

# Searches for Prompt Light Gravitino Production

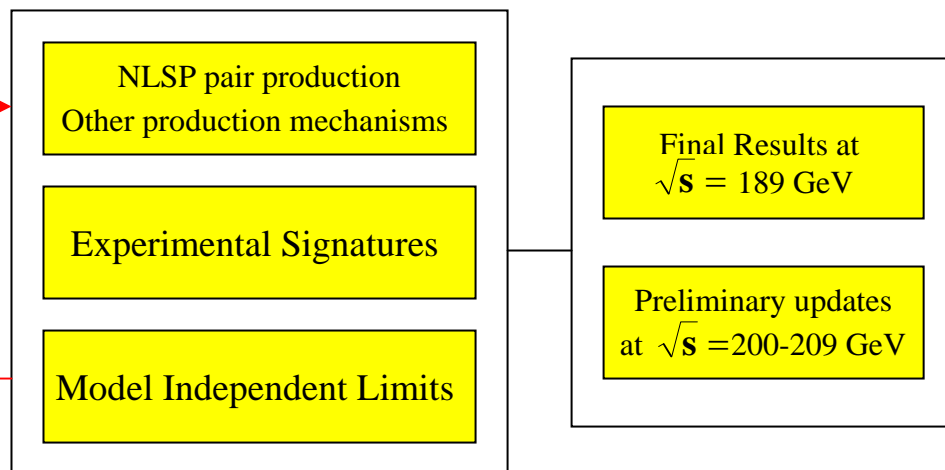


Peter Krieger, Carleton University, Ottawa Canada, for the OPAL Collaboration

- SUSY models with light gravitinos
- Gauge-Mediated Supersymmetry Breaking

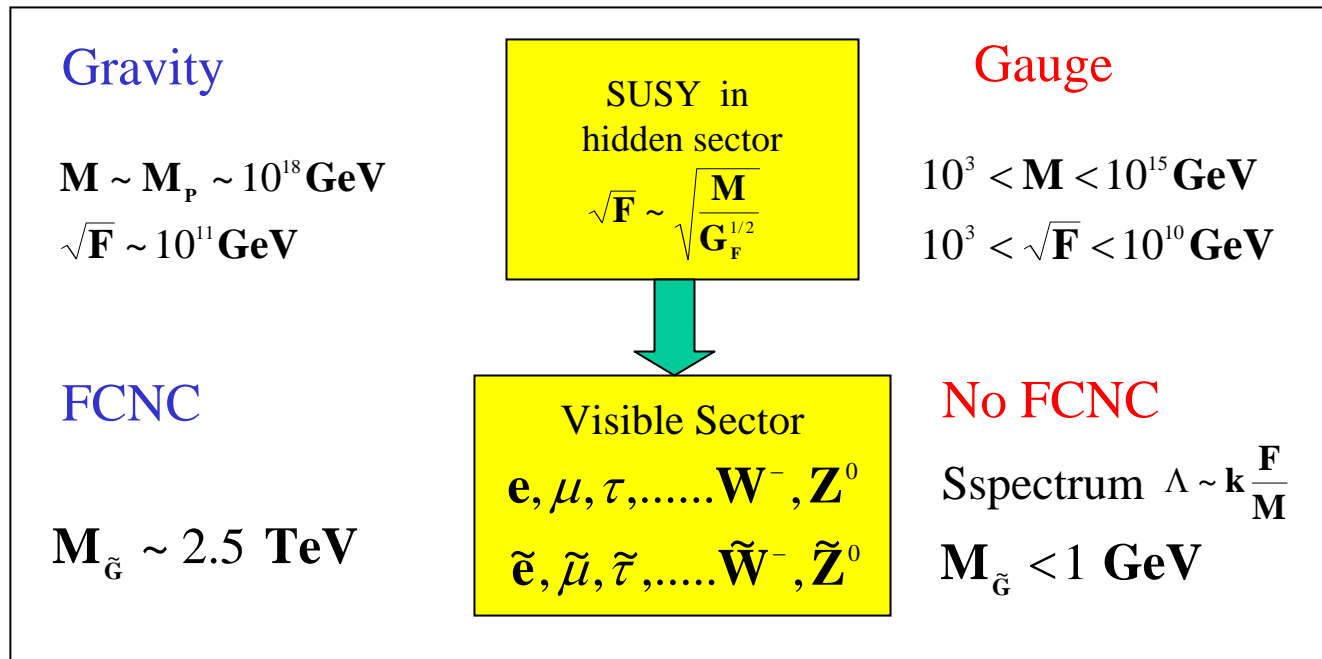
- Neutralino NLSP Scenario
- Slepton NLSP Scenario

- GMSB scan
- Gravitino Pair Production
- GMSB signatures with lifetime
- Summary



# SUSY models with a Light Gravitino

- Some SUSY models predict that the LSP is an almost massless gravitino (models with gauge-mediated supersymmetry breaking, no-scale supergravity)
- Richest phenomenology is from GMSB models



**R-parity conservation assumed throughout this talk**

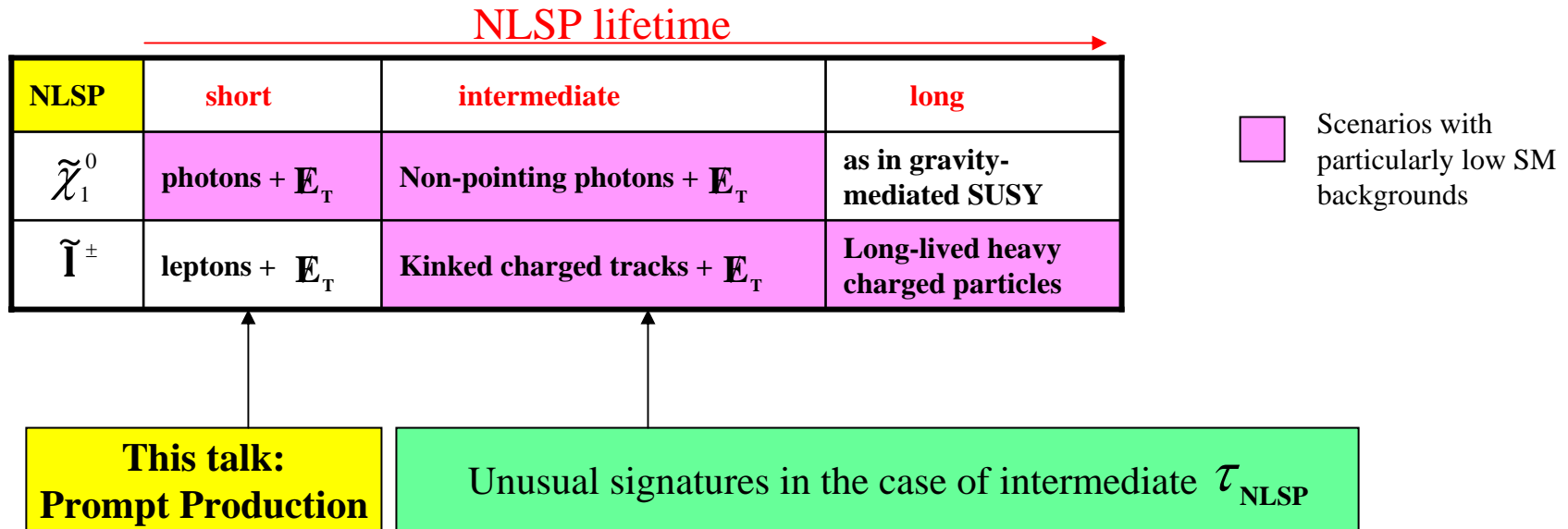


# Experimental Signatures

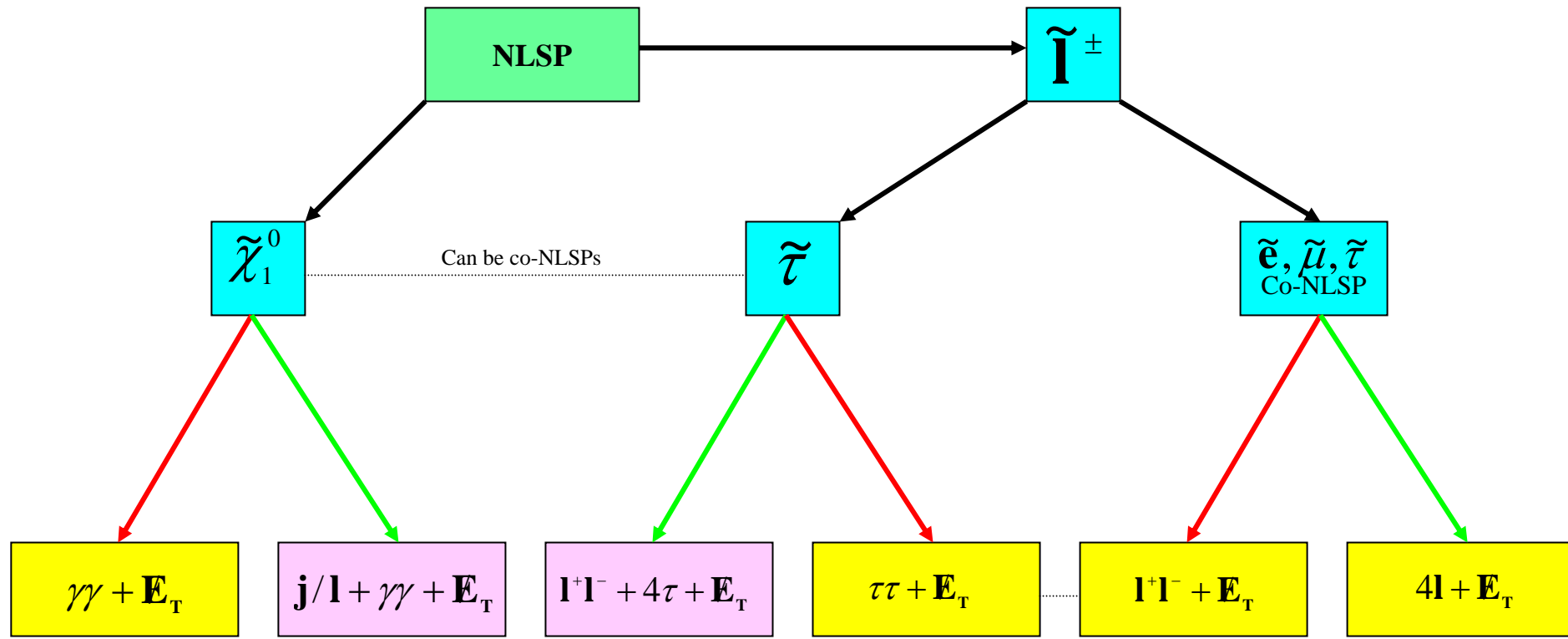
Scale of SUSY breaking  $\sqrt{\mathbf{F}}$  determines  $\tilde{\mathbf{G}}$  mass and NLSP lifetime

$$M_{\tilde{\mathbf{G}}} \cong 2.37 \times 10^{-2} \left( \frac{\sqrt{\mathbf{F}}}{10 \text{ TeV}} \right)^2 \text{ eV} \quad c\tau(\text{NLSP}) \cong \left( \frac{M_{\text{NLSP}}}{100 \text{ GeV}} \right)^{-5} \left( \frac{\sqrt{\mathbf{F}}}{10 \text{ TeV}} \right)^4 \mu\text{m}$$

Phenomenology dictated by the NLSP (usually either  $\tilde{\chi}_1^0$  or  $\tilde{\mathbf{I}}^\pm$ ) and on NLSP lifetime (decay length w.r.t detector volume)



# Experimental Signatures



Contriubute to exclusion regions from GMSB scan



**NLSP Pair Production**



**Other sparticle pair production,  
associated pair production**



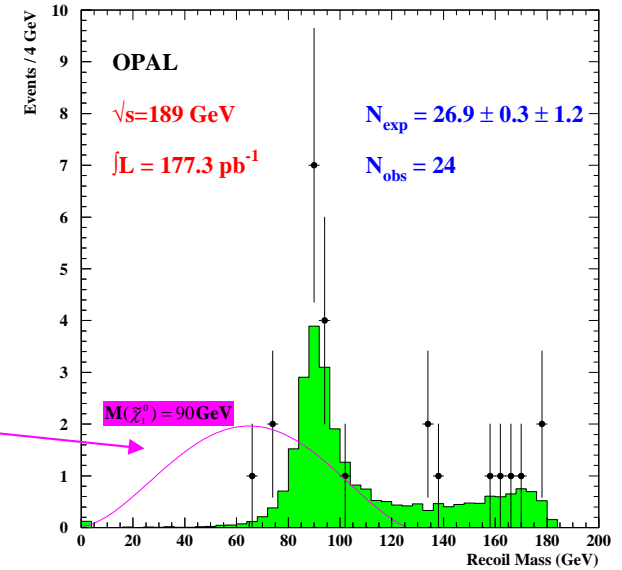
# Neutralino NLSP Pair Production

## Acoplanar photons

$$e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow (\tilde{G}\gamma)(\tilde{G}\gamma) \longrightarrow \gamma\gamma + \mathbf{E}_T$$

Irreducible SM background from  $e^+e^- \rightarrow \nu\bar{\nu}\gamma\gamma$

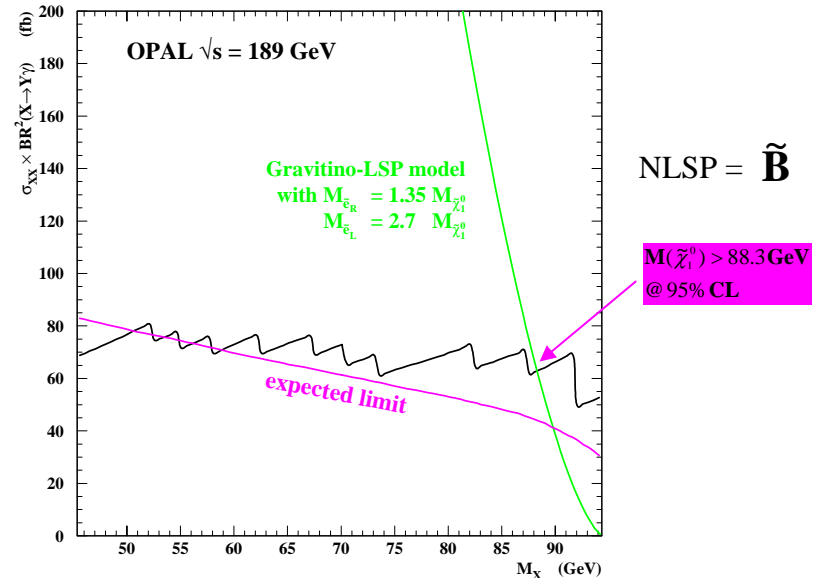
Large fraction of signal contribution expected at  $M_{\text{recoil}}$  values well below the Z peak



Selected events can be classified according to maximum mass for which they remain kinematically consistent with the above decay sequence ( $M_X^{\text{max}}$ )

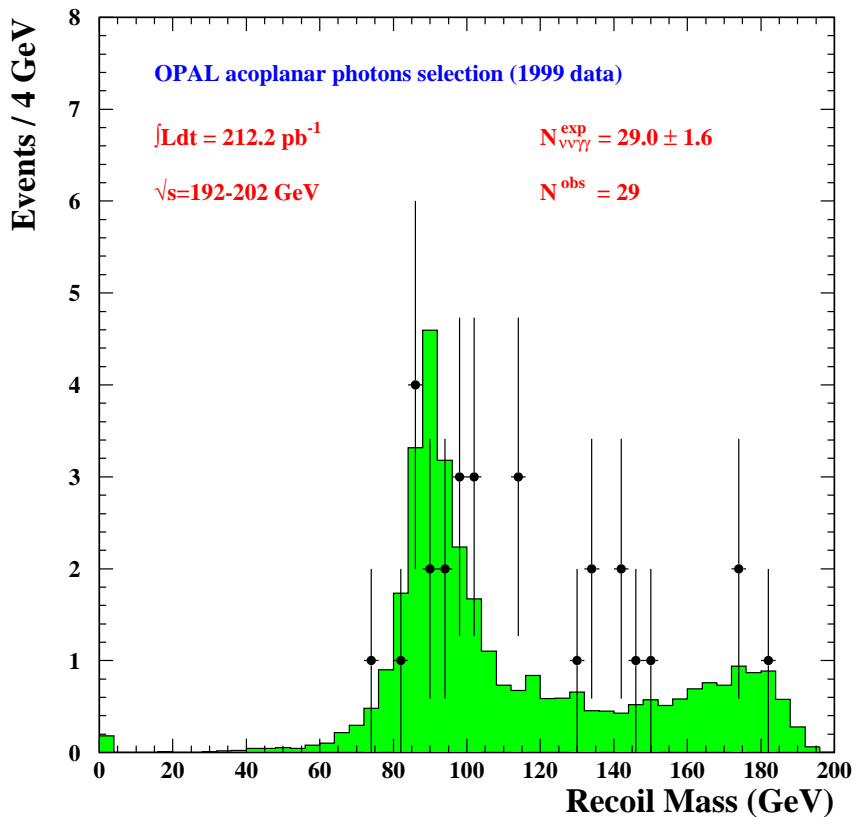
OPAL published results from  $\sqrt{s} = 189 \text{ GeV}$

$M_X(\text{GeV})$	Selection Efficiency (%)	Selection efficiency with $M_X^{\text{max}} > M_X - 5 \text{ GeV}$	$N_{\text{data}}$	$N_{\nu\bar{\nu}\gamma\gamma}$
50	70.2 +/- 1.2	67.7 +/- 1.3	14	13.67 +/- 0.20
60	74.0 +/- 1.1	71.1 +/- 1.2	11	10.05 +/- 0.18
70	71.4 +/- 1.2	69.2 +/- 1.2	9	7.22 +/- 0.15
80	72.3 +/- 1.2	68.7 +/- 1.4	7	4.81 +/- 0.13
90	71.3 +/- 1.2	67.5 +/- 1.3	5	2.40 +/- 0.09
94	72.2 +/- 1.2	70.4 +/- 1.2	3	1.34 +/- 0.07

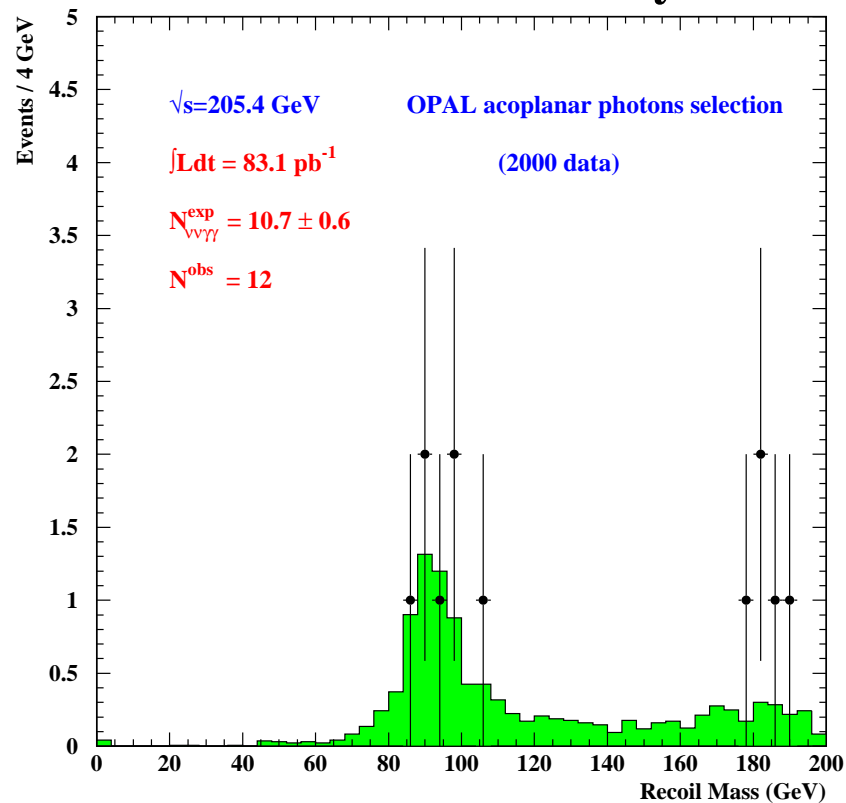


# Update of OPAL Acoplanar Photons Selection

## OPAL Preliminary



## OPAL Preliminary



No evidence for non-Standard Model contributions (especially in low recoil-mass region)

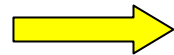


# Other Channels: Neutralino NLSP

## Other production channels:

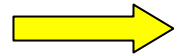
$$e^+e^- \rightarrow \tilde{I}^+\tilde{I}^- \rightarrow (I^+\tilde{\chi}_1^0)(I^-\tilde{\chi}_1^0) \rightarrow (I^+\gamma\tilde{G})(I^-\gamma\tilde{G}) \longrightarrow \boxed{I^+I^-\gamma\gamma + \mathbf{E}_T}$$

$$e^+e^- \rightarrow \chi_1^+\chi_1^- \rightarrow (W^{(*)+}\tilde{\chi}_1^0)(W^{(*)-}\tilde{\chi}_1^0) \rightarrow (W^{(*)+}\gamma\tilde{G})(W^{(*)-}\gamma\tilde{G})$$



Leptons and/or jets +  $\gamma\gamma$  +  $\mathbf{E}_T$

$$e^+e^- \rightarrow \chi_2^0\chi_1^0 \rightarrow Z^{(*)}\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow (Z^{(*)}\gamma\tilde{G})(\gamma\tilde{G})$$



Leptons or jets +  $\gamma\gamma$  +  $\mathbf{E}_T$

Separate selections into high/low multiplicity parts (depending on W,Z final states)

**Efficiencies depend on mass-difference of NLSP and produced particles**



Limits in  $[M(\mathbf{X}), M(\tilde{\chi}_1^0)]$  plane:  $\mathbf{X} = \tilde{I}^+, \tilde{\chi}_1^+, \tilde{\chi}_2^0$



# Other Channels: Neutralino NLSP

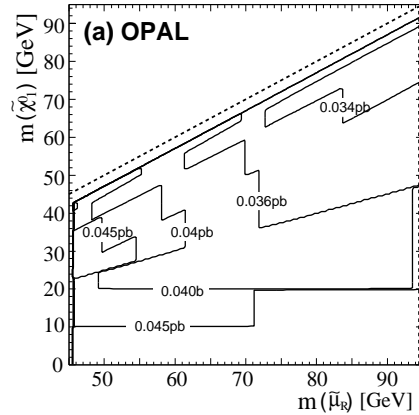
~ model independent cross-section limits at  $\sqrt{s} = 189 \text{ GeV}$

$$\sigma(e^+e^- \rightarrow \tilde{\mu}^+ \tilde{\mu}^-)$$

$$\tilde{\mu}^\pm \rightarrow \mu^\pm \tilde{\chi}_1^0 \rightarrow \mu^\pm \gamma \tilde{G}$$

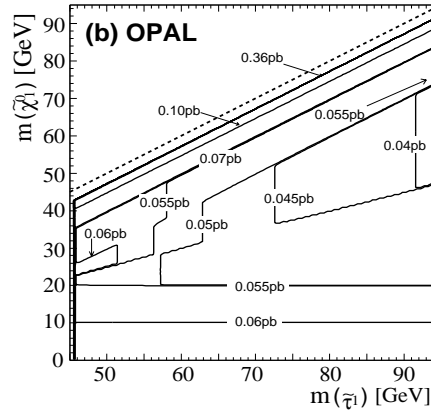
Efficiency ~ 30-50%

$$\mu^+ \mu^- \gamma\gamma + \mathbf{E}_T$$



34 - 45 fb

40 - 180 fb



40 - 360 fb

40 - 130 fb

$$\sigma(e^+e^- \rightarrow \tilde{\tau}^+ \tilde{\tau}^-)$$

$$\tilde{\tau}^\pm \rightarrow \tau^\pm \tilde{\chi}_1^0 \rightarrow \tau^\pm \gamma \tilde{G}$$

Efficiency ~ 20-40%

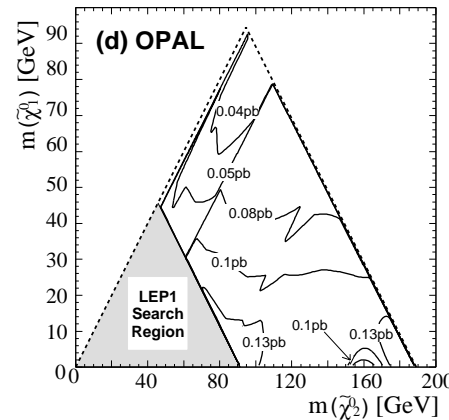
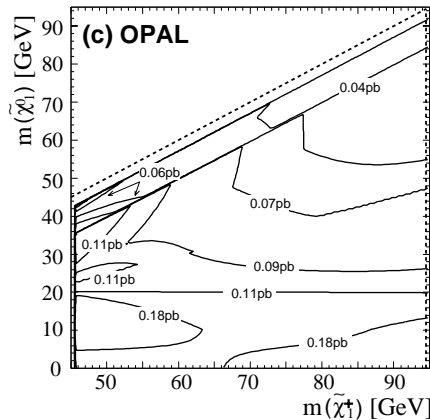
$$\tau^+ \tau^- \gamma\gamma + \mathbf{E}_T$$

$$\sigma(e^+e^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-)$$

$$\tilde{\chi}_1^\pm \rightarrow W^{(*)\pm} \tilde{\chi}_1^0 \rightarrow W^{(*)\pm} \gamma \tilde{G}$$

Efficiency ~ 20-50%

$$W^{(*)+} W^{(*)-} \gamma\gamma + \mathbf{E}_T$$



$$\sigma(e^+e^- \rightarrow \tilde{\chi}_2^0 \tilde{\chi}_1^0)$$

$$\tilde{\chi}_2^0 \rightarrow Z^{(*)} \tilde{\chi}_1^0 \rightarrow Z^{(*)} \gamma \tilde{G}$$

Efficiency ~ 20-50%

$$Z^{(*)} \gamma\gamma + \mathbf{E}_T$$

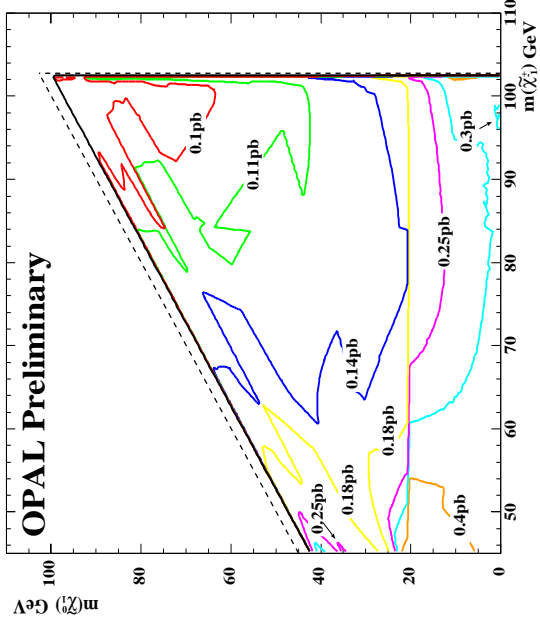




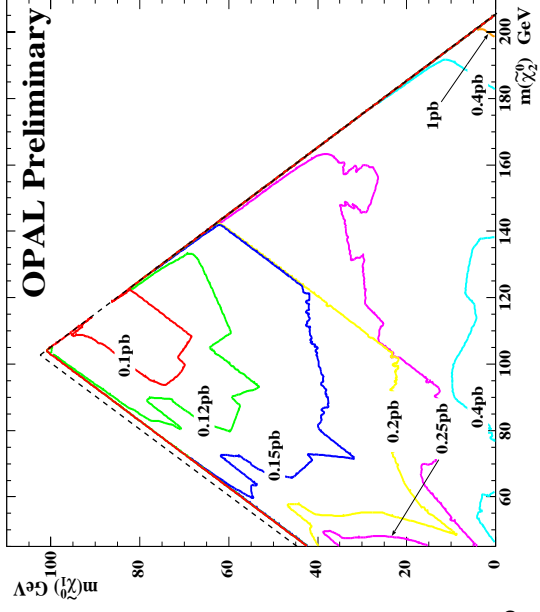
# Other Channels: Neutralino NLSP (Update)

~ model independent cross-section limits at  $\sqrt{s} = 205.5 \text{ GeV}$

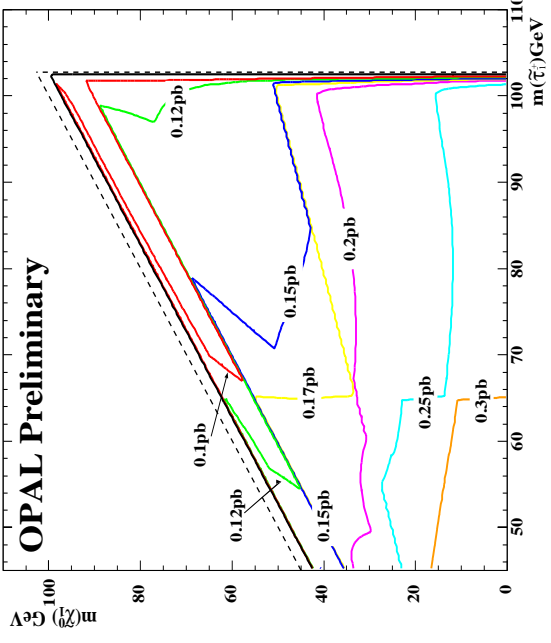
$$\mathbf{W}^{(*)+} \mathbf{W}^{(*)-} \gamma + \mathbf{E}_\tau$$



$$\mathbf{Z}^{(*)} \gamma + \mathbf{E}_\tau$$



$$\tau^+ \tau^- \gamma \gamma + \mathbf{E}_\tau$$



For slepton co-NLSP scenario with large  $\mathbf{M}(\tilde{\mathbf{I}}) - \mathbf{M}(\tilde{\chi}_1^0)$

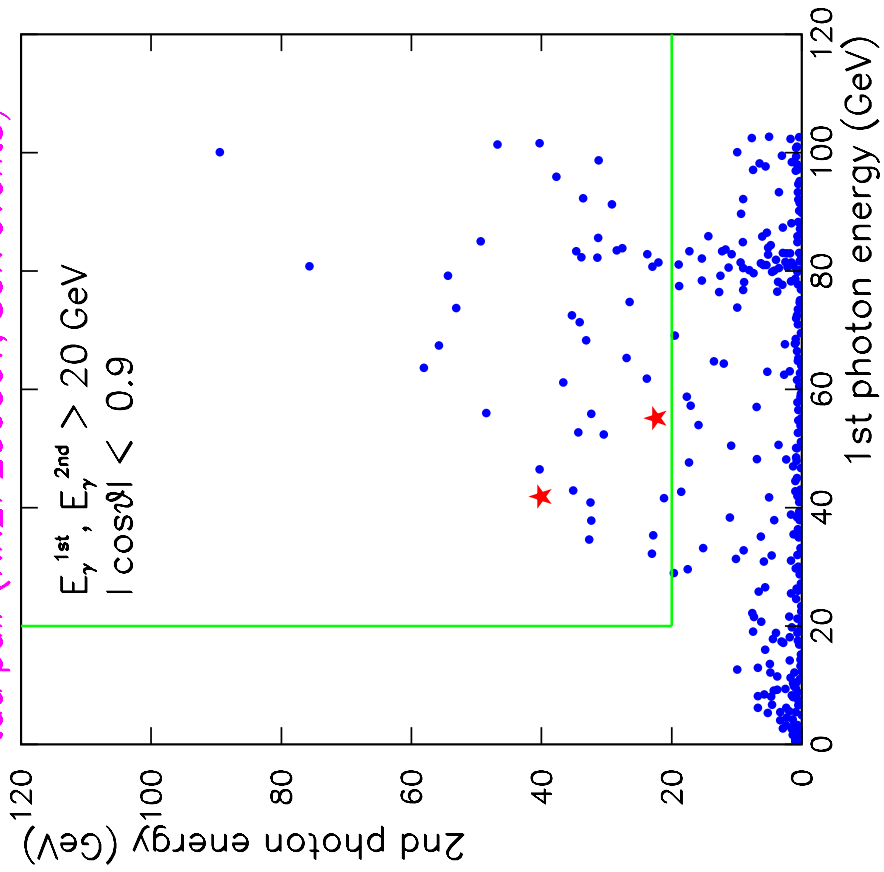
see 4 events where 1.3 +/- 0.3 are expected from SM

$\mathbf{M}(\tilde{\mathbf{I}}) - \mathbf{M}(\tilde{\chi})$	N(data)	$N_{\text{SM}}(\text{MC})$
3 – 10 GeV	1	0.5 +/- 0.2
10 GeV – $\mathbf{M}(\tilde{\mathbf{I}})/2$	2	0.7 +/- 0.2
$\mathbf{M}(\tilde{\mathbf{I}})/2 - \mathbf{M}(\tilde{\mathbf{I}})$	4	1.3 +/- 0.3



# $\mathbf{I}^+\mathbf{I}^-\gamma\gamma$ Events

tau pair (KK2f 206GeV, 80K events)



2/4 events have a single high-energy photon consistent with radiative return to the  $Z^0$

Other 2 each have 2 high energy photons  
**(55 GeV, 22 GeV) (42 GeV, 40 GeV)**

Background in this kinematic regime is essentially  $\mathbf{e}^+\mathbf{e}^- \rightarrow \tau^+\tau^-\gamma\gamma$

Prob of two events with 2 photons  $> 20 \text{ GeV}$  estimated from  $\mathbf{e}^+\mathbf{e}^- \rightarrow \tau^+\tau^-\gamma\gamma$  Monte Carlo (assuming only SM background)

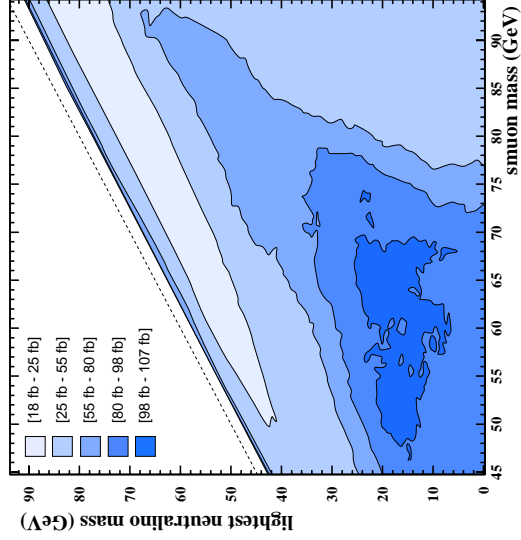
**Prob  $\sim 3 \times 10^{-4}$**



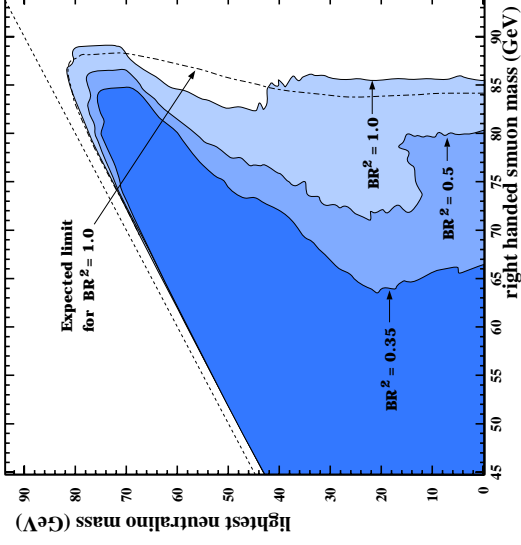


# Limits from Acoplanar Leptons

OPAL



OPAL

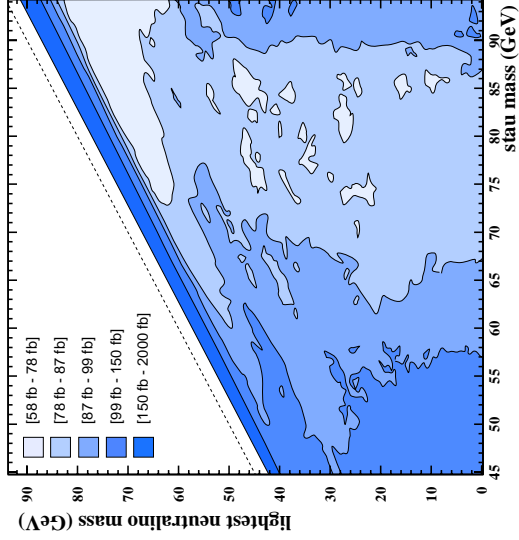


Limits along  $M_{LSP} = 0$  axis applicable to light gravitino scenario

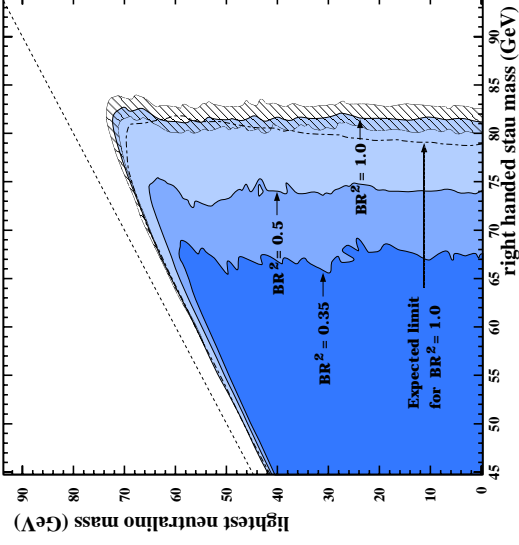
$$\sqrt{s} = 183 - 189 \text{ GeV}$$

$$\int \mathbf{L} = 237.4 \text{ pb}^{-1}$$

OPAL



OPAL



95% CL lower mass limits  
(for BR=1.0)

$$M(\tilde{\mu}_R) > 85.4 \text{ GeV}$$

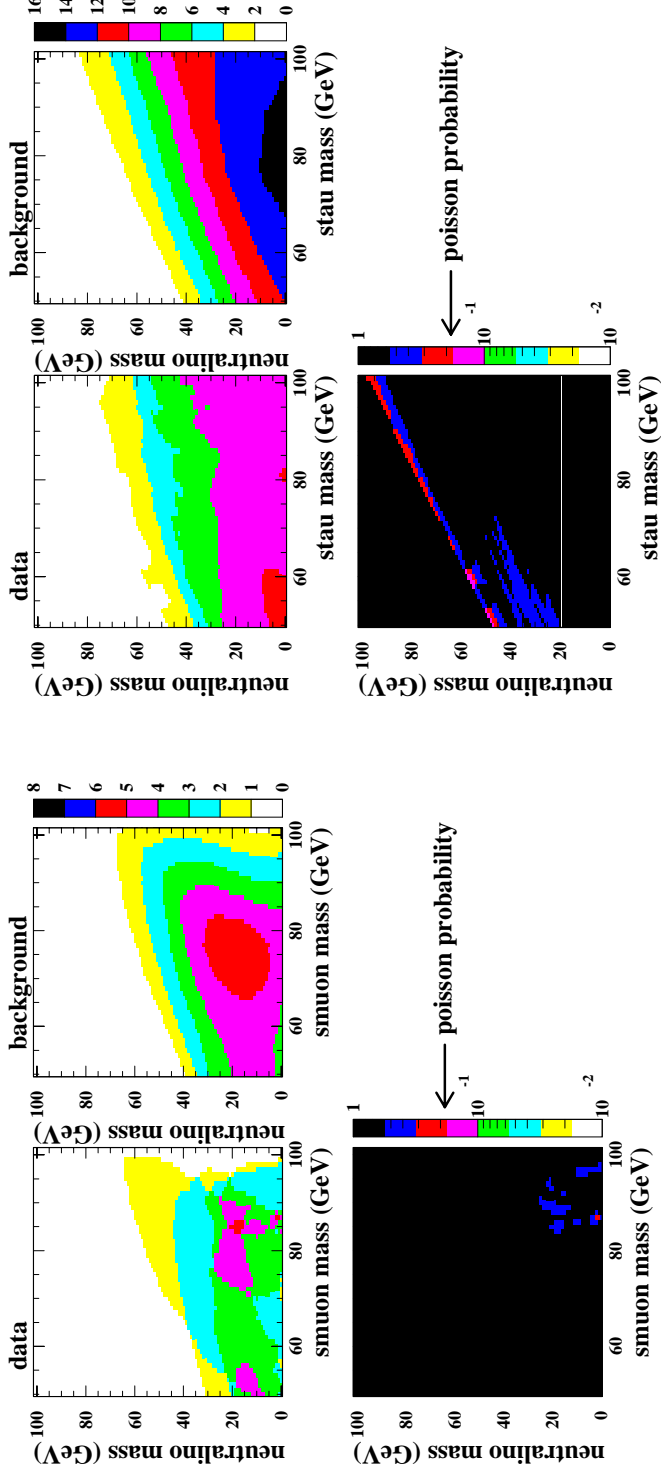
$$M(\tilde{\tau}_R) > 81.1 \text{ GeV}$$

$$M(\tilde{\tau}_1) > 80.0 \text{ GeV}$$

for any degree of stau mixing



# Slepton NLSP Pair Production @ 205.5 GeV



Preliminary results from OPAL 2000 data

Mass	selectrons			smuons			staus		
	N	$N_{SM}(MC)$	CL(%)	N	$N_{SM}(MC)$	CL(%)	N	$N_{SM}(MC)$	CL(%)
50 GeV	15	16.51 +/- 0.25	47.6	2	3.68 +/- 0.09	89.9	10	11.99 +/- 0.23	34.5
70 GeV	20	20.45 +/- 0.25	41.8	1	4.37 +/- 0.13	74.3	8	14.16 +/- 0.27	56.5
90 GeV	11	11.22 +/- 0.17	65.8	2	3.11 +/- 0.09	61.3	8	14.25 +/- 0.26	77.7
101 GeV	1	2.63 +/- 0.08	82.9	0	1.17 +/- 0.05	46.8	8	13.68 +/- 0.24	71.7

$\sqrt{s} = 200 - 209 \text{ GeV}$   
 $\int L = 83.1 \text{ pb}^{-1}$

Good consistency with Standard Model



# Other Channels: Slepton NLSP

## Other production channels

Final states with 4 or 6 leptons + missing energy

### Slepton co-NLSP

$$e^+ e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow (\tilde{I}^\pm I^\mp)(I'^\pm I'^\mp) \rightarrow I I I I' \tilde{G}\tilde{G}$$

$$I I I I' + E_T$$

$$I, I' \equiv e, \mu, \tau$$

$\tilde{\chi}_1^0$  is a Majorana fermion: can lead to same sign for two highest energy leptons (i.e. those from the slepton decays)

Equal branching fractions assumed

### Stau NLSP

$$e^+ e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow (\tilde{\tau}^\pm \tau^\mp)(\tilde{\tau}^\pm \tau^\mp) \rightarrow \tau\tau\tau\tau \tilde{G}\tilde{G}$$



$$\tau\tau\tau\tau + E_T$$

$$e^+ e^- \rightarrow \tilde{I}^+ \tilde{I}^- \rightarrow (I^+ \tilde{\tau} \tau)(I^- \tilde{\tau} \tau) \rightarrow I^+ I^- \tau\tau\tau\tau \tilde{G}\tilde{G}$$



$$I^+ I^- \tau\tau\tau\tau + E_T$$

$$I \equiv e, \mu$$



# Other Channels: Slepton NLSP

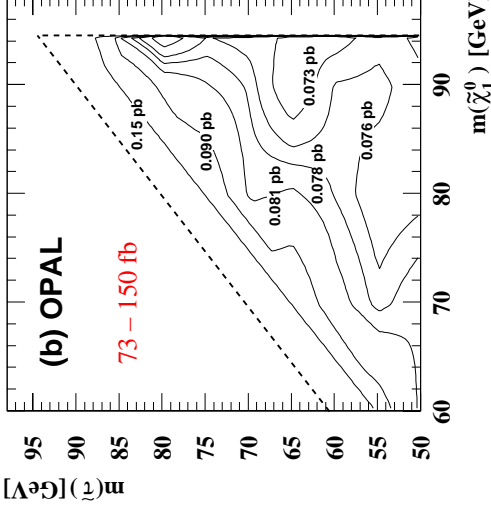
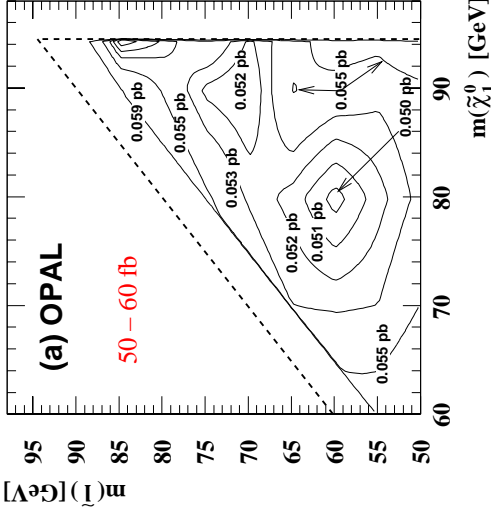
~ model independent cross-section limits at  $\sqrt{s} = 189\text{GeV}$

$$l^+ l^- l^+ l^- + \mathbf{E}_T$$

$$\sigma(e^+ e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)$$

$$\tilde{\chi}_1^0 \rightarrow \tilde{l} l, \tilde{l} \rightarrow l \tilde{G}$$

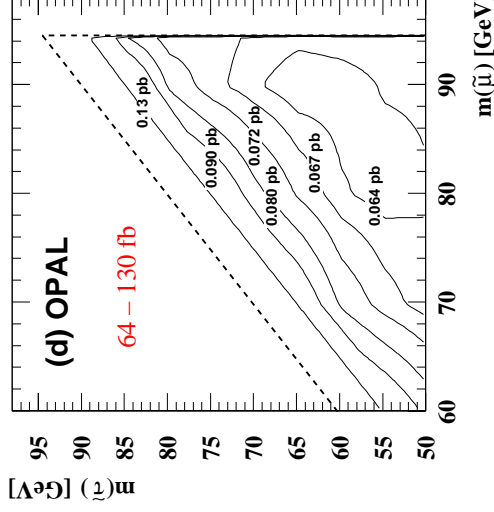
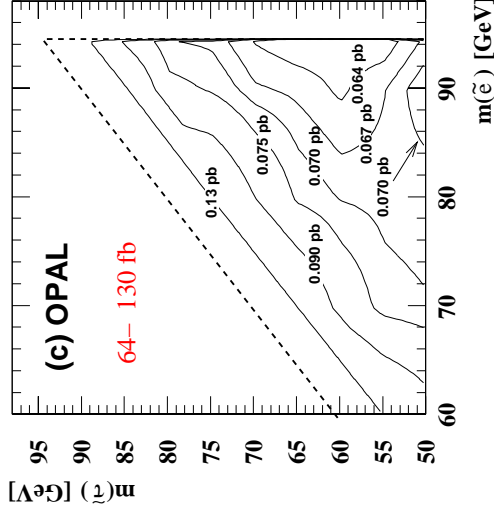
Assumes equal  
BRs to  $e, \mu, \tau$



$$e e \tau \tau \tau \tau + \mathbf{E}_T$$

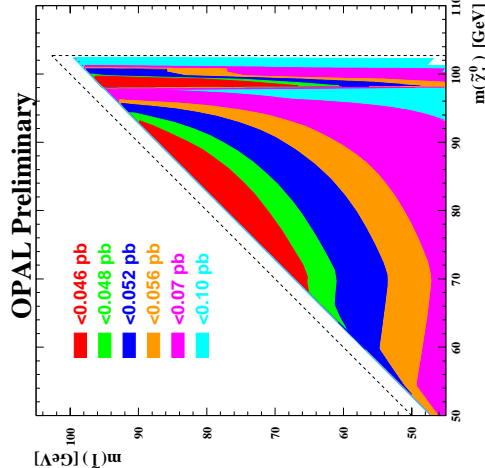
$$\sigma(e^+ e^- \rightarrow \tilde{e}^+ \tilde{e}^-)$$

$$\tilde{e} \rightarrow \tilde{\tau} \tau e, \tilde{\tau} \rightarrow \tau \tilde{G}$$



# Slepton NLSP Results at 200-209 GeV

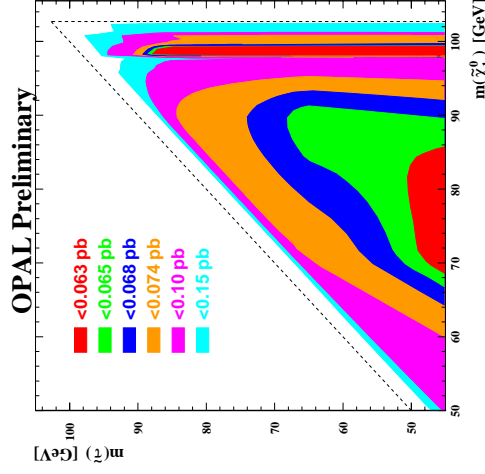
Cross-section limits at  $\sqrt{s} = 205.5 \text{ GeV}$



$$\sigma(e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)$$

$$\tilde{\chi}_1^0 \rightarrow \tilde{l}l, \tilde{l} \rightarrow l\tilde{G}$$

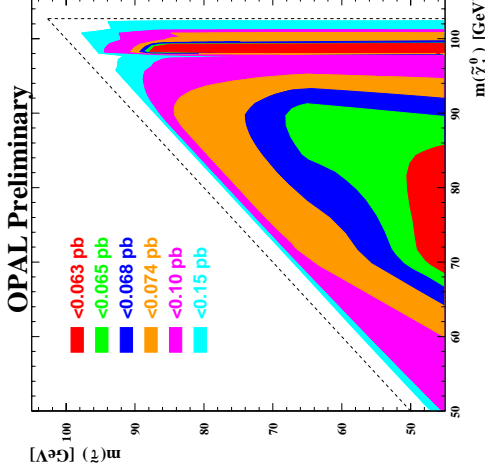
$$l^+l^+l^- + \mathbf{E}_T$$



$$\sigma(e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)$$

$$\tilde{\chi}_1^0 \rightarrow \tilde{\tau}^+ \tau^-, \tilde{\tau} \rightarrow \tau\tilde{G}$$

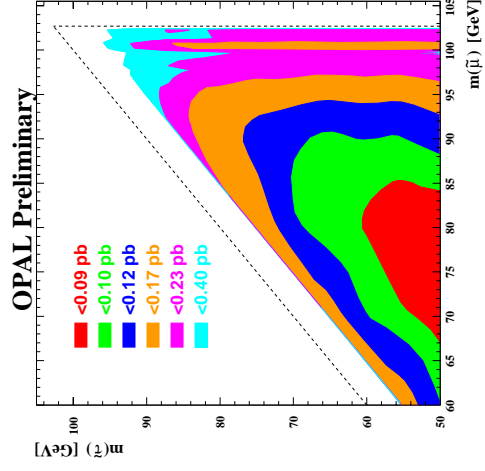
$$\tau^+ \tau^- \tau^+ \tau^- + \mathbf{E}_T$$



$$\sigma(e^+e^- \rightarrow \tilde{e}^+ \tilde{e}^-)$$

$$\tilde{e} \rightarrow \tilde{\tau} \tau, \tilde{\tau} \rightarrow \tau\tilde{G}$$

$$ee\tau\tau\tau + \mathbf{E}_T$$



$$\sigma(e^+e^- \rightarrow \tilde{\mu}^+ \tilde{\mu}^-)$$

$$\tilde{\mu} \rightarrow \tilde{\tau} \tau \mu, \tilde{\tau} \rightarrow \tau\tilde{G}$$

$$\mu\mu\tau\tau\tau + \mathbf{E}_T$$

channel	data	SM background
$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow (l\tilde{G})(l'\tilde{G})$	0	0.6 +/- 0.1
$\tilde{l}^+ \tilde{l}^- \rightarrow (l^+ \tau \tau \tilde{G})(l^- \tau \tau \tilde{G})$	6	2.1 +/- 0.2

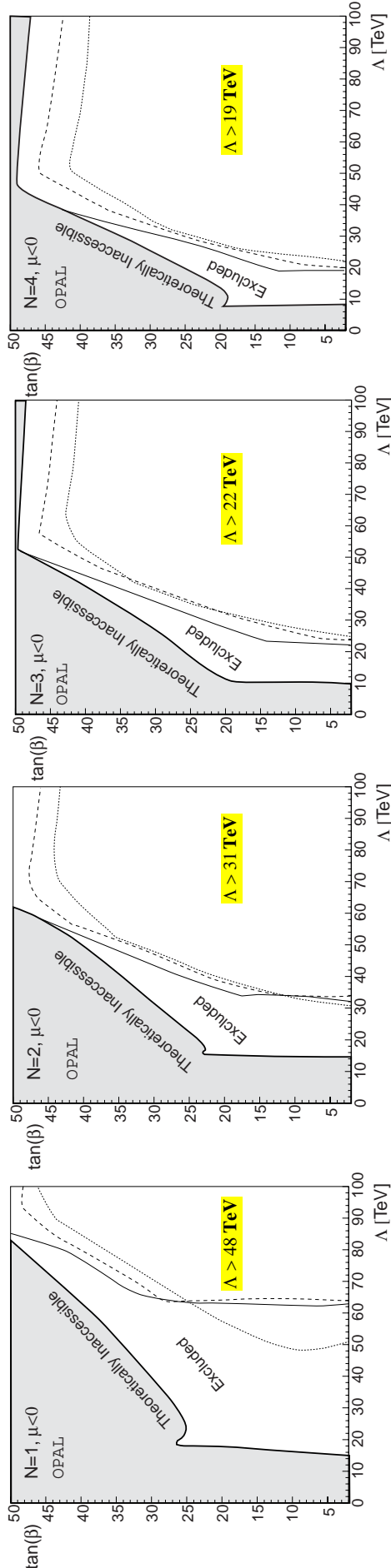
Prob ~ 2%





# GMSB Scan: Exclusion regions in $(\tan\beta, \Lambda)$ plane

Parameter	Description	Range for scan
$\sqrt{F}$	SUSY breaking scale	fixed $\longleftrightarrow$ we are assuming prompt decays
$\Lambda$	sets mass scale for sparticles	5 – 200 TeV
$M$	mass scale of messengers	$1.01\Lambda - 10^6$ TeV
$N$	number of sets of messenger particles	1 – 4
$\tan\beta$	as usual	2 – 50
$\text{sgn}(\mu)$	as usual	+1 / -1

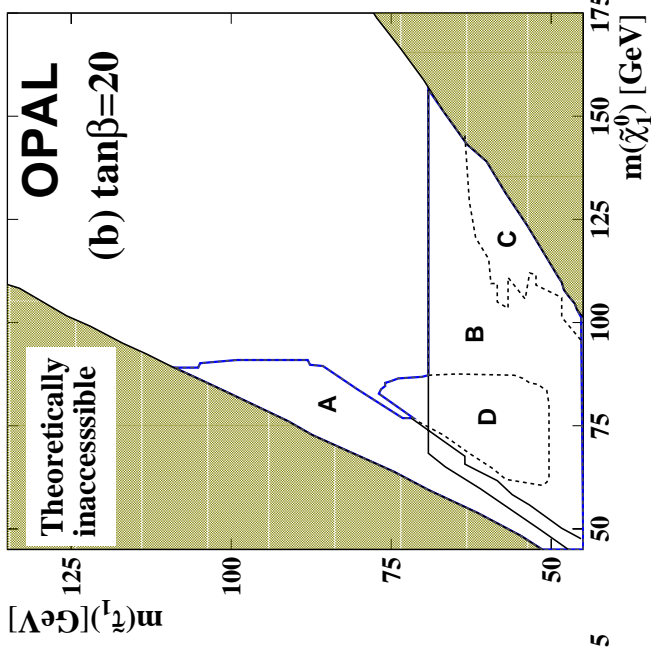
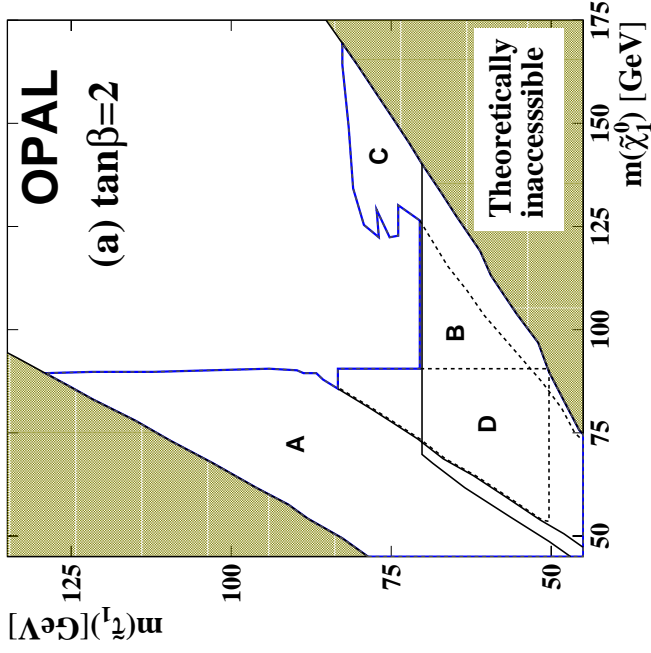


—  $M = 10^6$  TeV  
 - - -  $M = 250$  TeV  
 .....  $M = 1.01\Lambda$

Exclusions for  $\mu > 0$  are somewhat stronger



# GMSB Scan: Exclusion Regions in $(M_{\tilde{\tau}}, M_{\tilde{\chi}_1^0})$ plane



- A**  $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \gamma \tilde{G} \gamma \tilde{G}$
- B**  $\tilde{\tau}^+ \tilde{\tau}^- \rightarrow \tau^+ \tilde{G} \tau^- \tilde{G}$
- C**  $\tilde{\mu}^+ \tilde{\mu}^- \rightarrow \mu^+ \tilde{G} \mu^- \tilde{G}$
- D**  $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow I^+ I^- \tilde{G} I^+ I^- \tilde{G}$

sleptons  $\sim$  degenerate

$$\tan\beta = 2$$

$$M_{\text{NLSP}} > 70 \text{ GeV} \quad \tilde{I}^\pm$$

$$> 85 \text{ GeV} \quad \tilde{\chi}_1^0$$

$\tilde{\tau}_1^\pm$  NLSP

$$\tan\beta = 20$$

$$M(\tilde{\tau}) > 69 \text{ GeV}$$

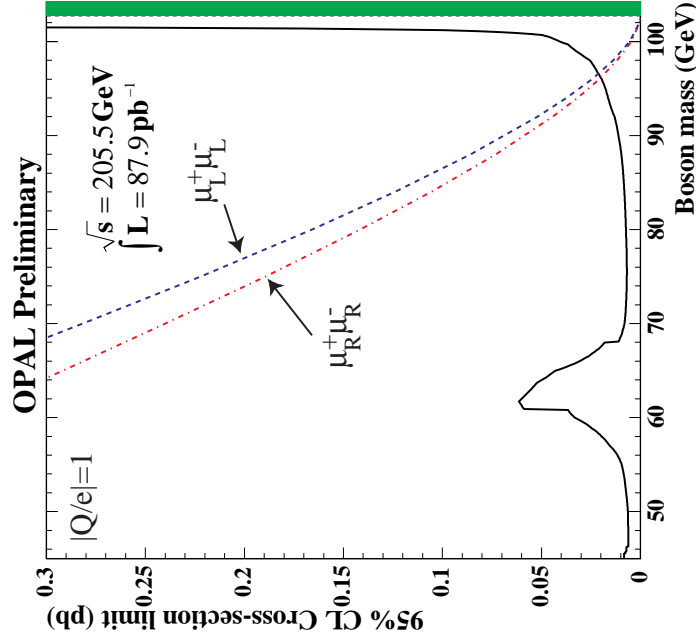
$$M(\tilde{e}, \tilde{\mu}) > 88 \text{ GeV}$$

$$M(\tilde{\chi}_1^0) > 76 \text{ GeV}$$



# Slepton NLSP with lifetime

- Decay length  $\sim$  detector size: **kinked charged tracks (analysis in progress)**
- Decay length  $\gg$  detector size: **long lived stable charged particles**



Includes results from lower energies

$M(\tilde{\Gamma}_R^+) > 96.0 \text{ GeV @ 95\% CL}$

$M(\tilde{\Gamma}_L^+) > 96.5 \text{ GeV @ 95\% CL}$

Results valid for lifetimes  $> 1 \text{ us}$

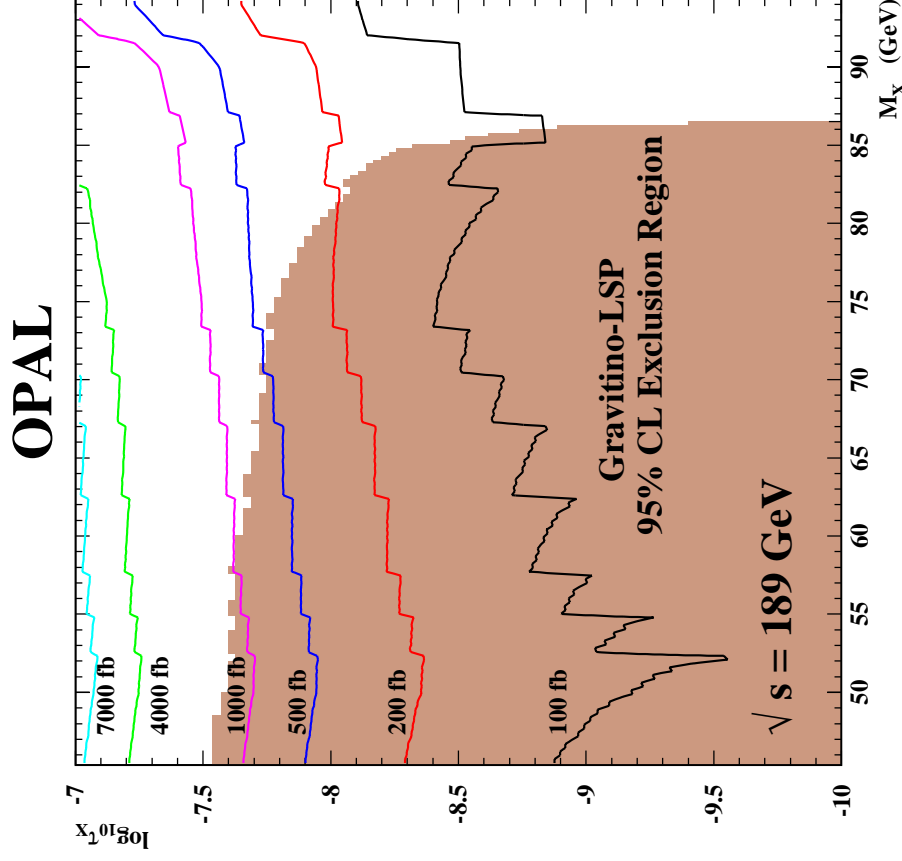


# Neutralino NLSP with Lifetime

- Decay length  $\sim$  detector size: **non-pointing photons (in progress)**
- Decay length  $\gg$  detector size: **conventional SUSY signatures**

Can quantify sensitivity of  $\gamma + \mathbf{E}_T$  analysis to finite lifetime

Evaluate selection efficiency as function of lifetime for  $\tau = 10^{-15} - 10^{-7} \text{ s}$  ( $c\tau \sim 30 \text{ m}$ )



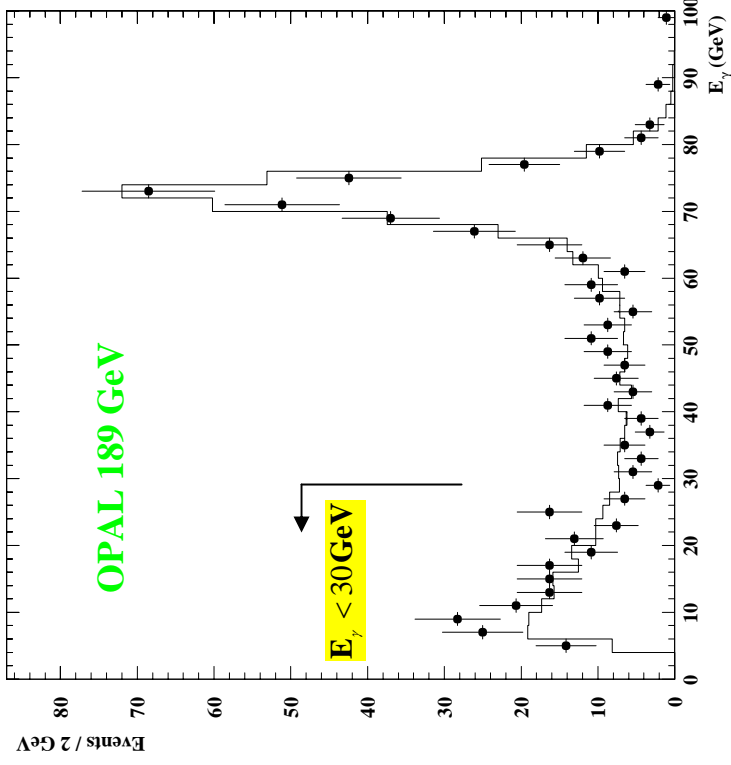
# Gravitino Pair Production $e^+e^- \rightarrow \tilde{G}\tilde{G}\gamma$

Brignole, Feruglio and Zwirner: models with superlight gravitino

→ single photon +  $\mathbf{E}_T$  signature

$$\frac{d^2\sigma}{d\mathbf{x}_\gamma d\cos\vartheta} = \left(\frac{\alpha G_N^2}{45}\right) \frac{s^3}{m_{\tilde{G}}^4} \mathbf{f}_{\tilde{G}\tilde{G}\gamma}(\mathbf{x}_\gamma, \cos\theta)$$

$$\mathbf{f}_{\tilde{G}\tilde{G}\gamma}(\mathbf{x}, \mathbf{y}) = 2(1-\mathbf{x})^2 \left[ \frac{(1-\mathbf{x})(2-2\mathbf{x}+\mathbf{x}^2)}{\mathbf{x}(1-\mathbf{y}^2)} + \frac{\mathbf{x}(-6+6\mathbf{x}+\mathbf{x}^2)}{16} - \frac{\mathbf{x}^3(1-\mathbf{y}^2)}{32} \right]$$

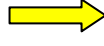


Soft photon spectrum:  
event counting in region

$E_\gamma < 30 \text{ GeV}$

195 events observed

$179.6 \pm 5.4$  expected  $\nu\bar{\nu}\gamma(\gamma)$



from KORALZ

$\sigma_{95} \rightarrow m_{\tilde{G}} > 8.7 \mu\text{eV @ 95\% CL}$

# Conclusions

- The end of LEP draws near: still no signs of SUSY
- Lots of work to still be done:
  - Update to highest energies and maximum luminosity
  - Results on intermediate / long lifetimes in progress
- Still some hope for discovery at LEP ?
  - Wait a few months .....
  - Or wait a few years

