial \beta_k}{\partial \lambda_\ell} \frac{\partial \beta_j}{\partial \lambda_k} \frac{\partial^2 \beta}{\partial \lambda_j \partial \lambda_i} \\
 & \quad + \frac{1}{6}\beta_i \beta_j \beta_\ell \frac{\partial^2 \beta_k}{\partial \lambda_\ell \partial \lambda_j} \frac{\partial^2 \beta}{\partial \lambda_k \partial \lambda_i} + \frac{1}{24}\beta_\ell \beta_k \beta_j \frac{\partial^3 \beta_i}{\partial \lambda_j \partial \lambda_k \partial \lambda_\ell} \frac{\partial \beta}{\partial \lambda_i} \\
 & \quad \left. + \frac{1}{8}\beta_\ell \beta_i \frac{\partial \beta_k}{\partial \lambda_\ell} \frac{\partial^2 \beta}{\partial \lambda_k \partial \lambda_j} \frac{\partial \beta_j}{\partial \lambda_i} + \frac{1}{24}\beta_\ell \beta_k \frac{\partial^2 \beta_j}{\partial \lambda_\ell \lambda_k} \frac{\partial \beta_i}{\partial \lambda_j} \frac{\partial \beta}{\partial \lambda_i} \right) \\
 & \quad \times (h^6).
 \end{aligned}$$

Higgs Quartic coupling $\beta(\lambda)$

Renormalization Group Equation:

$$\frac{d\lambda(v)}{d\ln(v)} = \frac{12}{16\pi^2} (\lambda^2 + \lambda g^2 - g^4) = \beta$$

g : top Yukawa cc



(I am ignoring EW contributions
for simplicity of discussion)

Higgs Quartic coupling $\beta(\lambda)$

Renormalization Group Equation:

$$\frac{d\lambda(v)}{d\ln(v)} = \frac{12}{16\pi^2} (\lambda^2 + \lambda g^2 - g^4) = \beta$$

approximate physical values:

$$\left. \begin{array}{l} \text{Higgs quartic cc: } \lambda = 1/4 \\ \text{Top Yukawa cc: } g = 1 \end{array} \right\} \longrightarrow \beta = -5.2244 \times 10^{-2}$$

β is small and negative in standard model
No solution !

Higgs Quartic coupling $\beta(\lambda)$

Renormalization Group Equation:

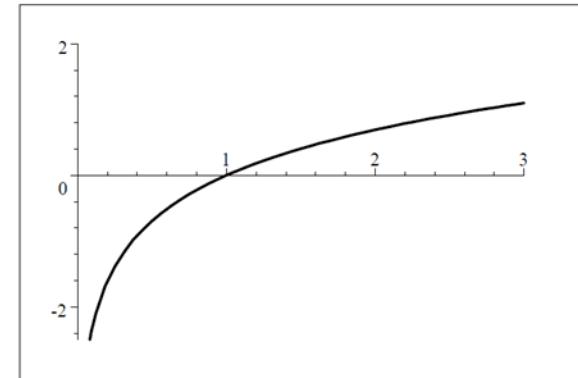
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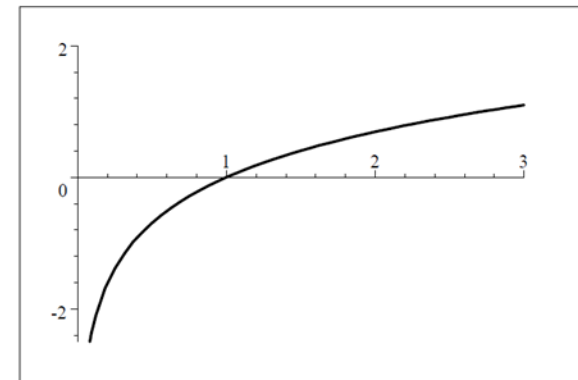
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β is small and negative in standard model

We require positive β to have a Coleman-Weinberg potential

Requires New Bosonic physics beyond the standard model



Higgs Quartic coupling $\beta(\lambda)$

Introduce a new field: S

Higgs-Portal Interaction $\lambda' |H|^2 |S|^2$

Higgs Quartic coupling $\beta(\lambda)$

Introduce a new field: S

Higgs-Portal Interaction $\lambda' |H|^2 |S|^2$

Two possibilities:

(1) Modifies RG equation to make $\beta > 0$:

$$\frac{d\lambda(v)}{d\ln(v)} = \frac{12}{16\pi^2} (\lambda^2 + \lambda g^2 - g^4 + c \lambda'^2)$$

(2) S develops its own CW potential, and VEV $\langle S \rangle = V'$ and Higgs gets mass, $\lambda' V'$

Simplest hypotheses:

S may be:

(1) A singlet field with or without VEV

e.g., Ultra-weak sector, Higgs boson mass, and the dilaton
Kyle Allison, Christopher T. Hill, Graham G. Ross. : [arXiv:1404.6268](#) [hep-ph]

(2) A new doublet NOT coupled to
 $SU(2) \times U(1)$ (inert) w or wo VEV

Hambye and Strumia Phys.Rev. D88 (2013) 055022

S sector is
Dark Matter

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Hambye and Strumia Phys.Rev. D88 (2013) 055022

(3) New doublet COUPLED to
 $SU(2) \times U(1)$ with no VEV (dormant)

Is the Higgs Boson Associated with
Coleman-Weinberg Dynamical Symmetry Breaking?
CTH, [arXiv:1401.4185](#) [hep-ph]. [Phys Rev D.89.073003](#).

S sector is
Dark Matter

S sector is
visible

Simplest hypotheses

S may be:

A new doublet NOT coupled to $SU(2) \times U(1)$ (inert) w or wo VEV

e.g., Hambye and Strumia [Phys.Rev. D88 \(2013\) 055022](#);
[S. Iso, and Y. Orikasa, PTEP \(2013\) 023B08](#); ...
“Ultra-weak sector, Higgs boson mass, and the dilaton,”
[K. Allison, C. T. Hill, G. G. Ross. arXiv:1404.6268 \[hep-ph\]](#)
[Light Dark Matter, Naturalness, and the Radiative Origin of the Electroweak Scale, W. Altmannshofer, W. Bardeen, M Bauer, M. Carena, J. Lykken e-Print: arXiv:1408.3429 \[hep-ph\]](#) ...

Many, many papers on this approach!

A New doublet COUPLED to $SU(2) \times U(1)$ with no VEV (dormant)

e.g., Is the Higgs Boson Associated with Coleman-Weinberg Dynamical Symmetry Breaking?
CTH, [arXiv:1401.4185 \[hep-ph\]](#). [Phys Rev D.89.073003](#)....

S sector is
Dark Matter

S sector is
visible at
LHC

Massless
two doublet
potential

$$V(H_1, H_2) = \frac{\lambda_1}{2}|H_1|^4 + \frac{\lambda_2}{2}|H_2|^4 + \lambda_3|H_1|^2|H_2|^2 + \lambda_4|H_1^\dagger H_2|^2 + \frac{\lambda_5}{2} \left[(H_1^\dagger H_2)^2 e^{i\theta} + h.c. \right]$$

$$16\pi^2 \frac{d\lambda_1(\mu)}{d\ln(\mu)} = 12\lambda_1^2 + 4\lambda_3^2 + 4\lambda_3\lambda_4 + 2\lambda_4^2 + 2\lambda_5^2 - 3\lambda_1(3g_2^2 + g_1^2) + \frac{3}{2}g_2^4 + \frac{3}{4}(g_1^2 + g_2^2)^2 + 12\lambda_1 g_t^2 - 12g_t^4$$

$$16\pi^2 \frac{d\lambda_2(\mu)}{d\ln(\mu)} = 12\lambda_2^2 + 4\lambda_3^2 + 4\lambda_3\lambda_4 + 2\lambda_4^2 + 2\lambda_5^2 - 3\lambda_2(3g_2^2 + g_1^2) + \frac{3}{2}g_2^4 + \frac{3}{4}(g_1^2 + g_2^2)^2 + 12\lambda_2 g_b^2 - 12g_b^4$$

$$16\pi^2 \frac{d\lambda_3(\mu)}{d\ln(\mu)} = (\lambda_1 + \lambda_2)(6\lambda_3 + 2\lambda_4) + 4\lambda_3^2 + 2\lambda_4^2 + 2\lambda_5^2 - 3\lambda_3(3g_2^2 + g_1^2) + \frac{9}{4}g_2^4 + \frac{3}{4}g_1^4 - \frac{3}{2}g_1^2 g_2^2 + 6\lambda_3(g_t^2 + g_b^2) - 12g_t^2 g_b^2$$

$$16\pi^2 \frac{d\lambda_4(\mu)}{d\ln(\mu)} = 2(\lambda_1 + \lambda_2)\lambda_4 + 4(2\lambda_3 + \lambda_4)\lambda_4 + 8\lambda_5^2 - 3\lambda_4(3g_2^2 + g_1^2) + 3g_1^2 g_2^2 - 12g_t^2 g_b^2$$

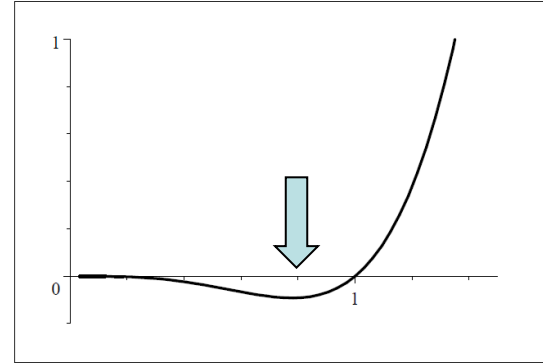
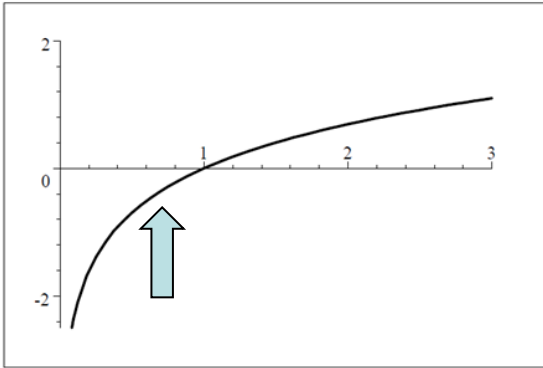
$$16\pi^2 \frac{d\lambda_5(\mu)}{d\ln(\mu)} = \lambda_5 [2(\lambda_1 + \lambda_2) + 8\lambda_3 + 12\lambda_4 - 3(3g_2^2 + g_1^2) + 2(g_t^2 + g_b^2)]$$

Two doublet
RG equations

The observed Higgs boson mass implies:

$$m_h^2 = -4\lambda v^2 = \beta v^2 > 0 \quad \Rightarrow \quad \beta = \left(\frac{126}{175}\right)^2 = 0.5184$$

Note that λ is negative: $\lambda = -(0.25)(0.5184) = -0.1296$



Can now solve for λ_3 :
$$\beta = \frac{1}{16\pi^2} (12\lambda^2 + 12\lambda g^2 - 12g^4 + 4\lambda_3^2)$$

$$g = g_{\text{top}} \approx 1$$

Solution is: $\lambda_3 = 4.8789$

Mass of New Doublet: $\sqrt{4.8789} \times (175) = 386.54 \text{ GeV}$

M^2 is determined \longrightarrow heavy "dormant" Higgs doublet

No VeV but coupled to $SU(2) \times U(1)$:

"Dormant" Higgs Doublet (vs. "Inert")

Production, mass, and decay details are model dependent

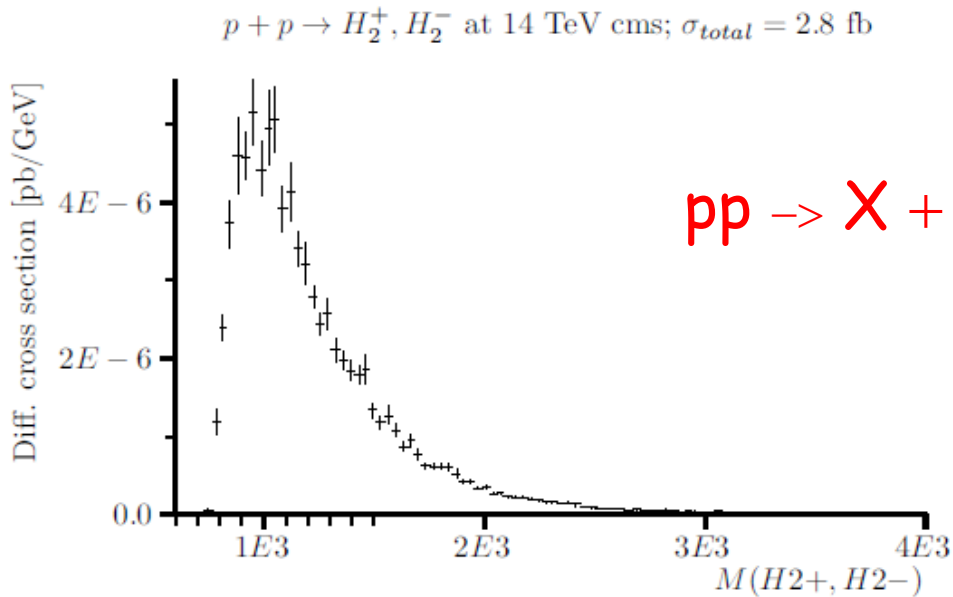
If Dormant Higgs couples to $SU(2) \times U(1)$ but not fermions

Parity $H_2 \rightarrow -H_2$ implies stability: $H_2^+ \rightarrow H_2^0 + (e^+\nu)$ if $M^+ > M^0$

Then H_2^0 is stable dark matter WIMP

The Dormant Doublet is pair produced above threshold near $2M_H \approx 800 \text{ GeV}$

CalcHEP estimates



$pp \rightarrow X + (\gamma^*, Z^*, W^*, h^*) \rightarrow X + H H^*$

FIG. 1: H^+H^- production at LHC.

$pp \rightarrow H^0 H^0$

$\sigma = 1.4 \text{ fb}$

$\Gamma_{H^0 \rightarrow bb} = 45 \text{ GeV}$

Assume $g_b' = 1$

$pp \rightarrow H^+ H^-$

$\sigma = 2.8 \text{ fb}$

$\Gamma_{H^+ \rightarrow tb} = 14 \text{ GeV}$

$pp \rightarrow H^+ H^0$

$\sigma = 0.9 \text{ fb}$

Maybe Run II?

TABLE I: Predicted decay widths and production cross-sections for the dormant Higgs bosons. We used CalcHep, and production runs CTEQ61 proton structure functions, 1.64×10^5 calls. All cross-sections are evaluated at 14 TeV cms energy with the mass of H_2 doublet set to 380 GeV/ c^2 . Model dependent processes have rates or cross-sections that are indicated as $\propto (g'_b)^2$.

Process	value	comments
$\Gamma(H^+ \rightarrow t + \bar{b}) = \Gamma(H^- \rightarrow b + \bar{t})$	$14.5 (g'_b)^2 \pm 5 \times 10^{-5}\%$ GeV	
$\Gamma(H^0 \rightarrow b + \bar{b}) = \Gamma(A^0 \rightarrow b + \bar{b})$	$5.67 (g'_b)^2 \pm 5 \times 10^{-5}\%$ GeV	
$\Gamma(H^0 \rightarrow 2h, 3h) = \Gamma(A^0 \rightarrow 2h, 3h)$		absent in model
$pp \rightarrow (\gamma, Z^0) \rightarrow H^+ H^-$	$\sigma_t = 1.4$ fb	
$pp \rightarrow (\gamma, Z) \rightarrow H^0 H^0$		absent in model
$pp \rightarrow (\gamma, Z) \rightarrow A^0 H^0$	$\sigma_t = 1.3$ fb	
$pp \rightarrow (\gamma, Z) \rightarrow A^0 A^0$		absent in model
$pp(gg) \rightarrow h \rightarrow H^0 H^0$ or $A^0 A^0$	$\sigma_t = 1.7 \times 10^{-5}$ fb	
$pp \rightarrow W^+ \rightarrow H^0 H^+$	$\sigma_t = 1.8$ fb	
$pp \rightarrow W^+ \rightarrow A^0 H^+$	$\sigma_t = 1.8$ fb	
$pp \rightarrow W^- \rightarrow H^0 H^-$	$\sigma_t = 0.74$ fb	
$pp \rightarrow W^- \rightarrow A^0 H^-$	$\sigma_t = 0.74$ fb	
$pp \rightarrow b + \bar{b} + H^0$ or A^0	$\sigma_t = 1.8 (g'_b)^2$ pb $\pm 2.4\%$	No p_T cuts
	$\sigma_t = 67 (g'_b)^2$ fb $\pm 5\%$	$p_T(b)$ and $p_T(\bar{b}) > 50$ GeV
	$\sigma_t = 9.6 (g'_b)^2$ fb $\pm 3.5\%$	$p_T(b)$ and $p_T(\bar{b}) > 100$ GeV
$pp \rightarrow t + \bar{b} + (H^-)$	$\sigma_t = 220 (g'_b)^2$ fb	No cuts
	$\sigma_t = 44 (g'_b)^2$ fb	$p_T(t), p_T(\bar{b}) > 50$ GeV
	$\sigma_t = 14 (g'_b)^2$ fb	$p_T(t), p_T(\bar{b}) > 100$ GeV
$pp \rightarrow \bar{t} + b + (H^+)$	$\sigma_t = 270 (g'_b)^2$ fb	No cuts
	$\sigma_t = 46 (g'_b)^2$ fb $p_T(\bar{t})$	$p_T(b) > 50$ GeV
	$\sigma_t = 14 (g'_b)^2$ fb $p_T(\bar{t})$	$p_T(b) > 100$ GeV

The trilinear, quartic and quintic Higgs couplings will be significantly different than in SM case

$$V_{CW}(H) = \frac{1}{2}m_h^2 h^2 + \frac{5}{6\sqrt{2}v} h^3 \left(\beta_1 + \frac{9}{20}\beta_3 \frac{\partial\beta_1}{\partial\lambda_3} \right) + \frac{11}{48v^2} h^4 \left(\beta_1 + \frac{35}{44}\beta_3 \frac{\partial\beta_1}{\partial\lambda_3} \right) + \frac{1}{40\sqrt{2}v} h^5 \left(\beta_1 + \frac{25}{12}\beta_3 \frac{\partial\beta_1}{\partial\lambda_3} \right) + \dots$$

$$\text{trilinear} = \frac{5}{3} \left(1 + \frac{v^2}{5m_h^2} \frac{\lambda_3^3}{8\pi^4} \right) \approx 1.75$$

$$\text{quadrilinear} = \frac{11}{3} \left(1 + \frac{35v^2}{44m_h^2} \frac{\lambda_3^3}{8\pi^4} \right) \approx 4.43$$

$$\text{quintic} = \frac{3}{5} \left(\frac{\beta_1}{\hat{\beta}} + \frac{25}{12\hat{\beta}} \frac{\lambda_3^3}{6\pi^4} \right) \approx -8.87$$

This may be doable at LHC !

Problem: UV Landau Pole implying strong scale

$$\lambda_3(175 \text{ GeV}) = 4.79 \quad (\text{black})$$

$$\lambda_1(175 \text{ GeV}) = -0.1 \quad (\text{red})$$

$$\lambda_2(175 \text{ GeV}) = 0.1 \quad (\text{green})$$

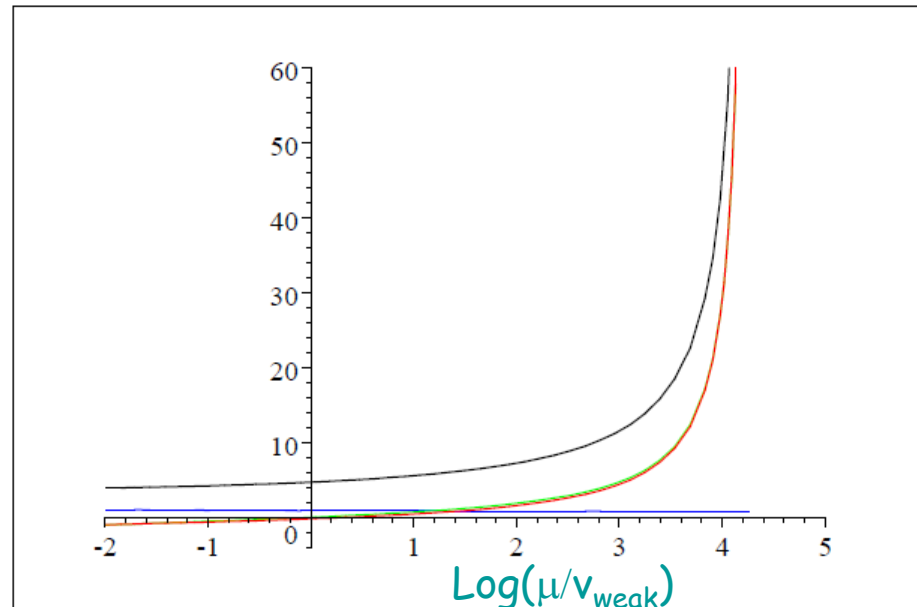
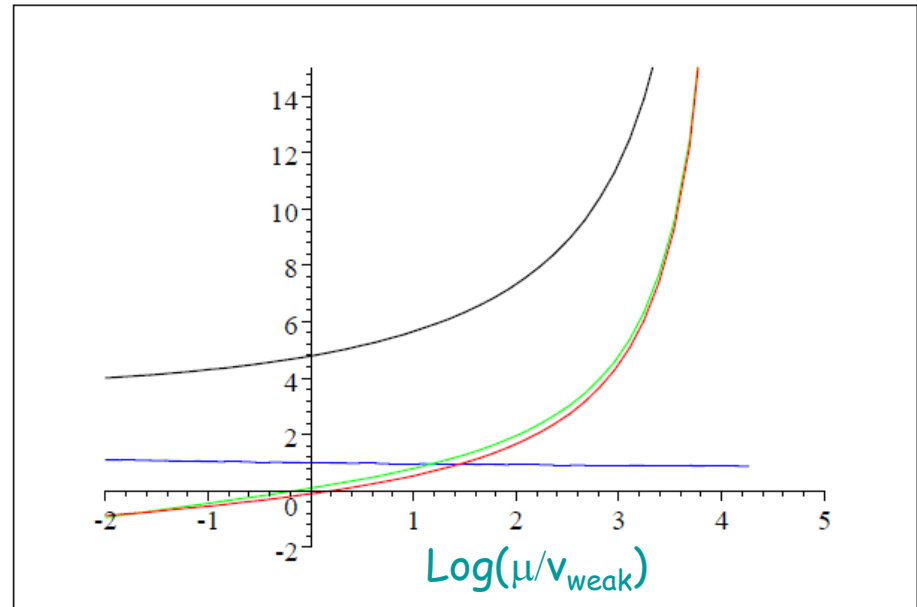
$$g_{\text{top}} = 1 \quad (\text{blue})$$

$$\lambda_4 = \lambda_5 = 0$$

Landau Pole = 10 - 100 TeV

Landau Pole ->
Composite H_2
New Strong Dynamics

e.g. [Higgs mass from compositeness at a multi-TeV scale](#),
[Hsin-Chia Cheng Bogdan Dobrescu, Jiayin Gu](#)
e-Print: [arXiv:1311.5928](#)



Hambye-Strumia Dark Matter Portal Model

[hep-ph]1306.2329

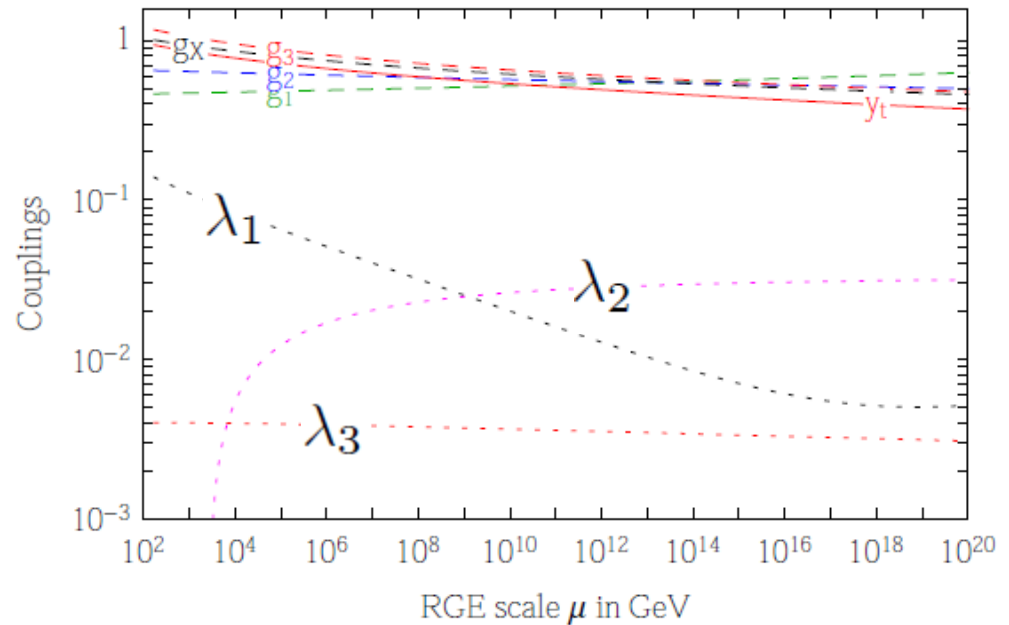
S develops a Coleman-Weinberg potential and VEV v_2

λ_3 is negative and gives the Higgs boson mass $-m^2 = \lambda_3 |S|^2$

The model does not require large quartic cc's, has sensible UV behavior

H_2 and associated gauge fields become viable dark matter

But, hard to detect !



Ultra-Weak Sector, dilaton, axion,

K. Allison, CTH, G. G. Ross, PL B738 191 (2015), NP B891, 613 (205)

$$V(H, \sigma) = \frac{\lambda}{2}(H^\dagger H)^2 + \frac{\zeta_1}{2}\sigma^2 H^\dagger H + \frac{\zeta_2}{4}\sigma^4 \quad V(\sigma, \phi_i, \lambda_i, \zeta_i) = V_1(\phi_i, \lambda_i) + V_2(\sigma_i, \phi_i, \zeta_i)$$

σ is a complex singlet

Here the full potential decomposes into components V_1 and V_2 where $\frac{\delta}{\delta\sigma_i}V_1 = \frac{\delta}{\delta\zeta_i}V_1 = 0$, and $\frac{\delta}{\delta\lambda_i}V_2 = 0$.

$$\beta_\lambda = \frac{d\lambda(\mu)}{d\ln(\mu)} = \frac{1}{16\pi^2} \left(12\lambda^2 - 3\lambda(3g_2^2 + g_1^2) + \frac{3}{4}(g_1^2 + g_2^2)^2 + \frac{3}{2}g_2^4 + 12\lambda g_t^2 - 12g_t^4 + \zeta_1^2 \right),$$

The ζ_i are technically naturally small: shift symmetry

$$\left\{ \begin{array}{l} \beta_1 = \frac{d\zeta_1(\mu)}{d\ln(\mu)} = \frac{1}{16\pi^2} \left(6\zeta_1\zeta_2 + 6\zeta_1\lambda + 4\zeta_1^2 - \frac{3}{2}\zeta_1(3g_2^2 + g_1^2) + 6\zeta_1g_t^2 \right), \\ \beta_2 = \frac{d\zeta_2(\mu)}{d\ln(\mu)} = \frac{1}{16\pi^2} (18\zeta_2^2 + 2\zeta_1^2). \end{array} \right.$$

$\langle\sigma\rangle = f$ can be very large,
e.g. GUT scale

Ultra-weak sector, dilaton, axion,

$$f \gtrsim 10^{10} \text{ GeV.}$$

Incorporates the axion,
GUT scale breaking, f ,
yields the Higgs boson mass

$$m_{\text{axion}} \sim \Lambda_{\text{QCD}}^2 / f$$

$$m_{\text{Dilaton}} \sim m_{\text{Higgs}}^2 / f$$

$$m_{\sigma} \approx 0.179 \left(\frac{10^{10} \text{ GeV}}{f} \right) \text{ keV.} \quad (20)$$

The model therefore predicts a low mass 0^+ particle for $f \gtrsim 10^{10} \text{ GeV}$.

The ζ_i are technically naturally small.
Extend to include right-handed neutrinos;

(σ, ν_R) can form an $N = 1$ SUSY multiplet;
SUSY broken by ζ_i allows tiny neutrino Dirac masses.

A Conjecture:



Max Planck

All mass is a quantum phenomenon.

$\hbar \rightarrow 0 \rightarrow$ Classical scale symmetry

Conjecture on the physical implications of the scale anomaly:
M. Gell-Mann 75th birthday talk: [C. T. Hill hep-th/0510177](#)

Musings: What if it's true?

All mass scales in physics are intrinsically quantum mechanical and associated with scale anomalies. The $\hbar \rightarrow 0$ limit of nature is exactly scale invariant.



(a heretic)

“Predictions” of the Conjecture:

We live in $D=4!$
$$T_{\mu}^{\mu} = \text{Tr} G_{\mu\nu} G^{\mu\nu} - \frac{D}{4} \text{Tr} G_{\mu\nu} G^{\mu\nu}$$

Cosmological constant is zero in classical limit

QCD scale is generated in this way; Hierarchy is naturally generated

Testable in the Weak Interactions!



“Predictions” of the Conjecture:

We live in D=4! $T_{\mu}^{\mu} = \text{Tr } G_{\mu\nu} G^{\mu\nu} - \frac{D}{4} \text{Tr } G_{\mu\nu} G^{\mu\nu}$

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Testable in the Weak Interactions!

Does the Planck Mass Come From Quantum Mechanics?

Can String Theory be an effective theory?

... or Weyl Gravity? (A-gravity?)

Weyl Gravity is Renormalizable!

Weyl Gravity is QCD-like:

$$\frac{1}{h^2} \sqrt{-g} (R_{\mu\nu} R^{\mu\nu} - \frac{1}{3} R^2)$$



“Predictions” of the Conjecture:

We live in D=4!
$$T_{\mu}^{\mu} = \text{Tr} G_{\mu\nu} G^{\mu\nu} - \frac{D}{4} \text{Tr} G_{\mu\nu} G^{\mu\nu}$$

Cosmological constant is zero in classical limit

D scale is generated in this way; Hierarchy naturally generated

in the Weak Interactions!



→ String Theory RULED OUT (classical string scale)

→ Weyl Gravity?

Weyl Gravity is Renormalizable in D=4!

Weyl Gravity in D=4 is QCD ($\nu R^{\mu\nu} - \frac{1}{3}R^2$)

→ The Planck Mass Comes From Quantum Mechanics!

See: **Conjecture on the physical implications of the scale anomaly.**
Christopher T. Hill (Fermilab) . hep-th/0510177 (and refs.therein)

“Predictions” of the Conjecture:

We live in D=4!
$$T_{\mu}^{\mu} = \text{Tr } G_{\mu\nu} G^{\mu\nu} - \frac{D}{4} \text{Tr } G_{\mu\nu} G^{\mu\nu}$$

Cosmological constant is zero in classical limit

QCD scale is generated in this way; Hierarchy naturally generated



... in the Weak Interactions!



String Theory RULES (classical string scale)

Asymptotically Safe Gravity:



Asymptotically Safe Gravity is Renormalizable! Predicts D=4!

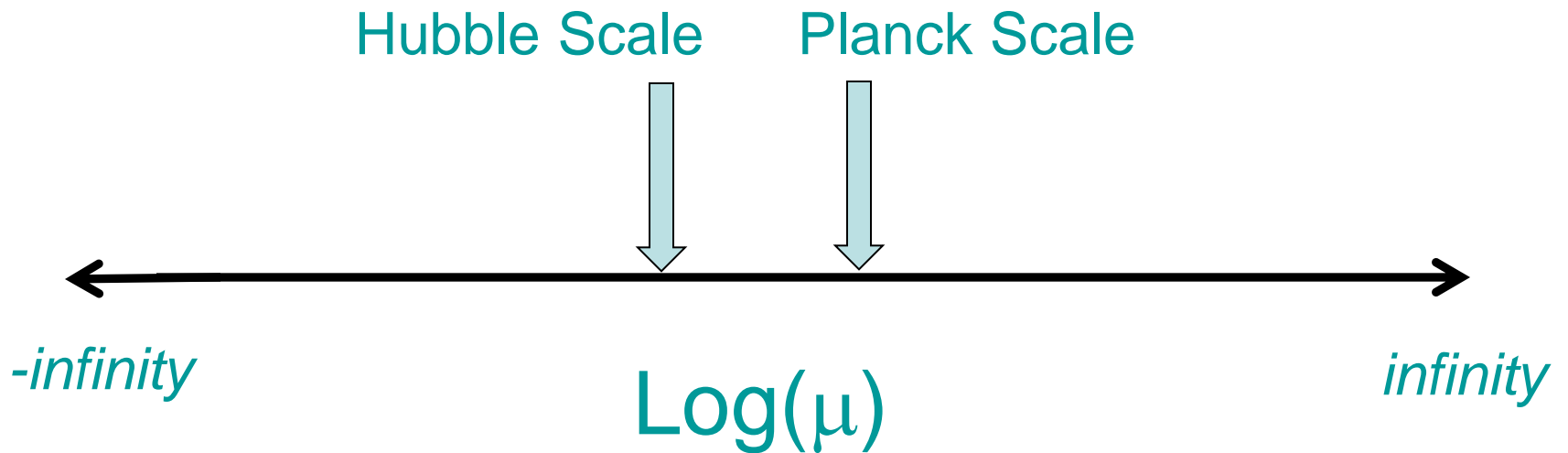
Asymptotically Safe Gravity in D=4 is QCD-like:
$$\frac{1}{h^2} \sqrt{-g} (R_{\mu\nu} R^{\mu\nu} - \frac{1}{3} R^2)$$

The Planck Mass Comes From Quantum Mechanics!

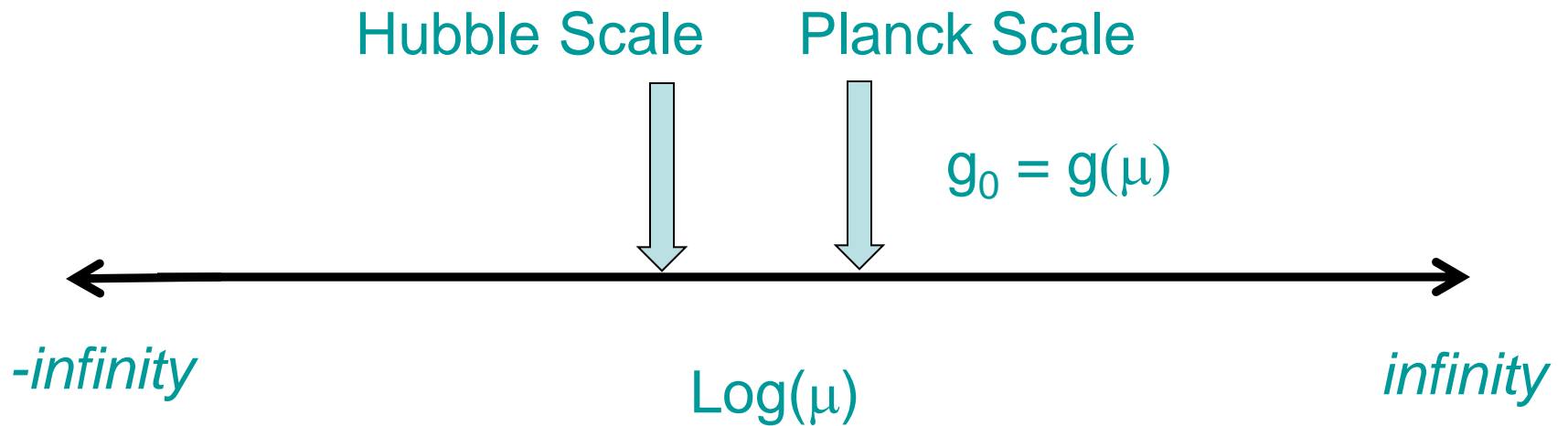
We Live in a Scaloplex !!!



The "Scaloplex" (scale continuum)
infinite, uniform, and classically isotropic



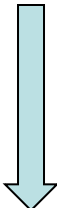
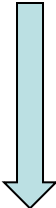
Physics is determined by **local values** of dimensionless coupling constants



Physics is determined by local values of dimensionless coupling constants

an equivalent universe $10^{1000} \times$

Hubble Scale' Planck Scale'



$$g_0 = g(10^{1000}\mu)$$



-infinity

Log(μ)

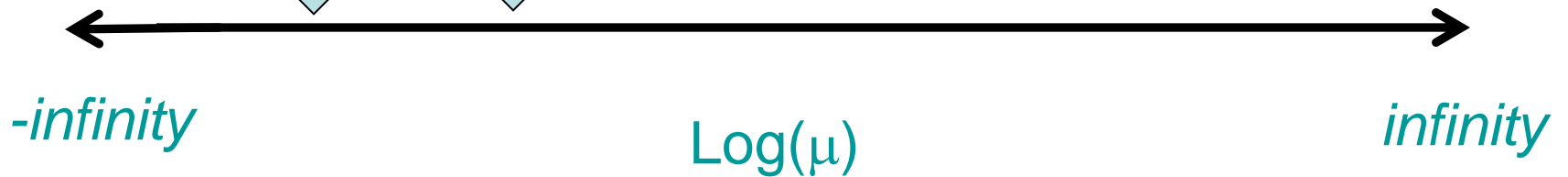
infinity

Physics is determined by local values of dimensionless coupling constants

an equivalent universe $10^{-1000} \times$

Hubble Scale” Planck Scale”

$$g_0 = g(10^{-1000}\mu)$$



Lack of additive scales:

Is the principle of scale recovery actually a "Principle of Locality" in Scale?

Physical Mass Scales, generated by e.g. Coleman-Weinberg or QCD-like mechanisms, are Local in scale, and do not add to scales far away in the scaloplex

E.g, "shining" with Yukawa suppression in extra dimensional models.

Does Coleman-Weinberg mechanism provide immunity from additive scales?

Conjecture on a solution to the Unitarity Problem of Weyl Gravity

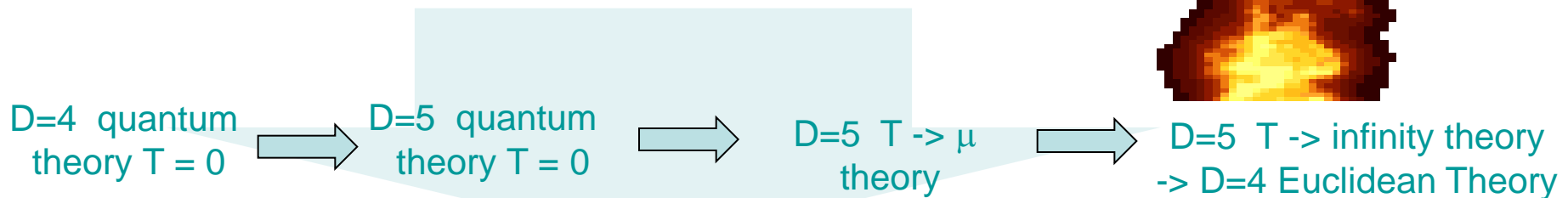
CTH, P. Agrawal

M_{Planck} arises via QCD-like mechanism.

Theory becomes Euclidean for $\mu > M_{\text{Planck}}$
(infinite temperature or instanton dominated)

Time is emergent for $\mu \ll M_{\text{Planck}}$

Passage through the Planck Scale



Time emerges
 $T < M_{\text{Planck}}$

$\text{Log}(\mu)$

Euclidean $D=4$

Hawking-Hartle Boundary Condition?

I think this is a profoundly important scientific question:

Is the Higgs potential Coleman-Weinberg?

- Examined a "maximally visible" scheme
- Dormant Higgs Boson from std 2-doublet scheme
 $M \approx 400 \text{ GeV}$
 - May be observable, LHC run II, III?
- Higgs trilinear and quartic couplings non-standard
 - UV problem \rightarrow new strong scale $< 100 \text{ TeV}$
 - or New bosons may be dark matter

Perhaps we live in a world where all
Mass comes from quantum effects
No classical mass input parameters.

Conclusions:

An important answerable scientific question:
Is the Higgs potential Coleman-Weinberg?

- We examined a "maximally visible" scheme
- Dormant Higgs Boson from std 2-doublet scheme
 $M \approx 386 \text{ GeV}$
 - May be observable, LHC run II, III?
 - Higgs trilinear ... couplings non-standard
or New bosons may be dark matter

Perhaps we live in a world where all
mass comes from quantum effects
No classical mass input parameters.

Everyone is still missing the
solution to the scale recovery problem!

End

Why do couplings run with field VEV?

$$S = \int d^4x \mathcal{L} = \int d^4x \left(\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right)$$

Equation of motion

$$\partial^\mu \frac{\delta S}{\delta \partial_\mu \phi} - \frac{\delta S}{\delta \phi} = \partial^2 \phi + V'(\phi) = 0 \quad V'(\phi) = \frac{\delta}{\delta \phi} V(\phi)$$

$$\begin{aligned} x^{\mu'} &= x^\mu - \zeta^\mu, & \delta dx^\mu &= -d\zeta^\mu(x) = -(\partial_\lambda \zeta^\mu) dx^\lambda \\ \phi'(x') &= \phi(x) & \delta \partial_\mu &= (\partial^\nu \zeta_\mu) \partial_\nu \\ & & \delta d^4x &= -(\partial_\mu \zeta^\mu) d^4x \end{aligned}$$

$$\begin{aligned} \delta S &= \int d^4x \left[-\frac{1}{2} (\partial_\rho \zeta^\rho) \partial_\mu \phi \partial^\mu \phi + (\partial^\rho \zeta_\mu) \partial_\rho \phi \partial^\mu \phi \right. \\ &\quad \left. + (\partial_\mu \zeta^\mu) V(\phi) \right] \equiv -\frac{1}{2} \int d^4x [(\partial_\mu \zeta_\nu) T^{\mu\nu}] \end{aligned}$$

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Stress tensor:
$$T_{\mu\nu} = \partial_\mu \phi \partial_\nu \phi - \eta_{\mu\nu} \left(\frac{1}{2} \partial_\rho \phi \partial^\rho \phi - V(\phi) \right)$$

$$\partial^\mu T_{\mu\nu} = \partial^2 \phi \partial_\nu \phi + \partial_\mu \phi \partial^\mu \partial_\nu \phi - \partial_\nu \left(\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) \right)$$

Equation of motion:
$$= \partial_\nu \phi (\partial^2 \phi + V'(\phi))$$

can choose $\zeta^\mu = -\epsilon x^\mu$

$$\delta S = \frac{1}{2} \int d^4x [(\partial_\mu \epsilon x_\nu) T^{\mu\nu}]$$

Scale Current:

$$\frac{\delta S}{\partial_\mu \epsilon} \equiv S^\mu = x_\nu T^{\mu\nu} \qquad \partial_\mu S^\mu = T^\mu_\mu$$

Scale Current not conserved with canonical stress tensor:

$$\partial_\mu S^\mu = T^\mu{}_\mu \qquad T^\mu{}_\mu = -\partial_\rho \phi \partial^\rho \phi + 4V(\phi)$$

The “Improved Stress Tensor”

$$S \rightarrow S + S_2 \qquad S_2 = \xi \int d^4x \partial^2 \phi^2$$

$$\delta S_2 = \xi \int d^4x [-(\partial_\mu \zeta^\mu) \partial^2 \phi^2 + \partial^\mu ((\partial^\nu \zeta_\mu) \partial_\nu \phi^2)] \equiv \int d^4x (\partial_\mu \zeta_\nu) [Q^{\mu\nu}]$$

$$Q_{\mu\nu} = \xi (\partial_\mu \partial_\nu \phi^2 - \eta_{\mu\nu} \partial^2 \phi^2) \qquad \tilde{T}_{\mu\nu} = T_{\mu\nu} + Q_{\mu\nu} \qquad \xi=1/6$$

$$= \frac{2}{3} \partial_\mu \phi \partial_\nu \phi - \frac{1}{6} \eta_{\mu\nu} \partial_\rho \phi \partial^\rho \phi - \frac{1}{3} \phi \partial_\mu \partial_\nu \phi + \frac{1}{3} \eta_{\mu\nu} \phi \partial^2 \phi + \eta_{\mu\nu} V(\phi)$$

Trace of improved stress tensor

$$\tilde{T}^\mu_\mu = \phi \partial^2 \phi + 4V(\phi) = -\phi \frac{\delta}{\delta \phi} V(\phi) + 4V(\phi)$$

Traceless for a scale invariant theory

$$\phi \frac{\delta}{\delta \phi} V(\phi) = DV(\phi) \quad \mathbf{D = 4} \rightarrow \quad V(\phi) = \frac{\lambda}{4} \phi^4,$$

In general, $D = 4 + \gamma$ $\frac{\delta}{\delta \phi} \lambda(\phi) = \beta(\lambda)$

$$\tilde{T}^\mu_\mu = -\frac{\beta(\lambda)}{\lambda} V(\phi)$$

Trace anomaly associate
with running coupling

A Canonical Picture of Scale Breaking

Stress tensor:
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Classical Standard Model Higgs Potential

$$m_{Higgs}^2 \approx \frac{1}{2} m_{top}^2 \quad m_{top} \approx v_{weak}$$

$$v_{weak} \approx 175 \text{ GeV}$$

$$\mathcal{L} = \mathcal{L}_{kinetic} + g_t \bar{\psi}_L t_R H + h.c. - \frac{\lambda}{2} (H^\dagger H - v_{weak}^2)^2$$

$$g_t \approx 1, \quad \lambda \approx \frac{1}{4}$$

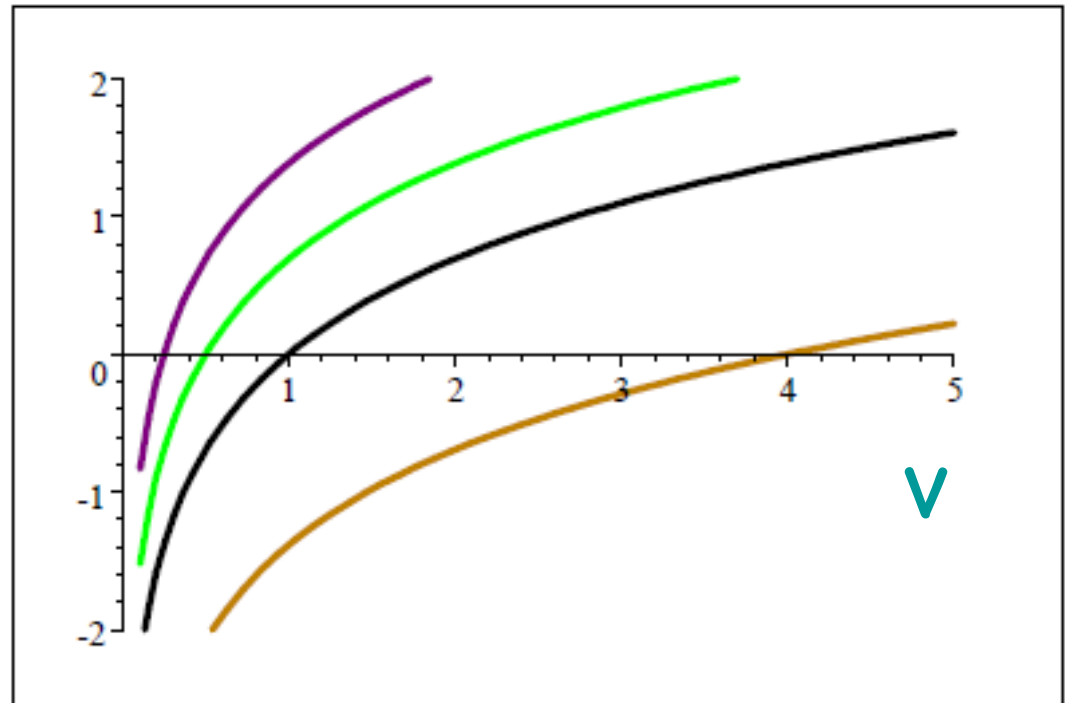
It is possible that we need
only the strongest coupled SUSY
partners to the Higgs Boson to be nearby in mass

e.g., "Natural SUSY" : A Light Stop

The More minimal supersymmetric standard model
A, G. Cohen , D.B. Kaplan, A.E. Nelson Phys.Lett. B388 (1996) 588-598

Quantum loops generate a logarithmic "running" of the quartic coupling

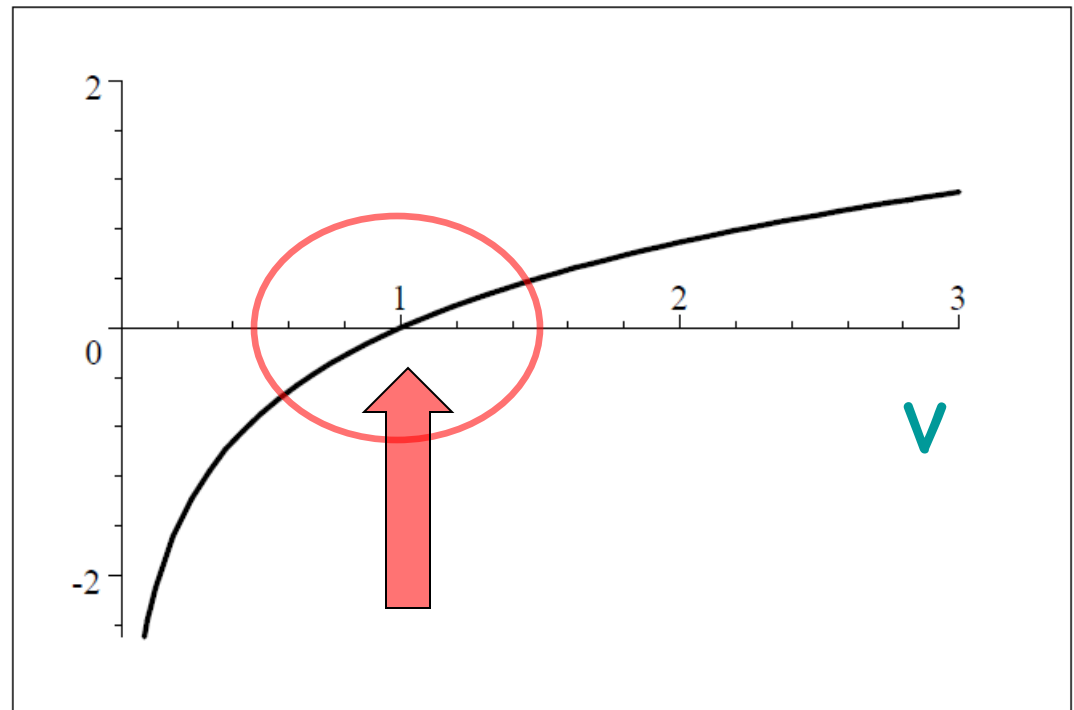
$$\lambda(v) \propto \hbar \beta \log(v/M)$$



running couplings have many possible trajectories, each parameterized by some M

Quantum loops can generate a logarithmic "running" of the quartic coupling

$$\lambda(v) \propto \hbar \beta \log(v/M)$$



this is the relevant behavior

$\tilde{\lambda}$ passing from < 0 to > 0 requires $\beta > 0$

