PHY385-H1F Introductory Optics

Last Class!! – Outline:

- Diffraction Gratings
- Things to know about the final exam
- A look ahead: PHY485: "Modern Optics"
- Some Course Review



Diffraction Gratings

Bright orders at: $m\lambda = a(\sin\theta_m - \sin\theta_i)$



Diffraction Gratings

- The m = 0 order corresponds to regular reflection: $\theta_m = \theta_r$: No dispersion.
- Other orders are dispersed: rainbows. Usually the first order lines (m = 1 or m = -1) are the most intense.



Diffraction Gratings

• Reflection gratings obey a similar equation as transmission gratings:



Diffraction Gratings

- The grating below is "blazed", meaning its surface is a reflective saw-tooth shape. Blazing can increase the efficiency for a particular order.
- It appears to be blazed for 1st order



Echelle Gratings

- An echelle grating is blazed for extremely high order. The purpose is to increase dispersion, which is proportional to *m*.
- In practice, adjacent orders always overlap, so a second "cross-grating" must be used to separate the orders on the detector.



Final Exam

- Thursday Dec. 13 from 2:00 to 4:00pm, in the Exam Centre at 255 McCaul, Room EX320
- First 2 Pages with "helpful information" is posted on the course web-site, under "Materials".
- Allowed aids include a calculator, and up to 2pages of hand-written notes.
- There are 8 problems, similar to the style you have seen on the mid-terms and Practicals presentation-sessions.
- No multiple choice.
- The material includes the entire course, spread evenly.

Echelle Spectrum on a CCD detector



PHY1485H-F / 485

- "Lasers and Modern Optics" is a cross-listed course (grad students and undergrads), currently taught by Professor Thywissen. He recommends Hecht 4e.
- Current Pre-requisites are "Electromagnetic Theory" (PHY350) and "Quantum I" (PHY356)
- Topics include:
 - Gaussian beam modes and their relation to optical resonators
 - Fibre and Slab waveguides
 - Laser Cooling
 - Photonic bandgap structures
 - Extreme optics
 - Quantum information
- There are End-of-year student seminars.



Some Review... Visible Light



 Hecht says: 455 to 780 nm. Personally, I can't see light beyond about 700 nm. And I am able to see violet down to about 420 nm.





A light wave travels, as a plane wave, from air (n = 1.0)into glass (n = 1.5). Which diagram shows the correct wave fronts? 1814 – Jean Fresnel used the idea of polarization to predict amplitudes of reflected and transmitted light from glass interfaces.



Partial transmission and reflection amplitudes of a wave travelling from a low to high refractive index medium.





f is negative for a diverging lens

 s_o is positive means object is real, to the left of lens

 s_o is positive means image is real, to the right of lens

 s_i is negative means image is virtual, to the left of lens

Lensmaker's Formula



$$\frac{1}{f} = \frac{n_l - n_m}{n_m} \left(\frac{1}{R_1} - \frac{1}{R_2} \right)$$

- The total amount of light collected by the lens is proportional to D².
- The image area of an extended object is proportional to f^2 .
- So the flux density at the image plane varies as $(D/f)^2$.
- *D*/*f* is called "relative aperture"
- f/D is called the "f-number" (ie F1.4, F2, F16, etc)
- (*D/f*)² is called the "speed". The higher the speed, the shorter an exposure time you need for the same image brightness.
- That's why f-numbers tend to increase by factors of $\sqrt{2}$ on cameras for each step you have to double the exposure time.

FIGURE 23.52 A real image formed by a concave mirror.



The Mirror Equation

For a spherical mirror with negligible thickness, the object and image distances are related by

$$\frac{1}{s_i} = \frac{1}{f}$$
 (thin-mirror equation)

where the focal length *f* is related to the mirror's radius of curvature by

1

s_

$f = \frac{R}{2}$ Sign convention for spherical mirrors		
	Positive	Negative
R and f	Concave toward the object	Convex toward the object
Si	Real image, same side as object	Virtual image, opposite side from object

Modal Dispersion In Optical Fibres



Compound Microscope



Telescope

- Invented by somebody in Holland (1608), later used by Galileo of Italy (1610).
- Objective forms a real, inverted image in the tube.
- Eyepiece is used as a magnifier to view this image.

$$MP = \frac{\alpha_a}{\alpha_u}$$







A photon with energy 2.0 eV is incident on an atom that is in the p state. Which can happen?

E(eV)

- 1. Absorption
- 2. Stimulated emission
- 3. Spontaneous Emission
- 4. Either 1 or 2
- 5. Either 2 or 3







Spectrum of a He-Ne Laser



What best describes "dichroism"?

- 1. Absorption coefficient depends on the linear polarization state (vertical or horizontal)
- 2. Absorption coefficient depends on the circular polarization state (L or R)
- 3. Index of refraction depends on the linear polarization state (vertical or horizontal)
- 4. Index of refraction depends on the circular polarization state (L or R)

What best describes "birefringence"?

- 1. Absorption coefficient depends on the linear polarization state (vertical or horizontal)
- 2. Absorption coefficient depends on the circular polarization state (L or R)
- 3. Index of refraction depends on the linear polarization state (vertical or horizontal)
- 4. Index of refraction depends on the circular polarization state (L or R)

What best describes "optical activity"?

- 1. Absorption coefficient depends on the linear polarization state (vertical or horizontal)
- 2. Absorption coefficient depends on the circular polarization state (L or R)
- 3. Index of refraction depends on the linear polarization state (vertical or horizontal)
- 4. Index of refraction depends on the circular polarization state (L or R)

Dichroism



"Unpolarized" light incident on a linear polarizer tilted at an angle θ with respect to the vertical



A Young Double Slit Experiment screen with pinholes ith interference pattern Measures Spatial Coherence



- πD^2 is the "Coherence Area"
- Coherence area for filtered sunlight: A ~ 10⁻² mm²
- Coherence area for filtered starlight: A ~ 6 m²

Source: http://skullsinthestars.com/2008/09/03/optics-basics-coherence/

Diffraction





- · Sometimes light does not travel in straight lines, or rays. The rays bend around edges, because light is actually a wave.
- · This was first noted and discussed by Grimaldi in 1640.

