PHY489 Lecture 12

Fermi's Golden Rule for Transitions

transition probability

$$W = \frac{2\pi}{\hbar} \left| \mathcal{M}_{if} \right|^2 \rho_f(E)$$

 Density of final states available for energy *E*.

matrix element

$$\left|\mathcal{M}_{if}\right|^{2} = \left|\left\langle \psi_{f} \left| V_{if} \left| \psi_{i} \right\rangle \right|^{2}\right.$$

Matrix element contains the fundamental physics (e.g. the dynamics).

1

There are cases where \mathcal{M}_{if} = constant, in which case the reaction is entirely governed by the kinematics, described by $\rho_f(E)$. Need this to be Lorentz invariant (we will see that \mathcal{M} is always a Lorentz scalar).

Reminder: $\langle \Psi_f | H_{int} | \Psi_i \rangle$ for $H_{TOT} = H_0 + H_{int}$ with $H_{int} << H_0$ treated with perturbation theory, starting with free particle wave-functions (plane-wave solutions). Conventional wave-function normalization is such that

$$\psi = \frac{1}{\sqrt{V}} e^{-i(\vec{p}\cdot\vec{r} - Et)} \qquad \int \psi^* \psi \, dV = 1$$

e.g. normalize to a particle density of 1/V.

Lorentz Invariant Wave-function Normalization²

Consider a particle in an interaction to be at rest w.r.t. a box of volume *V* containing it and only it. Now consider this system viewed in a reference frame S[/] in which the box is moving with relative velocity *v*. In this frame (in which we want to calculate *W*) the volume *V* is Lorentz contracted by a factor of $\gamma = E/mc^2$ (assuming the particle has mass).

$$V' = \frac{mc^2}{E}V \qquad \int \psi^* \psi \, dV' = \frac{E}{mc^2} \quad (\text{in S'})$$

e.g. the particle density has increased by a factor of γ . To ensure a reference-frame independent particle density normalization need to incorporate a factor of $\sqrt{2E}$ in the wave-function normalization:

$$\Psi = \sqrt{\frac{2E}{V}} e^{-i(\vec{p}\cdot\vec{r}-Et)}$$

This wave-function normalization is Lorentz invariant.

Density of States

The state of a single-particle with momentum between 0 and p, confined to a volume V, is specified by a point in 6D phase space (x,y,z,p_x,p_y,p_z)

The extent to which spatial and momentum coordinates along each axis can be simultaneously specified is limited by the uncertainty principle, to h, so the "elemental volume" of phase space is h^3 .

The number of states available to a single particle is thus

$$N = \frac{\text{total phase space available}}{\text{elemental volume}} = \frac{1}{\left(2\pi\hbar\right)^3} \int dx \, dy \, dz \, dp_x \, dp_y \, dp_z = \frac{V}{\left(2\pi\hbar\right)^3} \int d^3\vec{p} \, dx \, dy \, dz \, dp_y \, dp_z$$

For a system of *n* particles, the number of available final states is the product of the individual factors for each particle:

$$N_n = \left[\frac{V}{\left(2\pi\hbar\right)^3}\right]^n \int \prod_{i=1}^n d^3 \vec{p}_i$$

Phase space factor defined as the number of states / unit energy / unit volume

$$\rho(E) = \frac{dN_n}{dE} = \frac{1}{\left(2\pi\hbar\right)^{3n}} \frac{d}{dE} \int \prod_{i=1}^n d^3 \vec{p}_i \qquad \text{[ignore factors of V for now]}$$

Note that not all momenta are independent (conservation of momentum). Could integrate over n-1 momenta, but simpler to use Dirac δ -function to impose the constraint:

Total three momentum of final state

$$\vec{P} - \sum_{i=1}^{n} \vec{p}_{i} = 0 \qquad \Rightarrow \qquad \int d^{3}\vec{p}\,\delta\left(\vec{P} - \sum_{i=1}^{n} \vec{p}_{i}\right) = 1$$

$$\rho(E) = \frac{1}{\left(2\pi\hbar\right)^{3n}} \frac{d}{dE} \int \prod_{i=1}^{n} d^{3}\vec{p}_{i}\delta\left(\vec{P} - \sum_{i=1}^{n} \vec{p}_{i}\right)$$

Energy conservation gives us a further constraint $\left(\sum_{i=1}^{n} E_{i}\right) - E = 0$

4

and thus
$$\int dE \,\delta \left(\sum_{i=1}^{n} E_i - E\right) = 1$$
 so that we can write

$$\rho(E) = \frac{1}{\left(2\pi\hbar\right)^{3n}} \frac{d}{dE} \int \prod_{i=1}^{n} d^{3}\vec{p}_{i} dE \,\delta\left(\vec{P} - \sum_{i=1}^{n}\vec{p}_{i}\right) \delta\left(\sum_{i=1}^{n}E_{i} - E\right)$$

Using the relation $\frac{d}{dE}\int f(E)dE = f(E)$ this becomes:

$$\rho(E) = \frac{1}{\left(2\pi\hbar\right)^{3n}} \int \prod_{i=1}^{n} d^{3}\vec{p}_{i} \,\delta\left(\vec{P} - \sum_{i=1}^{n}\vec{p}_{i}\right) \delta\left(\sum_{i=1}^{n}E_{i} - E\right)$$

so each particle has a phase space factor (before integration) of $\frac{d^3\vec{p}}{(2\pi\hbar)^3}$

Unfortunately, this is NOT Lorentz invariant. Note that the factors of V (which we've been ignoring) will cancel with factors from the wave-function normalization in the matrix element.

Lorentz Invariance of Phase Space Factor

For Lorentz invariance now need additional factors of 2E from the Lorentz invariant wave-function normalization. The phase space factor of the form $\frac{d^3 \vec{p}_i}{E}$ is Lorentz invariant. Show this:

$$p'_{x} = \gamma(p_{x} - \beta \frac{E}{c}) \qquad \frac{E'}{c} = \gamma(\frac{E}{c} - \beta p_{x}) \qquad p'_{y} = p_{y} \qquad p'_{z} = p_{z}$$

$$\frac{dp'_{x}}{dp_{x}} = \frac{d}{dp_{x}} \left(\gamma(p_{x} - \beta \frac{E}{c})\right) = \gamma \left(1 - \frac{\beta}{c} \frac{dE}{dp_{x}}\right)$$

$$\frac{dE}{dp_{x}} = \frac{d}{dp_{x}} \left(p^{2}c^{2} + m^{2}c^{4}\right)^{1/2} = \left(2p_{x}c^{2}\right) \left(\frac{1}{2}\right) \left(p^{2}c^{2} + m^{2}c^{4}\right)^{-1/2} = \frac{p_{x}}{E}c^{2}$$

$$\frac{dp'_{x}}{dp_{x}} = \gamma \left(1 - \frac{\beta}{c} \frac{p_{x}}{E}c^{2}\right) = \gamma \left(1 - \beta \frac{p_{x}}{E}c\right) = \gamma \left(1 - \beta \frac{p_{x}}{E}c\right) = \frac{\gamma \left((E/c) - \beta p_{x}\right)}{(E/c)} = \frac{E'}{E}c^{2}$$

$$\frac{dp'_{x}}{dp_{x}} = \gamma \left(1 - \frac{\beta}{c} \frac{p_{x}}{E}c^{2}\right) = \gamma \left(1 - \beta \frac{p_{x}}{E}c\right) = \gamma \left(1 - \beta \frac{p_{x}}{E}c\right) = \frac{\gamma \left((E/c) - \beta p_{x}\right)}{(E/c)} = \frac{E'}{E}c^{2}$$

$$\frac{dp_x}{dp_x} = \frac{E}{E} \implies \frac{dp_x}{E'} = \frac{dp_x}{E} \text{ or } \left[\frac{dp}{E'} = \frac{dp}{E}\right]$$

Golden Rule for Decays

Consider decays of the form $1 \rightarrow 2+3+\ldots+N$

$$d\Gamma = \left|\mathcal{M}\right|^{2} \frac{S}{2\hbar m_{1}} \left[\left(\frac{cd^{3}\vec{p}_{2}}{\left(2\pi\right)^{3} 2E_{2}} \right) \left(\frac{cd^{3}\vec{p}_{3}}{\left(2\pi\right)^{3} 2E_{3}} \right) \dots \left(\frac{cd^{3}\vec{p}_{n}}{\left(2\pi\right)^{3} 2E_{n}} \right) \right] \times \left(2\pi\right)^{4} \delta^{4} \left(p_{1} - p_{2} - \dots - p_{n} \right)$$

(note that some the factors of \hbar have vanished.....)

δ-function ensures overall energy and momentum conservation. This is the differential decay rate for the case in which the momentum of final state particle *i* lies in the range $d^3 \vec{p}_i$ about \vec{p}_i for all *i*.

Need to integrate over all outgoing momenta. Decaying particle is at rest. For the two-body decay $1 \rightarrow 2+3$

$$\Gamma = \frac{S}{\hbar m_1} \left(\frac{c}{4\pi}\right)^2 \frac{1}{2} \int \frac{|\mathcal{M}|^2}{E_2 E_3} \delta^4 (p_1 - p_2 - p_3) d^3 \vec{p}_2 d^3 \vec{p}_3$$

In general the amplitude \mathcal{M} can depend on both \vec{P}_2 and \vec{P}_3 , so it cannot be pulled out of the integration. For two body decays, the integration can be done.

Two Body Decay to Massless Particles

spinless

Consider the decay of a massive \wedge particle (mass $m_1 = m$) into a final state consisting of two massless particles ($m_2 = m_3 = 0$) so $E_2 = |\vec{p}_2|c$ and $E_3 = |\vec{p}_3|c$.

Start with assumption that $\mathcal{M} = \mathcal{M}(\vec{p}_2, \vec{p}_3)$ and find Γ :

It should be clear from the start that $\vec{p}_2 = -\vec{p}_3$ so that what we really have here is $\mathcal{M} = \mathcal{M}(\vec{p}_2)$. One can also note that in the absence of spin, there is no direction against which to measure components of \vec{P}_2 , so expect that $\mathcal{M} = \mathcal{M}(|\vec{p}_2|)$.

Rewrite the δ -function in terms of $E_1 = mc^2$, $\vec{p}_1 = 0$, $E_2 = \left| \vec{p}_2 \right| c$, $E_3 = \left| \vec{p}_3 \right| c$

$$\delta^{4}(p_{1}-p_{2}-p_{3}) = \delta\left(mc - \frac{E_{2}}{c} - \frac{E_{3}}{c}\right)\delta^{3}(-\vec{p}_{2} - \vec{p}_{3})$$

$$\Gamma = \frac{S}{\hbar m} \left(\frac{1}{4\pi}\right)^2 \frac{1}{2} \int \frac{\left|\mathcal{M}\right|^2}{\left|\vec{p}_2\right| \left|\vec{p}_3\right|} \delta\left(mc - \left|\vec{p}_2\right| - \left|\vec{p}_3\right|\right) \delta^3\left(-\vec{p}_2 - \vec{p}_3\right) d^3\vec{p}_2 d^3\vec{p}_3$$

Now integrate over \vec{P}_3 using $\delta^3 \left(-\vec{p}_2 - \vec{p}_3\right)$

$$\Gamma = \frac{S}{2(4\pi)^2 \hbar m} \int \frac{|\mathcal{M}|^2}{\left|\vec{p}_2\right|^2} \delta\left(mc - 2\left|\vec{p}_2\right|\right) d^3 \vec{p}_2 \qquad \qquad \mathcal{M} = \mathcal{M}(\left|\vec{p}_2\right|).$$

N.B. $d^{3}\vec{p}_{2} = \left|\vec{p}_{2}\right|^{2} d\left|\vec{p}_{2}\right| \sin\theta d\theta d\phi$ $\int \sin\theta d\theta d\phi = 4\pi$ $\delta(kx) = \frac{1}{|k|}\delta(x)$

$$\Gamma = \frac{S}{8\pi\hbar m} \int_{0}^{\infty} |\mathcal{M}|^{2} \frac{1}{2} \delta \left(\left| \vec{p}_{2} \right| - \frac{mc}{2} \right) d \left| \vec{p}_{2} \right| \qquad \Rightarrow \qquad \Gamma = \frac{S}{16\pi\hbar m} |\mathcal{M}|^{2}$$

where \mathcal{M} is evaluated at $\left| \vec{p}_{2} \right| = \frac{mc}{2} \qquad \left[E_{2} = E_{3} = \frac{mc^{2}}{2} \Rightarrow \left| \vec{p}_{2} \right| = \left| \vec{p}_{3} \right| = \frac{mc}{2} \right]$

We have not discussed S yet, but it is a statistical factor related to the identity of the final state particles. If they are different, then S=1; if they are identical, then S=1/2 (we will see why later).

See §6.2.1.1 for the general case with arbitrary m_2 , m_3 .

$$\Gamma = \frac{S\left|\vec{p}\right|}{8\pi\hbar m_1^2 c} \left|\mathcal{M}\right|^2$$

Where \mathcal{M} is again evaluated at the momentum dictated by conservation of energy and momentum (here \vec{p}).

N.B. for
$$m_2 = m_3 = 0$$
 (as before)

$$\left|\vec{p}\right| = \frac{m_1 c}{2} \quad \Rightarrow \quad \Gamma = \frac{Sm_1 c}{16\pi\hbar m_1^2 c} \left|\mathcal{M}\right|^2 \quad \Rightarrow \quad \Gamma = \frac{S}{16\pi\hbar m_1} \left|\mathcal{M}\right|^2$$

as we just saw.

So for 1
$$\rightarrow$$
 2 + 3, we have for the general case that $\Gamma = \frac{S |\vec{p}|}{8\pi \hbar m_1^2 c} |\mathcal{M}|^2$

Where \vec{p} is the momentum dictated by the conservation of energy and momentum and the matrix element \mathcal{M} is evaluated at that momentum.

Note the "phase space factor" $|\vec{p}|$. This is (for the case of two body decays) the origin of the "rule of thumb" stating that, for a given system, the decay with the largest energy release will proceed more rapidly (*e.g.* has a larger width Γ).

But again, this does not typically trump dynamical effects. For instance, we have seen that $\Gamma(D^0 \to K^- \pi^+) > \Gamma(D^0 \to \pi^- \pi^+)$ because the effect of the CKM matrix elements dominates:

$$\frac{\Gamma(D^0 \to \pi^- \pi^+)}{\Gamma(D^0 \to K^- \pi^+)} = \frac{\left|V_{cd}\right|^2}{\left|V_{cs}\right|^2} \left(\frac{p_{\pi^-}}{p_{K^-}}\right) \approx \lambda^2 \left(\frac{922}{861}\right)$$

e.g. kinematic enhancement factor is 1.07 but dynamical suppression factor is $\lambda^2 \sim .05$.

Golden Rule for Scattering

For the process $1 + 2 \rightarrow 3 + 4 + \dots + N$

$$d\sigma = \left|\mathcal{M}\right|^{2} \frac{\hbar^{2}S}{4\sqrt{\left(p_{1} \cdot p_{2}\right)^{2} - \left(m_{1}m_{2}c^{2}\right)^{2}}} \left[\left(\frac{cd^{3}\vec{p}_{3}}{\left(2\pi\right)^{3}2E_{3}}\right) \left(\frac{cd^{3}\vec{p}_{4}}{\left(2\pi\right)^{3}2E_{4}}\right) \dots \left(\frac{cd^{3}\vec{p}_{n}}{\left(2\pi\right)^{3}2E_{n}}\right) \right] \times \left(2\pi\right)^{4}\delta^{4}\left(p_{1} + p_{2} - \dots - p_{n}\right)$$

Kinematic term in front of phase space factor now contains p_1 and p_2 . [Was the mass of the decaying particle in the case of the GR for decays.] Phase space factors look the same.

This describes the differential cross-section for the process in which the momentum of the *i*th final state particle is in the range $d^3 \vec{p}_i$ about \vec{p}_i for all *i*.

Typically, one might measure only a single quantity, such as the angle at which particle 3 emerges from the scattering. Need to integrate over everything else.

We will only deal with the case of two body scattering $1 + 2 \rightarrow 3 + 4$. Consider this process in the CM frame where

$$\vec{p}_1 = -\vec{p}_2$$
 $p_1 \cdot p_2 = \frac{E_1 E_2}{c^2} + \left| \vec{p}_1 \right|^2$

Two Body Scattering in the CM frame

One part of problem 6.7 (may be on your next assignment) asks you to show that, for 2-body scattering in the CM frame,

$$\sqrt{\left(p_{1} \cdot p_{2}\right)^{2} - \left(m_{1} m_{2} c^{2}\right)^{2}} = \left(E_{1} + E_{2}\right) \left|\vec{p}_{1}\right| / c$$

In this case we have

$$d\sigma = \left(\frac{\hbar}{8\pi}\right)^{2} \frac{S\left|\mathcal{M}\right|^{2}c}{(E_{1} + E_{2})\left|\vec{p}_{1}\right|} \frac{d^{3}\vec{p}_{3}}{E_{3}} \frac{d^{3}\vec{p}_{4}}{E_{4}} \times \delta^{4}\left(p_{1} + p_{2} - p_{3} - p_{4}\right)$$
$$\frac{1}{4}\left(\frac{1}{(2\pi)^{3}} \cdot \frac{1}{2}\right)\left(\frac{1}{(2\pi)^{3}} \cdot \frac{1}{2}\right)(2\pi)^{4} = \frac{1}{16}\frac{(2\pi)^{4}}{(2\pi)^{6}} = \frac{1}{64\pi^{2}} = \frac{1}{(8\pi)^{2}}$$

Rewrite the \delta-function using $\vec{p}_1 = -\vec{p}_2$ This will

This will enforce
$$\vec{P}_3 = -\vec{P}_4$$
 in the final state.

1

$$\delta^{4}(p_{1}+p_{2}-p_{3}-p_{4}) = \delta\left(\frac{E_{1}+E_{2}-E_{3}-E_{4}}{c}\right)\delta^{3}(-\vec{p}_{3}-\vec{p}_{4})$$

Use
$$E_3 = c\sqrt{m_3^2c^2 + |\vec{p}_3|^2}$$
 and $E_4 = c\sqrt{m_4^2c^2 + |\vec{p}_4|^2}$ and integrate over $d^3\vec{p}_4$.

$$d\sigma = \left(\frac{\hbar}{8\pi}\right)^{2} \frac{S\left|\mathcal{M}\right|^{2}c}{(E_{1} + E_{2})\left|\vec{p}_{1}\right|} \times \frac{\delta\left(\frac{E_{1} + E_{2}}{c} - \sqrt{m_{3}^{2}c^{2} + \left|\vec{p}_{3}\right|^{2}} - \sqrt{m_{4}^{2}c^{2} + \left|\vec{p}_{3}\right|^{2}}\right)}{\sqrt{m_{3}^{2}c^{2} + \left|\vec{p}_{3}\right|^{2}}\sqrt{m_{4}^{2}c^{2} + \left|\vec{p}_{3}\right|^{2}}} d^{3}\vec{p}_{3}$$

At this stage when we discussed decays we concluded that $\mathcal{M} = \mathcal{M}(|\vec{p}_3|)$ but in this case we can have dependence also on the direction of \vec{p}_3 since the trajectory of the incoming particles defines an axis against which we can measure this (*i.e.* instead of having only one scalar $|\vec{p}_3|$ we also have the scalar $\vec{p}_1 \cdot \vec{p}_3$). So we cannot perform the angular integration (at least not over θ), so defining $d^3\vec{p}_3 = \rho^2 d\rho d\Omega$ $\rho \equiv |\vec{p}_3|$

$$\frac{d\sigma}{d\Omega} = \left(\frac{\hbar}{8\pi}\right)^2 \frac{Sc}{(E_1 + E_2)|\vec{p}_1|} \int_{0}^{\infty} |\mathcal{M}|^2 \frac{\delta\left(\frac{E_1 + E_2}{c} - \sqrt{m_3^2 c^2 + \rho^2} - \sqrt{m_4^2 c^2 + \rho^2}\right)}{\sqrt{m_3^2 c^2 + \rho^2} \sqrt{m_4^2 c^2 + \rho^2}} \rho^2 d\rho$$

To do the integration, use
$$E = c \left[\left(m_3^2 c^2 + \rho^2 \right)^{1/2} + \left(m_4^2 c^2 + \rho^2 \right)^{1/2} \right]$$

(this is the total energy of the final state or the initial state)

$$\frac{d\sigma}{d\Omega} = \left(\frac{\hbar}{8\pi}\right)^2 \frac{Sc^2}{(E_1 + E_2)|\vec{p}_1|} |\mathcal{M}|^2 \left(\frac{\rho_0}{E_1 + E_2}\right) \qquad \rho_0 \equiv \begin{cases} \text{Magnitude of final state momentum}} \\ \text{dictated by conservation of energy and} \\ \text{momentum} \end{cases}$$

$$\frac{d\sigma}{d\Omega} = \left(\frac{\hbar c}{8\pi}\right)^2 \frac{S\left|\mathcal{M}\right|^2}{(E_1 + E_2)^2} \frac{\left|\vec{p}_f\right|}{\left|\vec{p}_i\right|} \qquad \text{where } \vec{p}_i \text{ and } \vec{p}_f \text{ are the initial and final state momenta (e.g. } \vec{p}_1 \text{ and } \vec{p}_3\text{).}$$

Now look at Griffiths Problem 6.8: Elastic scattering of the form A+B \rightarrow A+B in the lab frame. This is like 1+2 \rightarrow 3+4 with $m_1=m_3=m_A$, $m_2=m_4=m_B$ (with particle 2 initially at rest)

Consider the case where $m_B c^2 >> E_A$ so that we can neglect the recoil momentum of particle B.

The general case is problem 6.10 which may be on your next assignment.

$$d\sigma = \left|\mathcal{M}\right|^{2} \frac{\hbar^{2}S}{4\sqrt{\left(p_{1} \cdot p_{2}\right)^{2} - \left(m_{1}m_{2}c^{2}\right)^{2}}} \left[\left(\frac{cd^{3}\vec{p}_{3}}{\left(2\pi\right)^{3}2E_{3}}\right) \left(\frac{cd^{3}\vec{p}_{4}}{\left(2\pi\right)^{3}2E_{4}}\right) \right] \times \left(2\pi\right)^{4}\delta^{4}\left(p_{1} - p_{2} - p_{3} - p_{4}\right)$$

Note that this is Lorentz invariant. Now choose a particular reference frame:

Evaluate $\sqrt{(p_1 \cdot p_2)^2 - (m_1 m_2 c^2)^2}$ in the lab frame:

This may be on your next assignment.....

$$\Rightarrow \sqrt{\left(p_1 \cdot p_2\right)^2 - \left(m_1 m_2 c\right)^2} = m_2 c \left|\vec{p}_1\right|$$

$$d\boldsymbol{\sigma} = \left(\frac{\hbar c}{8\pi}\right)^2 \frac{\left|\mathcal{M}\right|^2}{m_B c \left|\vec{p}_1\right|} \left(\frac{d^3 \vec{p}_3}{E_3}\right) \left(\frac{d^3 \vec{p}_4}{E_4}\right) \times \delta^4 \left(p_1 + p_2 - p_3 - p_4\right)$$

Rewrite the δ -function

17

$$\delta^{4}(p_{1}+p_{2}-p_{3}-p_{4}) = \delta\left(\frac{E_{1}}{c}+m_{B}c-\frac{E_{3}}{c}-m_{B}c\right)\delta^{3}(\vec{p}_{1}+\vec{0}-\vec{p}_{3}-\vec{p}_{4}) = c\delta(E_{1}-E_{3})\delta^{3}(\vec{p}_{1}-\vec{p}_{3}-\vec{p}_{4})$$
18

Integrate over $d^3 \vec{p}_4$. In this case the integration doesn't do anything since we are ignoring \vec{p}_4 but the $\delta^3(\vec{p}_1 - p_3 - p_4)$ disappears with the integration.

$$d\sigma = \left(\frac{\hbar c}{8\pi}\right)^2 \frac{\left|\mathcal{M}\right|^2}{m_B c \left|\vec{p}_1\right|} \left(\frac{d^3 \vec{p}_3}{E_3}\right) \left(\frac{1}{E_4}\right) \times c\delta\left(E_1 - E_3\right)$$
$$= m_B c^2 \text{ in the limit in which we are working}$$

Now integrate over $d^3 \vec{p}_3$: $d^3 \vec{p}_3 = \left| \vec{p}_3 \right|^2 d \left| \vec{p}_3 \right| d\Omega$ [again, can't do the angular part]

$$\frac{d\sigma}{d\Omega} = \left(\frac{\hbar c}{8\pi}\right)^2 \frac{1}{m_B^2 c^2 \left|\vec{p}_1\right|} \int_0^\infty \left|\mathcal{M}\right|^2 \frac{\left|\vec{p}_3\right|^2 d\left|\vec{p}_3\right|}{E_3} \times \delta\left(E_1 - E_3\right)$$

Write this in terms of E_3 so that we can integrate using the δ -function:

$$E_{3}^{2} = \left|\vec{p}_{3}\right|^{2}c^{2} + m_{3}^{2}c^{4} \qquad 2E_{3}dE_{3} = 2\left|\vec{p}_{3}\right|c^{2}d\left|\vec{p}_{3}\right| \implies d\left|\vec{p}_{3}\right| = \frac{E_{3}dE_{3}}{\left|\vec{p}_{3}\right|c^{2}} \quad \left|\vec{p}_{3}\right|^{2} = \frac{E_{3}^{2}}{c^{2}} - m_{3}^{2}c^{2}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = \left(\frac{\hbar c}{8\pi}\right)^2 \frac{1}{m_B^2 c^2 |\vec{p}_1|} \int_{m_A c^2}^{\infty} |\mathcal{M}|^2 \frac{\left(E_3^2 - m_3^2 c^4\right)}{c^2 E_3} \frac{E_3 dE_3}{c \left(E_3^2 - m_3^2 c^4\right)^{1/2}} \times \delta(E_1 - E_3)$$
$$= \left(\frac{\hbar c}{8\pi}\right)^2 \frac{1}{m_B^2 c^2} \frac{\left(E_1^2 - m_A^2 c^4\right)^{1/2}}{|\vec{p}_1| c^3} |\mathcal{M}|^2$$
$$= \left(\frac{\hbar}{8\pi m_B c}\right)^2 |\mathcal{M}|^2 \qquad \text{since} \qquad \frac{\left(E_1^2 - m_A^2 c^4\right)^{1/2}}{|\vec{p}_1| c} = \frac{\sqrt{|\vec{p}_1|^2 c^2}}{|\vec{p}_1| c} = 1$$

As before $|\mathcal{M}|^2$ must be evaluated at the value of the final state momentum that is dictated by the conservation of energy and momentum,