

# lattice studies of dark matter

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1. Lattice QCD studies of  $\langle N | m_s \bar{s} s | N \rangle$ .  
*collected in* [Junnarkar and Walker-Loud, Phys Rev D87, 114510 \(2013\)](#)
2. Putting the dark matter directly on the lattice:
  - SU(2) gauge theory. [Lewis,Pica,Sannino, Phys Rev D85, 014504 \(2012\)](#)  
[Hietanen,Lewis,Pica,Sannino JHEP 07\(2014\)116](#)  
[Hietanen,Lewis,Pica,Sannino, arXiv:1308.4130](#)  
[Detmold,McCullough,Pochinsky, arXiv:1406.2276 and 1406.4116](#)
  - SU(3) gauge theory. [Appelquist et al \(LSD collab\), Phys Rev D88, 014502 \(2013\)](#)
  - SU(4) gauge theory. [Appelquist et al \(LSD collab\), Phys Rev D89, 094508 \(2014\)](#)
  - SO(4) gauge theory. [Hietanen,Pica,Sannino,Søndergaard, Phys Rev D87, 034508 \(2013\)](#)

# lattice QCD studies of $\langle N | m_s \bar{s} s | N \rangle$

Perhaps dark matter is a WIMP (weakly-interacting massive particle).  
WIMP detection requires knowledge of WIMP-nucleon interactions.

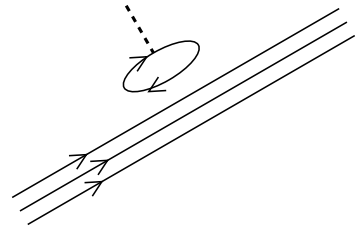
The low-energy limit of a spin-independent interaction is scalar.  
The scalar coupling to strangeness in a nucleon has been a challenge for theory.

Lattice QCD can determine the necessary matrix element,

$$f_s = \frac{\langle N | m_s \bar{s} s | N \rangle}{m_N}$$

either directly or via the Feynman-Hellman theorem:

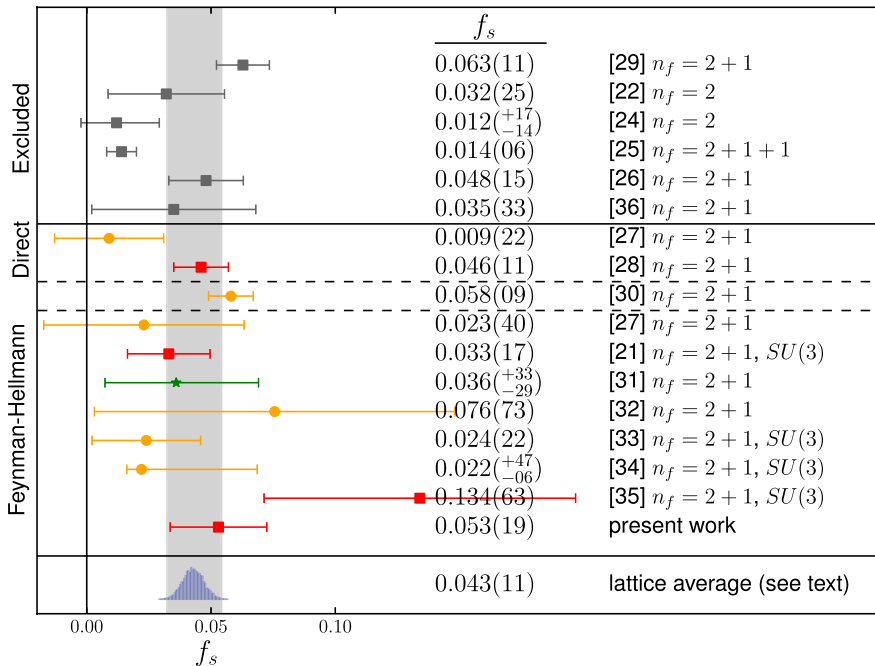
$$m_s \langle N | \bar{s} s | N \rangle = m_s \frac{\partial m_N}{\partial m_s}$$



Recent lattice results indicate that  $f_s$  is smaller than some previous estimates.

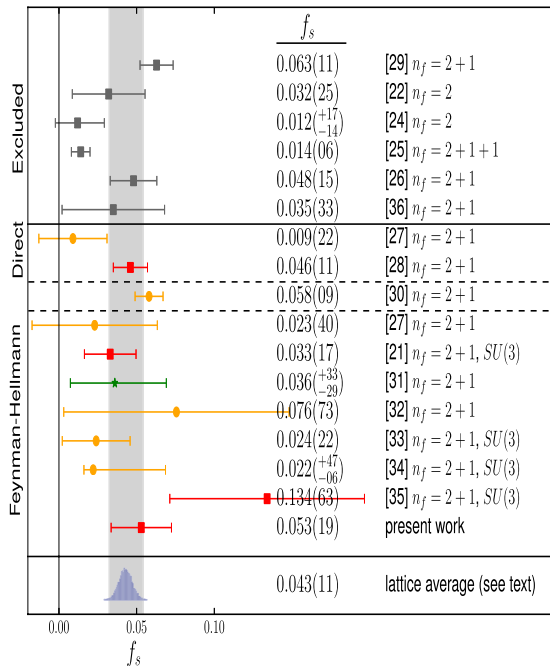
# lattice QCD studies of $\langle N | m_s \bar{s} s | N \rangle$

Graph taken from Junnarkar and Walker-Loud, Phys Rev D87, 114510 (2013)



# lattice QCD studies of $\langle N | m_s \bar{s} s | N \rangle$

Graph taken from Junnarkar and Walker-Loud, Phys Rev D87, 114510 (2013)



ASSESSING RELIABILITY:

Were u,d quarks light enough?  
Was the continuum limit taken?  
Was the  $\infty$  volume limit taken?



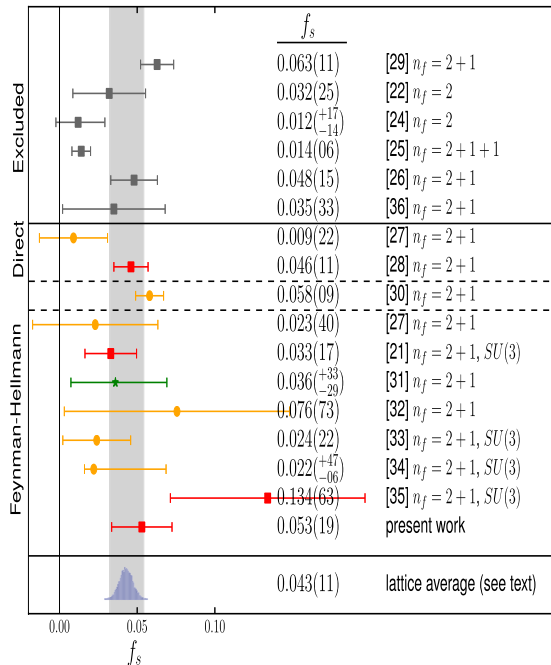
room for significant improvement

room for improvement

meets strictest constraints

# lattice QCD studies of $\langle N | m_s \bar{s} s | N \rangle$

Graph taken from Junnarkar and Walker-Loud, Phys Rev D87, 114510 (2013)



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# putting dark matter directly on the lattice

Dark matter is a BSM particle. Suppose it comes with a new strong interaction.

SU(2) gauge theory with 2 fundamental fermions is a minimal example.

- contains a dark matter candidate.
- produces electroweak symmetry breaking.
- accommodates a 125 GeV scalar.

Dynamical symmetry breaking,  $SU(4) \rightarrow Sp(4)$ , gives 5 Goldstone bosons:

$$\left. \begin{array}{l} \bar{U}\gamma_5 D \\ \bar{D}\gamma_5 U \\ \frac{1}{\sqrt{2}}(\bar{U}\gamma_5 U - \bar{D}\gamma_5 D) \end{array} \right\} \text{ eaten by } W^\pm \text{ and } Z$$
  

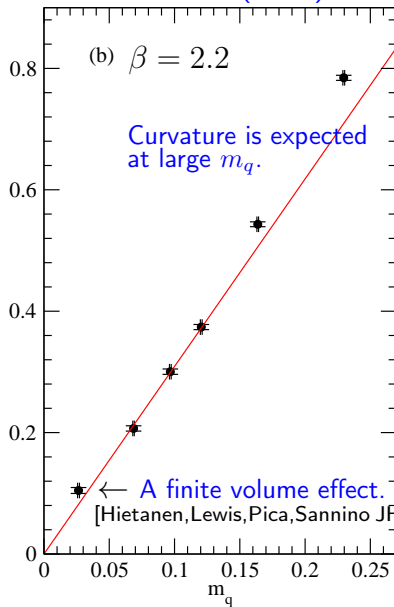
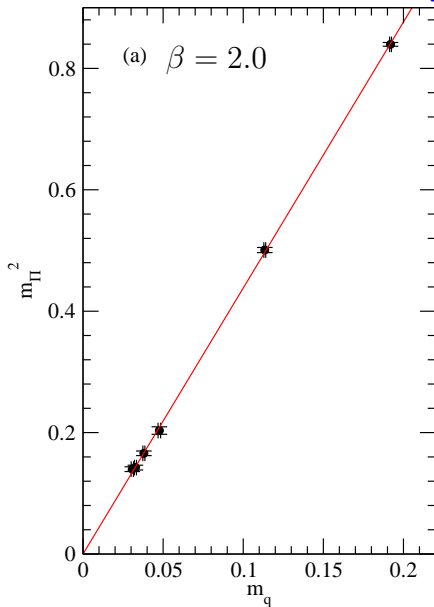
$$\left. \begin{array}{l} U^T(-i\sigma^2 C)\gamma_5 D \\ \bar{U}(-i\sigma^2 C)\gamma_5 \bar{D}^T \end{array} \right\} \begin{array}{l} \text{either } \textit{light} \text{ asymmetric dark matter (technicolor limit)} \\ \text{or Higgs + } \textit{heavier} \text{ dark matter (little Higgs limit)} \\ \text{or an interpolation between these two limits} \end{array}$$

## observing the Goldstone bosons in $N_c=N_f=2$

The expected behavior,  $m_\Pi^2 \propto m_q$  for small  $m_q$ , is observed.

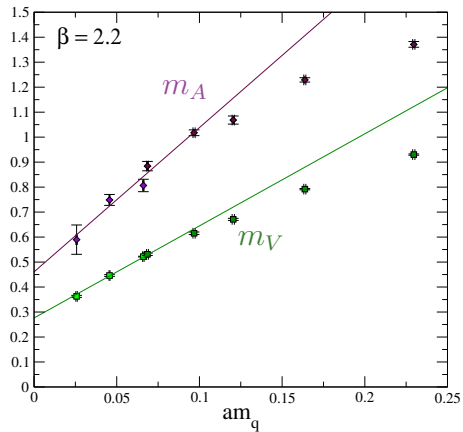
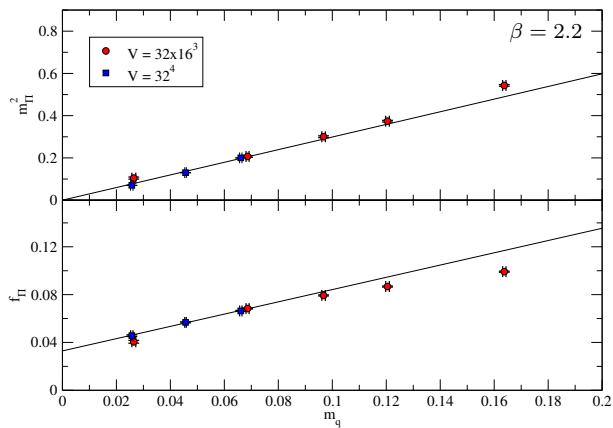
These plots apply to all five Goldstone bosons.

Lewis,Pica,Sannino, Phys Rev D85, 014504 (2012)



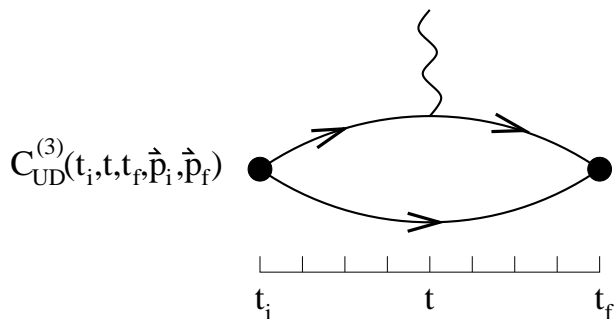
## observing light hadrons in $N_c=N_f=2$

Hietanen, Lewis, Pica, Sannino JHEP 07(2014)116





## relationships among Goldstone vector form factors in $N_c=N_f=2$



$$C_{UD}^{(3)}(t_i, t, t_f, \vec{p}_i, \vec{p}_f) = T^U - T^D$$

$$C_{\overline{UD}}^{(3)}(t_i, t, t_f, \vec{p}_i, \vec{p}_f) = -T^U + T^D$$

$$C_{U\overline{D}}^{(3)}(t_i, t, t_f, \vec{p}_i, \vec{p}_f) = T^U + T^D$$

$$C_{\overline{UD}}^{(3)}(t_i, t, t_f, \vec{p}_i, \vec{p}_f) = -T^U - T^D$$

$$C_{\overline{UU}+\overline{DD}}^{(3)}(t_i, t, t_f, \vec{p}_i, \vec{p}_f) = 0$$

$$T^X = \sum_{\vec{x}_i, \vec{x}, \vec{x}_f} e^{-i(\vec{x}_f - \vec{x}) \cdot \vec{p}_f} e^{-i(\vec{x} - \vec{x}_i) \cdot \vec{p}_i} \left\langle 0 \left| \mathcal{O}_{UD}^{(\gamma_5)}(x_f) V_\mu^X(x) \mathcal{O}_{UD}^{(\gamma_5)\dagger}(x_i) \right| 0 \right\rangle$$

## recalling resonance saturation in QCD

Lattice simulations with  $m_U \neq m_D$  are expensive  
 (photon hitting a vacuum loop doesn't cancel),  
 but with  $m_U = m_D$  the dark matter form factor vanishes.

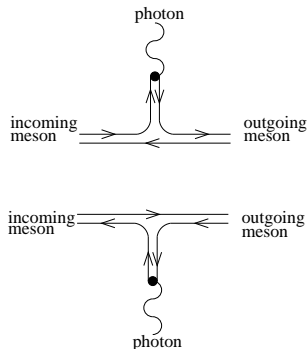
What to do?

Notice that vector meson dominance relates  $T^U$  to  $T^D$  in the large  $N_c$  limit  
 (and is successful for  $N_c = 3$  QCD):

$$F_{\pi^+} \approx \frac{2}{3} \left( \frac{m_\rho^2}{m_\rho^2 + Q^2} \right) + \frac{1}{3} \left( \frac{m_\rho^2}{m_\rho^2 + Q^2} \right)$$

$$F_{K^+} \approx \frac{2}{3} \left( \frac{m_\rho^2}{m_\rho^2 + Q^2} \right) + \frac{1}{3} \left( \frac{m_\phi^2}{m_\phi^2 + Q^2} \right)$$

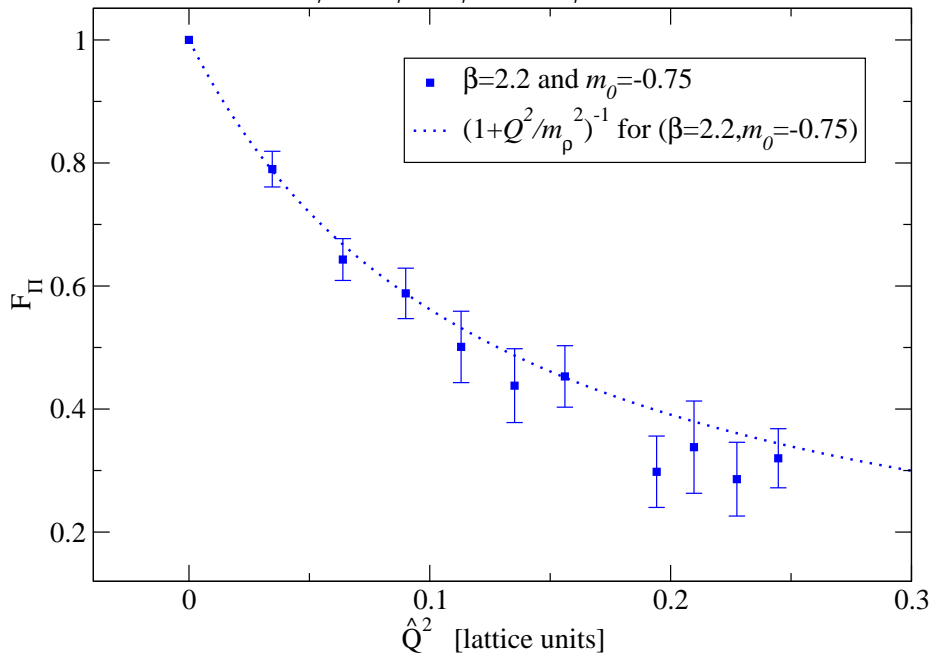
$$F_{K^0} \approx -\frac{1}{3} \left( \frac{m_\rho^2}{m_\rho^2 + Q^2} \right) + \frac{1}{3} \left( \frac{m_\phi^2}{m_\phi^2 + Q^2} \right)$$



We will test this VMD behaviour in SU(2) simulations with  $m_U = m_D$ .

**observing resonance saturation in  $N_c=N_f=2$** 

Hietanen, Lewis, Pica, Sannino, arXiv:1308.4130



## dark matter scattering by photon exchange in $N_c=N_f=2$

The coupling is due to the charge radius,

$$\mathcal{L}_B = ie \frac{d_B}{\Lambda^2} \phi^* \overleftrightarrow{\partial}_\mu \phi \partial_\nu F^{\mu\nu}$$

and we can calculate explicitly,

$$\frac{d_B}{\Lambda^2} = \lim_{Q^2 \rightarrow 0} \frac{1}{Q^2} \left( \frac{1}{2} \frac{m_{\rho U}^2}{m_{\rho U}^2 + Q^2} - \frac{1}{2} \frac{m_{\rho D}^2}{m_{\rho D}^2 + Q^2} \right) = \frac{m_{\rho U}^2 - m_{\rho D}^2}{2m_{\rho U}^2 m_{\rho D}^2}$$

Therefore

$$\boxed{\Lambda = m_\rho} \quad \text{and} \quad \boxed{d_B = \frac{m_{\rho U} - m_{\rho D}}{m_\rho}}$$

The cross section for scattering from a proton is

$$\sigma_p^\gamma = \frac{\mu^2}{4\pi} \left( \frac{8\pi\alpha d_B}{\Lambda^2} \right)^2 \quad \text{where} \quad \mu = \frac{m_\phi m_N}{m_\phi + m_N}$$

Given  $m_\phi > m_N$  and  $|d_B| < 1$ , we find  $\boxed{\sigma_p^\gamma < 2.3 \times 10^{-44} \text{ cm}^2}$ .

## adding the exchange of a composite Higgs

The dark matter candidate couples to a composite Higgs as follows:

$$\delta\mathcal{L} = \frac{d_1}{\Lambda} h \partial_\mu \phi^* \partial^\mu \phi + \frac{d_2}{\Lambda} m_\phi^2 h \phi^* \phi$$

We expect  $d_1$  and  $d_2$  to be of order unity.

The cross section for scattering from a proton is

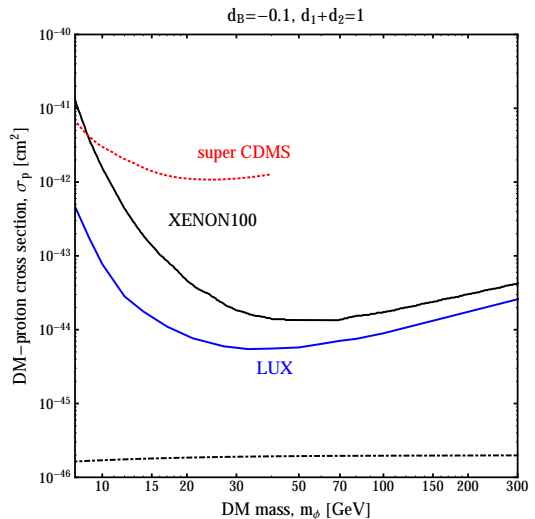
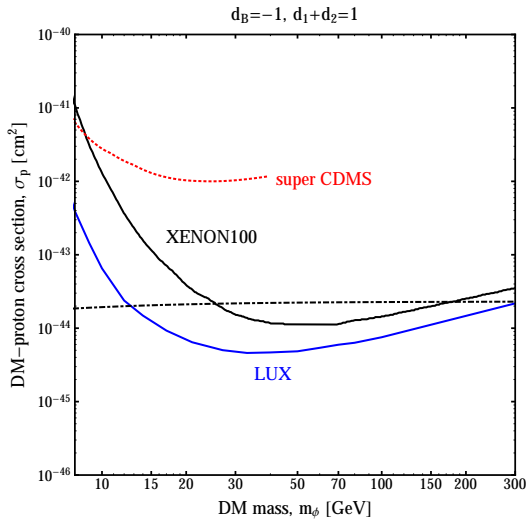
$$\sigma_p = \frac{\mu^2}{4\pi} \left( \underbrace{\frac{(d_1 + d_2) f m_N m_\phi^2}{m_H^2 m_\phi v_{EW} \Lambda}}_{f_n} + \frac{8\pi\alpha d_B}{\Lambda^2} \right)^2 \quad \text{where} \quad \mu = \frac{m_\phi m_N}{m_\phi + m_N}$$

The Higgs to nucleon coupling is parametrized by  $f \sim 0.3$ .

This cross section is thus a function of  $m_\phi$  and  $d_B$ . Compare to experiment...

## comparison of $N_c=N_f=2$ to experiments

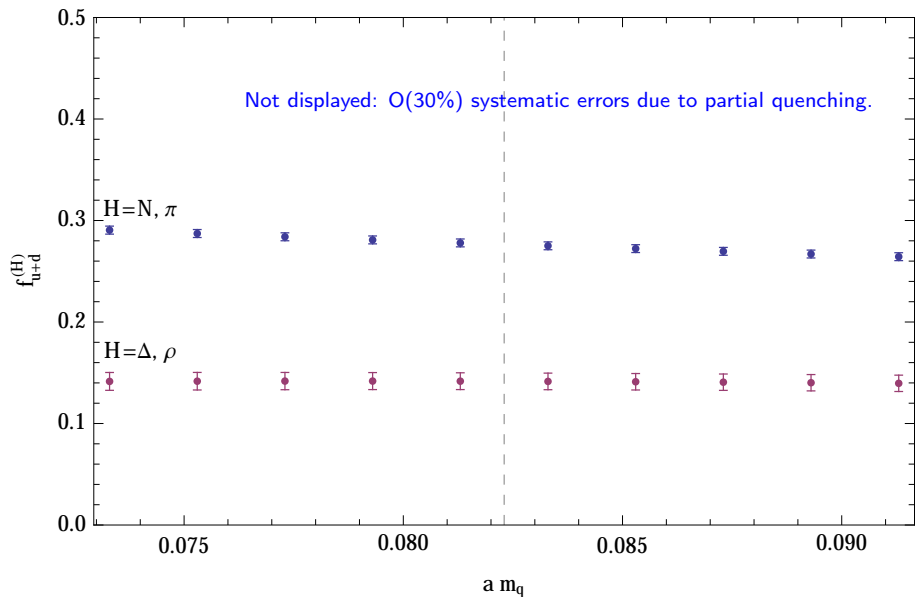
Hietanen, Lewis, Pica, Sannino, arXiv:1308.4130



## scalar couplings in $N_c=N_f=2$

Detmold,McCullough,Pochinsky, arXiv:1406.4116

$$f_q^{(H)} = \frac{m_q}{M_H} \frac{\partial M_H}{\partial m_q} = \frac{\langle H | m_q \bar{q} q | H \rangle}{M_H}$$



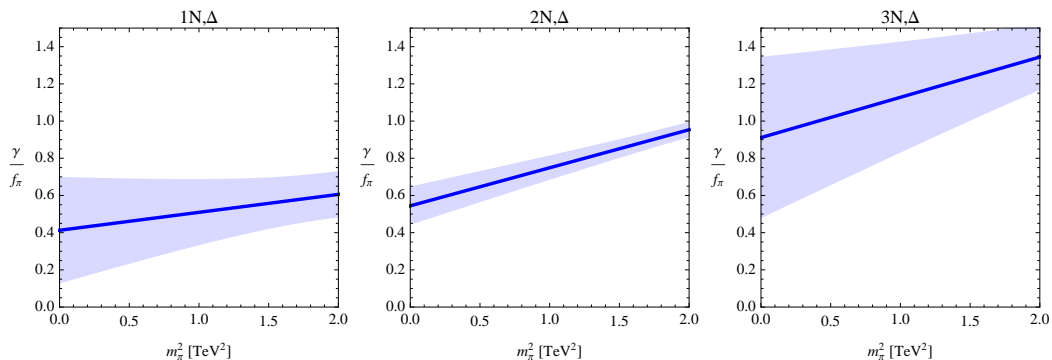
## dark nuclei in $N_c=N_f=2$

Detmold,McCullough,Pochinsky, arXiv:1406.4116

For scattering states,  $\Delta E(L) \propto 1/L^3 + \dots$

For bound states, 
$$\Delta E(L) = -\frac{\gamma^2}{2\mu} \left[ 1 + \frac{12\hat{C}}{\gamma L} e^{-\gamma L} \right]$$

Bound states are observed for  $J^P = 1^+$  in  $N\Delta$  and  $NN\Delta$  and perhaps  $NNN\Delta$ :

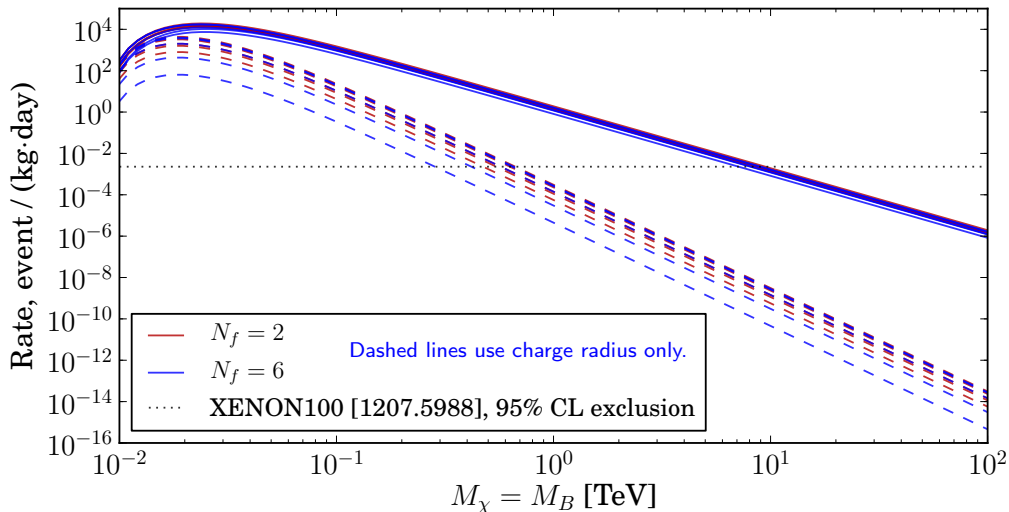


(This study uses  $m_\rho/2 < m_\pi < m_\rho$  and  $f_\pi = 246$  GeV.)



## Event rate for XENON100 from a $N_c=3$ dark matter model

Appelquist et al (LSD collab), Phys Rev D88, 014502 (2013)



All dark quarks are weak singlets.  $Q_U = \frac{2}{3}$ ,  $Q_D = -\frac{1}{3}$ . Disconnected lines omitted.

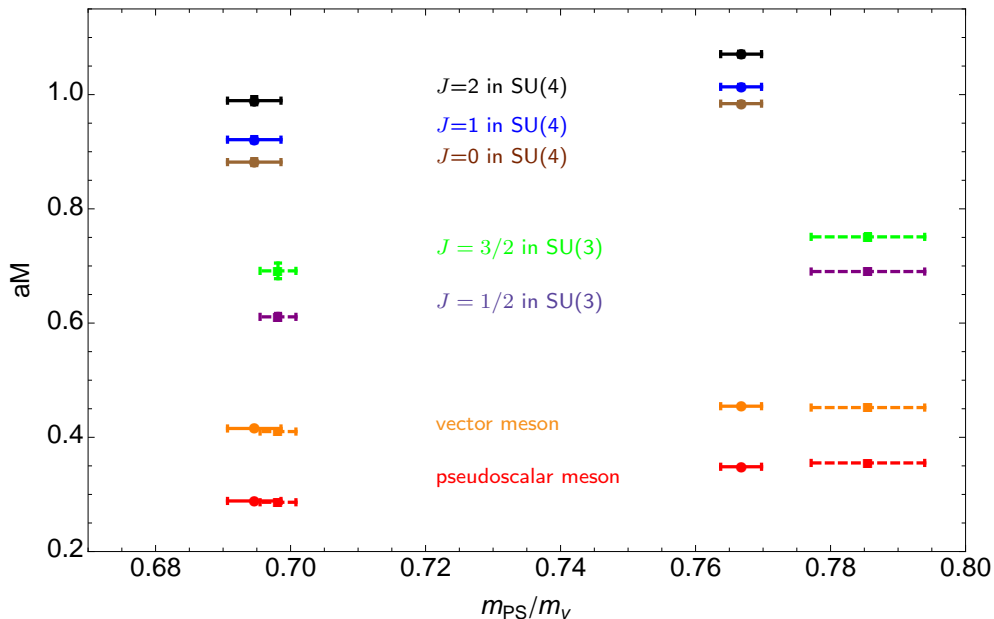
The  $N_f^2 - 1$  Goldstones are assumed unstable so “neutron” is the DM candidate.

Caution:  $\langle r_E^2 \rangle_{\text{neutron}} \approx \text{experiment}/10$ . Decreasing  $m_q$  might clarify this.

## hadron mass spectrum in $N_c = 4$ dark matter model

Appelquist et al (LSD collab), Phys Rev D89, 094508 (2014)

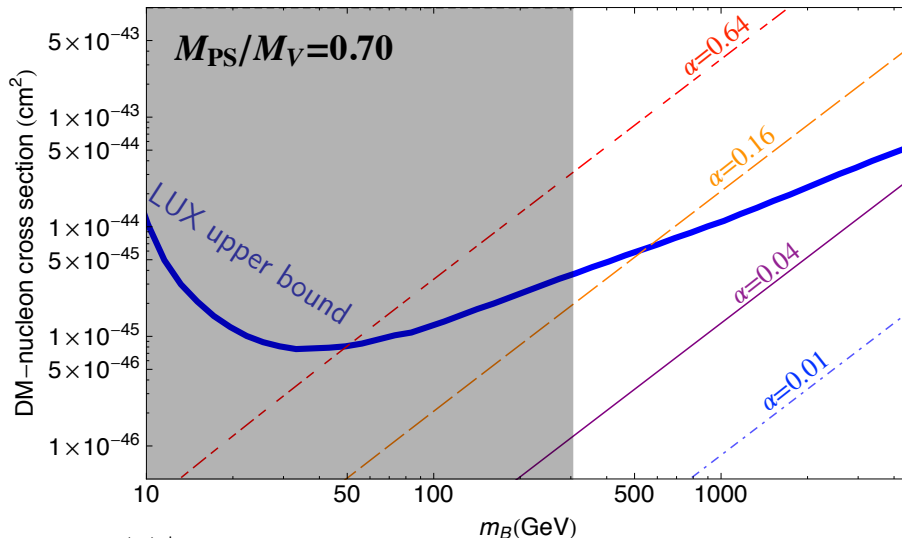
This is a quenched exploration. It has  $m_f \sim \Lambda_4$ .



## bounds on fermion-Higgs coupling in $N_c = 4$ dark matter model

Appelquist et al (LSD collab), Phys Rev D89, 094508 (2014)

This is a quenched exploration. It requires  $m_{\text{PS}} > 100$  GeV due to LEP.



$$\alpha = \left. \frac{v}{m_f} \frac{\partial m_f(h)}{\partial h} \right|_{h=v}$$

is the coupling of a dark baryon to the Higgs.

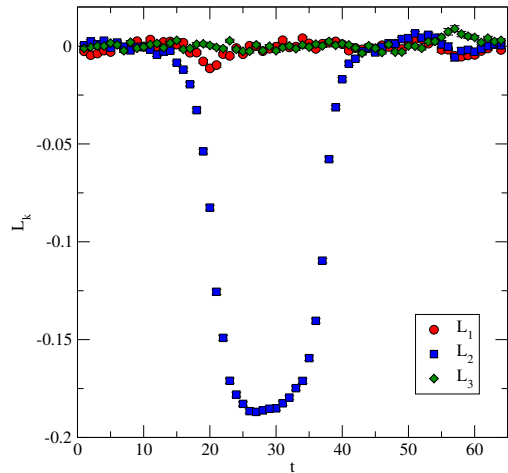
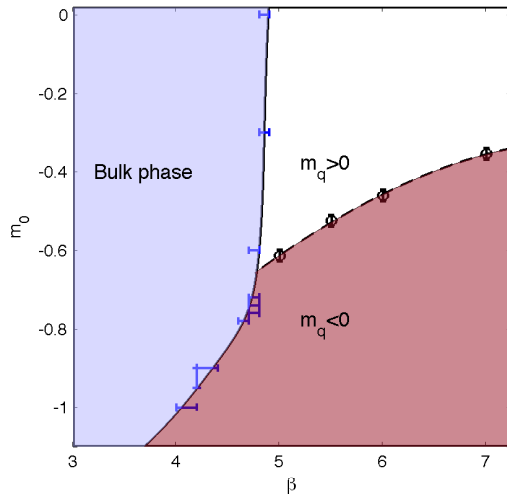
Lattice spacing, volume, and some range of  $m_{\text{PS}}/m_V$  were studied.

## phase structure of SO(4) with 2 fermions

Hietanen,Pica,Sannino,Søndergaard, Phys Rev D87, 034508 (2013)

Lattice dark matter beyond SU(N): step one is to explore the phases.

This study uses 2 (Wilson) Dirac fermions in the vector representation.

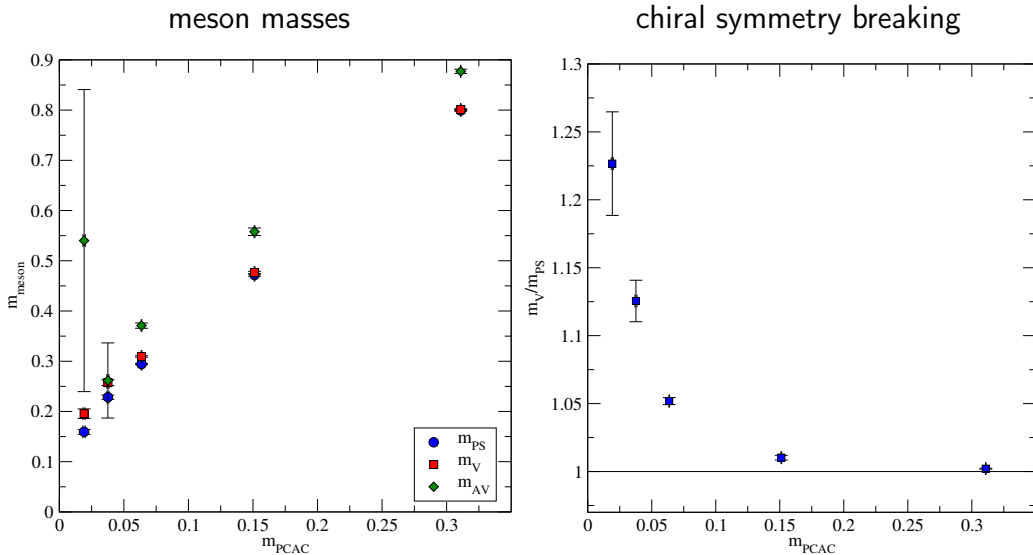


Surprising Polyakov multi-phase phenomenon not observed for larger volumes.

## hadron masses in SO(4) with 2 fermions

Hietanen,Pica,Sannino,Søndergaard, Phys Rev D87, 034508 (2013)

Expected global symmetry breaking is  $SU(4) \rightarrow SO(4)$ . Therefore 9 Goldstones.  
The isospin=0 Goldstone boson is the **dark matter candidate**.



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  - SO(4) gauge theory. Hietanen,Pica,Sannino,Søndergaard, Phys Rev D87, 034508 (2013)

***All of this is just the beginning...***