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## 1 MODULE 1 - TENSORS AND EXTERIOR CALCULUS ON MANIFOLDS

This module starts by emphasising a view of vectors as objects that can be discussed without explicit reference to a basis (or cobasis), as is implicit in the elementary notation $\mathbf{u}$, or $\vec{u}$. An alternative description in terms of components in a so-called cobasis, or dual basis will be introduced and its meaning explored. Our powerful geometric approach will then allow a conceptually simple generalisation of vectors to tensors. While everyone agrees that vectors are indispensable to the mathematical description of many physical quantities, the equally great importance of tensors is not always fully appreciated. For example, it is difficult to understand electromagnetism if one insists on regarding the electric and magnetic fields as just two vector fields connected by Maxwell equations, instead of the six non-zero components of the rank-2, antisymmetric Faraday tensor ${ }^{\dagger}$, or 2 -form, $\mathbf{F}$. The need to describe how vectors and $p$-forms change in time and space will lead to the exterior derivative, of which gradient, divergence and curl are but special cases. We will also see that differential $p$-forms in fact are the only objects that can be meaningfully integrated. The concise language of $p$-forms can illuminate many other areas of physics, such as mechanics, thermodynamics, general relativity and quantum field theory.

### 1.1 Vector Spaces and Linear Mappings

### 1.1.1 Vector spaces in a nutshell

Definition 1.1. A vector space $\mathcal{V}$ over a field $\mathbb{F}$ is a (possibly infinite) set of objects on which an operation called addition and another called s-multiplication (multiplication by a scalar) are defined, and which is is closed under these operations. That is, any two elements $\mathbf{u}$ and $\mathbf{v}$ of $\mathcal{V}$ satisfy:

$$
(a+b)(\mathbf{u}+\mathbf{v})=(a \mathbf{u}+a \mathbf{v}+b \mathbf{u}+b \mathbf{v}) \in \mathcal{V}
$$

$\forall a, b \in \mathbb{F}$; in what follows, $\mathbb{F}=\mathbb{R}$. This addition operation is commutative and associative; it can be composed with an inverse, the result being the zero element. As for s-multiplication, it obeys: $(a b) \mathbf{u}=a(b \mathbf{u})$. We will call (informally) elements of a vector space vectors, keepng in mind that deciding whether an object is a vector can only be done by reference to the vector space of which it is an element.

Example 1.1. $\mathbb{R}^{n}$, the set of all ordered $n$-tuples of real numbers, with addition defined as adding entries with the same place in the $n$-tuple, and s-multiplication by $\lambda$ defined as multiplying each entry by $\lambda$, is perhaps the best-known and most important vector space.

Let $\mathcal{V}$ and $\mathcal{W}$ be two vector spaces that share the same field of scalars. We shall be interested in the set of all linear mappings, $\operatorname{Hom}(\mathcal{V}, \mathcal{W}):=\{\mathbf{T}: \mathcal{V} \rightarrow \mathcal{W}\}$, such that, $\forall \mathbf{T}_{i} \in \operatorname{Hom}(\mathcal{V}, \mathcal{W})$ :

$$
\left(a \mathbf{T}_{i}+\mathbf{T}_{j}\right)(\mathcal{V})=a \mathbf{T}_{i}(\mathcal{V})+\mathbf{T}_{j}(\mathcal{V})
$$

where the operations on the left are defined in $\operatorname{Hom}(\mathcal{V}, \mathcal{W})$, and those on the right on $\mathcal{W}$. Henceforth, we will use the notation $\mathcal{L}(\mathcal{V}, \mathcal{W})$ instead of Hom for our space of linear mappings. Of course, one can define linear mappings on $\mathcal{L}(\mathcal{V}, \mathcal{W})$, ie., we can compose linear mappings, and the composition of two linear maps is itself a linear map.

### 1.2 Where Do Vectors Live?

From what we emphasised above, the answer has to be: in a vector space! Therefore, we should learn how to identify (or construct) such vector spaces. This is anything but trivial: in what follows, we shall see that such a familiar space as spacetime cannot be endowed with a vector-space structure.

[^0]
### 1.2.1 Manifolds and coordinates

Definition 1.2. Let $M$ be a set of elements, or "points", with a so-called topology that introduces the notion of open balls (or neighbourhoods) around points, in turn allowing a definition of continuity. If

- $M$ can be entirely covered by a union $U$ of possibly overlapping open (without boundary) subsets $U_{1}, U_{2}$, etc., (as given by the topology), each mapped in a one-to-one way to an open subset of $\mathbb{R}^{n}$ by a coordinate map: $x: U_{i} \longrightarrow \mathbb{R}^{n}, y: U_{j} \longrightarrow \mathbb{R}^{n}$, etc.;
- each coordinate map is differentiable, in the sense that a neighbourhood around any point $\mathcal{P}$ in $M$, which we know must exist by virtue of the topology, is mapped to a neighbourhood of the image via $x$ of $\mathcal{P}$ in $\mathbb{R}^{n}$;
- when there is an overlapping region $U_{i} \cap U_{j}$ in $M$, only coordinate maps for which the transition map (aka coordinate transformation) $y \circ x^{-1}: \mathbb{R}^{n} \longrightarrow U_{i} \cap U_{j} \longrightarrow \mathbb{R}^{n}$ between them is (once) differentiable are allowed;
then we say that $M$ is a differentiable manifold, usually abbreviated to just manifold. The minimum number $n$ of parameters-each a map $x^{i}: U \longrightarrow \mathbb{R}(i=1, \ldots, n)$-that uniquely specify every point in an open subset is the dimension of the manifold. Real differentiable $\left(C^{1}\right)$ manifolds can be shown to be smooth, ie., $C^{\infty}$.

Each open subset (itself a manifold) together with its coordinate map is called a coordinate chart, ( $U_{i}, x$ ), or local coordinate system, on $M$. Any collection of coordinate charts that covers the whole of $M$ is called an atlas.

Example 1.2. $\quad \mathbb{R}^{n}$ can be promoted to a manifold; it is mapped into iself and can be covered with just one coordinate chart, Cartesian coordinates, or standard (natural) coordinates. Other charts are of course possible, eg. polar coordinates on $\mathbb{R}^{2}$, but then more than one may be needed to cover the manifold.

- A conical surface, even a semi-infinite one, can never be a manifold because of its tip.
- A vector space $\mathcal{V}$ can be made into a manifold that can be covered with one chart $(\mathcal{V}, \Phi)$, where $\Phi$ maps elements of $\mathcal{V}$ to their components in $\mathbb{R}^{n}$ in that basis. Conversely, however, a manifold is not in general a vector space! Considering the manifold of points on Earth's surface, there is no meaning to adding the position of Toronto to that of London.
- Even though $\mathbb{R}^{n}$ can be endowed with a manifold structure, a unit ball in $\mathbb{R}^{n}$, defined in Cartesian coordinates by $\sum x_{i}^{2} \leq 1$, is not a manifold because it has an edge on which it is not differentiable. The open unit ball, $\sum x_{i}^{2}<1$, is a manifold. So is the unit sphere, $S^{n}$, defined by $\sum^{n} x_{i}^{2}=1$ and embedded in $\mathbb{R}^{n+1}$.
Let us look at $S^{1}$, the unit circle in the plane $\mathbb{R}^{2}$, and $S^{2}$, the 2-dim sphere in $\mathbb{R}^{3}$. These are the archetypal examples of (closed) curves in $\mathbb{R}^{2}$ and (closed) surfaces in $\mathbb{R}^{3}$.
$S^{1}$ being a 1-dim manifold, we wish to build an atlas for it. One way of doing this is with two open patches, $y= \pm \sqrt{1-x^{2}}$ with the two points at $x= \pm 1$ excluded (why?), and the $+/-$ sign corresponding to the submanifold in the upper/lower half-plane. Then each point of any of the two submanifolds is in one-to-one correspondence with some $x \in \mathbb{R}$, with $x<1$. To cover all of $S^{1}$, we can then repeat the procedure with two submanifolds in correspondence with $x>0$ and $x<0$, and an atlas with four charts has been constructed.
$S^{1}$ also has another local coordinate, $\theta$, related to $x$ by the multi-valued coordinate transformation: $\theta=$ $\tan ^{-1}(y / x)=\tan ^{-1}\left(\sqrt{1 / x^{2}-1}\right)$. To avoid a given point being in correspondence with more than one value of $\theta$, the interval mapped to in $\mathbb{R}$ must be $[0,2 \pi)$.
Similarly, an atlas can be constructed for $S^{2}$ out of patches corresponding to, first, $z>0$ and $z<0$ (leaving out the circle $z=0$ ), then $y>0$ and $y<0$ (leaving out the points corresponding to ( $y=0, z=0$ ),
and finally $x>0$ and $x<0$. Each point in each patch can be mapped unambiguously to $\mathbb{R}^{2}$, and there exist functions $f(u, v)$ which map $(x, y),(x, z)$ or $(y, z)$ into coordinates $u$ and $\pm \sqrt{1-u^{2}-v^{2}}$, both differentiable functions provided that $u^{2}+v^{2}<1$, which is always the case in any given patch.
On $S^{2}$ we could also use the local spherical coordinates $\theta=\cos ^{-1} z$ and $\phi=\tan ^{-1}(y / x)$. The region of $\mathbb{R}^{2}$ mapped to is $(0, \pi)$ and $[0,2 \pi]$, with the poles removed. This time, the problem comes from a failure of 1 -to- 1 correspondence at $\theta=0$ and $\theta=\pi$ : for each of these the value of $\phi$ is undetermined. More patches are needed to cover $S^{2}$.

Notice that we have looked at $S^{1}$ and $S^{2}$ as being embedded, or immersed, in a higher-dimensional manifold, $\mathbb{R}^{2}$ and $\mathbb{R}^{3}$. Whitney's embedding theorems guarantee that any smooth $M^{n}$ is a smooth submanifold of $\mathbb{R}^{m>2 n}$; there are stronger results in important restricted cases. Embedding curves and surfaces in, eg., $\mathbb{R}^{3}$ is great for visualisation purposes, but we are really interested in the intrinsic properties of a manifold which should be independent of whatever higher-dimensional manifold in which it may or may not be immersed.

Beyond the technicalities, it is enough almost all the time to view a manifold as a set which can be parametrised in a smooth way.

### 1.2.2 Curves, directional derivatives and vectors

The naïve notion of a vector as a straight arrow from one point to another in $\mathbb{R}^{n}$ cannot be extended to arbitrary manifolds $M$, on which straightness will in general have no well-defined meaning (think of straight arrows on a sphere). As noted before, manifolds are not vector spaces; so where do vectors that we want to attach to a point in $M$ actually live? And is it possible to think of a vector as a local object that involves only that point, in a way that is independent of any coordinate chart?

Definition 1.3. A curve $\Gamma$ on a manifold $M$ is a mapping, at least $C^{1}$ (no kinks!), that sends each value of a real parameter $\lambda$ to a unique point $\mathcal{P}$ in $M$. More succintly, $\Gamma: \mathbb{R} \longrightarrow M$. In effect, $\lambda$ is a coordinate on $\Gamma$, but note the direction of the mapping that, following tradition, is opposite the one used in definition 1.2. The curve at $\mathcal{P}$ is given by $\Gamma(\lambda)=\mathcal{P}$.

Definition 1.4. Now introduce the vector space, $C^{\infty}(M):=\{f: M \longrightarrow \mathbb{R}\}$, of all smooth, real-valued functions $f$ that map a point in $M$, such that:

$$
\begin{equation*}
(f+g)(\mathcal{P})=f(\mathcal{P})+g(\mathcal{P}) \quad \forall f, g \in C^{\infty}(M) \quad(a f)(\mathcal{P})=a f(\mathcal{P}) \quad a \in \mathbb{R} \tag{1.1}
\end{equation*}
$$

where the addition and s-multiplication on the left-hand side of the equations are operations on $C^{\infty}(M)$, while those on the right-hand side are on $\mathbb{R}$. We also have the composition: $f \circ \Gamma$ : $\mathbb{R} \underset{\Gamma}{\rightarrow} M \underset{f}{\longrightarrow}$. Thus, $f \circ \Gamma$ could be called $f(\lambda)$, with $\lambda \in \mathbb{R}$.
Definition 1.5. Let $\Gamma$ be a curve parametrised by a coordinate $\lambda$. The velocity at a point $\mathcal{P}$ with coordinate $\lambda_{0}$ on this curve, is the linear map $\mathbf{v}_{(\Gamma, \mathcal{P})}: C^{\infty}(M) \longrightarrow \mathbb{R}$, defined as:

$$
\begin{equation*}
\mathbf{v}_{(\Gamma, \mathcal{P})}(f):=\left.\mathrm{d}_{\lambda}(f \circ \Gamma)\right|_{\lambda_{0}} \quad\left(=\left.\mathrm{d}_{\lambda} f\right|_{\lambda_{0}}\right) \tag{1.2}
\end{equation*}
$$

where we use the notation $\mathrm{d}_{\lambda}$ as short-hand for $\mathrm{d} / \mathrm{d} \lambda$.
It should be clear that such a curve is only one of an infinite number containing $\mathcal{P}$ that have their own velocity at $\mathcal{P}$. Indeed, we could parametrise another curve $\Theta$ and write: $\mathbf{w}_{\Theta}(f)=\mathrm{d}_{\lambda} f$ at $\lambda_{1}$, where $\Theta\left(\lambda_{1}\right)=\mathcal{P}$. We say that the velocities are tangent to the manifold at $\mathcal{P}$.

Now, if we are willing to view the curve as part of an open subset $U \subset M$ parametrised by coordinate functions denoted collectively by $x: U \longrightarrow \mathbb{R}^{n}$, where $x$ describe what the curve "looks like" in $U$, we can derive an
alternate expression for the velocity's action on $f \in C^{\infty}$. Inserting the identity map into the defining equation for $\mathbf{v}_{(\Gamma, \mathcal{P})}(f)$, we obtain:

$$
\begin{align*}
\mathbf{v}_{(\Gamma, \mathcal{P})}(f) & :=\left.\mathrm{d}_{\lambda}(f \circ \Gamma)\right|_{\lambda_{0}}=\mathrm{d}_{\lambda}\left[\left(f \circ x^{-1}\right) \circ(x \circ \Gamma)\right]_{\lambda_{0}} \\
& =\sum_{\nu}^{n}\left[\left.\left.\mathrm{~d}_{\lambda}\left(x^{\nu} \circ \Gamma\right)\right|_{\lambda_{0}}\left[\partial_{\nu}\left(f \circ x^{-1}\right)\right]\right|_{x^{\nu}(\mathcal{P})}\right] \tag{1.3}
\end{align*}
$$

where the index $\nu$ in the chain rule runs over the number of parameters (or local coordinates) that specify each point in $U$. The parametrisation of the curve in $U$ can be written as a binary map from $\mathbb{R}$ to $\mathbb{R}^{n}:(x \circ \Gamma)(\lambda)$ (or, if one wishes, $x(\lambda)$ ).

This coordinate-dependent expression contains unexpected information. Taking $f=x^{\mu}$, we can write $\mathrm{d}_{\lambda}\left(x^{\nu} \circ\right.$ $\Gamma)\left.\right|_{0}=\left.\mathrm{d}_{\lambda} x^{\nu}\right|_{0}$. And we may suspect that the second factor has something to do with the partial derivative of $f$. Indeed, $f$ acts on $U$, not $\mathbb{R}$, without any reference to a local coordinate system on $U$. In order to have a calculable derivative, we take a detour via $\mathbb{R}^{n}: U \underset{x}{\longrightarrow} \mathbb{R}^{n} \underset{f \circ x^{-1}}{\longrightarrow} \mathbb{R}$. Since $f \circ x^{-1}$ maps $\mathbb{R}^{n}$ to $\mathbb{R}$, its usual derivatives can be calculated and behave for practical purposes like the standard $\partial_{x^{\nu}} f\left(x^{\mu}\right)$., that is: $\left.\partial_{x^{\nu}} f\right|_{\mathcal{P}}:=\left.\partial_{\nu}\left(f \circ x^{-1}\right)\right|_{\mathcal{P}}$. Although we don't always write the dependence of $f$ on $x$ explicitly, here it is essential.

Having made this technical point, we discard the arbitrary function $f$ and write the velocity vector as:

$$
\begin{equation*}
\mathbf{v}_{(\Gamma, \mathcal{P})}=\left.\sum_{\nu} \mathrm{d}_{\lambda} x^{\nu}\right|_{0}\left(\boldsymbol{\partial}_{\nu}\right)_{\mathcal{P}} \tag{1.4}
\end{equation*}
$$

We will interpret this important result a little later, after we have constructed the space where $\mathbf{v}_{(\Gamma, \mathcal{P})}$ lives.

### 1.2.3 The tangent space of a manifold

Definition 1.6. The tangent space $\mathcal{T}_{\mathcal{P}}$ to a manifold $M^{n}$ - even though $M^{n}$ itself is not a vector space! - at a point $\mathcal{P} \in M^{n}$, is a set that can be equipped with a vector-space structure, consisting of all the velocity vectors $\mathbf{v}_{\mathcal{P}}$ tangent to $M^{n}$ at $\mathcal{P}$, In fact, all vectors defined on $M^{n}$ at $\mathcal{P}$ live in $\mathcal{T}_{\mathcal{P}}$, not in $M^{n}$. $\mathcal{T}_{\mathcal{P}}$ is always of finite dimension. If $M^{n}$ is viewed as embedded in $\mathbb{R}^{N}(N \geq n), \mathcal{T}_{\mathcal{P}}=\mathbb{R}^{n}$.
The set of all points in a manifold, together with their tangent spaces, is called the tangent bundle of the manifold.

This definition rests on a very bold assertion, namely, that the velocities as defined above (definition 1.5) are in fact vectors. It is instructive to prove it, that is, to show that $\mathcal{T}_{\mathcal{P}}$ can indeed be made into a vector space. First, let us specify what is meant by addition and s-multiplication on $\mathcal{T}_{\mathcal{P}}$.

Definition 1.7. The addition operation on $\mathcal{T}_{\mathcal{P}}$ is a map, $\mathcal{T}_{\mathcal{P}}+\mathcal{T}_{\mathcal{P}} \longrightarrow \mathcal{L}\left(C^{\infty}(M), \mathbb{R}\right)$, such that, $\forall f \in C^{\infty}(M)$ and any two curves $(\Gamma, \Theta) \in M$ intersecting at $\mathcal{P} \in M$ :

$$
\left(\mathbf{v}_{(\Gamma, \mathcal{P})}+\mathbf{v}_{(\Theta, \mathcal{P})}\right)(f):=\mathbf{v}_{(\Gamma, \mathcal{P})}(f)+\mathbf{v}_{(\Theta, \mathcal{P})}(f)
$$

Again, the addition operation on the left is between mappings, whereas that on the right is on $\mathbb{R}$. As for s-multiplication, it is a map, $\mathbb{R} \times \mathcal{T}_{\mathcal{P}} \longrightarrow \mathcal{L}\left(C^{\infty}(M), \mathbb{R}\right)$, such that, $\forall a \in \mathbb{R}$ :

$$
\left(a \cdot \mathbf{v}_{(\Gamma, \mathcal{P})}\right)(f):=a \mathbf{v}_{(\Gamma, \mathcal{P})}(f)
$$

The question now is: do these operations close? In other words, can we find some curve $\Theta \in M$ such that: $a \cdot \mathbf{v}_{(\Gamma, \mathcal{P})}=\mathbf{v}_{(\Theta, \mathcal{P})}$, and perhaps another curve $\Sigma \in M$ such that: $\mathbf{v}_{(\Gamma, \mathcal{P})}+\mathbf{v}_{(\Theta, \mathcal{P})}=\mathbf{v}_{(\Sigma, \mathcal{P})}$ ?

To construct such a curve for s-multiplication, we first redefine the parameter of the curve $\Gamma$ as the linear function, $\mu: \mathbb{R} \longrightarrow \mathbb{R}$, of $\lambda: \mu=a \lambda+\lambda_{0}$, with $\lambda$ now the parameter of a curve $\Theta$ such that $\Theta(\lambda)=\Gamma(\mu)$

Therefore, $\Theta(0)=\Gamma\left(\lambda_{0}\right)=\mathcal{P}$. As in definition 1.3 we can write: $\Gamma(\mu)=\Gamma \circ \mu(\lambda)$. Insert this information into the expression for the velocity for $\Theta$ at $\mathcal{P}$ :

$$
\begin{aligned}
\mathbf{v}_{(\Theta, \mathcal{P})}(f) & =\left.\mathrm{d}_{\lambda}(f \circ \Theta)\right|_{\lambda=0}=\left.\mathrm{d}_{\lambda}(f \circ \Gamma \circ \mu)\right|_{\lambda=0} \\
& =\left.\left.\mathrm{d}_{\mu}(f \circ \Gamma)\right|_{\mu(\lambda=0)=\lambda_{0}} \mathrm{~d}_{\lambda} \mu\right|_{\lambda=0} \\
& =a \mathbf{v}_{(\Gamma, \mathcal{P})}(f)
\end{aligned}
$$

Therefore, we have found a curve $\Theta$ such that the operation $a \cdot \mathbf{v}_{(\Gamma, \mathcal{P})}$ gives the velocity for that curve at $\mathcal{P}$.
Up to now, in our discussion of tangent spaces, we have not needed any reference to coordinate charts. Unfortunately, when it comes to proving that addition of two velocities in $\mathcal{T}_{\mathcal{P}}$ gives a velocity in $\mathcal{T}_{\mathcal{P}}$, we cannot add the curve mappings directly since this has no meaning. Instead, as was done in the previous section, assume that both curves $\Gamma$ and $\Theta$ are in some open subset $U \subset M$ parametrised by coordinate functions $x$. Let $\Gamma$ and $\Theta$ go through point $\mathcal{P}$ at values $\lambda_{1}$ and $\lambda_{2}$ ot their respective parameter. Then construct a curve $\Sigma$ parametrised in $U$ by:

$$
(x \circ \Sigma)(\lambda)=(x \circ \Gamma)\left(\lambda_{1}+\lambda\right)+(x \circ \Theta)\left(\lambda_{2}+\lambda\right)-(x \circ \Gamma)\left(\lambda_{1}\right)
$$

Although there might appear to be an obvious cancellation in this expression, it is not allowed because the coordinate functions are not linear and thus do not distribute over the additions in $\mathbb{R}^{n}$ in the arguments on the right. At $\lambda=0$, however, the cancellation does occur, leaving $\Sigma_{x}(0)=\Theta\left(\lambda_{2}\right)=\mathcal{P}$, so that our curve $\Sigma_{x}$ runs through point $\mathcal{P}$ at $\lambda=0$.

We also need the derivative of the $\nu^{\text {th }} x$ coordinate of the curve $\Sigma$, evaluated at $\mathcal{P}$ :

$$
\begin{align*}
\left.\mathrm{d}_{\lambda}\left(x^{\nu} \circ \Sigma\right)\right|_{0} & =\left.\mathrm{d}_{\lambda}\left[\left(x^{\nu} \circ \Gamma\right)\left(\lambda_{1}+\lambda\right)+\left(x^{\nu} \circ \Theta\right)\left(\lambda_{2}+\lambda\right)-\left(x^{\nu} \circ \Gamma\right)\left(\lambda_{1}\right)\right]\right|_{0} \\
& =\left.\left.\mathrm{d}_{\lambda_{1}+\lambda}\left(x^{\nu} \circ \Gamma\right)\right|_{\lambda_{1}} \mathrm{~d}_{\lambda}\left(\lambda_{1}+\lambda\right)\right|_{0}+\left.\left.\mathrm{d}_{\lambda_{2}+\lambda}\left(x^{\nu} \circ \Theta\right)\right|_{\lambda_{2}} \mathrm{~d}_{\lambda}\left(\lambda_{2}+\lambda\right)\right|_{0} \\
& =\left.\mathrm{d}_{\lambda_{1}+\lambda}\left(x^{\nu} \circ \Gamma\right)\right|_{\lambda_{1}}+\left.\mathrm{d}_{\lambda_{2}+\lambda}\left(x^{\nu} \circ \Theta\right)\right|_{\lambda_{2}} \tag{1.5}
\end{align*}
$$

Now go back to ou expression (1.3) for the velocity in coordinates $x$. The first factor on the right has been evaluated in eq. (1.5) and, running the chain of equalities in eq. (1.3) backward, there comes:

$$
\begin{aligned}
\mathbf{v}_{\Sigma, \mathcal{P})}(f) & =\sum_{\nu}\left[\left.\left.\left[\partial_{\nu}\left(f \circ x^{-1}\right)\right]\right|_{x^{\nu}(\mathcal{P})} \mathrm{d}_{\lambda}\left(x^{\nu} \circ \Gamma\right)\right|_{\lambda_{1}}\right]+\sum_{\nu}\left[\left.\left.\left[\partial_{\nu}\left(f \circ x^{-1}\right)\right]\right|_{x^{\nu}(\mathcal{P})} \mathrm{d}_{\lambda}\left(x^{\nu} \circ \Theta\right)\right|_{\lambda_{2}}\right] \\
& =\left.\mathrm{d}_{\lambda}\left[\left(f \circ x^{-1}\right) \circ(x \circ \Gamma)\right]\right|_{\lambda_{1}}+\left.\mathrm{d}_{\lambda}\left[\left(f \circ x^{-1}\right) \circ(x \circ \Theta)\right]\right|_{\lambda_{2}} \\
& =\mathbf{v}_{(\Gamma, \mathcal{P})}(f)+\mathbf{v}_{(\Theta, \mathcal{P})}(f)
\end{aligned}
$$

Thus, adding the velocities for two curves meeting at some point yields the velocity for some other curve intersecting the others at that same point, and the tangent space of a curve at a point can indeed support a vector space structure! Do note that the result does not depend on whatever coordinate chart we might have used in the intermediate steps of the proof.

### 1.2.4 Vectors: components and bases

There still remains to interpret eq. (1.4). To do this, we need first to introduce a different perspective on vectors, one that does make reference to local coordinate systems.

Definition 1.8. If any $\mathbf{v} \in \mathcal{V}$ can be written as a linear combination ${ }^{\dagger}$ :
$\mathbf{v}=\sum_{\alpha}^{n<\infty} v^{\alpha} \mathbf{e}_{\alpha} \equiv v^{\alpha} \mathbf{e}_{\alpha} \quad$ Einstein's convention : summation over repeated indices implied!

[^1]of a set $\left\{\mathbf{e}_{\alpha} \in \mathcal{V}\right\}$, then that set is said to span, or to be a set of generators of, $\mathcal{V}$.
If, furthermore, this set is linearly independent, in the sense that demanding that $\mathbf{v}=0$ forces all coefficients $v^{\alpha}$ in eq. (1.6) to vanish, then it is a basis of $\mathcal{V}$. The number $n$ of vectors in the largest linearly independent set defines the dimension of $\mathcal{V}$, and we often write $\mathcal{V}^{n}$. Conversely, the number of elements of every basis of $\mathcal{V}^{n}$ is the dimension of $\mathcal{V}^{n}$.
The (real, and unique!) coefficients $v^{\alpha}$ in eq. (1.6) are called the components of the vector $\mathbf{v}$ in this basis. This one-to-one correspondence between $\mathcal{V}^{n}$ and $\mathbb{V}^{n}$ can be represented by a $n \times 1$ matrix:
\[

\mathbf{v} \longmapsto\left($$
\begin{array}{c}
v^{1} \\
v^{2} \\
\vdots \\
v^{n}
\end{array}
$$\right)
\]

Warning! v and its components are different beasts and should never be confused. Byron and Fuller (BF) do not make this distinction clear enough. Also, always remember that the index on $\mathbf{e}_{\alpha}$ identifies the vector, not a component of the vector.

Example 1.3. The standard, or natural, basis $\mathbb{R}^{n}$ is the set $\left\{\mathbf{e}_{\alpha}\right\}(\alpha=1,2, \ldots, n)$, where each $n$-tuple labelled by a value of $\alpha$ has 1 in the $\alpha^{\text {th }}$ position and 0 in all other positions.
Now we can intrepret eq. (1.4). Acting on a function $f$ at point $\mathcal{P}: \mathbf{v}_{(\Gamma, \mathcal{P})}(f)=\left.\left.\mathrm{d}_{\lambda} x^{\nu}\right|_{0} \boldsymbol{\partial}_{\nu}(f)\right|_{\mathcal{P}}$, we notice that it looks like the directional derivative of $f$ in the direction of $\mathbf{v}$, which in basic calculus is written $\partial_{\mathbf{v}} f:=\mathbf{v} \cdot \nabla f$. This motivates us to identify any tangent vector $\mathbf{t} \in \mathcal{T}_{\mathcal{P}}$ with the directional derivative at $\mathcal{P}$ in the direction of t . Thus:

Definition 1.9. Given an arbitrary differentiable function $f$ on a manifold $M^{n}$, parametrised by a local coordinate system by $f\left(x^{1} \ldots, x^{n}\right)$, then the action of a vector $\mathbf{v}$ on $f$ at a point $\mathcal{P}$ is defined as:

$$
\begin{equation*}
\left.\mathbf{t}(f)\right|_{\mathcal{P}}:=\partial_{\mathbf{t}} f=\left.t^{\nu} \boldsymbol{\partial}_{\nu} f\right|_{\mathcal{P}} \tag{1.7}
\end{equation*}
$$

where the components $t^{\nu}=\mathbf{t}\left(x^{\nu}\right)=\left.\mathrm{d}_{\lambda} x^{\nu}\right|_{0}$.
It is natural to think of $\boldsymbol{\partial}_{\nu}$ in eq. (1.4) as basis vectors for the tangent space of $M$ at $\mathcal{P}$. But are they linearly independent? Take $f=x^{\nu}$, the coordinate functions for $U$; then $\left.a^{\mu} \partial_{x^{\mu}} x^{\nu}\right|_{\mathcal{P}}=\left.a^{\mu} \partial_{\mu}\left(x^{\nu} \circ x^{-1}\right)\right|_{\mathcal{P}}=a^{\mu} \delta_{\mu}{ }^{\nu}=a^{\nu}$, where $a^{\nu} \in \mathbb{R}$. If $\left.a^{\mu} \boldsymbol{\partial}_{x^{\mu}} x^{\nu}\right|_{\mathcal{P}}=0, a^{\nu}=0$, which shows that the $\boldsymbol{\partial}_{\nu}$ do form a basis of the tangent space. Thus:

Definition 1.10. The tangent space $\mathcal{T}_{\mathcal{P}}$ to a manifold $M^{n}$, at a point $\mathcal{P} \in M^{n}$, admits a basis $\left\{\left.\boldsymbol{\partial}_{\nu}\right|_{\mathcal{P}}\right\}$ ( $\nu=1, \ldots, n$ ) called the coordinate (aka chart-induced) basis for the $n$ local coordinates $x^{\nu}$ that parametrise $M^{n}$.
To find these coordinate basis vectors, we freeze all the variables that parametrise the manifold, except one. Varying that remaining parameter generates a so-called coordinate curve whose tangent at a point is the partial derivative with respect to one of the parameters on which the coordinate functions $x^{\mu}$ of a point depend, Then the components of $\boldsymbol{\partial}_{\mu}$ are simply the partial derivatives with respect to the parameters (coordinates) on the manifold of the position vector $\mathbf{x} \in \mathbb{R}^{n}$ of a point. An example should make this procedure clearer:

Example 1.4. On $S^{2}$ (embedded in $\mathbb{R}^{3}$ ), a point is mapped into the spherical coordinates $(\theta, \phi)$, with $\theta \neq(0, \pi)$; it can also be described by the $\mathbb{R}^{3}$ coordinates $(\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)$. Freezing one of these generates a great circle on the sphere. Then these coordinates, with $\theta$ fixed, describe a circle of radius $\sin \theta$ at "colatitude" $\theta$, and $\boldsymbol{\partial}_{\phi}$ is a coordinate vector which can be visualised in $\mathbb{R}^{3}$ by the vector with components:

$$
\partial_{\phi}(\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)=(-\sin \theta \sin \phi, \sin \theta \cos \phi, 0)
$$

At each point on $S^{2}$ parametrised by $(\theta, \phi)$, this is a vector tangent to the circle at colatitude $\theta$.
Similarly, there is a spherical-coordinate vector, $\boldsymbol{\partial}_{\theta}$, tangent to a meridian going through that same point, with components $(\cos \theta \cos \phi, \cos \theta \sin \phi,-\sin \theta) . \boldsymbol{\partial}_{\theta}$ and $\boldsymbol{\partial}_{\phi}$ together form a basis for vectors in the plane tangent to $S^{2}$ at that point. These vectors do not live in $S^{2}!$ Instead, any vector on $S^{2}$ attached to that point lives in the tangent $\mathbb{R}^{2}$ plane. Also, each point on $S^{2}$ has its own tangent plane.

### 1.2.5 The space dual to a vector space

Given a basis $\left\{\mathbf{e}_{\alpha}\right\}$ for $\mathcal{V}^{m}$ and a basis $\left\{\mathbf{e}_{\beta}\right\}$ for $\mathcal{W}^{n}$, we can easily construct a basis for the space of linear mappings, $\mathcal{L}\left(\mathcal{V}^{m}, \mathcal{W}^{n}\right)$, introduced in section 1.1.1. This will be the set $\left\{\mathbf{E}_{\beta}^{\gamma}\right\}$ such that:

$$
\begin{equation*}
\mathbf{E}_{\beta}^{\gamma}\left(\mathbf{e}_{\alpha}\right)=\delta_{\alpha}^{\gamma} \mathbf{e}_{\beta} \quad \alpha, \gamma=1, \ldots, m ; \beta=1, \ldots, n \tag{1.8}
\end{equation*}
$$

where $\delta_{\alpha}{ }^{\gamma}$ is the Kronecker delta, and the left-hand side should be understood as the action (here, matrix multiplication) of the mappings $\mathbf{E}_{\beta}^{\gamma}$ on the vectors $\mathbf{e}_{\alpha}$.

Do keep in mind that indices on a bold-character object will always label the object itself, not its components, which will never be bold. Thus, a particular linear mapping in the basis we have constructed takes the form:

$$
\begin{equation*}
\mathbf{T}=T^{\beta}{ }_{\alpha} \mathbf{E}_{\beta}^{\alpha} \tag{1.9}
\end{equation*}
$$

where the matrix $\mathbf{T}$ with elements $T^{\beta}{ }_{\alpha}$ (with the row index on the left) represents the mapping $\mathbf{T}$ in the basis $\left\{\mathbf{E}_{\beta}^{\alpha}\right\}$. Thus, the action of $\mathbf{T}$ on some vector $\mathbf{v}$ is:

$$
\mathbf{T}(\mathbf{v})=T^{\beta}{ }_{\alpha} v^{\nu} \mathbf{E}_{\beta}^{\alpha}\left(\mathbf{e}_{\nu}\right)=T^{\beta}{ }_{\alpha} v^{\nu} \delta_{\nu}{ }^{\alpha} \mathbf{e}_{\beta}=T^{\beta}{ }_{\alpha} v^{\alpha} \mathbf{e}_{\beta}
$$

where the $T^{\beta}{ }_{\alpha} v^{\alpha}$ are seen to be the components of the vector $\mathbf{T}(\mathbf{v})$ in basis $\mathbf{e}_{\beta}$ of $\mathcal{W}$. Whenever $\mathcal{W}=\mathcal{V}$, that basis would be the same as the initial basis.

In this last equation, as well as in eq. (1.9) and in $\mathbf{v}=v^{\nu} \mathbf{e}_{\nu}$, the left-hand side is explicitly basis-independent; this notation we shall call index-free, or geometric. The right-hand side, in so-called index notation, makes explicit reference to a basis even though, taken as a whole, it is still basis-independent. Both notations have advantages and disadvantages which we shall discuss later. Fluency in both is highly recommended.

A very interesting subset of the set of linear mappings is $\operatorname{Hom}(\mathcal{V}, \mathbb{R})=\mathcal{L}(\mathcal{V}, \mathbb{R})$, that is, the set of all linear functions on a vector space which takes their value in the same scalar field over which $\mathcal{V}$ is defined, here the real numbers. It forms a space $\mathcal{V}^{*}$ dual to $\mathcal{V}$. Since $\mathcal{L}\left(\mathcal{V}^{m}, \mathcal{W}^{n}\right)$ has dimension $m \times n, \mathcal{V}^{*}$ and $\mathcal{V}$ have the same ${ }^{\dagger}$ dimension. The elements of $\mathcal{V}^{*}$ are called covectors, or linear functionals (in linear algebra), or $\mathbf{1}$-forms. An example would be the definite integral of a polynomial.

Go back to the expansion of a general mapping over a basis, eq. (1.9). If such a mapping is an element of $\mathcal{V}^{*}$, $\beta=1$ and, dropping that index, we write: $\mathbf{T}=T_{\alpha} \mathbf{E}^{\alpha}$, where $\mathbf{T}$ is represented by the elements $T_{\alpha}(\alpha=1, \ldots, n)$ of a row vector, and $\mathbf{T}(\mathbf{v})=T_{\alpha} v^{\alpha} \in \mathbb{R}$. Notation change alert: from now on, we switch to greek letters to denote 1-forms.

Eq. (1.8) also undergoes considerable simplification. Putting $\beta=1$, and noting that $\mathbf{e}_{1}$ on the right-hand side, being a basis for $\mathbb{R}$, is just a number that can be set to 1 , we obtain the following important definition:

Definition 1.11. If $\left\{\mathbf{e}_{\alpha}\right\}$ is a basis of a vector space $\mathcal{V}^{n}$, then its unique dual basis (cobasis) in $\mathcal{V}^{*},\left\{\boldsymbol{\omega}^{\alpha}\right\}$, satisfies:

$$
\begin{equation*}
\boldsymbol{\omega}^{\alpha}\left(\mathbf{e}_{\beta}\right)=\delta^{\alpha}{ }_{\beta} \quad \alpha, \beta=1, \ldots, n \tag{1.10}
\end{equation*}
$$

where $\boldsymbol{\omega}^{\alpha} \equiv \mathbf{E}^{\alpha}$, and the left-hand side is just classic matrix multiplication. Other basis for $\mathcal{V}^{*}$ are of course possible, but this is the most useful one.

From this we derive the action (sometimes used to define $\boldsymbol{\omega}^{\alpha}$ instead of eq. (1.10)) of an element $\boldsymbol{\omega}^{\alpha}$ of the cobasis of $\mathcal{V}^{*}$ on a vector $\mathbf{v} \in \mathcal{V}$ :

$$
\boldsymbol{\omega}^{\alpha}(\mathbf{v})=\boldsymbol{\omega}^{\alpha}\left(v^{\beta} \mathbf{e}_{\beta}\right)=v^{\beta} \boldsymbol{\omega}^{\alpha}\left(\mathbf{e}_{\beta}\right)=v^{\beta} \delta^{\alpha}{ }_{\beta}=v^{\alpha}
$$

from which we conclude that the cobasis element $\omega^{\alpha}$ projects out, or picks out, the corresponding component of $\mathbf{v}$. This will probably come as some surprise to many, who are used to think of $v^{\alpha}$ as the projection of $\mathbf{v}$ on $\mathbf{e}_{\alpha}$.

[^2]What happens if we act on some $\mathbf{e}_{\alpha}$ with a 1-form (covector) $\boldsymbol{\sigma}=\sigma_{\beta} \boldsymbol{\omega}^{\beta}$ ? Well,

$$
\boldsymbol{\sigma}\left(\mathbf{e}_{\alpha}\right)=\sigma_{\beta} \boldsymbol{\omega}^{\beta}\left(\mathbf{e}_{\alpha}\right)=\sigma_{\beta} \delta_{\alpha}^{\beta}=\sigma_{\alpha}
$$

This shows that the $\alpha^{\text {th }}$ component of a 1-form in a cobasis $\left\{\boldsymbol{\omega}^{\alpha}\right\}$ is calculated by acting with the 1 -form on the corresponding element of the basis to which the cobasis is dual.

Recall the one-to-one correspondence between a vector $\mathbf{v}$ and the $n$-tuple of its components in a basis $\left\{\mathbf{e}_{\alpha}\right\}$, $\left(v^{1}, \ldots, v^{n}\right) \in \mathbb{R}^{n}$. An analog correspondence exists between a 1 -form, $\sigma$, and its components $\sigma_{\alpha}$ :

$$
\mathbf{v} \longmapsto\left(v^{1} v^{2} \ldots v^{n}\right)^{\mathrm{T}} \quad \boldsymbol{\sigma} \longmapsto\left(\begin{array}{llll}
\sigma_{1} & \sigma_{2} & \ldots & \sigma_{n}
\end{array}\right)
$$

Therefore, we can also think of $\sigma$ as a linear mapping from $\mathbb{R}^{n}$ to $\mathbb{R}$, or as a procedure to obtain the number $\boldsymbol{\sigma}(\mathbf{v})=\sigma_{\alpha} v^{\alpha}$ out of the vector $\mathbf{v}$ via standard multiplication of a row vector with components $\sigma_{\alpha}$ by a column vector with components $v_{\beta}$ :

$$
\left(v^{1} v^{2} \ldots v^{n}\right)^{\mathrm{T}} \underset{\boldsymbol{\sigma}}{\longmapsto}\left(\begin{array}{llll}
\sigma_{1} & \sigma_{2} & \ldots & \sigma_{n} \tag{1.11}
\end{array}\right)\left(v^{1} v^{2} \ldots v^{n}\right)^{\mathrm{T}}=\sigma_{\alpha} v^{\alpha}
$$

### 1.2.6 Differential of a function and basis dual to a coordinate basis

Definition 1.12. Let $x^{\mu} \in \mathbb{R}^{n}(\mu=1, \ldots, n)$ be the coordinate functions of arbitrary coordinates at a point $\mathcal{P} \in M^{n}$, and $f \in C^{\infty}(M)$ a real-valued diffeentiable function on $M^{n}$. Let also $\mathbf{t} \in \mathcal{T}_{\mathcal{P}}$ be a vector tangent to $M^{n}$ at $\mathcal{P}$. Then the differential of $f$ at $\mathcal{P}, \mathbf{d} f$, is defined as the 1 -form in $\mathcal{T}_{\mathcal{P}}^{*}$ which, when $\mathbf{t}$ is inserted in its input slot, yields the basis-independent action of $\mathbf{t}$ on $f$ at $\mathcal{P}$ from eq. (1.7):

$$
\begin{equation*}
[\mathbf{d} f](\mathbf{t}):=\mathbf{t}(f)=\partial_{\mathbf{t}} f \tag{1.12}
\end{equation*}
$$

Now, if we take $\mathbf{t}_{\mathcal{P}}=\left(\boldsymbol{\partial}_{\nu}\right)_{\mathcal{P}}$, the vector tangent to the coordinate curve for $x^{\nu}$, and also $f=x^{\mu}$, eq. (1.12) at $\mathcal{P}$ gives immediately:

$$
\left(\mathbf{d} x^{\mu}\right)_{\mathcal{P}}\left(\boldsymbol{\partial}_{x^{\nu}}\right)_{\mathcal{P}}=\boldsymbol{\partial}_{x^{\nu}}\left(x^{\mu}\right)=\partial_{\nu}\left(x^{\mu} \circ x^{-1}\right)\left(x_{\mathcal{P}}\right)=\delta^{\mu}{ }_{\nu}
$$

Choosing $\left\{\boldsymbol{\partial}_{\mu}\right\}$ as basis for $\mathcal{T}_{\mathcal{P}}$, we conclude from eq. (1.10) that $\left\{\mathbf{d} x^{\mu}\right\}$ is the basis, dual to $\left\{\boldsymbol{\partial}_{\mu}\right\}$, of the cotangent space, $\mathcal{T}_{\mathcal{P}}^{*}$, dual to $\mathcal{T}_{\mathcal{P}}$. In other words, when $\mathbf{e}_{\mu}=\partial_{\mu}$, then $\boldsymbol{\omega}^{\mu}=\mathbf{d} x^{\mu}$. A general 1-form, when written in a coordinate cobasis as $\boldsymbol{\sigma}=\sigma_{\alpha} \mathbf{d} x^{\alpha}$, is often called a differential form. Warning: in general, a differential 1-form is not the differential of a function!

To find the components of $\mathbf{d} f$ in the $\left\{\mathbf{d} x^{\mu}\right\}$ coordinate cobasis, recall that the action of a cobasis on a vector $\mathbf{t}$ outputs the corresponding component of the vector: $\mathbf{d} x^{\mu}(\mathbf{t})=t^{\mu}$. Then, from eq. (1.7) and eq. (1.12):

$$
\begin{equation*}
[\mathbf{d} f](\mathbf{t})=\mathbf{t}(f)=t^{\mu} \partial_{\mu} f=\partial_{\mu} f \mathbf{d} x^{\mu}(\mathbf{t}) \tag{1.13}
\end{equation*}
$$

so that the components are $\partial_{\mu} f$. If we hink of $f$ as a 0 -form, the differential of $f$ is the gradient 1 -form $\mathbf{d} f$ :

$$
\begin{equation*}
\mathbf{d} f=\partial_{\mu} f \mathbf{d} x^{\mu} \tag{1.14}
\end{equation*}
$$

We recognise the well-known expression for the differential of a function in calculus, where it is taken to be a scalar, a number. But $\mathbf{d} f$, interpreted as the infinitesimal change of $f$ does not know in which direction this change should be evaluated. Only when a vector is inserted in its input slot, as in eq. (1.13), can it output a number, the change of $f$ in the direction of the vector.

As for the usual calculus interpretation of $\mathbf{d} x^{\mu}$ as the difference between the components of two coordinate vectors at infinitesimally close points, this may be acceptable in $\mathbb{R}^{n}$, but not if $x^{\mu}$ are coordinates on an arbitrary manifold, since $\mathbf{d} x^{\mu}$, like all vectors and $p$-forms at a point, does not live on the manifold, but in the tangent (co)space. Only in $\mathbb{R}^{n}$ can one ignore with impunity this crucial distinction between a base manifold and its tangent space at a point.

### 1.2.7 Vectors as linear mappings

Since $\mathcal{L}(\mathcal{V}, \mathbb{R})$, or $\mathcal{V}^{*}$, is technically a vector space, it has its own dual space, $\mathcal{L}\left(\mathcal{V}^{*}, \mathbb{R}\right)$, or $\mathcal{V}^{* *}$. We realise that nothing prevents us (with finite-dimensional spaces) from considering the elements $\mathbf{v} \in \mathcal{V}$ as themselves linear mappings on $\mathcal{V}^{*}$, and identifying $\mathcal{V}^{* *}$ with $\mathcal{V}$ ! Then $\mathbf{e}_{\alpha}\left(\boldsymbol{\omega}^{\beta}\right)=\delta_{\alpha}{ }^{\beta}$, and we would find: $\mathbf{v}(\boldsymbol{\sigma})=v^{\alpha} \sigma_{\alpha}$, exactly as in eq. (1.11) above.

These considerations suggest that we can also view a 1 -form (covector), as a kind of machine ${ }^{\dagger}$ whose input is a vector and whose output is a number; similarly, a vector can function as a device with a 1 -form as input and a number as output. To summarise, and noting that indices on vectors and covectors are labels specifying which of them, not which of their components, we construct a table that tells us how to calculate the output of a 1-form acting on a vector, and of a vector acting on a 1 -form:

| 1-form | Input vector | Output | Vector | Input 1-form | Output |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Cobasis element $\boldsymbol{\omega}^{\alpha}$ | Basis element $\mathbf{e}_{\beta}$ | $\boldsymbol{\omega}^{\alpha}\left(\mathbf{e}_{\beta}\right)=\delta^{\alpha}{ }_{\beta}$ | $\mathbf{e}_{\alpha}$ | $\boldsymbol{\omega}^{\beta}$ | $\mathbf{e}_{\beta}\left(\boldsymbol{\omega}^{\alpha}\right)=\delta^{\alpha}{ }_{\beta}$ |
| Cobasis element $\boldsymbol{\omega}^{\alpha}$ | $\mathbf{v}$ | $\boldsymbol{\omega}^{\alpha}(\mathbf{v})=v^{\alpha}$ | $\mathbf{e}_{\alpha}$ | $\boldsymbol{\sigma}$ | $\mathbf{e}_{\alpha}(\boldsymbol{\sigma})=\sigma_{\alpha}$ |
| $\boldsymbol{\sigma}$ | Basis element $\mathbf{e}_{\alpha}$ | $\boldsymbol{\sigma}\left(\mathbf{e}_{\alpha}\right)=\sigma_{\alpha}$ | $\mathbf{v}$ | $\boldsymbol{\omega}^{\alpha}$ | $\mathbf{v}\left(\boldsymbol{\omega}^{\alpha}\right)=v^{\alpha}$ |
| $\boldsymbol{\sigma}$ | $\mathbf{v}$ | $\boldsymbol{\sigma}(\mathbf{v})=\sigma_{\alpha} v^{\alpha}$ | $\mathbf{v}$ | $\boldsymbol{\sigma}$ | $\mathbf{v}(\boldsymbol{\sigma})=v^{\alpha} \sigma_{\alpha}$ |

Note that $\boldsymbol{\sigma}(\mathbf{v})=\mathbf{v}(\boldsymbol{\sigma})=\sigma_{\alpha} v^{\alpha}$ is basis-independent, but only if $\boldsymbol{\sigma}$ is referred to the cobasis of the basis in which $\mathbf{v}$ is written. At this stage, there is no natural (basis-independent) isomorphism between a vector space and its dual space, because there is as yet no unique connection between a given $\mathbf{v} \in \mathcal{V}$ and an element of $\in \mathcal{V}^{*}$. So, tempting as it is to identify it with the scalar product of two vectors, let us resist that urge. The $\sigma_{\alpha}$ are components of a 1 -form, not of a vector! As we shall soon discover, vector and 1 -form components behave differently under the same linear transformation. For the moment $\sigma_{\alpha} v^{\alpha}$ is the real number resulting from the 1 -form $\sigma$ acting on the vector $\mathbf{v}$ or, equivalently, $\mathbf{v}$ on $\boldsymbol{\sigma}$. This mapping is represented by a linear combination of the components of $\mathbf{v}$ with the coefficients of the combination the components of $\sigma$ or, equivalently, by matrix multiplication.

For a given vector $\mathbf{v}$, there exists a unique set of parallel ( $n-1$ )-dimensional hyperplanes that can provide a geometric picture of 1 -forms. This is very easy to do when $n=2$ (vector in a plane). Then any number $a=\sigma_{1} v^{1}+\sigma_{2} v^{2}$ determines a straight line perpendicular to $\mathbf{v}$ with equation $\sigma_{2}=a / v^{2}-\sigma_{1} v^{1} / v^{2}$. The lines generated by different values of $a$ all have identical slope $-v^{1} / v^{2}$.

### 1.2.8 Transformations on bases, cobases, and components

Let $\left(U_{1}, x\right)$ and $\left(U_{2}, y\right)$ be two overlapping charts (see definition 1.2) on a manifold $M$, with $x$ and $y$ their coordinate functions, respectively. Consider a point $\mathcal{P} \in U_{1} \cap U_{2}$. Let us obtain the relation between $\left.\boldsymbol{\partial}_{x^{\mu}}\right|_{x_{\mathcal{P}}}$ and $\left.\partial_{y^{\nu}}\right|_{y_{\mathcal{P}}}$, the coordinates bases for the two charts. These are maps, which we let act on some arbitrary differentiable function $f$. We remember that because $f$ acts on the manifold, we must write $\left.\partial_{x^{\mu}} f\right|_{x_{\mathcal{P}}}=\left.\partial_{\mu}\left(f \circ x^{-1}\right)\right|_{x_{\mathcal{P}}}$. Insert $y^{-1} \circ y$ and use the multidimensional version of the chain rule $(f \circ g)^{\prime}(\mathcal{P})=g^{\prime}(\mathcal{P}) f^{\prime}[g(\mathcal{P})]$ (written in the order opposite the usual one):

$$
\begin{align*}
\left.\partial_{x^{\mu}} f\right|_{x_{\mathcal{P}}} & =\partial_{\mu}\left[\left(f \circ y^{-1}\right) \circ\left(y \circ x^{-1}\right)\right]_{x_{\mathcal{P}}} \\
& =\left.\left.\partial_{x^{\mu}}\left(y \circ x^{-1}\right)^{\nu}\right|_{x_{\mathcal{P}}} \partial_{y^{\nu}}\left(f \circ y^{-1}\right)\right|_{\left(y \circ x^{-1}\right)\left(x_{\mathcal{P}}\right)} \\
& =\left.\left.\partial_{x^{\mu}} y^{\nu}\right|_{x_{\mathcal{P}}} \partial_{y^{\nu}} f\right|_{y_{\mathcal{P}}} \tag{1.15}
\end{align*}
$$

[^3]A vector $\mathbf{v} \in \mathcal{T}_{\mathcal{P}}$ must remain invariant under change of chart. That is: $\mathbf{v}=\left.v_{x}^{\mu} \boldsymbol{\partial}_{x^{\mu}}\right|_{x_{\mathcal{P}}}=\left.v_{y}^{\lambda} \boldsymbol{\partial}_{y^{\lambda}}\right|_{y_{\mathcal{P}}}$. Inserting the transformation law for the coordinate bases, we immediately find he transformation law for the components of $\mathbf{v}$ :

$$
\begin{equation*}
v_{y}^{\nu}=\left.\partial_{x^{\mu}} y^{\nu}\right|_{x_{\mathcal{P}}} v_{x}^{\mu} \tag{1.16}
\end{equation*}
$$

What is remarkable of both transformations is that they are linear and homogeneous, even though the transformations beween $\left(U_{1}, x\right)$ and $\left(U_{2}, y\right)$ can be non-linear. Thus, in coordinate bases, the coefficients, $\partial_{x^{\mu}} y^{\nu}$, in the transformation law are the elements of the Jacobian matrix of the transformation evaluated at $\mathcal{P}$. Conversely, if $v_{x}^{\nu}=v_{y}^{\mu} \partial_{y^{\mu}} x^{\nu}$, one shows easily (do it!), using the chain rule on partial derivatives, that $\mathbf{v}$ is indeed unchanged by the transformation.

In general bases, the transformations must be assumed homogeneous and linear, and take the form:

$$
\begin{equation*}
\mathbf{e}_{\mu}=L^{\alpha^{\prime}}{ }_{\mu} \mathbf{e}_{\alpha^{\prime}}=\mathbf{e}_{\alpha^{\prime}} L^{\alpha^{\prime}}{ }_{\mu} \tag{1.17}
\end{equation*}
$$

where the prime refers to the $y$ coordinates in eq. (1.15) and (1.16). Components transform as:

$$
\begin{equation*}
v^{\alpha^{\prime}}=v^{\mu} L^{\alpha^{\prime}}{ }_{\mu}=L^{\alpha^{\prime}}{ }_{\mu} v^{\mu} \tag{1.18}
\end{equation*}
$$

This is the more traditional definition of a vector.
The two ways of writing $v^{\alpha^{\prime}}$ in eq. (1.18) are equivalent, but the second one is a matrix product. The second expression in eq. (1.17) (and also eq. (1.15)), however, is not matrix multiplication, because the subscript of the basis vector is a label for a whole vector, not for a component of this vector.
$\mathbf{L}$ being non-singular, and therefore invertible, the action of the inverse transformation $\mathbf{L}^{-1}$ is represented by:

$$
\begin{equation*}
\mathbf{v}=\mathbf{L}^{-1} \mathbf{v}^{\prime} \quad \Longleftrightarrow \quad v^{\mu}=\left(\mathbf{L}^{-1}\right)^{\mu}{ }_{\nu^{\prime}} v^{\nu^{\prime}} \tag{1.19}
\end{equation*}
$$

Do not confuse matrix and index notation! Whereas matrix notation is readily translated into index notation, the reverse generally requires some rearrangement. This is because index notation does not care about ordering-one of its virtues-but matrix notation most certainly does.

Let $\left\{\mathbf{e}_{\mu}\right\}$ and $\left\{\mathbf{e}_{\mu^{\prime}}\right\}$ be two bases in $\mathcal{V}^{n}$, connected by $\mathbf{e}_{\mu}=\mathbf{e}_{\nu^{\prime}} L^{\nu^{\prime}}{ }_{\mu}$, where the $L^{\nu^{\prime}}{ }_{\mu}$ are the coefficients of the matrix $\mathbf{L}$ representing a linear transformation $\mathbf{L}$. Let $\left\{\boldsymbol{\omega}^{\alpha}\right\}$ and $\left\{\boldsymbol{\omega}^{\alpha}\right\}$ be their two respective cobases in $\mathcal{V}^{*}$. Then, writing $\boldsymbol{\omega}^{\alpha}=M^{\alpha}{ }_{\beta^{\prime}} \boldsymbol{\omega}^{\beta^{\prime}}$ where the $M^{\alpha}{ }_{\beta^{\prime}}$ are the matrix coefficients of the corresponding transformation $\mathbf{M}$ between the cobases, it can be shown (EXERCISE) that $\mathbf{M}$ is the inverse of $\mathbf{L}$, ie. $M_{\nu^{\prime}}^{\alpha} L^{\nu^{\prime}}{ }_{\beta}=\delta^{\alpha}{ }_{\beta}$ in index notation and $\mathbf{M}=\mathbf{L}^{-1}$ in matrix notation. This means that the transformation that takes the unprimed cobasis to the primed cobasis must be $\mathbf{L}: \boldsymbol{\omega}^{\alpha^{\prime}}=L^{\alpha^{\prime}}{ }_{\beta} \boldsymbol{\omega}^{\beta}$.

In the same way as for vector components, we can then obtain (EXERCISE) the transformation law of the components $\sigma_{\alpha}$ of a 1-form $\boldsymbol{\sigma}$. Since $\boldsymbol{\sigma}$ must be cobasis-independent, $\sigma_{\alpha} \boldsymbol{\omega}^{\alpha}=\sigma_{\beta^{\prime}} \boldsymbol{\omega}^{\beta^{\prime}}$ yields:

$$
\begin{equation*}
\sigma_{\alpha^{\prime}}=\sigma_{\mu}\left(\mathbf{L}^{-1}\right)^{\mu}{ }_{\alpha^{\prime}} \tag{1.20}
\end{equation*}
$$

while the inverse matrix, $\mathbf{M}=\mathbf{L}^{-1}$, takes the components in the opposite direction.
The following table summarises all the possible transformations, with the derivative expressions for components applying in coordinate bases:

| $\mathbf{e}_{\alpha^{\prime}}=\mathbf{e}_{\beta}\left(\mathbf{L}^{-1}\right)^{\beta}{ }_{\alpha^{\prime}}=\mathbf{e}_{\beta} \partial_{\alpha^{\prime}} x^{\beta}$ | $\mathbf{e}_{\alpha}=\mathbf{e}_{\beta^{\prime}} L^{\beta^{\prime}}{ }_{\alpha}=\mathbf{e}_{\beta^{\prime}} \partial_{\alpha} x^{\beta^{\prime}}$ |  |
| :---: | :---: | :---: |
| $v^{\alpha^{\prime}}=L^{\alpha^{\prime}}{ }_{\beta} v^{\beta}=\partial_{\beta} x^{\alpha^{\prime}} v^{\beta}$ | $v^{\alpha}=\left(\mathbf{L}^{-1}\right)^{\alpha}{ }_{{ }^{\prime}} v^{\beta^{\prime}}=\partial_{\beta^{\prime}} x^{\alpha} v^{\beta^{\prime}}$ |  |
| $\boldsymbol{\omega}^{\alpha^{\prime}}=L^{\alpha^{\prime}}{ }_{\beta} \boldsymbol{\omega}^{\beta}=\partial_{\beta} x^{\alpha^{\prime}} v^{\beta}$ | $\boldsymbol{\omega}^{\alpha}=\left(\mathbf{L}^{-1}\right)^{\alpha}{ }_{\beta^{\prime}} \boldsymbol{\omega}^{\beta^{\prime}}=\partial_{\beta^{\prime}} x^{\alpha} \boldsymbol{\omega}^{\beta^{\prime}}$ |  |
| $\sigma_{\alpha^{\prime}}=\sigma_{\beta}\left(\mathbf{L}^{-1}\right)^{\beta}{ }_{\alpha^{\prime}}$ | $\sigma_{\alpha}=\sigma_{\beta^{\prime}} L^{\beta^{\prime}}{ }_{\alpha}=\sigma_{\beta^{\prime}} \partial_{\alpha} x^{\beta^{\prime}}$ |  |
| $\sigma_{\alpha} v^{\alpha}=\sigma_{\beta^{\prime}} v^{\beta^{\prime}}$ |  |  |
|  |  |  |

Care should be exercised when comparing this table to the expressions given in $\S 2.9$ and in Box 8.4 of MTW which refer to Lorentz transformations. In their potentially confusing but standard notation, the matrix with elements $L^{\alpha}{ }_{\beta^{\prime}}$ is actually the inverse of the matrix with elements $L^{\beta^{\prime}}{ }_{\alpha}$; we prefer making this explicit by writing $\left(\mathbf{L}^{-1}\right)^{\alpha}{ }_{\beta^{\prime}}$.

Another word of caution: transformations in coordinate bases may well produce components in non-normalised bases, even if one starts from a normalised basis. This does not occur in the case of rotations and Lorentz boosts, but it will when we transform from Cartesian to curvilinear coordinates.

Also, if we wish to stick to a coordinate basis, we cannot call $\partial_{\mu} f$ the components of the gradient vector, $\nabla f$. These do not transform as the components of a vector, as can be seen by calculating $\partial_{\mu^{\prime}} f$ in terms of $\partial_{\nu} f$ using the chain rule (EXERCISE) and comparing with the form of the above transformation matrix in terms of derivatives. Moreover, everyone knows that in curvilinear coordinates (polar, spherical, cylindrical), the components of the gradient vector are not just the partial derivatives with respect to the coordinates. But even in Cartesian coordinates, which is a coordinate basis, $\partial_{\mu} f$ are not the components of a vector, but of a 1-form!

### 1.3 At Last, Tensors!

Our previous discussions make it straightforward to extend the concept of linear mappings to that of multilinear mappings, ie. mappings which are linear in each of their arguments, with the other arguments held fixed.

With $\mathcal{V}$ and its dual space $\mathcal{V}^{*}$, equipped respectively with coordinate basis $\left\{\boldsymbol{\partial}_{\nu_{i}}\right\}$ and cobasis $\left\{\mathbf{d} x^{\mu_{i}}\right\}(1 \leq i \leq$ $n$ ), we construct the space of multilinear mappings, $\left\{T: \mathcal{V}^{*} \times \ldots \times \mathcal{V}^{*} \times \mathcal{V} \times \ldots \times \mathcal{V} \longrightarrow \mathbb{R}\right\}$, in the following way:

Definition 1.13. Contravariant tensors $\mathbf{T} \in \mathcal{T}^{r}$ of $\operatorname{rank} r$ are real multilinear functions of $r$ 1-forms:

$$
\begin{equation*}
\mathbf{T}\left(\boldsymbol{\sigma}_{1}, \ldots, \boldsymbol{\sigma}_{r}\right)=\sigma_{\mu_{1}} \ldots \sigma_{\mu_{r}} \mathbf{T}\left(\mathbf{d} x^{\mu_{1}}, \ldots, \mathbf{d} x^{\mu_{r}}\right)=T^{\mu_{1} \ldots \mu_{r}} \sigma_{\mu_{1}} \ldots \sigma_{\mu_{r}} \tag{1.21}
\end{equation*}
$$

Covariant tensors $\mathbf{S} \in \mathcal{T}_{s}$ of rank $s$ are real multilinear functions of $s$ vectors:

$$
\begin{equation*}
\mathbf{S}\left(\mathbf{u}_{1}, \ldots, \mathbf{u}_{s}\right)=u^{\nu_{1}} \ldots u^{\nu_{s}} \mathbf{S}\left(\boldsymbol{\partial}_{\nu_{1}}, \ldots, \boldsymbol{\partial}_{\nu_{s}}\right)=S_{\nu_{1} \ldots \nu_{s}} u^{\nu_{1}} \ldots u^{\nu_{s}} \tag{1.22}
\end{equation*}
$$

Mixed tansors of type $(r, s)$ are real functions of $r$ covectors and $s$ vectors:

$$
\begin{align*}
\mathbf{Q}\left(\boldsymbol{\sigma}_{1}, \ldots, \boldsymbol{\sigma}_{r}, \mathbf{u}_{1}, \ldots, \mathbf{u}_{s}\right) & =\sigma_{\mu_{1} \ldots \sigma_{\mu_{r}}} u^{\nu_{1}} \ldots u^{\nu_{s}} \mathbf{Q}\left(\mathbf{d} x^{\mu_{1}}, \ldots, \mathbf{d} x^{\mu_{r}}, \boldsymbol{\partial}_{\nu_{1}}, \ldots, \boldsymbol{\partial}_{\nu_{s}}\right) \\
& =Q_{\nu_{1} \ldots \nu_{s}}^{\mu_{1} \ldots \mu_{r}} \sigma_{\mu_{1}} \ldots \sigma_{\mu_{r}} u^{\nu_{1}} \ldots u^{\nu_{s}} \tag{1.23}
\end{align*}
$$

$T^{\mu_{1} \ldots \mu_{r}}, S_{\nu_{1} \ldots \nu_{s}}$, and $Q^{\mu_{1} \ldots \mu_{r}}{ }_{\nu_{1} \ldots \nu_{s}}$ are the contravariant components of $\mathbf{T}$, the covariant components of $\mathbf{S}$, and the mixed components of $\mathbf{Q}$, respectively, in the chosen basis and cobasis.
Following the metaphor of tensors as machines, to output a number from a $(r, s)$ tensor, one must supply $r 1$-forms and $s$ vectors as input, one for each slot.

### 1.3.1 The tensor product

There is an important kind of multilinear mapping we can construct, this time out of known building blocks.

Definition 1.14. The Kronecker (tensor) product space of $\mathcal{V}_{1}^{*}$ and $\mathcal{V}_{2}^{*}$ is a set of bilinear mappings $\mathcal{L}\left(\mathcal{V}_{1}, \mathcal{V}_{2}, \mathbb{R}\right)$, denoted by $\mathcal{V}_{1}^{*} \times \mathcal{V}_{2}^{*}$, with as product elements the covariant tensor $\boldsymbol{\sigma} \otimes \boldsymbol{\tau}$ :

$$
\begin{equation*}
\boldsymbol{\sigma} \otimes \boldsymbol{\tau}(\mathbf{u}, \mathbf{v})=\boldsymbol{\sigma}(\mathbf{u}) \boldsymbol{\tau}(\mathbf{v}) \tag{1.24}
\end{equation*}
$$

for all $\mathbf{u} \in \mathcal{V}_{1}, \mathbf{v} \in \mathcal{V}_{2}, \boldsymbol{\sigma} \in \mathcal{V}_{1}^{*}$, and $\boldsymbol{\tau} \in \mathcal{V}_{2}^{*}$.
Similarly, the product space $\mathcal{L}\left(\mathcal{V}_{1}^{*}, \mathcal{V}_{2}^{*}, \mathbb{R}\right)=\mathcal{V}_{1} \times \mathcal{V}_{2}$, has as elements the contravariant tensor of rank 2:

$$
\begin{equation*}
\mathbf{u} \otimes \mathbf{v}(\boldsymbol{\sigma}, \boldsymbol{\tau})=\mathbf{u}(\boldsymbol{\sigma}) \mathbf{v}(\boldsymbol{\tau}) \tag{1.25}
\end{equation*}
$$

Finally, there are tensor product spaces $\mathcal{V}_{1} \otimes \mathcal{V}_{2}^{*}$ with product elements $\mathbf{u} \otimes \boldsymbol{\sigma}(\boldsymbol{\tau}, \mathbf{v})=\mathbf{u}(\boldsymbol{\tau}) \boldsymbol{\sigma}(\mathbf{v})$, and $\mathcal{V}_{1}^{*} \times \mathcal{V}_{2}$, with product elements $\boldsymbol{\sigma} \otimes \mathbf{v}(\mathbf{u}, \boldsymbol{\tau})=\boldsymbol{\sigma}(\mathbf{u}) \mathbf{v}(\boldsymbol{\tau})$. If a tensor is the tensor product of other tensors, we say that it is decomposable.
It is important to note that the tensor product is not commutative!.

Example 1.5. Let $P$ be the vector space whose elements are polynomials of some degree $n$. Such a space can be constructed provided we define addition and s-multiplication of polynomials. Then we can construct a map $\mathbf{T}: P \times P \rightarrow \mathbb{R}$ defined by $\int_{0}^{1} p(x) q(x) \mathrm{d} x$, where $p, q \in P$. This bilinear map-call it the inner product- which takes as inputs two vectors and outputs a number, is a $(0,2)$ tensor.

Now take $\mathcal{V}_{1}=\mathcal{V}_{2}=\mathcal{V}$. If $\left\{\boldsymbol{\partial}_{\mu}\right\}$ and $\left\{\boldsymbol{\partial}_{\nu}\right\}$ are coordinate bases for $\mathcal{V}$, then $\left\{\boldsymbol{\partial}_{\mu} \otimes \boldsymbol{\partial}_{\nu}\right\}$ is a coordinate basis for $\mathcal{V} \otimes \mathcal{V}$. Similarly, if $\left\{\mathbf{d} x^{\alpha}\right\}$ and $\left\{\mathbf{d} x^{\beta}\right\}$ are coordinate bases for $\mathcal{V}^{*}$, then $\left\{\mathbf{d} x^{\alpha} \otimes \mathbf{d} x^{\beta}\right\}$ is a coordinate basis for $\mathcal{V}^{*} \otimes \mathcal{V}^{*}$.

We assert that any contravariant tensor of rank 2 lives in $\mathcal{V} \times \mathcal{V}$, and any covariant rank-2 tensor lives in $\mathcal{V}^{*} \times \mathcal{V}^{*}$, that is:

$$
\begin{equation*}
\mathbf{A}=A^{\mu \nu} \boldsymbol{\partial}_{\mu} \otimes \boldsymbol{\partial}_{\nu}, \quad \mathbf{B}=B_{\alpha \beta} \mathbf{d} x^{\alpha} \otimes \mathbf{d} x^{\beta} \tag{1.26}
\end{equation*}
$$

Therefore, the action of $\mathbf{A}$ on pairs of one-forms and of $\mathbf{B}$ on pairs of vectors is given by:

$$
\begin{align*}
& \mathbf{A}(\boldsymbol{\sigma}, \boldsymbol{\tau})=A^{\mu \nu} \boldsymbol{\partial}_{\mu} \otimes \boldsymbol{\partial}_{\nu}(\boldsymbol{\sigma}, \boldsymbol{\tau})=A^{\mu \nu} \boldsymbol{\partial}_{\mu}(\boldsymbol{\sigma}) \boldsymbol{\partial}_{\nu}(\boldsymbol{\tau})=A^{\mu \nu} \sigma_{\mu} \tau_{\nu}  \tag{1.27}\\
& \mathbf{B}(\mathbf{u}, \mathbf{v})=B_{\alpha \beta} \mathbf{d} x^{\alpha} \otimes \mathbf{d} x^{\beta}(\mathbf{u}, \mathbf{v})=B_{\alpha \beta} \mathbf{d} x^{\alpha}(\mathbf{u}) \mathbf{d} x^{\beta}(\mathbf{v})=B_{\alpha \beta} u^{\alpha} v^{\beta}
\end{align*}
$$

As we have said before, both $\mathbf{A}$ and $\mathbf{B}$ can be viewed as operators, or devices, requiring two 1 -forms or two vectors, respectively, as ordered input, to output a product of numbers. But we can also input a single vector (1-form) and obtain a 1 -form (vector) as output, so long as we specify into which of the two input slots it should be inserted. For instance, we could write $\mathbf{B}(\mathbf{u}$, ), or $\mathbf{B}(, \mathbf{u})$, but just $\mathbf{B}(\mathbf{u})$ would (in general, but not always, as we shall see in a moment) be ambiguous. For instance:

$$
\begin{aligned}
\mathbf{B}(\mathbf{u},) & =B_{\alpha \beta} \mathbf{d} x^{\alpha}(\mathbf{u}) \mathbf{d} x^{\beta}=\sigma_{\beta} \mathbf{d} x^{\beta} \\
\mathbf{B}(, \mathbf{u}) & =B_{\alpha \beta} \mathbf{d} x^{\alpha} \mathbf{d} x^{\beta}(\mathbf{u})=\tau_{\alpha} \mathbf{d} x^{\alpha}
\end{aligned}
$$

where $\sigma_{\beta}=B_{\alpha \beta} u^{\alpha}$, and $\tau_{\alpha}=B_{\alpha \beta} u^{\beta}$. Unless the components $B_{\alpha \beta}$ happen to be symmetric in their indices, the two resulting 1 -forms are not the same!

With this in mind, $\mathbf{A}(\boldsymbol{\sigma}, \boldsymbol{\tau})$ is the same as $\boldsymbol{\sigma}(\mathbf{A}(, \boldsymbol{\tau}))$, where $\mathbf{A}(, \boldsymbol{\tau})$ is (exercise) a vector. In other words, the same machine can be put to quite different tasks.

More generally, if $\left\{\boldsymbol{\partial}_{\mu_{1}}\right\}, \ldots,\left\{\boldsymbol{\partial}_{\mu_{r}}\right\}$ are bases of $\mathcal{V}$, and $\left\{\mathbf{d} x^{\nu_{1}}\right\}, \ldots,\left\{\mathbf{d} x^{\nu_{s}}\right\}$ (co)bases of $\mathcal{V}^{*}$, then $\boldsymbol{\partial}_{\mu_{1}} \otimes \cdots \otimes$ $\boldsymbol{\partial}_{\mu_{r}} \otimes \mathbf{d} x^{\nu_{1}} \otimes \cdots \otimes \mathbf{d} x^{\nu_{s}}$ forms a basis for $\mathcal{T}_{s}^{r}$. Therefore, any tensor can be written in terms of its components as:

$$
\begin{equation*}
\mathbf{T}=T^{\mu_{1} \ldots \mu_{r}}{ }_{\nu_{1} \ldots \nu_{s}} \boldsymbol{\partial}_{\mu_{1}} \otimes \cdots \otimes \boldsymbol{\partial}_{\mu_{r}} \otimes \mathbf{d} x^{\nu_{1}} \otimes \cdots \otimes \mathbf{d} x^{\nu_{s}} \tag{1.28}
\end{equation*}
$$

To obtain a number, all input slots must be filled; but, as we saw for rank-2 tensors, we can also input one less vector and get a 1 -form as output, or one less 1 -form to get a vector. More generally, reducing the number of input vectors from $s$ to $s-m$ causes $\mathbf{T}$ to output a tensor of covariant rank $(0, m)$; reducing the number of input 1 -forms from $r$ to $r-q$ outputs a tensor of contravariant rank $(q, 0)$.

It is important to remember that, in general, interchanging vectors or 1 -forms in the input results in different output. Therefore, one should be mindful of the ordering of the vectors and 1 -forms that one feeds into a tensor.

### 1.3.2 Transposition, symmetric and skew-symmetric tensors

Interchanging any two contravariant or any two covariant slots of a tensor produces a transpose of this tensor. Strictly speaking, interchanging a covariant and a contravariant slot of a tensor does not make sense,

Definition 1.15. If a tensor remains unchanged under transposition of two of its input slots of the same type, we say that it is symmetric in these slots. Its components are unchanged under permutation of indices corresponding to those slots.

If a tensor switches sign under transposition of two of its input slots of the same type, we say that it is antisymmetric in these slots, and the components corresponding to these slots also switch sign. inserting the same 1 -form (in the contravariant slots) or vector (in the covariant slots outputs zero.
Symmetry and antisymmetry are basis-independent properties.
Example 1.6. Take a covariant antisymmetric tensor of rank 2: $\mathbf{F}=F_{[\mu \nu]} \mathbf{d} x^{\mu} \otimes \mathbf{d} x^{\nu}$, where square brackets around indices mean that they are antisymmetric. Then $\mathbf{F}(\mathbf{u}, \mathbf{u})=F_{\mu \nu} u^{\mu} u^{\nu}=0$. To see this in index notation, just rename $\mu$ as $\nu$ and $\nu$ as $\mu$ and transpose indices in the antisymmetric $F_{\mu \nu}$ and symmetric $u^{\mu} u^{\nu}$ to get $F_{\mu \nu} u^{\mu} u^{\nu}=-F_{\mu \nu} u^{\mu} u^{\nu}$. Similarly, inserting the same cobasis 1-forms in two antisymmetric contravariant slots, or basis vectors in two antisymmetric covariant slots, just makes the corresponding indices identical, and the corresponding components vanish.

Among important tensors are those which are completely symmetric in all their covariant indices and all their contravariant indices (eg. $A_{\mu_{1} \ldots \mu_{s}}=A_{\mu_{\pi_{1}} \ldots \mu_{\pi_{s}}}$ where $\mu_{\pi_{1}} \ldots \mu_{\pi_{s}}$ is any permutation of $\mu_{1} \ldots \mu_{s}$ ), and those which are completely antisymmetric (skew-symmetric, alternating) in all their covariant indices and all their contravariant indices.

A completely symmetric tensor of rank $r$ in $n$ dimensions has $\binom{n+r-1}{r}=(n+r-1)!/(n-1)!r$ ! independent components. A skew-symmetric tensor has $\binom{n}{r}=n!/(n-r)!r!$ independent non-zero components in all coordinate systems.

In three dimensions, many physically relevant tensors are symmetric, eg. examples 1.6, 1.7 and 1.8 (moment of inertia, electrical polarisation, multipole moment) in B\&F, as well as the Maxwell stress tensor. Antisymmetric 3-d rank-2 tensors are not usual, although I will argue toward the end of the module that in three dimensions a magnetic field is more naturally described by a rank-2 antisymmetric covariant tensor than by a vector.

In four dimensions, we also have symmetric tensors, such as the important energy-momentum tensor which carries all the information about the energy and momentum density at a point, plus the flux of these quantities at that point. And there is a famous antisymmetric, covariant, rank-2 tensor, the Faraday field tensor F, whose six independent non-zero components are the components of electric and magnetic field 3 -vectors.

Even when a tensor is neither completely symmetric nor skew-symmetric, it can be useful to construct a symmetrised or skew-symmetrised version of it. The simplest example is that of a rank-2 contravariant (or covariant) tensor. Its components can be written as:

$$
T^{\mu \nu}=\frac{1}{2}\left(T^{\mu \nu}+T^{\nu \mu}\right)+\frac{1}{2}\left(T^{\mu \nu}-T^{\nu \mu}\right) \equiv T^{(\mu \nu)}+T^{[\mu \nu]}
$$

with round brackets around symmetric indices. In this case, we say that $\mathbf{T}$ has been decomposed into a symmetrised tensor, with components $T^{(\mu \nu)}$, and a skew-symmetric tensor, with components $T^{[\mu \nu]}$.

More generally, $(r, 0)$ and $(0, s)$ tensors can be symmetrised and antisymmetrised, in the following ways:
Definition 1.16. To symmetrise the components of a covariant tensor $\mathbf{T}$, we can apply the following prescription to its components:

$$
\begin{equation*}
T_{\left(\mu_{1} \ldots \mu_{s}\right)}=\frac{1}{s!} \sum_{\pi} T_{\mu_{\pi_{1}} \ldots \mu_{\pi_{s}}} \tag{1.29}
\end{equation*}
$$

where the sums run over all permutations of $1, \ldots, s$. Contravariant components are symmetrised by the same procedure.

Definition 1.17. To antisymmetrise the components of a covariant (or contravariant) tensor $\mathbf{T}$, we can apply the following prescription to its components:

$$
\begin{equation*}
T_{\left[\mu_{1} \ldots \mu_{s}\right]}=\frac{1}{s!} \delta_{\mu_{1} \ldots \mu_{s}}^{\nu_{1} \ldots \nu_{s}} T_{\nu_{1} \ldots \nu_{s}} \tag{1.30}
\end{equation*}
$$

where the general permutation symbol, $\delta_{i_{1} \ldots i_{s}}^{j_{1} \ldots j_{s}}$, is defined as:

$$
\delta_{i_{1} \ldots i_{s}}^{j_{1} \ldots j_{s}}=\left\{\begin{array}{cl}
+1 & j_{1} \ldots j_{s} \text { an even permutation of } i_{1} \ldots i_{s}  \tag{1.31}\\
-1 & j_{1} \ldots j_{s} \text { an odd permutation of } i_{1} \ldots i_{s} \\
0 & j_{1} \ldots j_{s} \text { not a permutation of } i_{1} \ldots i_{s} \\
0 & j_{k}=j_{l} \text { or } i_{k}=i_{l} \text { for some } k, l
\end{array}\right.
$$

The permutation symbol is seen to be antisymmetric in its upper and lower indices. Clearly, as expected, $T_{\left[\mu_{1} \ldots \mu_{s}\right]}=0$ whenever any two of its indices are the same.
$s$ ! is the number of terms in all these summations, ie. the number of permutations of the indices of the tensor. The normalisation factor $1 / s$ ! ensures consistency in the event that the $T_{\mu_{1} \ldots \mu_{s}}$ should already be symmetric or skew-symmetric.
$(0, s)$ skew-symmetric tensors live in a space denoted by $\bigwedge^{s}\left(\mathcal{V}^{*}\right)$, whose elements are also called p-forms ( $p$ is traditionally used instead of $s$ ); very often, $\mathcal{V}^{*}=\mathbb{R}^{n}$.

EXERCISE: Symmetrise and antisymmetrise $\mathbf{F}(\boldsymbol{\sigma}, \boldsymbol{\tau}, \boldsymbol{\theta})$. Write $\mathbf{F}_{\mathrm{s}}$ and $\mathbf{F}_{\mathrm{a}}$ in an explicit basis such that $\mathbf{F}=$ $F^{\mu \nu \lambda} \mathbf{e}_{\mu} \otimes \mathbf{e}_{\nu} \otimes \mathbf{e}_{\lambda}$. How many components do $\mathbf{F}_{\mathrm{s}}$ and $\mathbf{F}_{\mathrm{a}}$ have when $\mathbf{F}$ is defined over a 3-dim space? a 4-dim space? Can you reconstruct $F^{\mu \nu \lambda}$ from $F^{(\mu \nu \lambda)}$ and $F^{[\mu \nu \lambda]}$ ?

### 1.3.3 Transformations on tensors

Using the transformations for a basis and a cobasis in the table of section 1.2.8, it is straightforward to generalise the transformation laws obeyed by tensor components. First, write $\mathbf{T}$ in the original basis and in the new (primed) basis:
$\mathbf{T}=T^{\mu_{1} \ldots \mu_{r}}{ }_{\nu_{1} \ldots \nu_{s}} \boldsymbol{\partial}_{\mu_{1}} \otimes \cdots \otimes \boldsymbol{\partial}_{\mu_{r}} \otimes \mathbf{d} x^{\nu_{1}} \otimes \cdots \otimes \mathbf{d} x^{\nu_{s}}=T^{\alpha_{1}^{\prime} \ldots \alpha_{r}^{\prime}}{ }_{\beta_{1}^{\prime} \ldots \beta_{s}^{\prime}} \boldsymbol{\partial}_{\alpha_{1}^{\prime}} \otimes \cdots \otimes \boldsymbol{\partial}_{\alpha_{r}^{\prime}} \otimes \mathbf{d} x^{\beta_{1}^{\prime}} \otimes \cdots \otimes \mathbf{d} x^{\beta_{s}^{\prime}}$
We obtain:

$$
\begin{equation*}
T^{\alpha_{1}^{\prime} \ldots \alpha_{r}^{\prime}}{ }_{\beta_{1}^{\prime} \ldots \beta_{s}^{\prime}}=T_{\nu_{1} \ldots \nu_{s}}^{\mu_{1} \ldots \mu_{r}} L_{\mu_{1}}^{\alpha_{1}^{\prime}} \ldots L_{\mu_{r}}^{\alpha_{r}^{\prime}}\left(L^{-1}\right)^{\nu_{1}}{ }_{\beta_{1}^{\prime}} \ldots\left(L^{-1}\right)^{\nu_{s}}{ }_{\beta_{s}^{\prime}} \tag{1.32}
\end{equation*}
$$

In traditional treatments, this transformation law actually defines a tensor. Scalars (tensors of type ( 0,0 )) must remain invariant; and we know how the components of vectors and 1-forms transform. What about, say, those of a tensor of type $(2,0)$ ?

$$
T^{\alpha^{\prime} \beta^{\prime}}=T^{\mu \nu} L_{\mu}^{\alpha^{\prime}} L_{\nu}^{\beta^{\prime}}=L_{\mu}^{\alpha^{\prime}} T^{\mu \nu} \tilde{L}_{\nu}^{\beta^{\prime}} \quad \Longleftrightarrow \quad \mathbf{T}^{\prime}=\mathbf{L} \mathbf{T} \tilde{\mathbf{L}}
$$

$\underset{\sim}{w}$ where $\tilde{\mathbf{L}}$ is the transpose of $\mathbf{L}$. Sometimes, as with 3-dim rotations, $\tilde{\mathbf{L}}=\mathbf{L}^{-1}$; sometimes, as with Lorentz boosts, $\tilde{\mathbf{L}}=\mathbf{L}$. Tensors of rank $2((2,0),(0,2),(1,1))$ can be represented by $n \times n$-dim matrices $\mathbf{T}$, where $n$ is the dimension of the spaces $\mathcal{V}$ and $\mathcal{V}^{*}$ on which they are defined. Unless they are decomposable, this is not true of tensors of higher rank. But the index notation itself does not care about the rank of a tensor.

An immediate consequence of eq. (1.32) is that a tensor that is zero in one basis will remain zero in any other transformed basis. Thus, any equation made of tensors that is valid in one basis must hold in any other basis. This might seem trivial in the geometric notation $\mathbf{T}$ of tensors, but the nice thing is that relations between their components are also basis-independent!

In the older view of tensors defined in terms of transformations, it is possible to restrict the tensor character to objects which may have it under certain transformations, but not others. For instance, 4-dim tensors might owe their tensor character to how they transform under Lorentz transformations, while 3-dim tensors might be tensors only under rotations. Then writing equations in terms of 4-tensors guarantees that they are consistent with Einstein's Relativity, ie. an equation valid in one Lorentz frame is valid in any other Lorentz-transformed frame.

The transformation rules can always be used to establish whether an object is a tensor. For instance, on a space of dimension $n$, the Kronecker delta, with components $\delta^{\nu}{ }_{\mu}$, is represented by the $n \times n$ identity matrix. It is a mixed rank-2 tensor. Indeed, from the transformation law, eq. (1.32):

$$
\delta_{\nu^{\prime}}^{\mu^{\prime}}=L_{\lambda}^{\mu^{\prime}}\left(L^{-1}\right)_{\nu^{\prime}}^{\rho} \delta_{\rho}^{\lambda}=L_{\lambda}^{\mu^{\prime}}\left(L^{-1}\right)_{\nu^{\prime}}^{\lambda}=\mathbb{I}_{\nu^{\prime}}^{\mu^{\prime}}
$$

which are the components of the identity matrix. Here we learn that there is something more to $\delta^{\mu}{ }_{\nu}$ than just being a tensor: its components remain the same under changes of basis!

### 1.3.4 The Levi-Civita symbol

Definition 1.18. In a Cartesian orthonormal basis of a $n$-dim space, the Levi-Civita symbol, $\epsilon_{\mu_{1} \ldots \mu_{n}}$, is defined in terms of the general permutation symbol, $\delta_{i_{1} \ldots i_{n}}^{j_{1} \ldots j_{n}}$ (eq. (1.31)), as:

$$
\epsilon_{\mu_{1} \ldots \mu_{n}}=\delta_{\mu_{1} \ldots \mu_{n}}^{1 \ldots n}
$$

It is skew-symmetric in its $n$ indices, with $\epsilon_{1 \ldots n}=+1$, where the indices are in ascending order. In pseudo-Riemannian manifolds, it is traditional to use $\epsilon_{0 \ldots n-1}$, the 0 index corresponding to time.

The determinant of a $n \times n$ matrix $\mathbf{L}$ is a product of its elements antisymmetrised with respect to rows (or columns):

$$
\begin{equation*}
\operatorname{det} \mathbf{L}=\epsilon_{\nu_{1} \ldots \nu_{n}} L^{\nu_{1}} \cdots L^{\nu_{n}}{ }_{n} \tag{1.33}
\end{equation*}
$$

This equation means that the Levi-Civita symbol is a tensor only when its components are transformed with a matrix whose Jacobian determinant is 1 (eg. rotations in 3-dim space, Lorentz boosts), in which case these components, like those of the Kronecker delta, are invariant under changes of basis. Indeed, if the Levi-Civita symbol is to be a tensor, the transformation laws on its components demand that:

$$
1=\epsilon_{1 \ldots n}=\epsilon_{\nu_{1}^{\prime} \ldots \nu_{n}^{\prime}} L^{\nu_{1}^{\prime}} \cdots L^{\nu_{n}^{\prime}}{ }_{n}=\operatorname{det} \mathbf{L}
$$

We shall discover a little later how the Levi-Civita general tensor (in all bases) can be constructed.
Fortunately, we can avoid using the transformation law (1.32) if we build tensors from other objects known to be tensors. The following sub-sections present some important examples.

### 1.4 Two More Ways to Construct Tensors

### 1.4.1 Contracted tensors

Definition 1.19. The contraction of a mixed-type tensor is a linear mapping $\mathcal{T}_{s}^{r} \rightarrow \mathcal{T}_{s-1}^{r-1},(r \geq 1$, $s \geq 1$ ). More precisely, going back to eq. (1.21), insert input (co)bases into only two tensor slots: $\boldsymbol{\sigma}^{i}=\mathbf{d} x^{\mu_{i}}(1 \leq i \leq r)$ and $\mathbf{u}_{j}=\boldsymbol{\partial}_{\nu_{j}}(1 \leq j \leq s)$, with $\nu_{j}=\mu_{i} \equiv \gamma$ :

$$
\begin{equation*}
\mathbf{T}\left(\ldots, \mathbf{d} x^{\gamma}, \ldots, \boldsymbol{\partial}_{\gamma}, \ldots\right)=T^{\ldots \gamma \ldots} \ldots \gamma \ldots \boldsymbol{\partial}_{\mu_{1}} \otimes \ldots \otimes \boldsymbol{\partial}_{\mu_{i-1}} \otimes \boldsymbol{\partial}_{\mu_{i+1}} \otimes \ldots \mathbf{d} x^{\nu_{1}} \otimes \ldots \otimes \mathbf{d} x^{\nu_{j-1}} \otimes \mathbf{d} x^{\nu_{j+1}} \otimes \ldots \tag{1.34}
\end{equation*}
$$

Note the need to keep track of the position of the tensor slots. In terms of the components of a tensor, it just involves making one contravariant index $\mu$ the same as one of the covariant indices, $\nu$, by multiplying the component by $\delta^{\nu}{ }_{\mu}$, thus forcing a summation over these indices.

For instance, consider $\mathbf{T} \in \mathcal{T}_{1}^{1}$. The contraction of $\mathbf{T}=T^{\alpha}{ }_{\beta} \boldsymbol{\partial}_{\alpha} \otimes \mathbf{d} x^{\beta}$ is a scalar, called its trace:

$$
\operatorname{Tr} \mathbf{T}=\mathbf{T}\left(\mathbf{d} x^{\mu}, \boldsymbol{\partial}_{\mu}\right)=T_{\beta}^{\alpha} \boldsymbol{\partial}_{\alpha}\left(\mathbf{d} x^{\mu}\right) \mathbf{d} x^{\beta}\left(\boldsymbol{\partial}_{\mu}\right)=T_{\mu}^{\mu}=T_{\nu}^{\mu} \delta_{\mu}^{\nu}
$$

When contracting tensors of type higher than 2 , it is important to specify which indices are being contracted. Thus, the tensor $T^{\mu \nu}{ }_{\lambda} \partial_{\mu} \otimes \partial_{\nu} \otimes \mathbf{d} x^{\lambda}$ has two possible contractions: the vectors $T^{\mu \nu}{ }_{\mu} \boldsymbol{\partial}_{\nu}$ and $T^{\mu \nu}{ }_{\nu} \boldsymbol{\partial}_{\mu}$.

### 1.4.2 Inner product

Up to now, in the space $\mathcal{T}_{s}^{r}$ of tensors, there has been no unique link between tensors of type $(r, 0),(0, r)$, or $(r-q, q)$, all of the same total rank. If we wish to establish such a link, a new object must be introduced: a bilinear covariant tensor denoted by $\mathfrak{g}=g_{\mu \nu} \mathbf{d} x^{\mu} \otimes \mathbf{d} x^{\nu}$ in a coordinate basis.

Indeed, let us insert only one vector in, say, the second slot of $\mathfrak{g}$ :

$$
\mathfrak{g}(\quad, \mathbf{u})=g_{\mu \nu} \mathbf{d} x^{\mu} \mathbf{d} x^{\nu}(\mathbf{u})=g_{\mu \nu} u^{\nu} \mathbf{d} x^{\mu} \equiv \tilde{\mathbf{u}}
$$

As expected, the result is a 1-form, which we have called $\tilde{\mathbf{u}}$. The correspondence will be unique if we demand that $\mathfrak{g}$ be symmetric, because then $\mathfrak{g}(\mathbf{u}, \quad)=\mathfrak{g}(\quad, \mathbf{u})$, or $g_{\nu \mu}=g_{\mu \nu}=\mathfrak{g}\left(\boldsymbol{\partial}_{\mu}, \boldsymbol{\partial}_{\nu}\right)$ when working in a coordinate basis. So, although sometimes called a form, $\mathfrak{g}$ is not a 2 -form because it is not antisymmetric. As befits any $(0,2)$ tensor, inserting vectors in its two input slots yields the number: $\mathfrak{g}(\mathbf{u}, \mathbf{v})=g_{\mu \nu} u^{\mu} v^{\nu}=\mathbf{u g} \mathbf{v}$, with the last expression in matrix form.

In effect, $\mathfrak{g}$ may be thought of as an invertible (ie. $\operatorname{det} \mathbf{g} \neq 0$ ) mapping from $\mathcal{V}$ to its dual space! Once defined, it establishes a unique correspondence between a vector $\mathbf{u}$ and a 1-form $\tilde{\mathbf{u}}$.

Definition 1.20. The inner product of two vectors $\mathbf{u}$ and $\mathbf{v},\langle\mathbf{u}, \mathbf{v}\rangle$, is defined as:

$$
\begin{equation*}
<\mathbf{u}, \mathbf{v}>=\mathfrak{g}(\mathbf{u}, \mathbf{v})=g_{\mu \nu} u^{\mu} v^{\nu} \tag{1.35}
\end{equation*}
$$

Often, $\langle\mathbf{u}, \mathbf{v}\rangle$ is written $\mathbf{u} \cdot \mathbf{v}$.
Definition 1.21. In a coordinate basis, we take: $g_{\mu \nu}=<\boldsymbol{\partial}_{\mu}, \boldsymbol{\partial}_{\nu}>$, which is just the naïve scalar product of the two basis vectors $\boldsymbol{\partial}_{\mu}$ and $\boldsymbol{\partial}_{\nu}$. In a general basis, $g_{\mu \nu}=<\mathbf{e}_{\mu}, \mathbf{e}_{\nu}>$.
$\mathfrak{g}(\mathbf{u}, \mathbf{u})=g_{\mu \nu} u^{\mu} u^{\nu}$ is called the norm of $\mathbf{u}$. If it is positive (negative) $\forall \mathbf{u}$, we say that $\mathfrak{g}$ is positive (negative) definite. But if $\mathfrak{g}(\mathbf{u}, \mathbf{u})=0$ for some non-zero vector (null vector) $\mathbf{u}$, then $\mathfrak{g}$ is indefinite.

As mentioned above, $\mathfrak{g}$ must be invertible, and we can always write: $\mathbf{u}=\mathfrak{g}^{-1} \mathfrak{g}(\mathbf{u})=\mathfrak{g}^{-1}(\tilde{\mathbf{u}})=u_{\mu} \mathfrak{g}^{-1}\left(\mathbf{d} x^{\mu}\right)$. Now $\mathfrak{g}^{-1}$ must take 1 -forms to vectors, which means it must be a $(2,0)$ tensor $\mathfrak{g}^{-1}=\left(\mathfrak{g}^{-1}\right)^{\mu \nu} \boldsymbol{\partial}_{\mu} \otimes \boldsymbol{\partial}_{\nu}$. Then:

$$
\mathbf{u}=u_{\mu}\left(\mathfrak{g}^{-1}\right)^{\alpha \beta} \boldsymbol{\partial}_{\alpha} \boldsymbol{\partial}_{\beta}\left(\mathbf{d} x^{\mu}\right)=u_{\mu}\left(\mathfrak{g}^{-1}\right)^{\alpha \beta} \boldsymbol{\partial}_{\alpha} \delta_{\beta}^{\mu}=u_{\mu}\left(\mathfrak{g}^{-1}\right)^{\alpha \mu} \boldsymbol{\partial}_{\alpha}
$$

As will be justified soon, we identify $\left(\mathfrak{g}^{-1}\right)^{\mu \nu}$ with the contravariant components of $\mathfrak{g}, g^{\mu \nu}$, and, comparing with $\mathbf{u}=u^{\alpha} \boldsymbol{\partial}_{\alpha}$, we conclude that $u^{\mu}=g^{\mu \nu} u_{\nu}, u^{\mu}$ being thought now as the contravariant components of the 1-form.

These mappings between $\mathcal{V}$ and $\mathcal{V}^{*}$ can be applied to any tensor $\mathbf{T} \in \mathcal{T}$; in other words, $\mathfrak{g}$ may be used to convert any contravariant index of a given tensor into a covariant one, while $\mathfrak{g}^{-1}$ may be used to convert any covariant index of a given tensor into a contravariant one. The total rank of the tensor is not changed.

Definition 1.22. Converting a contravariant index of a tensor $\mathbf{T}$ to a covariant one is called "lowering the index". It is always accomplished by contracting any one index of $g_{\mu \nu}$ with the index to be lowered, eg. $T^{\mu}{ }_{\nu}=g_{\nu \lambda} T^{\mu \lambda}=\mathbf{T} \mathbf{g}$. After lowering an index, the corresponding input slot now accepts a vector instead of a 1-form.
Converting a covariant index of a tensor $\mathbf{T}$ to a contravariant one is called "raising the index". It is accomplished by contracting any one index of $g^{\mu \nu}$ with the index to be raised, eg. $T^{\mu}{ }_{\nu}=g^{\mu \lambda} T_{\lambda \nu}=$ $\mathbf{g}^{-1} \mathbf{T}$. After raising an index, the corresponding input slot now accepts a 1-form instead of a vector.

It is usual to speak of $g_{\mu \nu} u^{\nu}$ as the covariant components $u_{\mu}$ of $\mathbf{u}$ as well as of $\tilde{\mathbf{u}}$, In that sense $\mathbf{u}$ and $\tilde{\mathbf{u}}$ can both have covariant and conravariant components. When $\mathbf{u}=u^{\mu} \boldsymbol{\partial}_{\mu}$, we think of it as a vector; when we write it as $u_{\nu} \mathbf{d} x^{\nu}$, we think of it as a covector. Thus, we can say that the inner product sets up an isomorphism between a vector space and its dual. Because of this relation, we also have: $\boldsymbol{\partial}_{\mu}=g_{\mu \nu} \mathbf{d} x^{\nu}$. Now, thanks to $\mathfrak{g}$, the same
tensor can have all-contravariant, all-covariant, or mixed components, with any one input slot accepting a vector or a covector, as appropriate for the corresponding index! In particular, $\mathfrak{g}$ also has mixed components:

$$
\begin{equation*}
g_{\nu}^{\mu}=\mathfrak{g}\left(\mathbf{d} x^{\mu}, \boldsymbol{\partial}_{\nu}\right) \equiv<\mathbf{d} x^{\mu}, \boldsymbol{\partial}_{\nu}>=\mathbf{d} x^{\mu}\left(\boldsymbol{\partial}_{\nu}\right)=\delta_{\nu}^{\mu} \tag{1.36}
\end{equation*}
$$

Since, as we have seen, $\delta^{\mu}{ }_{\nu}$ is basis-independent, so is $g^{\mu}{ }_{\nu}$, unlike $g_{\mu \nu}$ and $g^{\mu \nu}$. But $g^{\mu}{ }_{\nu}$ can also be obtained by index lowering (or raising): $g^{\mu \lambda} g_{\lambda \nu}=g^{\mu}{ }_{\nu}=\delta^{\mu}{ }_{\nu}$, which justifies our earlier assertion that $g^{\mu \nu}$ are the components of the inverse of $\mathfrak{g}$. We also see that on a $n$-dim space, $g^{\mu}{ }_{\mu}=\delta^{\mu}{ }_{\mu}=n$.

We also understand why, if $\delta^{\mu}{ }_{\nu}$ are the components of the identity matrix, $\mathbf{I}, \delta_{\mu \nu}=g_{\mu \rho} \delta^{\rho}{ }_{\nu}=g_{\mu \nu}$ will not in general be the entries of $\mathbf{I}$.

A final word of caution: before we introduced the inner product, we always wrote our matrices as $L_{\nu}^{\mu}$, with the left index a row index. Why do we not write the matrix of $\mathfrak{g}$ 's components the same way? Because $L^{\mu}{ }_{\nu}$ is a transformation between two bases in $\mathcal{V}^{n}$, whereas $g_{\mu \nu}$ transforms from a basis in $\mathcal{V}^{n}$ to the dual basis. For instance, in $\mathbb{R}^{3}$, the basis dual to $\left\{\mathbf{e}_{1}, \mathbf{e}_{2}, \mathbf{e}_{3}\right\}$ cannot be reached by any combination of rotations and translations. Also, $L^{\mu}{ }_{\nu}$ is not the component of a tensor, whereas $g_{\mu \nu}$ is.

### 1.4.3 The metric

The inner product that we have introduced plays another extremely important rôle: it allows us to define distances and lengths in $\mathcal{V}^{n}$ :

Definition 1.23. As a metric tensor (metric for short), $\mathfrak{g}$ tells us how to calculate lengths and distances in a vector space. The name is often extended (abusively) to its components $g_{\mu \nu}$. Thus,

$$
\begin{equation*}
\Delta s^{2}=g_{\mu \nu} \Delta x^{\mu} \Delta x^{\nu} \tag{1.37}
\end{equation*}
$$

gives the interval between two points labelled by $x^{\mu}$ and $x^{\mu}+\Delta x^{\mu}$.
In old-style notation, one often writes the metric in terms of an infinitesimal interval, or line element:

$$
\mathrm{d} s^{2}=g_{\mu \nu} \mathrm{d} x^{\mu} \mathrm{d} x^{\nu}
$$

with the $\mathrm{d} x^{\mu}$ the components of an infinitesimal displacement. In modern notation, however, one identifies the bilinear form $\mathbf{d} s^{2}=\mathfrak{g}=g_{\mu \nu} \mathrm{d} x^{\mu} \otimes \mathrm{d} x^{\nu}$ which then represents the interval $\Delta s^{2}$ for a $\Delta \mathbf{x}$ to be specified: $\Delta s^{2}=\langle\Delta \mathbf{x}, \Delta \mathbf{x}\rangle$, identical to the standard eq. (1.37).
Example 1.7. Consider the position vectors $\mathrm{x}_{1}$ and $\mathrm{x}_{2}$ in $\mathbb{R}^{3}$, with components:

$$
\mathbf{x}_{1} \longmapsto\left(x_{1}, y_{1}, z_{1}\right)^{\mathrm{T}}, \quad \mathbf{x}_{2} \longmapsto\left(x_{2}, y_{2}, z_{2}\right)^{\mathrm{T}}
$$

If we choose a positive-definite $\mathfrak{g}$ with matrix representation $\mathbf{g}=\mathbf{I}$ :

$$
\mathfrak{g}(\Delta \mathbf{x}, \Delta \mathbf{x})=g_{\mu \nu} \Delta \mathbf{x}^{\mu} \Delta \mathbf{x}^{\nu}=\left(\begin{array}{lll}
x_{1}-x_{2} & y_{1}-y_{2} & z_{1}-z_{2}
\end{array}\right)\left(\begin{array}{ccc}
1 & 0 & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{l}
x_{1}-x_{2} \\
y_{1}-y_{2} \\
z_{1}-z_{2}
\end{array}\right)
$$

The result, $(\Delta s)^{2}=\left(x_{1}-x_{2}\right)^{2}+\left(y_{1}-y_{2}\right)^{2}+\left(z_{1}-z_{2}\right)^{2}=(\Delta x)^{2}+\left(\Delta y^{2}\right)+(\Delta z)^{2}$, is recognised to be the "Pythagorean" distance squared between two points: $\left|\mathbf{x}_{1}-\mathbf{x}_{2}\right|^{2}$.

Example 1.8. In $\mathbb{R}^{4}$, let $\mathbf{x}_{i}(i=1,2)$ be two position vectors with $\left(c t_{i}, x_{i}, y_{i}, z_{i}\right)$ as contravariant and $\left(-c t_{i}, x_{i}, y_{i}, z_{i}\right)$ as covariant components. Then take the indefinite $\boldsymbol{\eta} \equiv \mathfrak{g}$ with matrix representation:

$$
\eta_{\mu \nu}=\left(\begin{array}{rrrr}
-1 & 0 & 0 & 0 \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1
\end{array}\right)=\operatorname{diag}(-1,1,1,1)
$$

Therefore:

$$
\mathfrak{g}(\Delta \mathbf{x}, \Delta \mathbf{x})=-c^{2}\left(t_{1}-t_{2}\right)^{2}+\left(x_{1}-x_{2}\right)^{2}+\left(y_{1}-y_{2}\right)^{2}+\left(z_{1}-z_{2}\right)^{2}
$$

is the spacetime distance between two events in Einstein's Special Relativity, with $c$ the speed of light. Because $\boldsymbol{\eta}$ is indefinite, there exist non-trivial null vectors (with zero norm) such that $\mathfrak{g}(\mathbf{x}, \mathbf{x})=0$. And, just as $\partial_{\nu} x^{\mu}=\delta^{\mu}{ }_{\nu}$, we must write $\partial_{\nu} x_{\mu}=\eta_{\mu \nu}$.

The metric, or line element, is said to define the geometry of the space. Two spaces of the same dimension can have different geometries, eg. $\mathbb{R}^{4}$ can be endowed with a positive-definite $\left(\Delta s^{2}>0\right)$ metric, but it would not be the metric of 4-dim spacetime.

Quite often, we will wish to work in bases other than coordinate bases. The formal properties of $\mathfrak{g}$ that we have reviewed still hold, but its covariant and contravariant components can be different, even in the same coordinates:

Definition 1.24. A basis $\left\{\mathbf{e}_{\mu}\right\}$ such that $\mathfrak{g}\left(\mathbf{e}_{\mu}, \mathbf{e}_{\nu}\right)= \pm 1$ when $\mu=\nu$ and 0 otherwise is said to be orthonormal. A useful notation to distinguish it from a coordinate basis is $\left\{\mathbf{e}_{\hat{\mu}}\right\}$. This is an extension of the usual definition of orthonormality which admits only +1 and 0 ; it is useful in the case of indefinite metrics.

An orthonormal basis can be found locally (over a sufficiently small coordinate patch) for any metric, if necessary by using the Gram-Schmidt procedure, with some modifications if the metric is indefinite.

Definition 1.25. Let $n_{+}\left(n_{-}\right)$denote the number of diagonal elements $\mathfrak{g}\left(\mathbf{e}_{\hat{\mu}}, \mathbf{e}_{\hat{\mu}}\right)$ which are equal to $+1(-1)$. The signature of the metric is defined by $s=n_{+}-n_{-}$. Since $n_{+}+n_{-}=n$, the dimension of the space, we also have $s=n-2 n_{-}$, and $\operatorname{det} \mathbf{g}=(-1)^{n_{-}} . n_{+}$and $n_{-}$are basis-independent, and therefore so is the signature.

The sign of the overall signature of an indefinite metric is arbitrary and must be set by convention, often a source of much confusion. Example 1.8 sets $s=+2$, a good choice of sign when the spatial indices 1,2 , and 3 are often raised or lowered. In the more general spacetimes of General Relativity, $s=-2$ is quite often (but not always... see Misner, Thorne and Wheeler's Gravitation) used. Thus, beware!

Definition 1.26. A $n$-dim space endowed with a metric of signature $\pm n$ is called Euclidean. If $n_{-}=1$ (or $n_{-}=n-1$ ), the space is pseudo-Euclidean, or Lorentzian (aka Minkowski when $n=4$ ).

Example 1.8 has a Minkowski metric in four-dimensional space.
EXERCISE: In Minkowski $\mathbb{R}^{4}$, take $(c t, x, y, z)$ to be Cartesian coordinates, and $(c t, r, \theta, \phi)$ to be the usual spherical coordinates (with $\phi$ the azimuthal angle). Consider the coordinate cobasis $\{\mathbf{d} t, \mathbf{d} r, \mathbf{d} \theta, \mathbf{d} \phi\}$ and the non-coordinate cobasis is $\{\mathbf{d} \hat{t}, \mathbf{d} \hat{r}, \mathbf{d} \hat{\theta}, \mathbf{d} \hat{\phi}\}=\{\mathbf{d}(c t), \mathbf{d} r, r \mathbf{d} \theta, r \sin \theta \mathbf{d} \phi\}$. First, write down the general form of the bases $\left\{\mathbf{e}_{0}, \mathbf{e}_{1}, \mathbf{e}_{2}, \mathbf{e}_{3}\right\}$ and $\left\{\hat{\mathbf{e}}_{0}, \hat{\mathbf{e}}_{1}, \hat{\mathbf{e}}_{2}, \hat{\mathbf{e}}_{3}\right\}$ of which the two cobases are duals (this should be trivial). Then obtain the coordinate-basis vectors by proceeding as in example 1.4, and use them to calculate the $g_{\mu \nu}$ directly.. Finally, show that the non-coordinate basis given is orthonormal. An easy way to do this is to expand $\mathfrak{g}$ in the coordinate cobasis and use the above relation between the two cobases.

Thanks to the metric, we recover the vector gradient of a function as defined in calculus. You may have noticed that throughout our discussion of manifolds and tangent spaces, no mention was made of an inner product on them, because none was needed-until now. An inner product, or metric $\mathfrak{g}$, pairs the 1 -form $\mathbf{d} f$ with a vector, $\nabla f$; indeed, from eq. (1.35):

$$
\begin{equation*}
<\boldsymbol{\nabla} f, \mathbf{v}>=\mathfrak{g}(\boldsymbol{\nabla} f, \mathbf{v})=g_{\mu \nu}(\boldsymbol{\nabla} f)^{\mu} v^{\nu}=\left(\partial_{\nu} f\right) v^{\nu}=[\mathbf{d} f](\mathbf{v}) \tag{1.38}
\end{equation*}
$$

where $\mathbf{v}$ is an arbitrary vector, and the components of $\boldsymbol{\nabla} f$ in a coordinate basis are given by: $(\boldsymbol{\nabla} f)^{\mu}=g^{\mu \nu} \partial_{\nu} f$. Only in a Euclidean metric with a standard basis are the components of the vector gradient the same as those of the 1 -form gardient.

Example 1.9. In a Minkowskian manifold with Cartesian coordinates $\left(x^{0}=c t, x^{1}, x^{2}, x^{3}\right)$, equiped with the metric $\mathfrak{g}=\boldsymbol{\eta}=\operatorname{diag}(-1,1,1,1)$ :

$$
\begin{aligned}
\mathbf{d} f & =\partial_{t} f \mathbf{d} t+\partial_{i} f \mathbf{d} x^{i} \\
\boldsymbol{\nabla} f & =-\partial_{c t} f \boldsymbol{\partial}_{c t}+\partial^{i} f \boldsymbol{\partial}_{i} \quad\left(\partial^{i}=g^{i j} \partial_{j}=\partial_{i}\right)
\end{aligned}
$$

There is something interesting about the determinant of the metric tensor which we find by writing the transformation law of the tensor: $g_{\mu \nu}^{\prime}=\partial_{\mu^{\prime}} x^{\alpha} g_{\alpha \beta} \partial_{\nu^{\prime}} x^{\beta}$, as a matrix equation, and taking the determinant. Defining $g=\operatorname{det} g_{\alpha \beta}$, we obtain:

$$
\begin{equation*}
g^{\prime}=\left|\frac{\partial x}{\partial x^{\prime}}\right|^{2} g \tag{1.39}
\end{equation*}
$$

where $\left|\partial x / \partial x^{\prime}\right|$ is the Jacobian of the transformation matrix from $x$ to $x^{\prime}$ coordinates. We discover that $g$ does not not transform as a scalar, ie. it is not invariant.

Definition 1.27. A quantity that has extra powers of $\left|\partial x / \partial x^{\prime}\right|$ as factors in its transformation law in addition to the usual $\partial_{\mu^{\prime}} x^{\alpha}$ and/or $\partial_{\alpha} x^{\mu^{\prime}}$ factors is called a tensor density. Thus, $g$ is a scalar density.

What might seem no more than an exotic property becomes more relevant when we consider the $n$-dim volume element as usually written in an integral. This, as we know from calculus, transforms as : $\mathrm{d}^{n} x^{\prime}=\left|\partial x^{\prime} / \partial x\right| \mathrm{d}^{n} x$ (note the position of the prime in the Jacobian!), so is not an invariant. As a result, the volume integral of a scalar function is not an invariant, yet there should be no memory of the variables of integration left after integrating. But if nstead we transform $\sqrt{|g|} \mathrm{d}^{n} x$, we obtain:

$$
\sqrt{\left|g^{\prime}\right|} \mathrm{d}^{n} x^{\prime}=\left|\frac{\partial x}{\partial x^{\prime}}\right| \sqrt{|g|}\left|\frac{\partial x^{\prime}}{\partial x}\right| \mathrm{d}^{n} x=\sqrt{|g|} \mathrm{d}^{n} x
$$

which is seen to be a scalar! Then integrals written as $\int \sqrt{|g|} f(\mathbf{x}) \mathrm{d}^{n} x$ are invariant, but what we thought we were integrating is in fact the scalar density $\sqrt{|g|} f(\mathbf{x})$. This concept of tensor density as a notational device has been widely used in General Relativity, although post-1970 literature largely dispenses with it when $p$-forms are involved. Indeed, later in section 1.5.2, we shall introduce a deeper definition of the volume element.

### 1.5 Exterior Algebra

### 1.5.1 The exterior product

Definition 1.28. The exterior (wedge) product of two 1-forms is the antisymmetrised tensor product:

$$
\boldsymbol{\sigma} \wedge \boldsymbol{\tau} \equiv \boldsymbol{\sigma} \otimes \boldsymbol{\tau}-\boldsymbol{\tau} \otimes \boldsymbol{\sigma}
$$

Being skew-symmetric, the exterior product of two 1-forms can be called a 2-form.
In general, $p 1$-forms can be used to construct a simple (or decomposable) $p$-form:

$$
\begin{equation*}
\boldsymbol{\sigma}^{1} \wedge \ldots \wedge \boldsymbol{\sigma}^{p}=\delta_{\mu_{1} \ldots \mu_{p}}^{1 \ldots p} \boldsymbol{\sigma}^{\mu_{1}} \otimes \cdots \otimes \boldsymbol{\sigma}^{\mu_{p}} \tag{1.40}
\end{equation*}
$$

Needless to say, this expression can be directly applied to a basis of $\mathcal{V}^{*},\left\{\mathbf{d} x^{\mu}\right\}(\mu=1, \ldots, p)$ :

$$
\begin{equation*}
\mathbf{d} x^{\rho_{1}} \wedge \cdots \wedge \mathbf{d} x^{\rho_{p}}\left(\boldsymbol{\partial}_{\nu_{1}}, \ldots, \boldsymbol{\partial}_{\nu_{p}}\right)=\delta_{\mu_{1} \ldots \mu_{p}}^{\rho_{1} \ldots \rho_{p}} \delta_{\nu_{1}}^{\mu_{1}} \cdots \delta^{\nu_{\nu_{\nu_{p}}}}=\delta_{\nu_{1} \ldots \nu_{p}}^{\rho_{1} \ldots \rho_{p}} \quad\left(1 \leq \rho_{1}<\ldots<\rho_{p} \leq n\right) \tag{1.41}
\end{equation*}
$$

Thus, from $\left\{\mathbf{d} x^{\rho}\right\}(1 \leq \rho \leq n)$ a basis of $\boldsymbol{\Lambda}^{p}\left(\mathcal{V}^{*}\right)$ can be constructed which contains $n!/(p!(n-p)!)$ elements. In particular, a $n$-form on a $n$-dimensional space is a one-component object, in the sense that it must be a multiple of the unique basis element, $\mathbf{d} x^{1} \wedge \mathbf{d} x^{2} \wedge \ldots \wedge \mathbf{d} x^{n}$, with indices in increasing order. If $p>n$, then at least two indices in the wedge product of the $p$ basis 1 -forms (or of the components of the $p$-form) must be identical. Therefore, skew-symmetry forces the maximum rank of a non-trivial $p$-form in $n$ dimensions to be $n$.

The exterior product of a basis of $\bigwedge^{p}$ and a basis of $\bigwedge^{q}$ is a basis, $\mathbf{d} x^{\rho_{1}} \wedge \ldots \wedge \mathbf{d} x^{\rho_{p}} \wedge \mathbf{d} x^{\rho_{p+1}} \wedge \ldots \wedge \mathbf{d} x^{\rho_{p+q}}$, of $\bigwedge^{p+q}$, again with indices in increasing order, and $p+q \leq n$.

By extension, a $(p+q)$-form can be constructed out of the antisymmetrised tensor product of $\boldsymbol{\sigma} \in \Lambda^{p}$ and $\boldsymbol{\tau} \in \Lambda^{q}$. In geometric and index notation:

$$
\begin{gather*}
\boldsymbol{\sigma} \wedge \boldsymbol{\tau}\left(\mathbf{u}_{\rho_{1}}, \ldots, \mathbf{u}_{\rho_{p+q}}\right)=\delta_{\rho_{1} \ldots \rho_{p+q}}^{\mu_{1} \ldots \mu_{p} \nu_{1} \ldots \nu_{q}} \boldsymbol{\sigma}\left(\mathbf{u}_{\mu_{1}} \ldots \mathbf{u}_{\mu_{p}}\right) \boldsymbol{\tau}\left(\mathbf{u}_{\nu_{1}} \ldots \mathbf{u}_{\nu_{q}}\right) \quad \mu_{1}<\mu_{2} \ldots<\mu_{p}, \quad \nu_{1}<\ldots<\nu_{q} \\
(\boldsymbol{\sigma} \wedge \boldsymbol{\tau})_{\rho_{1} \ldots \rho_{p+q}}=\delta_{\rho_{1} \ldots \rho_{p+q}}^{\mu_{1} \ldots \mu_{p} \nu_{1} \ldots \nu_{q}} \sigma_{\mu_{1} \ldots \mu_{p}} \tau_{\nu_{1} \ldots \nu_{q}} \quad \mu_{1}<\mu_{2} \ldots<\mu_{p}, \quad \nu_{1}<\nu_{2} \ldots<\nu_{q} \tag{1.42}
\end{gather*}
$$

The wedge-product defines a so-called exterior (Grassmann) algebra.
The exterior product, in contrast to the vector ("cross") product of vector analysis which it generalises, is associative: $\sigma \wedge(\tau \wedge \theta)=(\sigma \wedge \tau) \wedge \theta$.

Another very important property of the exterior product of a $p$-form and a $q$ form is that:

$$
\begin{equation*}
\boldsymbol{\sigma} \wedge \boldsymbol{\tau}=(-1)^{p q} \boldsymbol{\tau} \wedge \boldsymbol{\sigma} \tag{1.43}
\end{equation*}
$$

This follows directly from eq. (1.42) by noting that it takes $p q$ transpositions to get $\delta_{\rho_{1} \ldots \rho_{p+q}}^{\nu_{1} \ldots \nu_{q} \mu_{1} \ldots \mu_{p}}$ into $\delta_{\rho_{1} \ldots \rho_{p+q}}^{\mu_{1} \ldots \mu_{p} \nu_{1} \ldots \nu_{q}}$. This means that the exterior product commutes except when both forms have odd rank.

Eq. (1.42) is easier to use than it might seem. Here are three examples:
Example 1.10. Some people believe that we live in an 11 -dimensional world. Let us work out one component of the 3 -form that is the exterior product of a 2 -form, $\sigma$, and a 1 -form, $\tau$ :

$$
\begin{aligned}
(\boldsymbol{\sigma} \wedge \boldsymbol{\tau})_{1136} & =\delta_{1136}^{\mu \nu \lambda} \sigma_{\mu \nu} \tau_{\lambda} \quad \mu<\nu \\
& =\delta_{11366}^{3} 11 \sigma_{36} \tau_{11}+\delta_{1136}^{3116} \sigma_{311} \tau_{6}+\delta_{1136}^{6113} \sigma_{611} \tau_{3} \\
& =\sigma_{36} \tau_{11}-\sigma_{311} \tau_{6}+\sigma_{611} \tau_{3}
\end{aligned}
$$

Example 1.11. In two dimensions, the exterior product of two 1 -forms, $\boldsymbol{\sigma}^{1}$ and $\boldsymbol{\sigma}^{2}$, is:

$$
\begin{aligned}
\boldsymbol{\sigma}^{1} \wedge \boldsymbol{\sigma}^{2} & =\left(\sigma^{1}{ }_{1} \mathbf{d} x^{1}+\sigma^{1}{ }_{2} \mathbf{d} x^{2}\right) \wedge\left(\sigma^{2}{ }_{1} \mathbf{d} x^{1}+\sigma^{2}{ }_{2} \mathbf{d} x^{2}\right) \\
& =\sigma^{1}{ }_{1} \sigma^{2}{ }_{2} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}+\sigma^{1}{ }_{2} \sigma^{2}{ }_{1} \mathbf{d} x^{2} \wedge \mathbf{d} x^{1}=\left(\sigma^{1}{ }_{1} \sigma^{2}{ }_{2}-\sigma^{1}{ }_{2} \sigma^{2}{ }_{1}\right) \mathbf{d} x^{1} \wedge \mathbf{d} x^{2} \\
& =\operatorname{det} \mathbf{S} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}
\end{aligned}
$$

where $\mathbf{S}$ is the $2 \times 2$ matrix whose two rows are the components of $\boldsymbol{\sigma}^{1}$ and $\boldsymbol{\sigma}^{2}$, respectively.
Example 1.12. In three dimensions, $\left\{\mathbf{d} x^{1} \wedge \mathbf{d} x^{2}, \mathbf{d} x^{1} \wedge \mathbf{d} x^{3}, \mathbf{d} x^{2} \wedge \mathbf{d} x^{3}\right\}$ forms a basis of the space of 2 -forms, $\bigwedge^{2}(\mathcal{V})$. Therefore, the most general (not necessarily simple!) 2-form in a 3-dim dual space can be written as:

$$
\begin{equation*}
\boldsymbol{\tau}=\tau_{12} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}+\tau_{23} \mathbf{d} x^{2} \wedge \mathbf{d} x^{3}+\tau_{31} \mathbf{d} x^{3} \wedge \mathbf{d} x^{1}=\frac{1}{2} \tau_{\mu \nu} \mathbf{d} x^{\mu} \wedge \mathbf{d} x^{\nu} \in \bigwedge^{2} \tag{1.44}
\end{equation*}
$$

The summation on the right of the second equality is now unrestricted.
Three-dimensional simple 2 -forms $\boldsymbol{\sigma}^{1} \wedge \boldsymbol{\sigma}^{2}$, however, have the coordinate form (EXERCISE):

$$
\begin{equation*}
\left(\sigma_{1}^{1} \sigma^{2}{ }_{2}-\sigma_{2}^{1} \sigma^{2}{ }_{1}\right) \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}+\left(\sigma_{3}^{1} \sigma^{2}{ }_{1}-\sigma_{1}^{1} \sigma^{2}{ }_{3}\right) \mathbf{d} x^{3} \wedge \mathbf{d} x^{1}+\left(\sigma_{2}^{1} \sigma^{2}{ }_{3}-\sigma_{3}^{1} \sigma^{2}{ }_{2}\right) \mathbf{d} x^{2} \wedge \mathbf{d} x^{3} \tag{1.45}
\end{equation*}
$$

In Euclidean $\mathbb{R}^{3}$ with Cartesian coordinates, the components would be those of the vector product of the two vectors associated with $\sigma^{1}$ and $\sigma^{2}$. It suggests that we could associate with a vector $\mathbf{v}$ not just a 1-form with components $g_{\mu \nu} v^{\nu}$, but a 2-form $\tau$ with Cartesian components $\tau_{i j} \equiv \epsilon_{i j k} v^{k}$. $\bigwedge^{n-1}$ and $\bigwedge^{1}$ both have dimension $n$, which in turn suggests that, eg., a simple 2-form $\boldsymbol{\sigma}^{1} \wedge \boldsymbol{\sigma}^{2}$ might in some sense correspond to a 1 -form. But the fact that the vector product is a pseudo-vector means that we are not quite ready for this; we still have to introduce another important idea, that of orientation, in section 1.5.4.

In four dimensions, the basis for $\Lambda^{2}$ contains 6 elements. EXERCISE: What are the components of the exterior product of two 1-forms in three and four dimensions? (Given the basis elements $\mathbf{d} x^{\mu} \wedge \mathbf{d} x^{\nu}$, you should recognise that the components must be of the form (no pun intended) $\sigma_{\mu}^{1} \sigma_{\nu}^{2}-\sigma_{\mu}^{2} \sigma_{\nu}^{1}$.)

More generally, consider simple $p$-forms on a $n$-dimensional space. In terms of a basis $\left\{\mathbf{d} x^{\nu}\right\}$, we have for 1-forms $\boldsymbol{\sigma}^{\mu}: \boldsymbol{\sigma}^{\mu}=\sigma_{\nu}^{\mu} \mathbf{d} x^{\nu}$ (the superscripts on $\boldsymbol{\sigma}$ and $\mathbf{d} x$ being labels for the 1-forms). Thus:

$$
\begin{align*}
\boldsymbol{\sigma}^{1} \wedge \ldots \wedge \boldsymbol{\sigma}^{p} & =\sigma^{1}{ }_{\nu_{1}} \mathbf{d} x^{\nu_{1}} \wedge \ldots \wedge \sigma_{\nu_{p}}^{p} \mathbf{d} x^{\nu_{p}}=\sigma_{\nu_{1}}^{1} \ldots \sigma_{\nu_{p}}^{p} \mathbf{d} x^{\nu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\nu_{p}} \\
& \left.=\left[\sigma^{1}{ }_{\mu_{1}} \cdots \sigma^{p}{ }_{\mu_{p}} \delta_{\nu_{1} \ldots \nu_{p}}^{\mu_{1} \ldots \mu_{p}}\right] \mathbf{d} x^{\nu_{1}} \otimes \ldots \otimes \mathbf{d} x^{\nu_{p}} \quad \text { (unrestricted sum over } \nu_{i}\right) \\
& =\left[\sigma^{1}{ }_{\mu_{1}} \cdots \sigma^{p}{ }_{\mu_{p}} \delta_{\nu_{1} \ldots \nu_{p}}^{\mu_{1} \ldots \mu_{p}}\right] \mathbf{d} x^{\nu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\nu_{p}} \quad \nu_{1}<\ldots<\nu_{p} \tag{1.46}
\end{align*}
$$

where the summation over each $\nu_{i}(1 \leq i \leq p)$ runs from 1 to $n$,
If we construct a $p \times n$ matrix $\mathbf{S}$ whose $i^{\text {th }}$ row is the $n$ components of the 1 -form $\boldsymbol{\sigma}^{i}$, we may notice that the expression inside the square backets in eq. (1.46) is nothing but the determinant of the $p \times p$ submatrix extracted from column indices $\nu_{1} \ldots \nu_{p}$ of $\mathbf{S}$, with $\nu_{1}<\ldots<\nu_{p}$. Therefore, in eq. (1.46), each term in the sum over the $\nu$ indices has as coefficient a $p \times p$ determinant. Each row of a determinant contains $p$ out of the $n$ components of the 1 -forms $\sigma$, and these components, labelled by $\nu_{1}<\ldots<\nu_{p}$, must be the same as the ones on the wedge product of basis covectors $\mathbf{d} x^{\nu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\nu_{p}}$ in that term.

The output (a number!) resulting from inputting $\mathbf{u}_{1}, \ldots, \mathbf{u}_{p}$ into the simple $p$-form $\boldsymbol{\sigma}^{1} \wedge \ldots \wedge \boldsymbol{\sigma}^{p}$ is:

$$
\begin{equation*}
\boldsymbol{\sigma}^{1} \wedge \ldots \wedge \boldsymbol{\sigma}^{p}\left(\mathbf{u}_{1}, \ldots, \mathbf{u}_{p}\right)=\operatorname{det}\left[\boldsymbol{\sigma}^{i}\left(\mathbf{u}_{j}\right)\right] \tag{1.47}
\end{equation*}
$$

ie. the determinant of the $p \times p$ matrix $\mathbf{S}$ whose entries are $S^{i}{ }_{j}=\boldsymbol{\sigma}^{i}\left(\mathbf{u}_{j}\right)$. As we have seen, these are simply $\sigma_{\mu}^{i} u_{j}^{\mu}$, with $\mu$ running from 1 to $n$.

Example 1.13. For a 3 -dim $\mathcal{V}^{*}$ of which the 2 -forms $\mathbf{d} x^{i} \wedge \mathbf{d} x^{j}$ are basis elements, we have:

$$
\mathbf{d} x^{i} \wedge \mathbf{d} x^{j}(\mathbf{u}, \mathbf{v})=\mathbf{d} x^{i}(\mathbf{u}) \mathbf{d} x^{j}(\mathbf{v})-\mathbf{d} x^{j}(\mathbf{u}) \mathbf{d} x^{i}(\mathbf{v})=\left|\begin{array}{cc}
u^{i} & u^{j} \\
v^{i} & v^{j}
\end{array}\right|
$$

In $\mathbb{R}^{n}$ with Cartesian coordinates, we interpret this (up to a sign—see 1.5 .2 below!) as the area of the parallelogram whose defining sides are the projections of $\mathbf{u}$ and $\mathbf{v}$ on the $x^{i}-x^{j}$ plane.

We have noted that when $p=n, \boldsymbol{\sigma}^{1} \wedge \ldots \wedge \boldsymbol{\sigma}^{n}$ is a one-component object. We can also see that the summation in the second line in eq. (1.46) can have only one term with the $\nu_{i}$ indices increasing from 1 to $n$ :

$$
\begin{equation*}
\boldsymbol{\sigma}^{1} \wedge \ldots \wedge \boldsymbol{\sigma}^{n}=\left[\sigma_{\mu_{1}}^{1} \cdots \sigma_{\mu_{n}}^{n} \delta_{1 \ldots n}^{\mu_{1} \ldots \mu_{n}}\right] \mathbf{d} x^{1} \wedge \ldots \wedge \mathbf{d} x^{n}=\operatorname{det} \mathbf{S} \mathbf{d} x^{1} \wedge \ldots \wedge \mathbf{d} x^{n} \tag{1.48}
\end{equation*}
$$

where the matrix $\mathbf{S}$ is now $n \times n$. Note that the indices $1 \ldots n$ are not summed over.
Example 1.14. There is another useful definition of the permutation symbol, $\delta_{\mu_{1} \cdots \mu_{n}}^{\nu_{1} \ldots \nu_{n}}$, equivalent to the one given by eq. (1.31), and which follows directly from (1.41):

$$
\delta_{\mu_{1} \ldots \mu_{n}}^{\nu_{1} \ldots \nu_{n}}=\mathbf{d} x^{\nu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\nu_{n}}\left(\boldsymbol{\partial}_{\mu_{1}}, \ldots, \boldsymbol{\partial}_{\mu_{n}}\right)
$$

Then eq. (1.47) becomes:

$$
\delta_{\mu_{1} \ldots \mu_{n}}^{\nu_{1} \ldots \nu_{n}}=\left|\begin{array}{ccc}
\delta^{\nu_{1}}{ }_{\mu_{1}} & \cdots & \delta^{\nu_{1}}{ }_{\mu_{n}}  \tag{1.49}\\
\vdots & & \vdots \\
\delta^{\nu_{n}}{ }_{\mu_{1}} & \cdots & \delta^{\nu_{n}}{ }_{\mu_{n}}
\end{array}\right|
$$

Another application provides an easy test for the linear independence of $p 1$-forms: if their exterior product is different from zero, those $p$ 1-forms are linearly independent. If they were not, one of them at least could be written as a linear combination of the others and antisymmetry would force the exterior product to vanish. Conversely, if the $p$ 1-forms are linearly independent, $\boldsymbol{\sigma}^{1} \wedge \cdots \wedge \boldsymbol{\sigma}^{p}$ cannot vanish.

Example 1.15. In general a $p$-form $\mathbf{F}$ is not simple. If it is, then the $2 p$-form $\mathbf{F} \wedge \mathbf{F}$ must vanish by antisymmetry. But to what extent does the converse hold?
Let $\mathbf{F} \wedge \mathbf{F}=0$, and take $\mathbf{F} \in \bigwedge^{2}$. When $n=2, \mathbf{F}=F_{12} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}$ is by inspection always simple, with no need to invoke $\mathbf{F} \wedge \mathbf{F}=0$.
For $n=3, \mathbf{F} \wedge \mathbf{F}$ still trivially vanishes $\forall p \leq 3$. That vanishing will lead to $\mathbf{F}$ being simple, but the argument is a little more involved. We can always write:

$$
\mathbf{F}=F_{12} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}+F_{13} \mathbf{d} x^{1} \wedge \mathbf{d} x^{3}+F_{23} \mathbf{d} x^{2} \wedge \mathbf{d} x^{3} \equiv \boldsymbol{\sigma}+\boldsymbol{\tau} \wedge \mathbf{d} x^{3}
$$

where $\boldsymbol{\sigma}$ is a 2 -form on the 2 -dim subspace, spanned by $\mathbf{d} x^{1} \wedge \mathbf{d} x^{2}$, of the 3 -dim space, and $\boldsymbol{\tau}$ is a 1 -form on the same 2 -dim subspace. Then $\boldsymbol{\sigma}$ is simple, ie., $\boldsymbol{\sigma}=\boldsymbol{\alpha} \wedge \boldsymbol{\beta}\left(\boldsymbol{\alpha}, \boldsymbol{\beta} \in \Lambda^{1}\right)$ because $n=2$.
Thus, $\mathbf{F} \wedge \mathbf{F}=\boldsymbol{\sigma} \wedge \boldsymbol{\sigma}+2 \boldsymbol{\sigma} \wedge \boldsymbol{\tau} \wedge \mathbf{d} x^{3}=0$. Also, $\boldsymbol{\sigma} \wedge \boldsymbol{\sigma}=0$ because $\boldsymbol{\sigma}$ is simple. Then $\boldsymbol{\alpha} \wedge \boldsymbol{\beta} \wedge \boldsymbol{\tau} \wedge \mathbf{d} x^{3}=0$. But the only possible linear dependence between the four 1-forms is between $\boldsymbol{\alpha}, \boldsymbol{\beta}$ and $\boldsymbol{\tau}$ since none of them depends on $\mathbf{d} x^{3}$. Therefore, $\lambda_{1} \boldsymbol{\alpha}+\lambda_{2} \boldsymbol{\beta}+\lambda_{3} \boldsymbol{\tau}=0$.
If $\lambda_{3}=0$, then $\boldsymbol{\beta}$ is a multiple of $\boldsymbol{\alpha}$, so $\boldsymbol{\sigma}=0$, leaving $\mathbf{F}=\boldsymbol{\tau} \wedge \mathbf{d} x^{3}$, a simple form. If $\lambda_{3} \neq 0$, $\boldsymbol{\tau}=a \boldsymbol{\alpha}+b \boldsymbol{\beta}$, and we obtain:

$$
\mathbf{F}=\boldsymbol{\alpha} \wedge \boldsymbol{\beta}+(a \boldsymbol{\alpha}+b \boldsymbol{\beta}) \wedge \mathbf{d} x^{3}=\left(\boldsymbol{\alpha}+\frac{b}{a} \boldsymbol{\beta}\right) \wedge\left(\boldsymbol{\beta}+a \mathbf{d} x^{3}\right)
$$

which is a simple form. Thus, 2 -forms (and in fact $p$-forms) on 3-dim space are always simple! EXERCISE: A 2 -form in $n=4$ is simple if, and only if, $\mathbf{F} \wedge \mathbf{F}=0$. Also, when $\mathbf{F} \wedge \mathbf{F} \neq 0$ in $n=4, \mathbf{F}$ can be written (EXERCISE) as the sum of two simple 2-forms. Note that these statements hold whether or not a metric has been introduced.

### 1.5.2 Oriented manifolds, pseudo-vectors, pseudo-forms and the volume form

Definition 1.29. Two bases are said to have the same (opposite) orientation if the determinant of the matrix of the transformation between the two bases is positive (negative). Therefore, bases fall into two classes, or orientations. Orienting a manifold then means arbitrarily specifying one of the two orientations to be positive (right-handed), and the other negative (left-handed). Manifolds on which transport of a basis around some closed loop leads to a reversal of orientation are said to be non-orientable (eg. the Möbius strip).
In $\mathbb{R}^{3}$, for instance, $\mathbf{e}_{x} \wedge \mathbf{e}_{y} \wedge \mathbf{e}_{z}, \mathbf{e}_{y} \wedge \mathbf{e}_{z} \wedge \mathbf{e}_{x}$ and $\mathbf{e}_{z} \wedge \mathbf{e}_{x} \wedge \mathbf{e}_{y}$ can be transformed into one another by matrices of determinant +1 . By convention, they are all taken to be right-handed. But $\mathbf{e}_{y} \wedge \mathbf{e}_{x} \wedge \mathbf{e}_{z}=-\mathbf{e}_{x} \wedge \mathbf{e}_{y} \wedge \mathbf{e}_{z}$ cannot be reached from $\mathbf{e}_{x} \wedge \mathbf{e}_{y} \wedge \mathbf{e}_{z}$ by a transformation with a matrix of positive determinant: it is an element of a left-handed basis.

Definition 1.30. An object that behaves in all respects as a vector or a $p$-form, except that its sign is reversed under a reversal of orientation of the manifold, is called a pseudovector or a pseudoform.

Example 1.16. Generalising example 1.13 above, the simple $n$-form $\mathbf{d} x^{1} \wedge \cdots \wedge \mathbf{d} x^{n}$, when acting on the vectors $\mathbf{v}_{1}, \ldots \mathbf{v}_{n}$ in that order, outputs a mumber of magnitude equal to the volume of the parallelopiped whose edges are the vectors $\mathbf{v}_{1}, \ldots \mathbf{v}_{n}$. Using eq. (1.48), this is readily computed as the determinant of all the vector components. But there is also a sign involved, with + corresponding to the orientation defined by the vectors being the same as the orientation of the basis in which the vectors are written. We then say that this volume is oriented, in the sense that it changes sign under interchange of any two basis vectors, and we recognise it as a pseudoform.

Definition 1.31. In general coordinates $u^{i}$ on a $n$-dim manifold, we define the volume pseudoform:

$$
\mathbf{d}^{n} u:=\left|\frac{\partial x}{\partial u}\right| \mathbf{d} u^{1} \wedge \cdots \wedge \mathbf{d} u^{n}=\sqrt{|g|} \mathbf{d} u^{1} \wedge \cdots \wedge \mathbf{d} u^{n}
$$

where the $x^{i}$ form an orthonormal basis, usually Cartesian, and we have used eq. (1.39) with $|g|=1$ for orthonormal bases. Both bases are connected by a transformation that preserves orientation.

### 1.5.3 The Levi-Civita pseudotensor

We have already remarked on the fact that the Levi-Civita symbol does not transform as a tensor. Consider, however, the volume pseudoform of definition 1.31. By inspection it is a $n$-form with the single independent component $\left(\mathbf{d}^{n} u\right)_{1 \ldots n}=\sqrt{|g|}$. Its other components are obtained by antisymmetrising with the Levi-Civita symbol, which we shall now denote by $\left[\mu_{1} \ldots \mu_{n}\right]$ to avoid any confusion later. That is:

$$
\left(\mathbf{d}^{n} u\right)_{\mu_{1} \ldots \mu_{n}}=\sqrt{|g|}\left[\mu_{1} \ldots \mu_{n}\right]
$$

The objects on the right are themselves the components of a covariant pseudotensor, $\boldsymbol{\epsilon}$, of rank $n$. Henceforth, whenever we write components $\epsilon_{\mu_{1} \ldots \mu_{n}}$, they are to be understood as $\sqrt{|g|}\left[\mu_{1} \ldots \mu_{n}\right]$, so that $\epsilon_{1 \ldots n}=\sqrt{|g|}$.

We obtain $\epsilon^{1 \ldots n}$ by raising the $n$ indices of $\epsilon_{1 \ldots n}$ with $\mathfrak{g}$. In general coordinates:

$$
\epsilon^{1 \ldots n}=g^{1 \mu_{1}} \cdots g^{n \mu_{n}} \epsilon_{\mu_{1} \ldots \mu_{n}}=g^{1 \mu_{1}} \cdots g^{n \mu_{n}} \sqrt{|g|} \delta_{\mu_{1} \ldots \mu_{n}}^{1 \ldots n}=\operatorname{det} g^{\alpha \beta} \sqrt{|g|}=\frac{1}{(-1)^{n_{-}}|g|} \sqrt{|g|}=\frac{(-1)^{n_{-}}}{\sqrt{|g|}}
$$

In orthonormal bases, this is simply: $\epsilon^{1 \ldots n}=(-1)^{n_{-}} \epsilon_{1 \ldots n}$.
Both $\epsilon^{\nu_{1} \ldots \nu_{n}}$ and $\epsilon_{\mu_{1} \ldots \mu_{n}}$ being antisymmetric, we can relate the permutation symbol to the Levi-Civita pseudotensor:

$$
\epsilon^{\nu_{1} \ldots \nu_{n}} \epsilon_{\mu_{1} \ldots \mu_{n}}=a \delta_{\mu_{1} \ldots \mu_{n}}^{\nu_{1} \ldots \nu_{n}}
$$

To determine $a$, we note that $\epsilon^{1 \ldots n} \epsilon_{1 \ldots n}=(-1)^{n_{-}}$. The relation is then:

$$
\epsilon^{\nu_{1} \ldots \nu_{n}} \epsilon_{\mu_{1} \ldots \mu_{n}}=(-1)^{n_{-}} \delta_{\mu_{1} \ldots \mu_{n}}^{\nu_{1} \ldots \nu_{n}}=(-1)^{n_{-}}\left|\begin{array}{ccc}
\delta^{\nu_{1}}{ }_{\mu_{1}} & \cdots & \delta^{\nu_{1}}{ }_{\mu_{n}}  \tag{1.50}\\
\vdots & & \vdots \\
\delta^{\nu_{n}}{ }_{\mu_{1}} & \cdots & \delta^{\nu_{n}}{ }_{\mu_{n}}
\end{array}\right|
$$

For instance, in a Euclidean 3-dim space, $n_{-}=0$, and the expanded product has six terms. When contracted over the last or first indices, we obtain (EXERCISE): $\epsilon^{i j k} \epsilon_{l n k}=\delta^{i}{ }_{l} \delta^{j}{ }_{n}-\delta^{j}{ }_{l} \delta^{i}{ }_{n}$. Other expressions for the product of Levi-Civita tensors in a 4-dim Minkowski space can be found in MTW, pp. 87-88.

### 1.5.4 The Hodge dual of a p-form

In section 1.5 .1 we had pointed out that it might be possible to associate a 2 -form with a vector. Now we can do this in all generality on $n$-dim metric-endowed spaces. To a vector $\mathbf{v}$ corresponds a pseudoform $\boldsymbol{\sigma}$ of rank $n-1$ :

$$
\begin{equation*}
\boldsymbol{\sigma}=v^{\nu} \epsilon_{\nu \mu_{1} \ldots \mu_{n-1}} \mathbf{d} u^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} u^{\mu_{n-1}} \quad\left(\mu_{1}<\ldots<\mu_{n-1}\right) \tag{1.51}
\end{equation*}
$$

which, like $\mathbf{v}$, has $n$ (independent!) components. In 3 - dim $\mathbb{R}^{3}$ we have the pseudo-2-form:

$$
\boldsymbol{\sigma}=\sqrt{|g|}\left(v^{3} \mathbf{d} u^{1} \wedge \mathbf{d} u^{2}-v^{2} \mathbf{d} u^{1} \wedge \mathbf{d} u^{3}+v^{1} \mathbf{d} u^{2} \wedge \mathbf{d} u^{3}\right)
$$

And there must also be a mapping between the 1 -form dual to $\mathbf{v}$ and the ( $n-1$ )-pseudoform. We now generalise to a linear mapping between $\bigwedge^{p}$ and $\bigwedge^{n-p}$.

Definition 1.32. Let $\mathcal{V}^{n}$ be endowed with a metric and a basis $\left\{\boldsymbol{\partial}_{\mu}\right\}$. With $\boldsymbol{\epsilon}$ the Levi-Civita pseudotensor, the Hodge dual ${ }^{\dagger}$ maps a $p$-form $\boldsymbol{\sigma}$ to a $(n-p)$-form $\downarrow \boldsymbol{\sigma}$ with components:

$$
\begin{equation*}
(\varkappa \boldsymbol{\varkappa})_{\nu_{1} \ldots \nu_{n-p}}=\frac{1}{p!} \sigma_{\mu_{1} \ldots \mu_{p}} \epsilon_{\nu_{1} \ldots \nu_{n-p}}^{\mu_{1} \ldots \mu_{p}} \quad\left(=\sigma^{\mu_{1} \ldots \mu_{p}} \epsilon_{\mu_{1} \ldots \mu_{p} \nu_{1} \ldots \nu_{n-p}} \quad \mu_{1}<\ldots<\mu_{p}\right) \tag{1.52}
\end{equation*}
$$

from which it is obvious that the Hodge dual of a $p$-form is a $p s e u d o$-form, and vice-versa. It can be shown that, given a metric $\mathfrak{g}$, $\mathfrak{\varkappa} \boldsymbol{\sigma}=(-1)^{n_{-}}(-1)^{p(n-p)} \boldsymbol{\sigma}$. So the mapping is idempotent in Euclidean spaces $\left(n_{-}=0\right)$ of odd dimension, such as $\mathbb{R}^{3}$. In a 4-dim Minkowski space ( $n_{-}=1$ ), Hodge duality is idempotent only on 1 - and 3 -forms.
One immediate application of eq. (1.52) is that the $n$-dim volume form is the Hodge dual of the 0 -form 1:

$$
\begin{aligned}
\star 1 & =\epsilon_{\mu_{1} \ldots \mu_{n}} \mathbf{d} u^{\mu_{1}} \wedge \cdots \wedge \mathbf{d} u^{\mu_{n}} \quad\left(\mu_{1}<\ldots<\mu_{n}\right) \\
& =\sqrt{|g|} \mathbf{d} u^{1} \wedge \cdots \wedge \mathbf{d} u^{n}
\end{aligned}
$$

A very important consequence of the fact that $ぇ \boldsymbol{\sigma}= \pm \boldsymbol{\sigma}$ is that a $p$-form and its Hodge dual contain exactly the same information! Thus, "dualising" a $p$-form (or an antisymmetric contravariant tensor) can remove some (or all!) the redundancy due to anrisymmetry while preserving its information. For instance, in 4-dim Minkowski space, one might be a bit intimidated by a 4-form with components $\sigma_{\mu \nu \lambda \rho}$ until one realises that it is dual to a pseudo-0-form, so one independent number instead of $4^{4}=256$. Or a 3 -form with a priori $4^{3}=64$ components can be Hodge-dualised, that is, written in terms of only the four components of its dual pseudo-1-form, which are (up to $\sqrt{|g|}$ ) the only independent components of the 3 -form.

Example 1.17. If $\mathbf{T}$ is a $(2,0)$ skew-symmetric tensor:

$$
\begin{aligned}
(\grave{\aleph} T)_{\lambda} & =\frac{1}{2} \epsilon_{\mu \nu \lambda} T^{\mu \nu} & \text { in } \mathbb{R}^{3} \\
(\grave{\aleph} T)_{\lambda \rho} & =\frac{1}{2} \epsilon_{\mu \nu \lambda \rho} T^{\mu \nu} & \text { in } \mathbb{R}^{4}
\end{aligned}
$$

In the first line, if the metric is Euclidean and Cartesian, it is not hard to work out that $\hbar T_{1}=T^{23}$, $\star T_{2}=T^{31}$, and $\star T_{3}=T^{12}$, so that the tensor (1-form) dual to $\mathbf{T}$ contains only the three independent components of the original tensor.

[^4]In example 1.12 we had noticed that in the 3 -dim Euclidean space of $p$-forms, both $\left\{\mathbf{d} x^{1}, \mathbf{d} x^{2}, \mathbf{d} x^{3}\right\}$ and $\left\{\mathbf{d} x^{2} \wedge \mathbf{d} x^{3}, \mathbf{d} x^{3} \wedge \mathbf{d} x^{1}, \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}\right\}$ are bases, the first being used to generate 1 -forms and the second 2-forms. Now we know that the two bases are each other's Hodge dual. In fact, instead of Hodge-dualising the components, we can Hodge-dualise the (co)basis: $\underset{\sim}{ }\left(\mathbf{d} u^{\mu_{1}} \wedge \cdots \wedge \mathbf{d} u^{\mu_{p}}\right)=\epsilon^{\mu_{1} \ldots \mu_{p}}{ }_{\mu_{p+1} \ldots \mu_{n}} \mathbf{d} u^{\mu_{p+1}} \wedge \cdots \wedge \mathbf{d} u^{\mu_{n}}$, with $\mu_{p+1}<$ $\cdots<\mu_{n}$ (or divide by $(n-p)$ ) if the summations are unrestricted). There are corresponding expressions for Hodgedualising coordinate bases or the components of contravariant tensors, ie., the Hodge dual of a contravariant tensor (or its components) also exists, as the above example shows.

Example 1.18. If $\boldsymbol{\sigma}$ and $\boldsymbol{\tau}$ are 3-dim 1-forms, the 2-form: $\boldsymbol{\sigma} \wedge \boldsymbol{\tau}=\left(\sigma_{2} \tau_{3}-\sigma_{3} \tau_{2}\right) \mathbf{d} x^{2} \wedge \mathbf{d} x^{3}+$ $\left(\sigma_{3} \tau_{1}-\sigma_{1} \tau_{3}\right) \mathbf{d} x^{3} \wedge \mathbf{d} x^{1}+\left(\sigma_{1} \tau_{2}-\sigma_{2} \tau_{1}\right) \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}$ has as its Hodge dual on a space with metric $\mathfrak{g}$ the pseudo-1-form:

$$
\mathfrak{\psi}(\boldsymbol{\sigma} \wedge \boldsymbol{\tau})=\sqrt{|g|}\left[\left(\sigma_{2} \tau_{3}-\sigma_{3} \tau_{2}\right) \mathbf{d} x^{1}+\left(\sigma_{3} \tau_{1}-\sigma_{1} \tau_{3}\right) \mathbf{d} x^{2}+\left(\sigma_{1} \tau_{2}-\sigma_{2} \tau_{1}\right) \mathbf{d} x^{3}\right]
$$

If $\boldsymbol{\sigma}$ corresponds to the vector $\mathbf{u}$ and $\boldsymbol{\tau}$ to $\mathbf{v}$ via the metric, this is the same as saying that $\mathfrak{z}(\mathbf{u} \wedge \mathbf{v})=$ $\mathbf{u} \times \mathbf{v}$, or, according to eq. (1.52), $(\mathbf{u} \times \mathbf{v})^{\mu}=\frac{1}{2} g^{\mu \rho} \epsilon_{\nu \lambda \rho}\left(u^{\nu} v^{\lambda}-v^{\lambda} u^{\nu}\right)=g^{\mu \rho} \epsilon_{\rho \nu \lambda} u^{\nu} v^{\lambda}$. In other words, when calculating a vector product, one is implicitly taking a Hodge dual; this is the only way that the result can be a pseudo-vector.
It is easy to recover all the relations of vector analysis in Cartesian $\mathbb{R}^{3}$. For instance:

$$
\begin{aligned}
\mathbf{u} \cdot(\mathbf{v} \times \mathbf{w}) & =\epsilon_{\mu \nu \rho} u^{\mu} v^{\nu} w^{\rho} \\
& =w^{\rho} \epsilon_{\rho \mu \nu} u^{\mu} v^{\nu} \quad(\text { cyclic permutation of indices on } \epsilon) \\
& =\mathbf{w} \cdot(\mathbf{u} \times \mathbf{v})
\end{aligned}
$$

Why not try your hand at a few if you have never done it, if only to practice index manipulation.

### 1.6 Exterior Calculus

Definition 1.33. A tensor field $\mathbf{T}(p)$ on a manifold $M^{n}$ is a function of points $p \in M^{n}$ whose components $T(p)$ are real-valued differentiable functions of coordinates on the manifold.
Examples: the coordinate vector field $\boldsymbol{\partial}_{\mu}$, the gravitational and electric fields, the metric tensor with components $<\boldsymbol{\partial}_{\mu}, \boldsymbol{\partial}_{\mu}>$ in a coordinate basis.

The components of a $(r, s)$ tensor field with respect to the coordinates $x^{\mu}$ in $n$ dimensions are the $n^{r+s}$ realvalued functions $\mathbf{T}\left(\mathbf{d} x^{\nu_{1}}, \ldots, \mathbf{d} x^{\nu_{r}}, \boldsymbol{\partial}_{\mu_{1}}, \ldots, \boldsymbol{\partial}_{\mu_{s}}\right)$.

How do we describe the change of a tensor field at a point? More precisely, how do we differentiate it? We already know from section 1.2 how to take the directional derivative of a $(0,0)$ tensor, ie. a function. On a "flat" (without curvature) manifold, directional derivatives of tensor components, which are functions, can be calculated in the same way.

For general $(r, s)$ tensors, however, because of point-dependent bases, defining differentiation requires extra structure, called a connection, or covariant derivative. Raising this important issue is like opening Pandora's box (aka can of worms), because there are a lot of ways to construct such a connection. A few, however, have gained favour as "natural". We shall discuss a particular type of differentiation that offers a neat unification of the ideas of gradient, divergence and curl in vector calculus, and for which a connection is actually not needed.

### 1.6.1 Exterior derivative

We introduce the exterior derivative operator ${ }^{\dagger}$, $\mathbf{d}$, which acts on $p$-forms $\boldsymbol{\sigma}=\sigma_{\left|\mu_{1} \ldots \mu_{p}\right|} \mathbf{d} x^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\mu_{p}}$ $\left(\left|\mu_{1} \ldots \mu_{p}\right| \equiv \mu_{1}<\ldots<\mu_{p}\right)$ defined over some manifold $M^{n}$ to give $p+1$-forms, also defined on $M^{n}$. Let $\sigma$ be a $p$-form and $\tau$ a $q$-form. The operator satisfies the following properties:

[^5](a)
$\mathbf{d} \boldsymbol{\sigma}=\partial_{\mu_{0}} \sigma_{\left|\mu_{1} \ldots \mu_{p}\right|} \mathbf{d} x^{\mu_{0}} \wedge \mathbf{d} x^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\mu_{p}}=\left(\mathbf{d} \sigma_{\left|\mu_{1} \ldots \mu_{p}\right|}\right) \wedge \mathbf{d} x^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\mu_{p}}$
In terms of components:
$$
(\mathbf{d} \boldsymbol{\sigma})_{\left|\mu_{1} \ldots \mu_{p+1}\right|}=\delta_{\mu_{1} \ldots \mu_{p+1}}^{\nu_{0} \nu_{1} \ldots \nu_{p}} \partial_{\nu_{0}} \sigma_{\left|\nu_{1} \ldots \nu_{p}\right|}=\partial_{\mu_{1}} \sigma_{\mu_{2} \ldots \mu_{p+1}}-\partial_{\mu_{2}} \sigma_{\mu_{1} \ldots \mu_{p+1}}+\ldots
$$
(b) $\mathbf{d}(\boldsymbol{\sigma}+\boldsymbol{\tau})=\mathbf{d} \boldsymbol{\sigma}+\mathbf{d} \boldsymbol{\tau}(p=q)$.
(c) If $\boldsymbol{\sigma}$ is a 0 -form, ie. just a function, then $\mathbf{d} \boldsymbol{\sigma}$ is the 1-form gradient of that function.
(d) $\mathbf{d}(\boldsymbol{\sigma} \wedge \boldsymbol{\tau})=\mathbf{d} \boldsymbol{\sigma} \wedge \boldsymbol{\tau}+(-1)^{p} \boldsymbol{\sigma} \wedge \mathbf{d} \boldsymbol{\tau}$ (aka the antiderivation property of $\mathbf{d}$ with respect to the exterior product).
(e) $\mathbf{d}^{2} \boldsymbol{\sigma}=0$ (Poincaré's lemma).

We shall not prove the antiderivation property (you can do it as an EXERCISE), but Poincaré's lemma is so famous and important that it deserves some proof.

First, for an arbitrary function $f$ (0-form):

$$
\mathbf{d}^{2} f=\mathbf{d}\left(\partial_{\nu} f \mathbf{d} x^{\nu}\right)=\partial_{\mu} \partial_{\nu} f \mathbf{d} x^{\mu} \wedge \mathbf{d} x^{\nu}=0
$$

since $\partial_{\mu} \partial_{\nu}$ is symmetric in $\mu$ and $\nu$. If $g$ is another function, $\mathbf{d}(\mathbf{d} f \wedge \mathbf{d} g)=\mathbf{d}^{2} f \wedge \mathbf{d} g-\mathbf{d} f \wedge \mathbf{d}^{2} g$ from the antiderivation property; this must vanish since $\mathbf{d}^{2}$ applied to any function gives zero. By extension, the exterior derivative of the exterior product of any number of arbitrary differential 1-forms also vanishes. Now, from properties (a) and (d) above:

$$
\begin{aligned}
\mathbf{d}^{2} \boldsymbol{\sigma} & =\mathbf{d}\left[\mathbf{d} \sigma_{\left|\mu_{1} \ldots \mu_{p}\right|} \wedge \mathbf{d} x^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\mu_{p}}\right] \\
& =\mathbf{d}^{2}\left[\sigma_{\left|\mu_{1} \ldots \mu_{p}\right|}\right] \wedge \mathbf{d} x^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\mu_{p}}+(-1)^{\mu_{p}+1} \mathbf{d} \sigma_{\left|\mu_{1} \ldots \mu_{p}\right|} \wedge \mathbf{d}\left[\mathbf{d} x^{\mu_{1}} \wedge \ldots \wedge \mathbf{d} x^{\mu_{p}}\right]
\end{aligned}
$$

The first term vanishes because we have already shown that $\mathbf{d}$ is nilpotent on functions; the second vanishes because we have also shown that the exterior derivative of a wedge product of differential 1-forms vanishes.

Example 1.19. In $\mathbb{R}^{3}$, with $u, v$, and $w$ as arbitrary coordinates, the differential of a function $f$ in the coordinate basis $\{\mathbf{d} u, \mathbf{d} v, \mathbf{d} w\}$ is the 1 -form:

$$
\begin{equation*}
\mathbf{d} f=\partial_{u} f \mathbf{d} u+\partial_{v} f \mathbf{d} v+\partial_{w} f \mathbf{d} w \tag{1.53}
\end{equation*}
$$

Only if the coordinates are Cartesian is this also valid for an orthonormal basis in the same coordinates. In a spherical coordinate basis $\{\mathbf{d} r, \mathbf{d} \theta, \mathbf{d} \phi\}$, for instance, $\mathbf{d} f$ would keep the above simple form. If instead we insist on using a basis whose components are normalised to unity, such as $\{\mathbf{d} \hat{r}, \mathbf{d} \hat{\theta}, \mathbf{d} \hat{\phi}\}=\{\mathbf{d} r, r \mathbf{d} \theta, r \sin \theta \mathbf{d} \phi\}-$ as is almost always the case in vector analysis applied to physics - consistency demands that we write:

$$
\begin{equation*}
\mathbf{d} f=\partial_{r} f \mathbf{d} \hat{r}+\frac{1}{r} \partial_{\theta} f \mathbf{d} \hat{\theta}+\frac{1}{r \sin \theta} \partial_{\phi} f \mathbf{d} \hat{\phi} \tag{1.54}
\end{equation*}
$$

Example 1.20. As an example of the first property of $\mathbf{d}$, the exterior derivative of a 1 -form $\sigma$ is the 2-form:

$$
\begin{align*}
\boldsymbol{\theta}=\mathbf{d} \boldsymbol{\sigma} & =\partial_{\mu} \sigma_{\nu} \mathbf{d} x^{\mu} \wedge \mathbf{d} x^{\nu} \\
& =\left(\partial_{\mu} \sigma_{\nu}-\partial_{\nu} \sigma_{\mu}\right) \mathbf{d} x^{\mu} \wedge \mathbf{d} x^{\nu} \quad(\mu<\nu) \tag{1.55}
\end{align*}
$$

where we have used the fact that the exterior derivative of the 0 -forms (functions) $\sigma_{\nu}$ are 1 -forms. The components of $\boldsymbol{\theta}$ are: $\theta_{\mu \nu}=\partial_{\mu} \sigma_{\nu}-\partial_{\nu} \sigma_{\mu}$. (EXERCISE: what would be the exterior derivative of a 2-form? What would its components be?)

Example 1.21. What about the exterior derivative of a 1 -form $\boldsymbol{\sigma}=\sigma_{u} \mathbf{d} u+\sigma_{v} \mathbf{d} v+\sigma_{w} \mathbf{d} w$ in $\mathbb{R}^{3}$ ? Using the equivalent form: $\mathbf{d} \boldsymbol{\sigma}=\mathbf{d} \sigma_{\nu} \wedge \mathbf{d} x^{\nu}$, we get:

$$
\begin{align*}
\mathbf{d} \boldsymbol{\sigma} & =\left(\partial_{v} \sigma_{u} \mathbf{d} v+\partial_{w} \sigma_{u} \mathbf{d} w\right) \wedge \mathbf{d} u+\left(\partial_{u} \sigma_{v} \mathbf{d} u+\partial_{w} \sigma_{v} \mathbf{d} w\right) \wedge \mathbf{d} v+\left(\partial_{u} \sigma_{w} \mathbf{d} u+\partial_{v} \sigma_{w} \mathbf{d} v\right) \wedge \mathbf{d} w \\
& =\left(\partial_{v} \sigma_{w}-\partial_{w} \sigma_{v}\right) \mathbf{d} v \wedge \mathbf{d} w+\left(\partial_{w} \sigma_{u}-\partial_{u} \sigma_{w}\right) \mathbf{d} w \wedge \mathbf{d} u+\left(\partial_{u} \sigma_{v}-\partial_{v} \sigma_{u}\right) \mathbf{d} u \wedge \mathbf{d} v \tag{1.56}
\end{align*}
$$

Taking the Hodge dual gives the pseudo-1-form:

$$
\begin{equation*}
\dot{\psi} \boldsymbol{\sigma}=\sqrt{|g|}\left[\left(\partial_{v} \sigma_{w}-\partial_{w} \sigma_{v}\right) \mathbf{d} u+\left(\partial_{w} \sigma_{u}-\partial_{u} \sigma_{w}\right) \mathbf{d} v+\left(\partial_{u} \sigma_{v}-\partial_{v} \sigma_{u}\right) \mathbf{d} w\right] \tag{1.57}
\end{equation*}
$$

By analogy with tensor algebra results, we recover the covariant version of the components of the 3 -dim curl of a vector, but only in Cartesian coordinates! Only in those coordinates is $\sqrt{|g|}=1$, with covariant and contravariant components the same in Cartesian $\mathbb{R}^{n}$.
As we know all too well, the vector components of the curl of a vector in curvilinear coordinates can be quite complicated, but this complication is largely due to our insisting on working with objects which are less natural. Indeed, when a curl is called for, it always is the curl of a 1 -form, and there is little need for its contravariant components which are sure to be more complicated because of the curvilinear metric being used to raise indices.
It is interesting that, in vector calculus with Cartesian coordinates, we could write $\boldsymbol{\sigma}$ as $\mathbf{A} \cdot \mathrm{d} \mathbf{x}$, with A a corresponding vector. Then the right-hand side of eq. (1.56) would correspond to $\boldsymbol{\nabla} \times \mathbf{A} \cdot \mathrm{d} \mathbf{S}$, where $\mathrm{d} \mathbf{S}$ is a surface element with Cartesian components $\mathbf{d} y \wedge \mathbf{d} z, \mathbf{d} z \wedge \mathbf{d} x$, and $\mathbf{d} x \wedge \mathbf{d} y$. Then we could write $\mathrm{d}(\mathbf{A} \cdot \mathrm{d} \mathbf{x})=\boldsymbol{\nabla} \times \mathbf{A} \cdot \mathrm{d} \mathbf{S}$.

Example 1.22. Here is an intriguing example: the exterior derivative of a pseudo-2-form $\boldsymbol{\tau}$ in $\mathbb{R}^{3}$ with some metric $\mathfrak{g}$. Since this will be a pseudo-3-form, we expect it to be a one-component object. Indeed:

$$
\begin{align*}
\mathbf{d} \tau & =\left(\partial_{u} \tau_{v w} \mathbf{d} u\right) \wedge \mathbf{d} v \wedge \mathbf{d} w+\left(\partial_{v} \tau_{w u} \mathbf{d} v\right) \wedge \mathbf{d} w \wedge \mathbf{d} u+\left(\partial_{w} \tau_{u v} \mathbf{d} w\right) \wedge \mathbf{d} u \wedge \mathbf{d} v \\
& =\left(\partial_{u} \tau_{v w}+\partial_{v} \tau_{w u}+\partial_{w} \tau_{u v}\right) \mathbf{d} u \wedge \mathbf{d} v \wedge \mathbf{d} w \tag{1.58}
\end{align*}
$$

Now, in three-dimensions $\boldsymbol{\tau}$ can be viewed as the Hodge dual, $\boldsymbol{\tau}=\underset{\boldsymbol{\tau}}{\boldsymbol{\sigma}} \boldsymbol{\sigma}$, of the 1-form $\boldsymbol{\sigma}=\sigma_{u} \mathbf{d} u+$ $\sigma_{v} \mathbf{d} v+\sigma_{w} \mathbf{d} w$. In terms of components, $\tau_{\mu \nu}=\epsilon_{\mu \nu \lambda} \sigma^{\lambda}$. Inserting and then taking the Hodge dual of the last expression, using $\underset{\sim}{ }(\mathbf{d} u \wedge \mathbf{d} v \wedge \mathbf{d} w)=\epsilon^{123}=(-1)^{n_{-}} / \sqrt{|g|}$ from section 1.5.3, gives:

$$
\begin{equation*}
(-1)^{n_{-}} \underset{\mathbf{d}}{\boldsymbol{\eta} \boldsymbol{\sigma}}=\frac{1}{\sqrt{|g|}} \partial_{\mu}\left(\sqrt{|g|} \sigma^{\mu}\right) \tag{1.59}
\end{equation*}
$$

Definition 1.34. Extending to $n$ dimensions, we call the right-hand side the divergence, $\operatorname{div} \mathbf{B}$, of the $n$-dim vector $\mathbf{B}$ with components $B^{\mu}=\sigma^{\mu}$. It is valid in general coordinates in any metric-endowed space.

In vector calculus with Cartesian coordinates, $\boldsymbol{\tau}=B^{1} \mathbf{d} y \wedge \mathbf{d} z+B^{2} \mathbf{d} z \wedge \mathbf{d} x+B^{3} \mathbf{d} x \wedge \mathbf{d} y=\mathbf{B} \cdot \mathrm{d} \mathbf{S}$, and eq. (1.58) could be written as: $\mathrm{d}(\mathbf{B} \cdot \mathrm{d} \mathbf{S})=\boldsymbol{\nabla} \cdot \mathbf{B} \mathrm{d}^{3} x$.

The operator $\sharp \mathbf{d} \sharp$ sends a $p$-form into a $(p-1)$-form. In mathematical references, this operator is introduced (up to a sign!) as the codifferential operator, $\delta$. We quote without proof the relation between them: When acting on a $p$-form in a Euclidean manifold, $\boldsymbol{\delta} \boldsymbol{\sigma}=(-1)^{n(p+1)+1} \leadsto \mathbf{d} \boldsymbol{\sim} \boldsymbol{\sigma}$, and $\boldsymbol{\delta} \boldsymbol{\sigma}=(-1)^{n(p+1)} \underset{\sharp}{ } \mathbf{d} \boldsymbol{\sigma}$ in a pseudo-Euclidean manifold. Actually, these expressions happen to hold also in a Riemannian (curved) or pseudoRiemannian manifold!

Like the exterior derivative, the codifferential operator is nilpotent. Indeed, $\boldsymbol{\delta}^{2}=\sharp \mathbf{d} \boldsymbol{d} \boldsymbol{\sharp} \mathbf{d}= \pm \sharp \mathbf{d}^{2} \leadsto=0$.

Definition 1.35. We define the divergence of any $p$-form: $\operatorname{div} \boldsymbol{\sigma}:=-\boldsymbol{\delta} \boldsymbol{\sigma}=(-1)^{n(p+1)+n_{-}} \boldsymbol{\varkappa} \mathbf{d} \boldsymbol{\sim} \boldsymbol{\sigma}$. This ensures consistency between eq. (1.59) and the conversion between $\approx \mathbf{d} \dot{\boldsymbol{\psi}}$ and $\boldsymbol{\delta}$. We extend eq. (1.59) to the divergence of any $p$-form $\sigma$ on a $n$-dim space:

$$
\begin{equation*}
(\operatorname{div} \boldsymbol{\sigma})_{\mu_{1} \ldots \mu_{p-1}}=\frac{1}{\sqrt{|g|}} \partial_{\nu}\left(\sqrt{|g|} \sigma_{\mu_{1} \ldots \mu_{p-1}}^{\nu}\right)=\frac{1}{\sqrt{|g|}} \partial_{\nu}\left(\sqrt{|g|} g^{\nu \rho} \sigma_{\rho \mu_{1} \ldots \mu_{p-1}}\right) \tag{1.60}
\end{equation*}
$$

From eq. (1.59) also follows the definition of the 3-dim Laplacian of a scalar function $f$ in coordinates $u^{i}$ :

$$
\begin{equation*}
\nabla^{2} f=\frac{1}{\sqrt{|g|}} \partial_{i}\left(\sqrt{|g|} \partial^{i} f\right)=\frac{1}{\sqrt{|g|}} \partial_{i}\left(\sqrt{|g|} g^{i j} \partial_{j} f\right) \tag{1.61}
\end{equation*}
$$

### 1.6.2 Laplace-de Rham operator, harmonic forms, and the Hodge decomposition

Definition 1.36. The Laplace-de Rham operator is defined as $\boldsymbol{\Delta}=\boldsymbol{\delta} \mathbf{d}+\mathbf{d} \boldsymbol{\delta}=(\mathbf{d}+\boldsymbol{\delta})^{2}$. It is a mapping $\boldsymbol{\Delta}: \bigwedge^{p} \longrightarrow \bigwedge^{p}$. When acting on a 0 -form (scalar function), $\boldsymbol{\Delta}=\boldsymbol{\delta} \mathbf{d}$ always; then we also speak of the Laplace-Beltrami operator.

It is not hard to show that it reduces to the negative of the Laplacian operator of vector analysis, ie. $\boldsymbol{\Delta}=\boldsymbol{\delta} \mathbf{d}=$ $-\hbar \mathbf{d} \hbar \mathbf{d}=-\partial_{i} \partial^{i}=-\nabla^{2}$, when acting on 0 -forms on Euclidean $\mathbb{R}^{3}$ with Cartesian coordinates. We shall define $\nabla^{2}$ so that $\nabla^{2}=-\boldsymbol{\Delta}$ when acting on any $p$-form in Euclidean $\mathbb{R}^{3}$ equipped with a standard basis.

For instance, let it act on a 1-form $\boldsymbol{\sigma}$ in Euclidean $\mathbb{R}^{3}$. That is, take $\boldsymbol{\Delta} \boldsymbol{\sigma}=\boldsymbol{\sigma} \boldsymbol{d} \mathbf{d} \boldsymbol{\sigma}-\mathbf{d} \boldsymbol{d} \boldsymbol{d} \boldsymbol{\sigma}$ using the conversion formula between $\delta$ and $\nleftarrow \mathbf{d} \nsim$. Using eq. (1.57), the first term is the curl of a curl, whereas the second is the gradient of a divergence. Thus, we recover the expression well-known from vector calculus: $\boldsymbol{\nabla}^{2} \mathbf{A}=$ $\boldsymbol{\nabla}(\boldsymbol{\nabla} \cdot \mathbf{A})-\boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \mathbf{A}$, where $\mathbf{A}$ is the vector associated with the 1-form $\boldsymbol{\sigma}$.

When acting on functions (0-forms) in Minkowski space, the Laplace-de Rham operator becomes the negative of the d'Alembertian wave operator $\square=\partial_{\mu} \partial^{\mu}: \Delta=-\square$. In fact, this will define the d'Alembertian of any $p$-form in a Minkowski manifold.

Definition 1.37. A $p$-form $\boldsymbol{\sigma}$ is said to be harmonic if $\boldsymbol{\Delta} \boldsymbol{\sigma}=0$. This generalises the notion of functions being called harmonic when they satisfy the Laplace equation.

Definition 1.38. A closed form is one whose exterior derivative vanishes. A p-form that can be written as the exterior derivative of a $(p-1)$-form is said to be exact.

Clearly, Poincaré's lemma states that an exact form is closed. But is a closed form exact, ie. if $\mathbf{d} \boldsymbol{\sigma}=0$, does it follow that $\sigma=\mathbf{d} \tau$, with $\tau$ uniquely determined? The answer is no, if only because one can always add the exterior derivative of an arbitrary ( $p-2$ )-form $\boldsymbol{\theta}$ to $\boldsymbol{\tau}$ and still satisfy $\mathbf{d} \boldsymbol{\sigma}=0$. Also, the converse of Poincarés lemma (not proved) states that only in a submanifold in which all closed curves can be shrunk to a point does $\mathbf{d} \boldsymbol{\sigma}=0$ entail the existence in that submanifold of a non-unique ( $p-1$ )-form whose exterior derivative is $\boldsymbol{\sigma}$. In topology, we say that the submanifold must be simply connected (eg. no doughnuts!).

We quote without proof an important result of Hodge: On finite-volume (compact) manifolds without boundaries, such as $S^{n}$, or on a torus, $\boldsymbol{\Delta} \boldsymbol{\sigma}=0$ if, and only if, $\mathbf{d} \boldsymbol{\sigma}=0$ and $\mathbf{d}^{*} \boldsymbol{\sigma}=0$ (or $\boldsymbol{\delta} \boldsymbol{\sigma}=0$ ). We then say that harmonic forms are both closed and co-closed. This property also holds on open manifolds (eg. $\mathbb{R}^{n}$ ) if $\sigma$ has compact support (it vanishes outside a bounded closed region), or if it goes to zero sufficiently fast at infinity. This is often the case for physically useful quantities.

Definition 1.39. Assuming a compact manifold without boundaries or, failing that, compact support or sufficiently fast fall-off at infinity, the Hodge decomposition of a $p$-form is:

$$
\begin{equation*}
\boldsymbol{\sigma}=\mathbf{d} \boldsymbol{\alpha}+\boldsymbol{\delta} \boldsymbol{\beta}+\text { harmonic form } \tag{1.62}
\end{equation*}
$$

where $\boldsymbol{\alpha}$ is a ( $p-1$ )-form and $\boldsymbol{\beta}$ is a $(p+1)$-form. This decomposition is unique, in the sense that each term in the sum is unique, but not $\boldsymbol{\alpha}$ or $\boldsymbol{\beta}$ themselves. $\mathbf{d} \boldsymbol{\alpha}, \boldsymbol{\delta} \boldsymbol{\beta}$ and the harmonic $p$-form in the decomposition live in orthogonal subspaces of $\bigwedge^{p}$.

Example 1.23. In non-compact Euclidean $\mathbb{R}^{3}$, if $\mathbf{A}$ is a vector field with compact support, then the Hodge decomposition says that its associated 1 -form can be written as the exterior derivative of a 0 form (ie. the gradient of a function), plus the divergence of a 2 -form, $\boldsymbol{\beta}$, plus some harmonic 1 -form. Now, since $\rightsquigarrow \boldsymbol{\beta}$ is a pseudo- 1 -form in $\mathbb{R}^{3}, \boldsymbol{\delta} \boldsymbol{\beta}=\sharp \mathbf{d} \boldsymbol{\xi} \boldsymbol{\beta}$ is a 1 -form, and we find from eq. (1.57) that this term in the Hodge decomposition corresponds to the curl of a vector. Therefore, we obtain in terms of vectors the Helmholtz decomposition for a vector field with compact support or sufficiently fast fall-off at infinity:

$$
\begin{equation*}
\mathbf{A}=\boldsymbol{\nabla} \phi+\boldsymbol{\nabla} \times \mathbf{M}+\mathbf{H} \tag{1.63}
\end{equation*}
$$

where $\phi$ is a scalar field, $\mathbf{M}$ a vector field, and $\mathbf{H}$ another vector field which satisfies $\boldsymbol{\nabla}^{2} \mathbf{H}=0$ everywhere. But if $\mathbf{H}$ vanishes at infinity in $\mathbb{R}^{3}$, then it must vanish everywhere.
The curl of the first term in the Helmholtz decomposition vanishes identically, and is often called the longitudinal projection of $\mathbf{A}$; the divergence of the second term vanishes identically, and we can call it the transverse projection of $\mathbf{A}$. It follows that $\boldsymbol{\nabla} \cdot \mathbf{A}$ contains no information whatsoever about the transverse part of $\mathbf{A}$, whereas $\boldsymbol{\nabla} \times \mathbf{A}$ knows nothing of its longitudinal part. This provides a very useful and powerful tool for analysing 3 -dim first-order field equations (eg. Maxwell's equations) which are usually statements about the divergence and the curl of the fields. Conversely, if $\boldsymbol{\nabla} \cdot \mathbf{A}=0$ everywhere, we can conclude that $\mathbf{A}$ is purely transverse, since then the 3 -scalar field $\phi$ in eq. (1.63) satisfies the Laplace equation everywhere, which means it must vanish if it has compact support or fast enough fall-off.

### 1.6.3 Exterior derivative and codifferential operator of a 2-form in Minkowski spacetime

Let $\mathbf{F} \in \Lambda^{2}$ on Minkowski (pseudo-Euclidean) $\mathbb{R}^{4}$. Demand that $\mathbf{F}$ be exact, with compact support or sufficiently fast fall-off at infinity. Then there exists a 1 -form $\mathbf{A}$ such that $\mathbf{F}=\mathbf{d} \mathbf{A}$, and $F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu}$, in any metric. This means that $\mathbf{d F}=0$. It is clear from Poincaré's lemma, $\mathbf{d}^{2} \mathbf{A}=0 \forall \mathbf{A}$, that any information about $\mathbf{F}$ contained in $\mathbf{d F}=0$, is of no use whatsoever in determining $\mathbf{A}$ : we say that it is an identity on $\mathbf{A}$.

In addition, we give the exterior derivative of the Hodge dual of $\mathbf{F}$, the pseudo-2-form $\mathbf{d} \boldsymbol{\jmath} \mathbf{F}$, as a "source" pseudo-3-form $\mathcal{J}$, with corresponding Hodge dual 1-form $\mathbf{J}=\sharp \mathcal{J}$. Then we have the inhomogeneous equation:

$$
\begin{equation*}
\mathbf{d} \mathfrak{H}=4 \pi \mathcal{J} \tag{1.64}
\end{equation*}
$$

Interestingly enough, one can view eq. (1.64) as a statement that the source 3 -form is exact, and therefore closed. This is a very important property because, if we take the exterior derivative of the equation, the left-hand side vanishes identically, and the right-hand side becomes: $\mathbf{d} \mathcal{J}=0$ or, equivalently, the better-known statement that the 4 -divergence of the source 1 -form vanishes: $\leadsto \mathbf{d} \downarrow \mathbf{J}=0 . \mathbf{d} \mathcal{J}=0$ is actually a (metric-independent) conservation law for the source!

What we have just seen is Maxwell's theory, with $\mathbf{F}$ the Faraday 2-form (tensor), A the electromagnetic potential 1-form, and $\mathbf{J}$ the electromagnetic 4-current.

Our treatment assumes a mostly positive metric, as in MTW or Griffith's Introduction to Electrodynamics. With a mostly negative metric, there would be a minus sign on the right-hand side of eq. (1.64).

Because the differential operator $\mathbf{d}$ is metric-independent, we have given both the homogeneous and inhomogeneous first-order Maxwell equations in terms of exterior derivatives of $\mathbf{F}$ and its dual $ヶ \mathbf{F}$. It is easy to convert the inhomogeneous equation to a divergence, simply by taking its Hodge dual:

$$
\begin{equation*}
\hbar \mathbf{d} \boldsymbol{H}=4 \pi \text { 动 } \mathcal{J}=4 \pi \mathbf{J} \quad \Longleftrightarrow \quad \operatorname{div} \mathbf{F}=-4 \pi \mathbf{J} \tag{1.65}
\end{equation*}
$$

In terms of Cartesian components, this can be shown (EXERCISE) to be equivalent to ${ }^{\dagger}$

$$
\partial^{\mu} F_{\mu \nu}=-4 \pi J_{\nu} \quad \Leftrightarrow \quad \partial_{\mu} F^{\mu \nu}=-4 \pi J^{\nu}
$$

[^6]the latter form being more appropriate if we insist on thinking of the source term as a vector. I would argue, however, that the less conventional form eq. (1.64) is much the more natural. The exterior derivative is metricindependent, and its index form can be written entirely with covariant indices, the natural ones for $p$-forms. But to obtain its equivalent in divergence form, we have to Hodge-dualise the right-hand side, so that the vector $\mathbf{J}$ source depends on the metric (see the paragraph after eq. (1.52)), whereas its 3 -form version does not. The price, of course, is that the 3 -form version has 64 compnents.

It is worth noting that, although $\mathrm{dF}=0$ and the source equation (1.64) completely determine $\mathbf{F}, \mathbf{A}$ is determined only up to an additive term $\mathbf{d} f$, where $f$ is an arbitrary differentiable function.

As a 3 -form, the homogeneous equation $\mathbf{d F}=0$ also has a lot of components, and when it comes to solving the system, we may want to extract only the independent ones. Well, this equation is the same as $\mathbf{d} \hbar(\hbar \mathbf{F})=0$ for which the Hodge dual is $\boldsymbol{\delta} \boldsymbol{\hbar} \mathbf{F}=0$. In other words, the divergence of the dual of $\mathbf{F}$ vanishes, which represents only four equations. Actually, this is a general, easily shown property (EXERCISE): whenever the exterior derivative of a $p$-form in some manifold vanishes, so does the codifferential of its dual, and vice-versa.

Another great advantage of writing Maxwell's equations in the geometric, index-independent formalism is that, provided the source is smoothly varying, they are formally the same in curved spacetime! All the extra terms due to the covariant derivatives that one has to introduce in General Relativity cancel out, essentially because they are symmetric in the same indices in which the $p$-form equations are antisymmetric. In index notation, however, only the equations involving no Hodge dual remain the same. Thus, the homogeneous Maxwell equation in the form of a divergence becomes metric-dependent, and in index notation will not be the same in curved spacetime!

Finally, nothing prevents us from constructing an extended Maxwell-like theory (which does not describe electromagnetism) involving $\mathbf{F}$ as a 3 -form instead of a 2 -form. In the past few decades it has received a good deal of attention in some quarters.

### 1.7 Integrals of Differential (Pseudo)Forms

In the course of figuring out the meaning of $\int \boldsymbol{\sigma}^{p}$, where we now use the notation $\boldsymbol{\sigma}^{p}$ to show explicitly the rank of a $p$-form, we shall discover that pretty much any integral that makes sense in $n$-dim calculus can be written as the integral of some $p$-form or pseudo- $p$-form.

### 1.7.1 Integrals of (pseudo) $\boldsymbol{p}$-forms over a $\boldsymbol{p}$-dim submanifold

As a warm-up, consider the integral of the Hodge dual of a scalar function $f, \int \sharp f$, over a $n$-dim region $V$ in $\mathbb{R}^{n}$ (eg., over some volume in $\mathbb{R}^{3}$ ). The Hodge dual of a scalar function, of course, is a $n$-form. Then:

$$
\int_{V} \star f=\int_{V} f(\mathbf{u}) \sqrt{|g|} \mathbf{d} u^{1} \wedge \cdots \wedge \mathbf{d} u^{n}=\int_{V} f(\mathbf{x}) \mathbf{d} x^{1} \wedge \cdots \wedge \mathbf{d} x^{n}=\int_{V} f(\mathbf{x}) \mathbf{d}^{n} x
$$

where $u$ are general coordinates and $\mathbf{d}^{n} x$ is in the volume pseudo- $n$-form in Cartesian coordinates. If the coordinate basis $\left\{\boldsymbol{\partial}_{\mu}\right\}(1 \leq \mu \leq n)$ has the same orientation as the (arbitrarily chosen) orientation of the region, then we define:

## Definition 1.40.

$$
\begin{equation*}
\int_{V} f(\mathbf{x}) \mathbf{d} x^{1} \wedge \cdots \wedge \mathbf{d} x^{n}:=\int_{V} f(\mathbf{x}) \mathrm{d} x^{1} \cdots \mathrm{~d} x^{n}=\int_{V} f(\mathbf{x}) \mathrm{d}^{n} x \tag{1.66}
\end{equation*}
$$

ie. the ordinary multiple integral of a scalar function of $n$ variables in $n$ dimensions.
When a $p$-dim region $R$ is embedded in a $n$-dim manifold, it will be described with some coordinates $\mathbf{u}(\mathbf{x})$, that is, $n$ functions $u^{i}$ of the $p$ Cartesian coordinates $x^{j}$ that parametrise $\mathbb{R}^{p}$. Also, an orientation can be defined for the region. What is the meaning of the integral of a $p$-form over such a region? We give two examples in $\mathbb{R}^{3}$.

## Example 1.24. Integral of a 1-form over a curve or "line integral"

As we have seen before, a curve $C$ can be parametrised in terms of some parameter $t \in \mathbb{R}$, with $\mathrm{d}_{t}$ defining the positive orientation of the parameter space $\mathbb{R}$. Then, if $\boldsymbol{\alpha}$ is a 1 -form on $\mathbb{R}^{3}$, eq. (1.66) and the chain rule yield:

$$
\int_{C} \boldsymbol{\alpha}=\int_{C} \alpha_{i} \mathbf{d} u^{i}=\int_{a}^{b} \alpha_{i}[\mathbf{u}(t)] \mathrm{d}_{t} u^{i} \mathrm{~d} t=\int_{a}^{b} \boldsymbol{\alpha}\left(\mathrm{~d}_{t} \mathbf{u}\right) \mathrm{d} t
$$

Only if $\mathbb{R}^{3}$ is given a metric and the curve parametrised with Cartesian coordinates is this the usual integral of a vector $\mathbf{A}$ along the curve, $\int \mathbf{A} \cdot \mathrm{d} \mathbf{x}$. In general, to integrate a vector along a curve, a metric must be introduced so as to transform the vector components into its associated 1-form's components: $\int \mathbf{A} \cdot \mathrm{d} \mathbf{u}=\int g_{i j} A^{j} \mathrm{~d}_{t} u^{i} \mathrm{~d} t$. But no metric is needed to integrate a 1 -form along a curve, and this is the simpler and more natural operation.

If $\boldsymbol{\alpha}$ is exact, then we immediately have the fundamental theorem of calculus:

$$
\int_{C} \mathbf{d} f=\int_{a}^{b} \partial_{u^{i}} f \mathrm{~d}_{t} u^{i} \mathrm{~d} t=\int_{a}^{b} \mathrm{~d} f=f(b)-f(a)=\int_{\partial C} f
$$

where $\partial C$ is the boundary, ie. the end-points, of the curve.

## Example 1.25. Integral of a 2-form over a surface

Let $S$ be some surface embedded in $\mathbb{R}^{3}$ and described with three coordinate functions $u^{i}\left(x^{1}, x^{2}\right)$. The surface is parametrised with $\left(x^{1}, x^{2}\right) \in \mathbb{R}^{2}$, with basis vectors $\boldsymbol{\partial}_{x^{1}} \equiv \boldsymbol{\partial}_{1}$ along the $x^{1}$ direction and $\boldsymbol{\partial}_{x^{2}} \equiv \boldsymbol{\partial}_{2}$ along the $x^{2}$ direction, for which some orientation has been defined as positive. What meaning can we give to the integral of a 2 -form $\boldsymbol{\beta}$ over $S$ ? From the chain rule, example 1.13, and eq. (1.66) we find:

$$
\int_{S} \boldsymbol{\beta}=\int_{S} \beta_{j k} \mathbf{d} u^{j} \wedge \mathbf{d} u^{k}=\int \beta_{j k}\left[\mathbf{u}\left(x^{1}, x^{2}\right)\right]\left(\partial_{1} u^{j} \partial_{2} u^{k}-\partial_{2} u^{j} \partial_{1} u^{k}\right) \mathrm{d} x^{1} \mathrm{~d} x^{2} \quad(j<k)
$$

The integrals in $\mathbb{R}^{2}$ on the right are over a rectangular region in parameter space. The two coordinate vectors (see section 1.2.3), $\partial_{1} \mathbf{u}$ and $\partial_{2} \mathbf{u}$, are tangent to the surface at every point, and are usually linearly independent, so form a basis for the space of (pseudo)vectors tangent to the surface. Here again, no metric is required.

Now the Hodge dual of $\boldsymbol{\beta}$, which is a pseudo-1-form, has an associated pseudo-vector $\mathbf{B}$ with, as components, the contravariant components of the Hodge dual, $B^{i}=\epsilon^{i j k} \beta_{j k}(j<k)$, eg., $B^{1}=$ $\beta_{23} / \sqrt{|g|}$, etc. Then:

$$
\beta_{j k}\left(\partial_{1} u^{j} \partial_{2} u^{k}-\partial_{2} u^{j} \partial_{1} u^{k}\right)=\epsilon_{i j k} B^{i}\left(\partial_{1} u^{j} \partial_{2} u^{k}-\partial_{2} u^{j} \partial_{1} u^{k}\right)=\sqrt{|g|}\left|\begin{array}{ccc}
B^{1} & B^{2} & B^{3} \\
\partial_{1} u^{1} & \partial_{1} u^{2} & \partial_{1} u^{3} \\
\partial_{2} u^{1} & \partial_{2} u^{2} & \partial_{2} u^{3}
\end{array}\right|
$$

From eq. (1.47), we recognise the last member of the equality as the output obtained from inserting the three vectors whose components are the rows of the determinant, into the three input slots of a simple 3 -form-more accurately, a pseudo-3-form which, from definition (1.31) can be identified with the volume pseudo-form $\mathbf{d}^{3} u$. Then our integral can be written:

$$
\int_{S} \boldsymbol{\beta}=\int \mathbf{d}^{3} u\left(\mathbf{B}, \partial_{1} \mathbf{u}, \partial_{2} \mathbf{u}\right) \mathrm{d} x^{1} \mathrm{~d} x^{2}
$$

This makes it obvious that the integral is independent of the orientation of $\mathbb{R}^{3}$, since switching it flips the sign of both $\mathbf{B}$ and $\mathbf{d}^{3} u$. At every point on the surface, we can choose the unit $\mathbf{n}$ normal to the
surface so that $\mathbf{n}$ and the vectors $\partial_{1} \mathbf{u}$ and $\partial_{2} \mathbf{u}$ tangent to the surface form a right-handed (positive orientation) system. We also note that only the normal component of $\mathbf{B}$ can contribute to the integral (WHY?).
Then the scalar functionl $\mathbf{d}^{3} u\left(\mathbf{B}, \partial_{1} \mathbf{u}, \partial_{2} \mathbf{u}\right)$ is the normal component of $\mathbf{B}$ multiplied by the surface of the parallelogram defined by the coordinate vectors. Defining the surface element $\mathrm{d} S \equiv\left|\partial_{1} \mathbf{u} \times \partial_{2} \mathbf{u}\right|$, there comes:

$$
\begin{equation*}
\int_{S} \boldsymbol{\beta}=\int B_{\mathrm{n}} \mathrm{~d} S=\int \mathbf{B} \cdot \mathrm{d} \mathbf{S} \tag{1.67}
\end{equation*}
$$

where the often used last expression is called the flux of the pseudo-vector $\mathbf{B}$ through the surface $S$. It does not depend on the parametrisation chosen for $S$ which is integrated out. The same result holds if $\boldsymbol{\beta}$ is a pseudo-2-form, with $\mathbf{B}$ now a vector.

### 1.7.2 Stokes Theorem

This famous theorem, which we shall not prove, equates the integral of the exterior derivative of a differentiable (pseudo) $p$-form, $\boldsymbol{\omega}$, over a bounded region $V$ in a manifold to the integral of $\boldsymbol{\omega}$ over the boundary $\partial V$ of $V$. A technicality is that both $V$ and $\partial V$ must have compatible orientations.

$$
\begin{equation*}
\int_{V} \mathbf{d} \omega=\int_{\partial V} \omega \tag{1.68}
\end{equation*}
$$

The boundary need not be connected, and it can be broken up into non-overlapping parts when it cannot be covered by a single coordinate patch. Then we simply sum the integrals over each part.

Example 1.26. At the end of example 1.24 we had already worked out an application when $\boldsymbol{\omega}$ is a 0 -form: the fundamental theorem of calculus. When $\boldsymbol{\omega}$ is a 1-form and $V$ a 2-dim surface in Euclidean $\mathbb{R}^{3}$ parametrised with Cartesian coordinates and bounded by a closed curve $C$, the same example gives immediately: $\int_{\partial V} \boldsymbol{\omega}=\oint_{C} \mathbf{A} \cdot$ du. From eq. (1.56) and example $1.25, \int_{S} \mathbf{d} \boldsymbol{\omega}=\int_{S} \boldsymbol{\nabla} \times \mathbf{A} \cdot \mathrm{d} \mathbf{S}$, and we recover the well-known Kelvin-Stokes formula. Only if we want to work with that usual vector calculus relation is a metric required.
Finally, when $\boldsymbol{\omega}$ is a pseudo-2-form in Euclidean $\mathbb{R}^{3}, V$ a volume and $S$ a surface enclosing the volume, we recover the relation $\int_{V} \boldsymbol{\nabla} \cdot \mathbf{B} \mathrm{~d} V=\oint_{S} \mathbf{B} \cdot \mathrm{~d} \mathbf{S}$ from examples 1.22 and 1.25. This time, a metric is needed.

### 1.8 Maxwell Differential Forms in Three + One Dimensions

With $\mathbf{F}$ the Maxwell 2-form, define two 3-dim $p$-forms: an electric field strength 1-form $\mathcal{E}$ and a magnetic field strength 2 -form $\mathcal{B}$, by:

$$
\begin{equation*}
\mathbf{F}=F_{|\mu \nu|} \mathrm{d} x^{\mu} \wedge \mathbf{d} x^{\nu}=\mathcal{E} \wedge \mathbf{d} t+\mathcal{B} \tag{1.69}
\end{equation*}
$$

where:

$$
\begin{equation*}
\mathcal{E}:=F_{10} \mathbf{d} x^{1}+F_{20} \mathbf{d} x^{2}+F_{30} \mathbf{d} x^{3} \quad \mathcal{B}:=F_{12} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}+F_{31} \mathbf{d} x^{3} \wedge \mathbf{d} x^{1}+F_{23} \mathbf{d} x^{2} \wedge \mathbf{d} x^{3} \tag{1.70}
\end{equation*}
$$

Now, formally, $\mathbf{d}=\overrightarrow{\mathbf{d}}+\mathbf{d} t \wedge \partial_{t}$. where $\overrightarrow{\mathbf{d}}$ denotes the 3-dimensional exterior derivative. Then the homogeneous Maxwell equation, $\mathbf{d F}=0$, becomes:

$$
\begin{aligned}
{\left[\overrightarrow{\mathbf{d}}+\mathbf{d} t \wedge \partial_{t}\right][\mathcal{E} \wedge \mathbf{d} t+\mathcal{B}] } & =\overrightarrow{\mathbf{d}} \mathcal{E} \wedge \mathbf{d} t+\overrightarrow{\mathbf{d}} \mathcal{B}+\mathbf{d} t \wedge \partial_{t} \mathcal{B} \\
& =\left(\overrightarrow{\mathbf{d}} \mathcal{E}+\partial_{t} \boldsymbol{\mathcal { B }}\right) \wedge \mathbf{d} t+\overrightarrow{\mathbf{d}} \mathcal{B} \\
& =0
\end{aligned}
$$

In the second line, the plus sign in the bracket results from the double transposition needed to bring $\mathbf{d} t$ to the right through the two basis 1 -forms of $\mathcal{B}$. We conclude that in three dimensions, the homogeneous Maxwell equation gives rise to:

$$
\begin{equation*}
\overrightarrow{\mathbf{d} \mathcal{B}=0} \quad \overrightarrow{\mathbf{d} \mathcal{E}}+\partial_{t} \mathcal{B}=0 \tag{1.71}
\end{equation*}
$$

Expand the potential 1-form $\mathbf{A}$ as $-\phi \mathbf{d} t+\mathcal{A}$, where $\phi \in \bigwedge^{0}$ and $\mathcal{A} \in \Lambda^{1}$, both in three dimensions. Expanding $\mathbf{F}=\mathbf{d A}$, or using the Poincaré lemma for $\overrightarrow{\mathbf{d}}$, one obtains (EXERCISE):

$$
\begin{equation*}
\mathcal{E}=-\mathbf{d} \phi-\partial_{t} \mathcal{A}, \quad \mathcal{B}=\overrightarrow{\mathbf{d}} \mathcal{A} \tag{1.72}
\end{equation*}
$$

I wish to reiterate that eq. (1.71) and (1.72) are metric-independent, and will thus hold in flat or curved spacetime, so long as we use a coordinate basis.

The Hodge duals of eq. (1.71) can be written as:

$$
\operatorname{div}^{*} \mathcal{B}=0 \quad{ }^{*} \overrightarrow{\mathbf{d}} \mathcal{E}+\partial_{t}^{*} \mathcal{B}=0
$$

If we identify the contravariant components of the pseudo-1-form ${ }^{*} \mathcal{B}$ with the usual components of the magneticfield pseudo-vector, and use eq. (1.57), we see that these are equivalent to the homogeneous Maxwell equations in their vector-calculus form: $\boldsymbol{\nabla} \cdot \mathbf{B}=0$ and $\boldsymbol{\nabla} \times \mathbf{E}+\partial_{t} \mathbf{B}=0$.

After Hodge-dualising the second equation (1.72), these are readily identified with the standard expressions for the fields in terms of the scalar and vector potential $* \overrightarrow{\mathbf{d}} \mathcal{A}$. But we also see that it is much more natural to view the 3-dim magnetic field as a 2 -form which is the exterior derivative of a 1-form, than as a pseudo-vector which is the curl of another vector. In the same vein, I contend that it is more natural to describe the electric field strength with the 1 -form $\mathcal{E}$ than with the vector $\mathbf{E}$. It is consistent with force and momentum also being more naturally 1-forms (consider $\mathrm{e}^{\mathrm{i} p_{\mu} x^{\mu}}$ !).

The inhomogeneous (source) Maxwell equation requires much more care, because the Hodge dual inevitably involves a metric, and because a 4-dim Hodge dual is not necessarily like a 3-dim Hodge dual! First, we must derive an expansion of ${ }^{*} \mathbf{F}$ in terms of $\mathcal{E}$ and $\mathcal{B}$. A safe, if somewhat inelegant, method is to expand it in terms of the components of $\mathbf{F}=\frac{1}{2} F_{\mu \nu} \mathbf{d} x^{\mu} \wedge \mathbf{d} x^{\nu}$ :

$$
\begin{aligned}
* \mathbf{F} & =\frac{1}{4} F^{\mu \nu} \epsilon_{\mu \nu \alpha \beta} \mathbf{d} x^{\alpha} \wedge \mathbf{d} x^{\beta} \\
& =-\sqrt{|g|}\left[F^{10} \mathbf{d} x^{2} \wedge \mathbf{d} x^{3}+F^{20} \mathbf{d} x^{3} \wedge \mathbf{d} x^{1}+F^{30} \mathbf{d} x^{1} \wedge \mathbf{d} x^{2}+\left(F^{12} \mathbf{d} x^{3}+F^{31} \mathbf{d} x^{2}+F^{23} \mathbf{d} x^{1}\right) \wedge \mathbf{d} t\right]
\end{aligned}
$$

Now we must write this in terms of the covariant components of $\mathbf{F}$, and this is where the metric must come in, since $F^{\mu \nu}=g^{\mu \alpha} g^{\nu \beta} F_{\alpha \beta}$ :

$$
F^{i 0}=\left(g^{00} g^{i j}-g^{i 0} g^{0 j}\right) F_{j 0}+g^{i j} g^{0 k} F_{j k}, \quad \quad F^{i j}=\left(g^{i 0} g^{j l}-g^{i l} g^{j 0}\right) F_{l 0}+g^{i k} g^{j l} F_{k l}
$$

We know that $F_{j 0}$ and $F_{j k}$ are the components of the 3 - $\operatorname{dim} p$-forms $\mathcal{E}$ and $\mathcal{B}$, respectively. If $g^{0 i} \neq 0$, each contravariant component of $\mathbf{F}$ will involve both $\mathcal{E}$ and $\mathcal{B}$, which will lead to very complicated results. When $g^{0 i}=0$, however, we are left with $F^{i 0}=g^{00} g^{i j} F_{j 0}$, and $F^{i j}=g^{i k} g^{j l} F_{k l}$, and lowering the spatial components of $\mathbf{F}$ involves only the spatial sector of the metric (ignoring the $g^{00}$ factor), the same sector that is used to raise indices on the Levi-Civita tensor. Also, if we take $g^{00}=-1$ (mostly positive) Minkowski metric, the $\sqrt{|g|}$ factor is the same for the three-dimensional metric determinant as for the 4-dim one. Because of all this, we can now write:

$$
{ }^{*} \mathbf{F}=-\left[\frac{1}{2} \epsilon_{i j k} F^{i 0} \mathbf{d} x^{j} \wedge \mathbf{d} x^{k}+\frac{1}{2} \epsilon_{i j k} F^{i j} \mathrm{~d} x^{k} \wedge \mathbf{d} t\right]
$$

where the roman indices run from 1 to 3 . Now we can relate the two terms to $\mathcal{E}$ and $\mathcal{B}$ :

$$
\frac{1}{2} \epsilon_{i j k} F^{i 0} \mathbf{d} x^{j} \wedge \mathbf{d} x^{k}=\frac{1}{2} \epsilon_{i j k} g^{00} g^{i l} F_{l 0} \mathbf{d} x^{j} \wedge \mathbf{d} x^{k}=\frac{1}{2} g^{00} \epsilon_{i j k} \mathcal{E}^{i} \mathbf{d} x^{j} \wedge \mathbf{d} x^{k}=g^{00 *} \mathcal{E}=-{ }^{*} \mathcal{E}
$$

Also:

$$
\frac{1}{2} \epsilon_{i j k} F^{i j} \mathbf{d} x^{k}={ }^{*} \mathcal{B}
$$

with no assumption needed for the spatial part of the 4-dim metric. Then our expansion is ${ }^{*} \mathbf{F}=-{ }^{*} \mathcal{B} \wedge \mathbf{d} t+{ }^{*} \mathcal{E}$ where it is understood that, on the right-hand side only, the 3-dim Hodge dual is taken. It is not difficult to show $($ EXERCISE $)$ that: $\mathbf{d}^{*} \mathbf{F}=-\left(\overrightarrow{\mathbf{d}}^{*} \mathcal{B}-\partial_{t}{ }^{*} \mathcal{E}\right) \wedge \mathbf{d} t+\overrightarrow{\mathbf{d}}^{*} \mathcal{E}$.

We define the Maxwell source pseudo-3-form as the expansion:

$$
\mathcal{J} \equiv \boldsymbol{\rho}-\mathbf{j} \wedge \mathbf{d} t \equiv \rho \epsilon_{i j k} \mathbf{d} x^{i} \wedge \mathbf{d} x^{j} \wedge \mathbf{d} x^{k}-{ }^{*} \mathbf{J} \wedge \mathbf{d} t \quad(i<j<k)
$$

where $\rho$ is the charge scalar density, $\boldsymbol{\rho}$ the three-dim charge-density pseudo-3-form and $\mathbf{J}$ the 3 -dim current density 1 -form. Inserting these expansions in eq. (1.64) yields the two 3 -dim Maxwell field equations:

$$
\begin{equation*}
\overrightarrow{\mathbf{d}}^{*} \mathcal{E}=4 \pi \boldsymbol{\rho}, \quad \overrightarrow{\mathbf{d}}^{*} \mathcal{B}=\mathbf{j}+\partial_{t}^{*} \mathcal{E} \tag{1.73}
\end{equation*}
$$

Taking the 3-dim Hodge dual of these equations recovers the vector-calculus form of Gauss's law for electricity and the Ampère-Maxwell equation.

## 2 MODULE II - GROUP THEORY 1: Discrete Groups

One of the most beautiful and useful concepts in physics and, dare I say, mathematics, is that of symmetry. Loosely speaking, it recognises the existence of patterns or, mathematically-speaking, a characteristic behaviour of objects under transformations. When a symmetry in a problem is detected, the problem is amenable to much simplification and might even be solvable. Useful information can be recovered even if the symmetry is only approximate, or is "broken" in a way that can be understood. Equally important, a symmetry often signals the existence of a conserved quantity. For instance, from space-translation invariance (aka homogeneity of space) follows linear-momentum conservation, whereas time-translation invariance generates energy conservation; isotropy of space (its invariance under rotations) leads to angular-momentum conservation. Conservation of electric charge is embodied in the gauge invariance of Maxwell's equations, which itself can be seen as generating constraints that automatically suppress potentially disastrous negative-energy modes.

In modern mathematics, the language of group theory provides a unified and systematic framework for classifying and describing symmetries. In part because it is jargon-heavy, group theory is often relegated to the fringes of most physicists' training. Yet much insight can be gained from at least a modicum of familiarity with it, and this is what we shall now attempt to acquire.

### 2.1 Groups: Definitions, Taxonomy and Examples (BF 10.1)

As much of what we are going to discuss involves mappings, it is worth saying a few words right at the outset about what kind of mappings will be of interest to us.

Most important, we will want to be able to compose mappings. For instance, if $f$ and $g$ are two mappings, we want to be able to apply one, and then the other; we shall denote ${ }^{\dagger}$ a binary composition by $f \circ g$ or $f \star g$, with the understanding that $g$ is applied first. In order that a string of such mappings be uniquely defined, they must be associative, and thus reducible to a succession of binary mappings. Moreover, we should like to be able to retrace any step we take in a unique way; this can be achieved by demanding that any mapping $f$ we consider be invertible with a mapping $f^{-1}$. Therefore, there must exist an identity mapping, $e$, such that $f \star f^{-1}=f^{-1} \star f=e$.

### 2.1.1 Basic definitions and nomenclature

The mappings of interest to us belong to a more general type of collection of objects called a group:
Definition 2.1. Let $G$ be a set of distinct objects endowed with an associative binary composition law or operation such that:

- $\forall a, b \in G, a \star b \in G$;
- there exists a unique element $e \in G$ such that, $\forall a \in G, e \star a=a \star e=a$;
- $\forall a \in G$, there exists a unique element $a^{-1} \in G$ such that $a^{-1} \star a=a \star a^{-1}=e$;
then we say that $G$ is a group.
Two remarks are in order. The composition law is often called group multiplication, a term we shall try to avoid because it almost irresistibly evokes ordinary multiplication. Also, although a group composition law definitely meets the requirements set out above for mappings, group elements themselves may be mappings, or transformations (eg. translations, rotations, permutations), but also numbers, vectors, matrices, etc.

The composition of two elements of $G$ is in general noncommutative: $a \star b \neq b \star a$.
Definition 2.2. When $a \star b$ is commutative, $\forall a, b \in G$, we call $G$ an Abelian group.
Definition 2.3. A group of $n$ elements $(n<\infty)$ is said to be finite and of order $n$.

[^7]Definition 2.4. A group is discrete when it contains a countable number of elements. All finite groups are discrete, but infinite discrete groups exist. Infinite groups which are not discrete are called continuous.

Definition 2.5. A group of the form $\left\{e, a, a^{2}, \ldots, a^{n-1}\right\}$, where $a^{i} \equiv a \star a \star \ldots i$ times, and where $n$ is the smallest integer for which $a^{n}=e$, is called $Z_{n}$ (sometimes $C_{n}$ ), the cyclic group of order $n$.
All $a^{i}$ in $Z_{n}$ are distinct for, supposing $a^{i}=a^{j}$, we would have $a^{i-j}=e$, with $i-j<n$, and $n$ would not be the smallest number of compositions of $a$ that yields $e$.

Given any element $a_{i}$ of a finite group, there will be an integer $m$ such that $a_{i}^{m}=e$. Then we say that $a_{i}$ itself is of order $m$. If $m$ is smaller than the order of the group, the group is not cyclic. One shows (EXERCISE) that groups whose elements all have order 2 are Abelian.

Here are a few straightforward examples of groups: $\mathbb{C}$ under addition, with $e=0$ and $a^{-1}=-a ; \mathbb{C}-\{0\}$ under multiplication, with $e=1$ and $z^{-1}=1 / z$; the set of complex $n \times n$ matrices with non-zero determinant, $G L(n, \mathbb{C})$, under matrix multiplication, with $e=\mathbf{I}$; the set of the $n$ complex roots of 1 under multiplication, with $e=1$.

It is important to keep in mind that a given set may be a group under one operation, but not under another. Thus, the set $\mathbb{Z}$ is not a group under multiplication (and neither is $\mathbb{Z}-\{0\}$ ); but is is a group under addition with $e=0$ and $a^{-1}=-a$.

Exercise: spot any discrete and cyclic groups in these examples.

### 2.1.2 Cayley tables

Let $a_{i}(i=1, \ldots, n)$ be an element of a finite group. By convention, $a_{1}=e$. We can construct a $n \times n$ composition table, or Cayley table, whose $i j^{\text {th }}$ element is $a_{i} \star a_{j}$. Then the first row and the first column must be $\left\{e, a_{2}, \ldots, a_{n}\right\}$. They are sometimes omitted by authors who are not nice to their readers.

A constraint (and check!) is that any column or row of the composition table of a finite group must contain all elements of the group, and only once. Indeed, suppose that in the row corresponding to $a_{i}$ we had $a_{i} \star a_{j}=a_{i} \star a_{k}$ for some $j, k$. But since $a_{i}$ must have a unique inverse, this would force $a_{j}=a_{k}$. A similar argument can be made for columns.

Another constraint is that $a_{i} \star a_{j}=a_{j}$ only if $a_{i}=e$, and $a_{i} \star a_{j}=a_{i}$ only if $a_{j}=e$, which occurs only in the first row and the first column, respectively. It follows that $a_{i} \star a_{j}=a_{j}$ and $a_{i} \star a_{j}=a_{i}$ are ruled out in the other entries.

Constructing Cayley tables for cyclic groups is easy. Let us do it for $n=2,3$, and 4 :

$$
\begin{array}{c|c}
e & a \\
\hline a & e \\
\{e, a\}
\end{array}
$$



| $e$ | $a$ | $b$ | $c$ |
| :---: | :---: | :---: | :---: |
| $a$ | $b$ | $c$ | $e$ |
| $b$ | $c$ | $e$ | $a$ |
| $c$ | $e$ | $a$ | $b$ |

$$
\left\{e, a, b=a^{2}, c=a^{3}\right\}
$$

Notice that, for $n=2$ and 3 , these tables are the only ones we can construct that meet the two constraints mentioned just above. Therefore, finite groups of order 1,2 and 3 are cyclic. The $n=4$ case, however, would seem to open up more possibilities: if $a \star a=b$, the constraints determine the other entries, and we obtain a table for the cyclic group $Z_{4}$. But we could also take $a \star a=c$ and $a \star a=e$ : In the former case, the constraints determine the rest of the table; in the latter case, we can choose $b \star b=a$ or $b \star b=e$, yielding:

| $e$ | $a$ | $b$ | $c$ |
| :---: | :---: | :---: | :---: |
| $a$ | $c$ | $e$ | $b$ |
| $b$ | $e$ | $c$ | $a$ |
| $c$ | $b$ | $a$ | $e$ |


| $e$ | $a$ | $b$ | $c$ |
| :---: | :---: | :---: | :---: |
| $a$ | $e$ | $c$ | $b$ |
| $b$ | $c$ | $a$ | $e$ |
| $c$ | $b$ | $e$ | $a$ |


| $e$ | $a$ | $b$ | $c$ |
| :---: | :---: | :---: | :---: |
| $a$ | $e$ | $c$ | $b$ |
| $b$ | $c$ | $e$ | $a$ |
| $c$ | $b$ | $a$ | $e$ |

By relabelling $b \longleftrightarrow c$ in the first table, and $a \longleftrightarrow b$ in the second, and re-ordering the rows and columns, we obtain tables which are identical to the cyclic table, and we conclude that in they are really those of the cyclic group of order 4.

The last table is genuinely different. it corresponds to a group of order 4, called the four-group-aka Felix Klein's Vierergruppe $V$ —in which every element is its own inverse (and thus of order 2), with the fourth element constructed out of the other two non-identity elements (otherwise $V$ would be cyclic!): $\{e, a, b, a \star b\}$. An example is the symmetry group, $D_{2}$, of a rectangle centered on the origin: the identity, one rotation by $\pi$ and two reflections its axes.

EXERCISE: Is it possible to construct the Cayley table of a group of order 5 from the requirement that all its elements be of order 2 ?

### 2.1.3 Generators of a group

Cyclic groups and the four-group illustrate a very useful feature of some groups: the fact that all their elements can be generated by a subset of those elements. More precisely:

Definition 2.6. A set of generators of a group $G$ is a subset of $G$ from which all other elements of $G$ can be obtained by repeated compositions of the generators among themselves.

Any element $g$ of a finite group generates a cyclic group $Z_{n}$, where $n$ is the order of $g$, whereas the four-group is obtained from two generators. EXERCISE: construct the Cayley table for the group of order 6: $\left\{e, a, b, b^{2}, a \star\right.$ $b, b \star a\}$.

As another example, we can think of a rotation by $\pi / 6$ about the $z$ axis as the generator of the finite group of rotations by $k \pi / 6(k=0, \ldots, 11)$.

### 2.1.4 Isomorphisms

We have just been introduced to the important idea that groups which look different may in some sense be the same because their composition tables are identical or can be made to be identical by relabelling. We now formalise this idea:

Definition 2.7. If there exists a one-to-one correspondence between all the elements of one group $\{G, \circ\}$ and all the elements of another group $\{H, \star\}$ such that under this mapping the two groups have identical composition tables, then the mapping is said to be an isomorphism, and $G$ and $H$ are isomorphic: $G \cong H$.

Another definition can be more helpful when, as happens with continuous groups, there is no composition table as such:

Definition 2.8. If there exists a one-to-one mapping $f$ between all the elements of one group $\{G, \circ\}$ and all the elements of another group $\{H, \star\}$ such that under this mapping, $f(a), f(b) \in H$ and $f(a \circ b)=f(a) \star f(b) \forall a, b \in G$, then $f$ is an isomorphism of $G$ onto $H$, and $G \cong H$.

Other examples of isomorphic groups are

- the group of permutations of two objects $\left(S_{2}\right)$, the group of rotations by $\pi$ around the $z$ axis, the group $\{1,-1\}$ (under multiplication);
- the group of complex numbers and the group of vectors in a plane, both under addition;
- the continuous groups $\{\mathbb{R},+\}$ and $\left\{\mathbb{R}^{+}, \times\right\}$with the exponential as the isomorphism. Later we will see that because $\mathrm{e}^{x} \mathrm{e}^{y}=\mathrm{e}^{x+y}, \mathrm{e}^{x} \in\left\{\mathbb{R}^{+}, \times\right\}$provides a one-dimensional matrix representation of $\{\mathbb{R},+\}$.

Definition 2.9. A homomorphism, like an isomorphism, preserves group composition, but it is not one-to-one (eg. it could be many-to-one). .

### 2.2 Special Subsets of a Group (BF10.3)

There are a number of useful ways to classify the elements of a given group.

### 2.2.1 Special Ternary Compositions: Conjugacy Classes

Definition 2.10. Given $a \in G$, any element $b \in G$ which can be obtained as $b=x \circ a \circ x^{-1}$, where $x \in G$, is called the conjugate of $a$ by $x$. This conjugation operation, which consists of two binary compositions, is analogous to similarity transformations on matrices.

Now conjugacy has the following properties:

- Reflexivity: $a=e \circ a \circ e^{-1}$, or $a$ is self-conjugate.
- Symmetry: let $b=x \circ a \circ x^{-1}$. Then $a=y \circ b \circ y^{-1}$, with $y=x^{-1} \in G$.
- Transitivity: let $b=x \circ a \circ x^{-1}$ and $a=y \circ c \circ y^{-1}$. Then:

$$
b=x \circ a \circ x^{-1}=x \circ y \circ c \circ y^{-1} \circ x^{-1}=(x \circ y) \circ c \circ(x \circ y)^{-1}
$$

and since $x \circ y \in G, b$ is conjugate to $c$.
This leads to the definition:

Definition 2.11. The subset of elements of a group which are conjugate to one another form a conjugacy, or equivalence ${ }^{\dagger}$, class, often abbreviated to just a class.

The systematic way of constructing the class for any element $a_{i}$ of a group is to form:

$$
\left(e \circ a_{i} \circ e^{-1}, a_{1} \circ a_{i} \circ a_{1}^{-1}, \ldots, a_{i-1} \circ a_{i} \circ a_{i-1}^{-1}, a_{i+1} \circ a_{i} \circ a_{i+1}^{-1}, \ldots\right)
$$

From this it follows that $e$ is always in a class by itself, and that each element of an Abelian group must also be in a class by itself. Since the cyclic group of order $n$ is Abelian, each of its elements is its own class, and the same goes for the four-group.

Classes are disjoint: they have no element in common (EXERCISE: show this).
Elements in the same class share some properties. In particular, they must all be of the same order (EXERCISE). In a particularly important type of group, matrix groups, conjugate matrices are similar to one another; they could represent the same "thing" in different bases.

EXERCISE: obtain the classes for the group: $\left\{e, a, b, b^{2}, a \star b, b \star a\right\}$.
Note that there is another way to form interesting sets of conjugates: $\forall a_{i} \in G$, form set $\left\{x \circ a_{i} \circ x^{-1}\right\}$ for a given $x \in G$.

### 2.2.2 Subgroups

Definition 2.12. A subset $H$ of a group $G$ that behaves as a group in its own right, and under the same composition law as $G$ is said to be a subgroup of $G$ : $H \subseteq G$. A subgroup $H$ of $G$ is proper if it is non-trivial (ie. not $e$ ) and if $H \subset G$ (ie. $H \neq G$ ). Unlike classes, the subgroups of a group are not dsjoint, and can have more elements than $e$ in common.

The four-group $V$ has the (non-disjoint) proper subgroups: $\{e, a\},\{e, b\}$, and $\{e, c=a \star b\}$. By inspection, the group of order $6\left\{e, a, b, b^{2}, a \star b, b \star a\right\}$ contains the proper cyclic subgroup $\left\{e, b, b^{2}\right\}$ of order 3 .

[^8]
## A Small Zoo of Famous (and Important!) Subgroups:

- The set of even integers is a subgroup of $\{\mathbb{Z},+\}$, which is itself a subgroup of $\{\mathbb{R},+\}$, which is a subgroup of $\{\mathbb{C},+\}$. All those subgroups are proper.
- The special subset $S L(n, \mathbb{C})$ of the General Linear group $G L(n, \mathbb{C})$ of transformations represented by $n \times n$ invertible complex matrices which have determinant one.
- The subset $U(n)$ of the General Linear group $G L(n, \mathbb{C})$ of transformations represented by $n \times n$ complex matrices which are unitary $\left(\mathbf{U}^{\dagger} \mathbf{U}=\mathbf{I}\right)$, and whose determinant has the form $\mathrm{e}^{\mathrm{i} \theta} . S U(n)$ is the special subgroup of $U(n)$ matrices with determinant 1 .
- $G L(n, \mathbb{R})$, which has $S L(n, \mathbb{R})$ as a proper subgroup. is a proper subgroup of $G L(n, \mathbb{C})$,
- The subset $O(n)$ of the General Linear group $G L(n, \mathbb{R})$ of transformations represented by $n \times n$ real matrices which are orthogonal $\left(\mathbf{O}^{\mathrm{T}} \mathbf{O}=\mathbf{I}\right)$, and whose determinant is 1 or $-1 . S O(n)$ is the subgroup of $O(n)$ matrices with unit determinant, isomorphic to the group of rotations on $\mathbb{R}^{n}$.


## New Notation:

1. It is now time to wean ourselves from the often cumbersome star (circle) notation for group composition. From now on, we shall omit them whenever there is little risk of confusion with ordinary multiplication.
2. If $H$ and $H^{\prime}$ are two subsets of $\{G, \star\}$, often we shall write $H H^{\prime}=H \star H^{\prime}$ for $\left\{h h^{\prime}=h \star h^{\prime}\right\} \forall h \in$ $H, h^{\prime} \in H^{\prime}$.

Let us try out our new notation on the following definition:
Definition 2.13. A subgroup $N \subseteq G$ is called invariant (or normal) if $N=G N G^{-1}$. More precisely, if $g h g^{-1} \in N \forall h \in N$ and $\forall g \in G, N$ is an invariant subgroup of $G$. Alternate notation: $N \triangleleft G, G \triangleright N$.

From this definition, we see that any invariant subgroup must be self-conjugate. As well, the identity and $G$ itself are always invariant subgroups. EXERCISE: Show that the set $\left\{g_{i}^{-1} g_{j}^{-1} g_{i} g_{j}\right\} \forall g_{i}, g_{j} \in G$, keeping only distinct objects, forms an invariant subgroup of $G$.

Definition 2.13 is sometimes written $G N=N G$, but it does not mean that an invariant subgroup must be Abelian (although it can be). Rather, it means that if you compose any element of $N$ with any group element from the left, there is some element of $N$ which, when composed with the group element from the right, gives the same result.

Definition 2.14. The centre of a group is the set of all its elements that commute with all elements of the group.

Example 2.1. From the composition table for the four-group, it is straightforward to work out that each of its non-trivial subgroups, $\{e, a\},\{e, b\},\{e, a b\}$, are invariant. Indeed, the four-group is Abelian, and all the subgroups of an Abelian group are invariant. Like all Abelian groups, the fourgroup is its own (trivial) centre.

A very powerful method for finding invariant subgroups is based on the following lemma : Let $H \subset G$. Then $H$ is invariant if and only if it contains complete classes, ie. if it is a union of classes of $G$. Indeed, if $H$ is invariant, all the conjugates (elements in the same class) of any $h \in H$ are also in $H$, and this holds for all classes, which are disjoint; so only complete classes can be in $H$. Conversely, let a subgroup $H$ of $G$ be a union of complete classes; this means that the conjugates of all elements of $H$ are in $H$, or $g h g^{-1} \in H \forall g \in G$, which is precisely the definition of an invariant subgroup.

Definition 2.15. A simple group has no invariant subgroup other than itself and the identity.

### 2.2.3 Cosets and factor groups (BF 10.3)

Definition 2.16. Let $H$ be a subgroup of $G$, and let $g \in G$. Then $g H$ is a left coset of $H$ for a given $g$, and $H g$ is a right coset of $H$. The set of all left (right) cosets of $H$ is called the left (right) coset space of $H$.

Every coset $g H$ of $H$ must contain the same number of elements, equal to the order of $H$.
If $H$ is invariant, to any ot its left cosets corresponds an identical right coset, and vice-versa, as follows immediately from Def. 2.13. In particular, the right and left cosets of any Abelian subgroup are identical.

Example 2.2. Let $G=\mathbb{R}^{3}$ under addition, and $H$ be a plane containing the origin. For a given vector $\mathbf{a}, \mathbf{a}+H \in H$ if $\mathbf{a} \in H$; otherwise, $\mathbf{a}+H$ is another plane $H^{\prime}$ parallel to $H$, and we would say in this language that it is a left (or right) coset of the plane through the origin. And $H$ itself would also be a coset.

The most important property of cosets is that they are either disjoint or else identical.
Indeed, let $g_{1} h_{1}=g_{2} h_{2}$ for some $h_{1}, h_{2} \in H$ and $g_{1}, g_{2} \in G$. Therefore, $g_{1}=g_{2} h_{2} h_{1}^{-1}$. Now consider some other element of the same coset, $g_{1} h_{3}\left(h_{3} \in H\right)$; then $g_{1} h_{3}=g_{2}\left(h_{2} h_{1}^{-1} h_{3}\right)=g_{2} h_{4}$, where $h_{4}=$ $h_{2} h_{1}^{-1} h_{3} \in H$. That is, if two elements of different cosets are the same, then any other element, say $g h_{3}$, in the first coset, must be equal to some element of the second coset. Since the same argument holds when we switch $g_{1}$ and $g_{2}$, we conclude that if $g_{1} H$ and $g_{2} H$ have one element in common, they have all their elements in common and are thus identical. The same proof applies to right cosets. Thus, we can say that the coset space of a subgroup $H \subset G$ provides a partition of $G$.

An immediate consequence of this property is that since $e H=H$ is a left coset of a subgroup $H$, that coset contains $e$ and is the only coset that is a group. No other coset of $H$ can be a group since it cannot contain $e$. For any other element $h$ of subgroup $H$ that is not the identity, we still have $h H=H$ because $H$ is a group, and all cosets $h H=H$ are in fact the same one, ie. $H$ itself.

If $H \subset G$, every element of $G$ must occur either in $H$ or one (and only one) of its other cosets.This forms the foundation of the proof of Lagrange's Theorem: The order $n$ of a finite group is an integer multiple of the order $m$ of any of its subgroups. Indeed, since every element of the group is either in the subgroup or in one of its other $k$ distinct cosets, each with $m$ elements, $(k+1) m=n$. The ratio $n / m$ is called the index of the subgroup.

Let $a \in G$. Clearly, it generates a cyclic subgroup of $G$ of order $m$ : $\left\{e, a, a^{2}, \ldots, a^{m-1}\right\}$, where $m \leq n$ is the order of $a$. Therefore, the order $n$ of $G$ must be an integer multiple of the order $m$ of any of its elements. If $n$ is prime, $m=n$ or $m=1$, and we have proved that the only non-trivial finite group of order prime is the cyclic group.

Now consider the set whose elements are the subgroup as a whole and all its other cosets, each also as a whole:
Definition 2.17. A factor space for $H \subset G$ is the set of all left cosets of $H$. Note that the elements of this space are the cosets themselves, each considered as a whole.

Factor spaces of a subgroup $H$ are not necessarily groups; but there is one important exception:

Definition 2.18. To an invariant subgroup $N$ of $G$ is associated a factor group of $G, G / N$, whose elements are $N$ and all its cosets. Again, $N$ and its cosets themselves are elements of the factor group, not the elements of $N$ or of the cosets. The order of $G / N$ is the order of $G$ divided by the order of $N$, ie. the index of $N$, hence the name quotient group often applied to $G / N$.

To show that the factor space of an invariant subgroup is a group, we note that for any coset $g N$,

$$
(g N) N=g N N=g N ; \quad N(g N)=g N N=g N
$$

where we have used the associativity of the group product and the invariant nature of $N$. This establishes that $N$ must be the identity of the factor group. The composition law follows from:

$$
\left(g_{1} N\right)\left(g_{2} N\right)=g_{1} g_{2} N N=\left(g_{1} g_{2}\right) N
$$

since $g N=N g \forall g \in G$, Lastly,

$$
(g N)\left(g^{-1} N\right)=g g^{-1} N N=N=e
$$

So $g^{-1} N$ is the inverse of $g N$.
Factor groups can be useful when, for one reason or another, we do not need to distinguish between the elements of subgroups of a group.

### 2.2.4 Direct Products

Definition 2.19. Let $H_{1}$ and $H_{2}$ be subgroups of $G$ which have only the identity $e$ as common element, and let $h_{1} h_{2}=h_{2} h_{1} \forall h_{1} \in H_{1}, \forall h_{2} \in H_{2}$. If it is possible to write $g=h_{1} h_{2} \forall g \in G$, then $G \equiv H_{1} \times H_{2}$ is said to be the internal direct product of its subgroups $H_{1}$ and $H_{2}$. The latter are invariant (EXERCISE).

Example 2.3. We can think of $O(3)$ as the direct product of $S O(3)$ and the subgroup consisting of the identity matrix and the reflection matrix - $\mathbf{I}$. This latter group is obviously Abelian and, since its elements commute with all 3 -dim rotations (ie. the elements of $S O(3)$ ), it is an invariant subgroup and the centre of $O(3)$. Also, the four-group introduced in section 2.1 .2 can be seen as $Z_{2} \times Z_{2}$, or $\{e, a\} \times\{e, b\}=\{e, a, b, a b\}$. However, $Z_{4} \neq Z_{2} \times Z_{2}$, even though $Z_{2}$ is an invariant subgroup of $Z_{4}$, and thus $Z_{2}=Z_{4} / Z_{2}$ !.

Another well-known way of constructing a (this time, external) direct product of, say, two a priori unrelated matrix groups with elements $\mathbf{A} \in H_{1}$ and $\mathbf{B} \in H_{2}$ would be:

$$
\left(\begin{array}{ll}
A & 0 \\
0 & B
\end{array}\right)
$$

Or we could construct $\{1,-1\} \times\{1,-1\}=\{(1,1),(1,-1),(-1,1),(-1,-1)\}$. the external direct product of $Z_{2}$ with itself, in this realisation. This, of course, is the four-group.

### 2.3 The Mother of All Finite Groups: the Group of Permutations

### 2.3.1 Definitions, cycles, products

The most important finite group is the group of permutations of $n$ objects, $S_{n}$, aka the symmetric group, which contains $n$ ! elements corresponding to the $n$ ! possible rearrangements of the objects. A permutation is by definition a bijective mapping. Following a standard convention, we notate, with $1 \leq k \leq n$ !:

Definition 2.20.

$$
\pi_{k}=\left(\begin{array}{ccccc}
1 & 2 & 3 & \ldots & n \\
\pi_{k}(1) & \pi_{k}(2) & \pi_{k}(3) & \ldots & \pi_{k}(n)
\end{array}\right)
$$

The horizontal ordering of the initial objects is immaterial. Also as a matter of convention, we agree that it is the objects in the slots which are rearranged, not the slots. It is quite possible to define $\pi_{k}(j)$ as the mapping of whatever object happens to sit in slot $j$; but, again following standard usage, eg. in BF, we define $\pi_{k}(j)$ as the mapping, in a given slot, of the object labelled by $j$ to another object labelled by a number between 1 and $n$. Finally, we do not have to use numbers as labels, but they offer the greatest range.

In a permutation, an object $i$ may be mapped into itself, ie. it stays in the same slot. But more typically object $i$ is mapped to $j$, while $j$ is mapped to $k$; and so on along a chain that ends back at object $a$ after $l$ steps. When this occurs, we speak of a $l$-cycle. More precisely:

Definition 2.21. Let $\pi_{k} \in S_{n}$, and let $l$ be the smallest integer for which $\left[\pi_{k}(j)\right]^{l}=j$, for some $1 \leq j \leq n$. Then the sequence of objects in $\left[\pi_{k}(j)\right]^{l}$ is called a $l$-cycle (sometimes a $r$-cycle...).

This suggests a much more compact notation for $\pi_{k}$, one in which we bother to write only the $l$-cycles $(l>1)$, and consider a given permutation as the product of simpler permutations.

As an example, we write:

$$
\left(\begin{array}{llllll}
1 & 2 & 3 & 4 & 5 & 6 \\
5 & 4 & 2 & 3 & 1 & 6
\end{array}\right)=\left(\begin{array}{cccccc}
1 & 2 & 3 & 4 & 5 & 6 \\
5 & 2 & 3 & 4 & 1 & 6
\end{array}\right)\left(\begin{array}{cccccc}
1 & 2 & 3 & 4 & 5 & 6 \\
1 & 4 & 2 & 3 & 5 & 6
\end{array}\right)\left(\begin{array}{llllll}
1 & 2 & 3 & 4 & 5 & 6 \\
1 & 2 & 3 & 4 & 5 & 6
\end{array}\right) \equiv(15)(243)
$$

It is easy to see the advantages of the cycle notation introduced at the end of the line! Note that the cycles are disjoint. Any permutation can be, and usually is, represented by a sequence of disjoint cycles.

Insight: if one imagines $n$ states of a system being evolved through discrete time increments by successive applications of $S_{n}$, the cyclic structure means that not all states are accessible from a given initial state, only those allowed by the cycle to which the initial state belongs. The time evolution of the system is therefore constrained. We could assign some quantity that takes some value corresponding to each particular cycle. Then we could say that this quantity is conserved, in the sense that the time evolution of an initial state is constrained to states that have the same value of the conserved quantity, because they are the only ones accessible from the initial state. Perhaps this gives us an inkling of a connection between groups and conservation laws.

Any $\pi_{k} \in S_{n}$ can always be written as the product ${ }^{\dagger}$ of transpositions, or two-cycles. Indeed, a $l$-cycle may always be decomposed as a product of $l-1$ transpositions, but these are not disjoint.

Definition 2.22. A permutation is even (odd) if it is equivalent to an even (odd) number of transpositions, or switches; thus, a $l$-cycle which contains an even number of symbols is equivalent to an odd permutation, and vice-versa. An even permutation is said to have parity 1 , and an odd permutation parity -1 . We expect that parity will put strong constraints on the group product table of $S_{n}$.

Transpositions always have odd parity. The mapping from $S_{n}$ to the parities $\{1,-1\}$ is a nice example of a homomorphism.

Definition 2.23. A permutation with only one cycle of length $l>1$ is a cyclic permutation of length $l$.

In cycle notation, $S_{2}=\{e,(12)\}$ and $S_{3}=\{e,(12),(13),(23),(132),(123)\} \equiv\left\{\pi_{1}, \pi_{2}, \pi_{3}, \pi_{4}, \pi_{5}, \pi_{6}\right\}$, are the smallest non-trivial symmetric groups. For $S_{3}$, note the three-cycles $(123)=\left(\begin{array}{ll}1 & 2 \\ 2 & 3\end{array}\right)$ and $(132)=$ $\left(\begin{array}{lll}1 & 2 & 3 \\ 3 & 1 & 2\end{array}\right)$. I have deliberately changed the order of the latter from what it is in BF, but if you write out the corresponding permutation in full notation for BF's (321), you will see that it is identical to mine. So long as we cycle through in the same direction (here, to the right), the order of the elements in a cycle does not matter!

Warning: do not confuse the symbols in a 3 -cycle with the outcome of a permutation in $S_{3}$ !

### 2.3.2 Subgroups of $S_{n}$

One obvious subgroup of $S_{n}$ is the so-called alternating group, $A_{n}$, of all its even permutations. Odd permutations cannot form a group, because their product is always an even permutation. Another, less obvious, but important subgroup of $S_{n}$ is the cyclic group of order $n$, generated by the permutation (123 $\ldots n$ ), which returns the initial state to itself after $n$ products.

[^9]Now for subgroups of $S_{3}$ : Lagrange's Theorem allows only non-trivial proper subgroups of order 2 or 3. The alternating subgroup $A_{3}$ is read off the list of the elements of $S_{3}$ : \{e, (132), (123), \} which must be cyclic because all groups of order 3 are isomorphic to $Z_{3}$. Note: this is not a general feature as the cyclic subgroups of higher order generated by odd permutations in $S_{n>3}$ contain permutations of both even and odd parity.

The product of a transposition by itself is the identity, so $\pi_{2} \pi_{2}=\pi_{3} \pi_{3}=\pi_{4} \pi_{4}=e$. Then the other (isomorphic!) subgroups of $S_{3}$ are $\left\{e, \pi_{2}\right\},\left\{e, \pi_{3}\right\}$, and $\left\{e, \pi_{4}\right\}$. The centre of $S_{3}$ —and of $S_{n>3}$ - is just $e$.

### 2.3.3 Cayley table of $S_{n}$

The group-product table of $S_{3}$ contains 36 entries, "only" 25 of which are non-trivial, from which we have just found three. But I claim that no more than one other product needs to be worked out with the explicit form of the permutations; the rest can all be found by a bit of astute reasoning.

The entries of the $2 \times 2$ sub-table for rows and columns corresponding to $\pi_{5}$ and $\pi_{6}$ are determined by the fact that they must be even permutations (since they are the group product of even permutations). The diagonals cannot be $e$; if they were, the other entry in the same row would be the same as the the first entry of its column. Alternatively, $\pi_{5}$ and $\pi_{6}$ being the only non-trivial elements in $A_{3}$, they must be each other's inverse.

Next, the non-diagonal elements of rows and columns corresponding to $\pi_{2}, \pi_{3}$ and $\pi_{4}$ must be either $\pi_{5}$ or $\pi_{6}$, the only even permutations other than $e$. To fill in this sector only requires calculating one group product, say, $\pi_{2} \pi_{3}$ :

$$
\pi_{2} \pi_{3}=\left(\begin{array}{lll}
1 & 2 & 3 \\
2 & 1 & 3
\end{array}\right)\left(\begin{array}{lll}
1 & 2 & 3 \\
3 & 2 & 1
\end{array}\right)=\left(\begin{array}{lll}
1 & 2 & 3 \\
3 & 2 & 1 \\
3 & 1 & 2
\end{array}\right)=\left(\begin{array}{lll}
1 & 2 & 3 \\
3 & 1 & 2
\end{array}\right)=\pi_{5}
$$

The other unfilled entries in rows and columns for $\pi_{5}$ and $\pi_{6}$ must be either $\pi_{2}, \pi_{3}$, or $\pi_{4}$. For columns $\pi_{5}$ and $\pi_{6}$, applying $\pi_{2}$ to $\pi_{2} \pi_{3}$ gives $\pi_{3}=\pi_{2} \pi_{5}$, which determines the rest from the general table-building rules. Similarly, $\pi_{2} \pi_{3} \pi_{3}=\pi_{2}=\pi_{5} \pi_{3}$, and the other entries in the $\pi_{5}$ and $\pi_{6}$ rows are determined. Here is the final result in two equivalent forms:

| $e$ | $\pi_{2}$ | $\pi_{3}$ | $\pi_{4}$ | $\pi_{5}$ | $\pi_{6}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\pi_{2}$ | $e$ | $\pi_{5}$ | $\pi_{6}$ | $\pi_{3}$ | $\pi_{4}$ |
| $\pi_{3}$ | $\pi_{6}$ | $e$ | $\pi_{5}$ | $\pi_{4}$ | $\pi_{2}$ |
| $\pi_{4}$ | $\pi_{5}$ | $\pi_{6}$ | $e$ | $\pi_{2}$ | $\pi_{3}$ |
| $\pi_{5}$ | $\pi_{4}$ | $\pi_{2}$ | $\pi_{3}$ | $\pi_{6}$ | $e$ |
| $\pi_{6}$ | $\pi_{3}$ | $\pi_{4}$ | $\pi_{2}$ | $e$ | $\pi_{5}$ |


$\equiv$| $e$ | $\pi_{5}$ | $\pi_{6}$ | $\pi_{2}$ | $\pi_{3}$ | $\pi_{4}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\pi_{5}$ | $\pi_{6}$ | $e$ | $\pi_{4}$ | $\pi_{2}$ | $\pi_{3}$ |
| $\pi_{6}$ | $e$ | $\pi_{5}$ | $\pi_{3}$ | $\pi_{4}$ | $\pi_{2}$ |
| $\pi_{2}$ | $\pi_{3}$ | $\pi_{4}$ | $e$ | $\pi_{5}$ | $\pi_{6}$ |
| $\pi_{3}$ | $\pi_{4}$ | $\pi_{2}$ | $\pi_{6}$ | $e$ | $\pi_{5}$ |
| $\pi_{4}$ | $\pi_{2}$ | $\pi_{3}$ | $\pi_{5}$ | $\pi_{6}$ | $e$ |

### 2.3.4 Cayley's Theorem

Why is $S_{n}$ so important? As so often, the Cayley table of a group $G$ of order $n$ gives the key to the answer. $\forall a_{i} \in G$, the row $\left\{a_{i} a_{j}\right\}(1 \leq j \leq n)$ is merely a bijective rearrangement of $\left\{a_{i}\right\}$, that is:
$a_{i} \longmapsto \pi_{a_{i}}=\left(\begin{array}{cccc}a_{1} & a_{2} & \ldots & a_{n} \\ a_{i} a_{1} & a_{i} a_{2} & \ldots & a_{i} a_{n}\end{array}\right), \quad \quad a_{i} a_{j} \longmapsto \pi_{a_{i} a_{j}}=\left(\begin{array}{cccc}a_{1} & a_{2} & \ldots & a_{n} \\ a_{i} a_{j} a_{1} & a_{i} a_{j} a_{2} & \ldots & a_{i} a_{j} a_{n}\end{array}\right)$
But we can also write:

$$
\left.\begin{array}{rl}
\pi_{a_{i}} & =\left(\begin{array}{cccc}
a_{1} & a_{2} & \ldots & a_{n} \\
a_{i} a_{1} & a_{i} a_{2} & \ldots & a_{i} a_{n}
\end{array}\right)=\left(\begin{array}{ccc}
a_{j} a_{1} & a_{j} a_{2} & \ldots \\
a_{i}\left(a_{j} a_{1}\right) & a_{i}\left(a_{j} a_{2}\right) & \ldots \\
a_{j} a_{n} \\
a_{i}\left(a_{j} a_{n}\right)
\end{array}\right) \\
& =\left(\begin{array}{ccccc}
a_{j} a_{1} & a_{j} a_{2} & \ldots & a_{j} a_{n} \\
a_{i} a_{j} a_{1} & a_{i} a_{j} a_{2} & \ldots & a_{i} a_{j} a_{n}
\end{array}\right)\left(\begin{array}{cccc}
a_{1} & a_{2} & \ldots & a_{n} \\
a_{j} a_{1} & a_{j} a_{2} & \ldots & a_{j}
\end{array}\right) \\
a_{n}
\end{array}\right)
$$

What we have shown is that $\pi_{a_{i}} \pi_{a_{j}}=\pi_{a_{i} a_{j}}$; in other words, by definition 2.8, permutations preserve the group product of $G$, and we have Cayley's Theorem:

Every group of order $n$ is isomorphic to a subgroup of $S_{n}$ whose elements (except for e) shuffle all objects in the set on which it acts. By Lagrange's theorem, the index of a finite group as a subgroup of $S_{n}$ is $n!/ n=(n-1)$ !.

We have already seen an example of this: the cyclic group of order 3 is a subgroup of $S_{3}$, and there is only one instance of $Z_{3} \subset S_{3}$. EXERCISE: How many distinct instances of $Z_{4} \subset S_{4}$ are there? How many of the four-group?

Example 2.4. Another interesting example of a relation between a group and $S_{n}$ is a permutation $\pi \in S_{3}$ that defines an action of $S_{3}$ on $\mathbb{R}^{3}$ which permutes the coordinates $x^{i}$ of a vector, $x^{i} \pi=x^{\pi_{i}}$. We can define an action of $\pi$ from the right such that, for instance, $\left(x_{1}, x_{2}, x_{3}\right) \star \pi_{5}=\left(x_{3}, x_{1}, x_{2}\right)$. Since we act from the right ${ }^{\dagger}$, the vectors are columns; with implicit summation over $j$, this action can then be represented by a matrix $\mathbf{T}(\pi)$ :

$$
x^{i} \pi_{5}=T^{i}{ }_{j}\left(\pi_{5}\right) x^{j} \quad T^{i}{ }_{j}\left(\pi_{5}\right)=\left(\begin{array}{ccc}
0 & 0 & 1 \\
1 & 0 & 0 \\
0 & 1 & 0
\end{array}\right), \quad x^{i} \pi_{6}=T^{i}{ }_{j}\left(\pi_{6}\right) x^{j} \quad T^{i}{ }_{j}\left(\pi_{6}\right)=\left(\begin{array}{ccc}
0 & 1 & 0 \\
0 & 0 & 1 \\
1 & 0 & 0
\end{array}\right)
$$

Note that $\pi_{5} \pi_{6}=e$, so that $\pi_{5}$ and $\pi_{6}$ are the inverse of one another. This is reflected in the $\mathbf{T}$ matrices: they are each other's inverse and transpose, and they have determinant +1 , so they are elements of $S O(3)$. EXERCISE: Is this true of $T^{i}{ }_{j}\left(\pi_{2,3,4}\right)$ ?

### 2.3.5 Conjugates and Classes of $S_{n}$

To find the classes of $S_{n}$, we must form, for each $\pi_{i} \in S_{n}$, all its conjugates $\pi_{j} \pi_{i} \pi_{j}^{-1}$. This seemingly daunting task can actually be performed fairly easily, thanks to the nature of $S_{n}$. To keep the following manipulations as uncluttered as possible, let us write $\pi_{i}=a$ and $\pi_{j}=b$, with $a=\left(\begin{array}{cccc}1 & 2 & \cdots & n \\ a_{1} & a_{2} & \ldots & a_{n}\end{array}\right)$ and $b=\left(\begin{array}{cccc}1 & 2 & \ldots & n \\ b_{1} & b_{2} & \ldots & b_{n}\end{array}\right)$. Then:

$$
\begin{align*}
b a b^{-1} & =\left(\begin{array}{cccc}
1 & 2 & \ldots & n \\
b_{1} & b_{2} & \ldots & b_{n}
\end{array}\right)\left(\begin{array}{cccc}
1 & 2 & \ldots & n \\
a_{1} & a_{2} & \ldots & a_{n}
\end{array}\right)\left(\begin{array}{cccc}
b_{1} & b_{2} & \ldots & b_{n} \\
1 & 2 & \ldots & n
\end{array}\right)=\left(\begin{array}{cccc}
1 & 2 & \ldots & n \\
b_{1} & b_{2} & \ldots & b_{n}
\end{array}\right)\left(\begin{array}{cccc}
b_{1} & b_{2} & \ldots & b_{n} \\
a_{1} & a_{2} & \ldots & a_{n}
\end{array}\right) \\
& =\left(\begin{array}{cccc}
a_{1} & a_{2} & \ldots & a_{n} \\
b_{a_{1}} & b_{a_{2}} & \ldots & b_{a_{n}}
\end{array}\right)\left(\begin{array}{cccc}
b_{1} & b_{2} & \ldots & b_{n} \\
a_{1} & a_{2} & \ldots & a_{n}
\end{array}\right)  \tag{2.1}\\
& =\left(\begin{array}{cccc}
b_{1} & b_{2} & \ldots & b_{n} \\
b_{a_{1}} & b_{a_{2}} & \ldots & b_{a_{n}}
\end{array}\right)
\end{align*}
$$

How did we obtain $\left(\begin{array}{cccc}a_{1} & a_{2} & \ldots & a_{n} \\ b_{a_{1}} & b_{a_{2}} & \ldots & a_{a_{n}}\end{array}\right)$ in the second line from $\left(\begin{array}{cccc}1 & 2 & \ldots & n \\ b_{1} & b_{2} & \ldots & b_{n}\end{array}\right)$ in the last member of the first line? Well, $a_{1}$ must occur in some slot on the top line of the latter; since the order of the slots in the permutation is arbitrary, we move that slot to first position and rename the upper element $a_{1}$. Then we do the same for $2 \rightarrow a_{2}$, etc. The bottom elements are then the outcome of permuting $a_{i}$ with permutation $b$ to get $b_{a_{i}}$.

Something very important can be deduced from this result: All permutations in a class have the same cycle structure, not only for $S_{n}$, but for all finite groups because of Cayley's theorem. Since classes are disjoint, classifying the elements of $S_{n}$ according to their cycle structure also uniquely yields the classes of $S_{n}$ ! In groups other than $S_{n}$, although all elements in a class have the same cycle structure, elements with the same cycle structure may belong to different classes (eg. $A_{4} \subset S_{4}, Z_{4} \subset S_{4}$ )!

Take $S_{3}$ as a simple example. As classes we only have $\mathcal{C}_{1}=\{e\}, \mathcal{C}_{2}=\{(12),(13),(23)\}$, and $\mathcal{C}_{3}=$ $\{(132),(123)\}$. Thus, because it is the only subgroup of $S_{3}$ (apart form $\{e\}$ and $S_{3}$ itself) that is the sum of complete classes, and whose order divides the order of $S_{3}(6), \mathcal{C}_{1}+\mathcal{C}_{3}=\{e,(123),(132)\}=A_{3}$ is the only invariant ${ }^{\ddagger}$ proper subgroup of $S_{3}$.

[^10]Now consider $S_{4}$. There are two other permutations with the same cycle structure as $(12)(34):(13)(24)$ and $(14)(23)$. Apart from this and the separate class $\{e\}$, the other classes of $S_{4}$ are easily obtained as (12) and its 5 conjugate transpositions, (123) and its seven conjugates, and (1234) and its five conjugates. All five classes are of course disjoint.

In the literature, classes of $S_{n}$ are routinely identified by partitions of $n$ reflecting their cycle structure. Thus, a given class will be written $\left(i^{\alpha_{i}} \ldots j^{\alpha_{j}}\right)$, with $(1 \leq(i, j) \leq n)$, where $\alpha_{i}$ is the number of $i$-cycles in the class.

Start with $e$, which is always a class of $S_{n}$. Its cycle structure can be written as a product of $n 1$-cycles: $e=(1)(2) \cdots(n)$. So the class would be denoted in this notation by $1^{n}$. A transposition must have one 2 -cycle and $n-21$-cycles, and $S_{n}$ must contain $n(n-1) / 2$ of them (eg., six for $S_{4}$ as above); it is denoted by $\left(21^{n-2}\right)$. An arbitrary permutation involves $\alpha_{i} i$-cycles, and $\sum_{i}^{n} i \alpha_{i}=n$. It is in that sense that we say that the cycle structure of a class corresponds to a partition of $n$.

Once we have noticed this correspondence, it becomes rather easy to find how many classes there are in $S_{n}$ and what their cycle structure is. We adopt the usual convention that represents the cycle structure of a class by $\left(\lambda_{1} \geq \lambda_{2} \geq \cdots \geq \lambda_{n}\right)$, where the $\lambda_{i}$ must sum up to $n$. Thus, the only possible partitions of $S_{3}$ would correspond to classes $\left(1^{3}\right),(21)$, and (3), ie. a class with three 1-cycles (the identity), a class with one 2 -cycle and one 1-cycle (the transpositions), and a class with one 3 -cycle. As for $S_{4}$, the only possible partitions of 4 give rise to the five classes $\left(1^{4}\right),\left(21^{2}\right),\left(2^{2}\right),(31)$, and (4).

It is important not to confuse the cycle notation we first introduced, which is suitable for individual permutations, with this standard notation which lists all the cycles in a class as a whole, including 1-cycles when they occur (whereas the other one ignores them).

The number of elements in a class of $S_{n}$ is the number of distinct ways of partitioning $n$ numbers into the cycle structure of the class:

$$
\begin{equation*}
\frac{n!}{\alpha_{1}!\ldots \alpha_{n}!1^{\alpha_{1}} \ldots n^{\alpha_{n}}} \tag{2.2}
\end{equation*}
$$

where $\alpha_{i}$ ! is the number of non-distinct ways of ordering $\alpha_{i}$ commuting cycles of a given length, and $i^{\alpha_{i}}$ is the number of equivalent orderings of the symbols inside each $i$-cycle occurring $\alpha^{i}$ times. From this expression it should be easy to recover the number of elements in each class of $S_{4}$ as given above.

Now we can identify (EXERCISE) the invariant subgroups of $S_{4}$ without writing down its $24 \times 24$ Cayley table!

### 2.3.6 Graphical representation of classes: Young frames

A graphical way of representing the classes of $S_{n}$ is to take $n$ squares and arrange them in rows and columns so that each column corresponds to an $i$-cycle and the number of boxes cannot increase from one column to the next on the right, and from one row to the one next below. The game then consists in building all possible arrangements that satisfy this constraint. For instance, with $S_{4}$, the possibilities are as follows:


Then we just read off the cycle structure for each: $\left(1^{4}\right),\left(21^{2}\right),\left(2^{2}\right),(31)$, and (4), respectively. Suddenly, finding the classes of such monsters as, say $S_{8}$, no longer seems so intimidating. These diagrams are known as Young frames.

### 2.3.7 Cosets of $S_{n}$

Now that we know the Cayley table of $S_{3}$, finding the left cosets of its subgroups is as easy as reading column elements in the table. Take the subgroup $\left\{e, \pi_{2}\right\}$; its cosets are $\pi_{k}\left\{e, \pi_{2}\right\}=\left\{\pi_{k}, \pi_{k} \pi_{2}\right\}(1 \leq k \leq 6)$. So each coset consists of the first entry and the entry in the $\pi_{2}$ column in each row of the table. There are thus only three distinct cosets: $\left\{e, \pi_{2}\right\},\left\{\pi_{3}, \pi_{6}\right\},\left\{\pi_{4}, \pi_{5}\right\}$. As stated earlier in Definition 2.17, this set of cosets is the factor space for $\left\{e, \pi_{2}\right\} \subset S_{3}$, which can be written as the partition $e\left\{e, \pi_{2}\right\}+\pi_{3}\left\{e, \pi_{2}\right\}+\pi_{4}\left\{e, \pi_{2}\right\}$. But the
right cosets (EXERCISE) are not the same as the left cosets, as expected since $\left\{e, \pi_{2}\right\}$ is not invariant. The same arguments apply to the subgroups $\left\{e, \pi_{3}\right\}$ and $\left\{e, \pi_{4}\right\}$.

Turn now to the remaining non-trivial proper subgroup of $S_{3}, A_{3}=\left\{e, \pi_{5}, \pi_{6}\right\}$, of all even permutations in $S_{3}$. Its left cosets are $\left\{\pi_{k}, \pi_{k} \pi_{5}, \pi_{k} \pi_{6}\right\}$. For instance, by inspection of the Cayley table of $S_{3}, \pi_{2}\left\{e, \pi_{5}, \pi_{6}\right\}=$ $\left\{\pi_{2}, \pi_{3}, \pi_{4}\right\}$, which is identical to the other cosets $\pi_{3}\left\{e, \pi_{5}, \pi_{6}\right\}$ and $\pi_{4}\left\{e, \pi_{5}, \pi_{6}\right\}$. Also, $e\left\{e, \pi_{5}, \pi_{6}\right\}=$ $\pi_{5}\left\{e, \pi_{5}, \pi_{6}\right\}=\pi_{6}\left\{e, \pi_{5}, \pi_{6}\right\}$, as expected. So another partition of $S_{3}$ is provided by $\left\{e, \pi_{5}, \pi_{6}\right\}+\pi_{2}\left\{e, \pi_{5}, \pi_{6}\right\}$. Note that the left and right cosets are now identical, another way of saying that $\left\{e, \pi_{5}, \pi_{5}\right\}$ is invariant, as we had found through simpler means. Then $\left\{\left\{e, \pi_{5}, \pi_{6}\right\},\left\{\pi_{2}, \pi_{3}, \pi_{4}\right\}\right\}$ is the factor group of $S_{3}$. From the Cayley table for $S_{3}$, we see that the element $\left\{e, \pi_{5}, \pi_{6}\right\}$ is the identity, and that this factor group $S_{3} / A_{3}$ is in fact isomorphic to $Z_{2}$ and can be identified with it. It is easy to show that $Z_{2}$ is a factor group of $S_{n} \forall n$. Equivalently, $A_{n}$ is always a normal subgroup of $S_{n}$.

### 2.4 Representations of Groups

We have already mentioned that groups can be associated with symmetries, but we have to make this connection explicit in the language of group theory. We wish to flesh out the rather abstract ideas and tools we have introduced. We shall find that linear operators on vector spaces (most often, on a Hilbert space) provide us with this connection.

### 2.4.1 What is a symmetry?

Let $G$ be a group of linear transformations that act on some $\mathbf{x} \in \mathbb{R}^{n}$. Let us also give ourselves functions $f(\mathbf{x})$ that are square-integrable, ie., that live in a Hilbert space, $\mathcal{L}^{2} \equiv \mathcal{H}$, which we call the carrier space. Consider a linear operator $A_{\mathbf{x}}$ such that, $\forall f \in \mathcal{H},\left[A_{\mathbf{x}} f\right](\mathbf{x})=h(\mathbf{x})$, where $h \in \mathcal{H}$. Introduce a set of operators, $\left\{T_{\mathrm{g}}\right\}$, with each $T_{\mathrm{g}}$ corresponding to a group element $g \in G$, that like $A_{\mathbf{x}}$ act on those functions $f$ themselves (not on $\mathbf{x}!$ ). Then we transform $A_{\mathrm{x}}$ under $G$ in the following way: $T_{\mathrm{g}} A_{\mathrm{x}} T_{\mathrm{g}}^{-1}$.

Definition 2.24. When $T_{\mathrm{g}} A_{\mathbf{x}} T_{\mathrm{g}}^{-1}=A_{\mathbf{x}}, \forall \mathrm{g} \in G, A_{\mathbf{x}}$ is said to be invariant under the action of the group. If also $\left[T_{\mathrm{g}} f\right](\mathbf{x})=f(\mathbf{x}), f$ is invariant under $G$ itself as well. (eg., $f(r)$ in polar coordinates under rotations)

Since the condition for invariance can also be written as $T_{\mathrm{g}} A_{\mathrm{x}}=A_{\mathrm{x}} T_{\mathrm{g}}, \forall \mathrm{g} \in G$, then an operator that is invariant under a group of operators must commute with all the operators in that group.

If $A_{\mathbf{x}}$ has eigenvalues and eigenfunctions. and if it is invariant under $G$, then there should exist a set of functions $\left\{f^{i}\right\}$ such that:

$$
\left[A_{\mathbf{x}} T_{\mathrm{g}} f^{i}\right](\mathbf{x})=\left[T_{\mathrm{g}} A_{\mathbf{x}} f^{i}\right](\mathbf{x})=\lambda_{\mathrm{g}}\left[T_{\mathrm{g}} f^{i}\right](\mathbf{x})
$$

which says that if $f$ is an eigenfunction of $A_{\mathbf{x}}$, so is $T_{\mathrm{g}} f^{i}$, with the same eigenvalue. If the eigenvalue is nondegenerate, ie. if $f$ is unique, then $T_{\mathrm{g}} f^{i}$ must be proportional to $f^{i}$, ie., $f^{i}$ is also an eigenfunction of $T_{\mathrm{g}}$, but with some a priori different eigenvalue also depending on $g$. In the degenerate case, however, given $N$ eigenfunctions, all we can say is that the $T_{g} f^{i}$ are linear combinations of $\left\{f^{j}\right\}$ :

$$
\begin{equation*}
T_{\mathrm{g}} f^{i}=f^{j} D_{j}{ }^{i}(\mathrm{~g}) \tag{2.3}
\end{equation*}
$$

with summation up to $N$ over repeated indices implied.
Definition 2.25. Let $\mathrm{g} \in G$. We distinguish between an action from the left, $\left[T_{\mathrm{g}} f\right](\mathbf{x}):=f\left(\mathrm{~g}^{-1} \mathbf{x}\right)$, and an action from the right, $\left[T_{\mathrm{g}} f\right](\mathrm{x}):=f(\mathbf{x} \mathrm{~g}), \forall f$.
Why did we define the left action of $g \in G$ as $\mathrm{g}^{-1} \mathbf{x}$, and not g x ? Well, let us check whether the $\left\{T_{\mathrm{g}}\right\}$ do form a group with, first, the group action from the left defined as $\mathrm{g}^{-1} \mathbf{x}$, and then, as $\mathrm{g} \mathbf{x}$. Denote by $T_{\mathrm{g}_{i} \mathrm{~g}_{j}}$ the transformation associated with the group product pair $\mathrm{g}_{i} \mathrm{~g}_{j} \in G$. Then, with $\mathrm{g}_{i}=i$ and $\mathrm{g}_{j}=j$ in subscripts so as to lighten up the formalism:

$$
\left[T_{i j} f\right](\mathbf{x})=f\left(\left(\mathrm{~g}_{i} \mathrm{~g}_{j}\right)^{-1} \mathbf{x}\right)=f\left(\mathrm{~g}_{j}^{-1} \mathrm{~g}_{i}^{-1} \mathbf{x}\right)=\left[T_{j} f\right]\left(\mathrm{g}_{i}^{-1} \mathbf{x}\right)=\left[T_{i} T_{j} f\right](\mathbf{x})
$$

which means that the $T$ operators form a group; but what if instead:

$$
\left[T_{i j} f\right](\mathbf{x})=f\left(\mathrm{~g}_{i} \mathrm{~g}_{j} \mathbf{x}\right)=\left[T_{i} f\right]\left(\mathrm{g}_{j} \mathbf{x}\right)=\left[T_{j} T_{i} f\right](\mathbf{x})
$$

Something awkward has happened: if we write the left action as $\mathbf{g X}$, the associated transformations do not form a group! And, as you should verify, neither do they if we write the right action as $\mathbf{x g}^{-1}$.

So, as a matter of notational consistency, we should always write $\mathbf{x} g$ for the right action and $g^{-1} \mathbf{x}$ for the left action, which is indeed what BF do (without much explanation) for the left action.

### 2.4.2 Matrix representations of a group (BF10.4)

Definition 2.26. A representation $\mathbf{D}$ of a group $G$ is a homomorphic mapping onto a set of finitedimensional invertible matrices such that $\mathbf{D}(e)=\mathbf{I}$, the identity matrix, and $\mathbf{D}\left(\mathrm{g}_{i}\right) \mathbf{D}\left(\mathrm{g}_{j}\right)=\mathbf{D}\left(\mathrm{g}_{i} \mathrm{~g}_{j}\right)$, in the sense that matrix multiplication preserves the group composition law.

If the homomorphism is one-to-one, a representation is faithful. The dimension of the representation is the rank of its matrices or, equivalently, the dimension of the carrier space on which it acts.

Whenever we find a set of degenerate eigenfunctions for some operator that is invariant under a group G, as in eq. (2.3), we expect to be able to connect these functions to a representation of the group.

Matrix representations arise in a much more general context than symmetry. The matrices $G L(n, \mathbb{C})$ of rank $n$ can be thought of as the set of all invertible linear transformations of a vector space of complex-valued functions $\mathcal{V}=\{f(\mathbf{x})\}$, where $\mathbf{x} \in \mathbb{R}^{n}$. If $\left\{\mathbf{e}_{\mathrm{i}}\right\}$ is a basis for $\mathcal{V}$, then $\mathbf{x}=x^{i} \mathbf{e}_{\mathbf{i}}$, where the $x^{i}$ are the components of $\mathbf{x}$ in the basis, and the subscript $i$ on the basis vectors specifies which vector, not which component of the vector.

Even addition can be represented by matrix multiplication: $\mathbf{D}_{\alpha} \mathbf{D}_{\beta}=\mathbf{D}_{\alpha+\beta}$, where $\alpha$ and $\beta$ are two values of a group parameter, eg. angle of rotation in a plane, Lorentz boost.

Now let us simplify things a bit by taking $f(\mathbf{x})=\mathbf{x}$. Then the left action of an element $\mathrm{g} \in G$, expressed in terms of the linear transformations $T_{\mathrm{g}}$, must be written as:

$$
\begin{equation*}
T_{\mathrm{g}}(\mathbf{x})=\mathrm{g}^{-1} \mathbf{x}=x^{i} \mathrm{~g}^{-1} \mathbf{e}_{\mathrm{i}}=x^{i}\left(\mathbf{e}_{\mathrm{j}} D^{j}{ }_{i}\left(\mathrm{~g}^{-1}\right)\right)=\mathbf{e}_{\mathrm{j}}\left(D^{j}{ }_{i}\left(\mathrm{~g}^{-1}\right) x^{i}\right) \tag{2.4}
\end{equation*}
$$

Only this exact definition of the associated $\mathbf{D}$ matrices preserves the group product of $G$. Indeed:

$$
\begin{gathered}
\mathrm{g}_{1}^{-1} \mathrm{e}_{\mathrm{i}}=\mathrm{e}_{\mathrm{j}} D^{j}{ }_{i}\left(\mathrm{~g}_{1}^{-1}\right) \\
\mathrm{g}_{2}^{-1} \mathrm{~g}_{1}^{-1} \mathbf{e}_{\mathrm{i}}=\mathrm{g}_{2}^{-1} \mathbf{e}_{\mathrm{j}} D^{j}{ }_{k}\left(\mathrm{~g}_{1}^{-1}\right)=\mathbf{e}_{\mathrm{k}} D^{k}{ }_{j}\left(\mathrm{~g}_{2}^{-1}\right) D^{j}{ }_{i}\left(\mathrm{~g}_{1}^{-1}\right)
\end{gathered}
$$

Comparing with $g_{2}^{-1} \mathrm{~g}_{1}^{-1} \mathrm{e}_{\mathrm{i}}=\mathbf{e}_{\mathrm{k}} D_{i}^{k}\left(\mathrm{~g}_{2}^{-1} \mathrm{~g}_{1}^{-1}\right)$, we see that $\mathbf{D}\left(\mathrm{g}_{2}^{-1} \mathrm{~g}_{1}^{-1}\right)=\mathbf{D}\left(\mathrm{g}_{2}^{-1}\right) \mathbf{D}\left(\mathrm{g}_{1}^{-1}\right)$, or $\mathbf{D}\left(\mathrm{g}_{1} \mathrm{~g}_{2}\right)=$ $\mathbf{D}\left(\mathrm{g}_{1}\right) \mathbf{D}\left(\mathrm{g}_{2}\right)$, as required for the $\mathbf{D}$ matrices to have the same product rule as the group. This is perfectly consistent with eq. (2.3) above, but now we know that eq. (2.3) corresponded to the left action of the group, $\mathrm{g}^{-1} f^{i}$, which was not so obvious because of the use of the $T_{\mathrm{g}}$ operators which always act from the left.

It is an instructive exercise to show that the proper way of expressing the right action of the same group, $\mathbf{x} \mathbf{g}$, in terms of its (right) representation $\mathbf{D}$ matrices is:

$$
\begin{equation*}
\mathbf{x} \mathrm{g}=\mathbf{e}_{\mathrm{i}} x^{i} \mathrm{~g}=\left(D^{i}{ }_{j}(\mathrm{~g}) x^{j}\right) \mathbf{e}_{\mathrm{i}} \tag{2.5}
\end{equation*}
$$

in which $\mathbf{D}$ acts on the the $x^{i}$ written as a column vector. Because of this, some people see the right action as the more "natural" one. For a given g, the right $\mathbf{D}$ matrices are in general different than the left ones. See the Appendix at the end of this module for more details on left and right action.

### 2.4.3 Non-unicity of group representations

One might hope to define an algorithm that would churn out the representation of a group. But there is no such thing as a unique representation! Indeed, suppose we have a set of $n$-dimensional matrices which represent a group. It is always possible to obtain another representation, of dimension 1 , by mapping these matrices to the number 1. This is called the identity representation, and it always exists. Also, the homomorphic map of the same matrices to their determinant preserves the group product (since $\operatorname{det}(A B)=(\operatorname{det} A)(\operatorname{det} B)$, which provides another one-dimensional representation. Of course, nobody claims that such representations are faithful...

Also, we can make a change of basis: $\mathbf{e}_{i}^{\prime}=\mathbf{e}_{j} S^{j}{ }_{i}$, or $\mathbf{e}_{i}=\mathbf{e}_{j}^{\prime}\left(S^{-1}\right)^{j}{ }_{i}$. Then we have the similarity transformation: $\mathbf{D}^{\prime}(\mathrm{g})=\mathbf{S} \mathbf{D}(\mathrm{g}) \mathbf{S}^{-1}$, and the $\mathbf{D}^{\prime}$ obey the same product rules as the $\mathbf{D}$ matrices.

Definition 2.27. Representations connected by a similarity transformation are said to be equivalent if the transformation matrix is the same for all group elements. They differ only by a choice of basis.
Example 2.5. Consider the continuous group, called $S O(2)$, of rotations around the $z$ axis embedded into the group of three-dimensional rotations. We focus on its left action and look for representations.
We parametrise a rotation by $g=R_{\alpha}$ such that $R_{\alpha} \phi=\phi-\alpha$. This would correspond to rotating the standard basis in $\mathbb{R}^{2}$ by $\alpha$ (passive transformation), so that a vector at angle $\phi$ is at angle $\phi-\alpha$ after the rotation; or, equivalently, rotating the vector by $-\alpha$ in the initial basis. One method for finding representations is to use definition 2.25 (with $T_{\mathrm{g}}=\mathcal{R}_{\alpha}$ ) and eq. (2.4), which here give:

$$
\left[\mathcal{R}_{\alpha} f_{i}\right](\phi)=f_{i}\left[R_{\alpha}^{-1} \phi\right]=f_{i}(\phi+\alpha)=D_{i}^{j}(-\alpha) f_{j}(\phi)
$$

We want to find a set of functions which transform into linear combinations of themselves under $\mathcal{R}_{\alpha}$. Try $f_{1}=\cos \phi, f_{2}=\sin \phi$. Then:

$$
\begin{gathered}
{\left[\mathcal{R}_{\alpha} f_{1}\right](\phi)=\cos (\phi+\alpha)=(\cos \alpha) \cos \phi-(\sin \alpha) \sin \phi=(\cos -\alpha) f_{1}(\phi)+(\sin -\alpha) f_{2}(\phi)} \\
{\left[\mathcal{R}_{\alpha} f_{2}\right](\phi)=\sin (\phi+\alpha)=(\sin \alpha) \cos \phi+(\cos \alpha) \sin \phi=-(\sin -\alpha) f_{1}(\phi)+(\cos -\alpha) f_{2}(\phi)}
\end{gathered}
$$

Compare this with $D_{i}{ }^{j}(-\alpha) f_{j}(\phi)$, and switch the sign of $\alpha$ to obtain the $\mathbf{D}(\alpha)$ matrix:

$$
\mathbf{D}^{(1)}\left(R_{\alpha}\right)=\left(\begin{array}{rr}
\cos \alpha & \sin \alpha \\
-\sin \alpha & \cos \alpha
\end{array}\right)
$$

Well, that's one two-dimensional representation for $S O(2)$, and it is probably the most often used. But it is not the only one! If we had instead chosen $f_{1}=\mathrm{e}^{\mathrm{i} \phi}, f_{2}=\mathrm{e}^{-\mathrm{i} \phi}$, going through the same procedure would yield another matrix:

$$
\mathbf{D}^{(2)}\left(R_{\alpha}\right)=\left(\begin{array}{cc}
\mathrm{e}^{\mathrm{i} \alpha} & 0 \\
0 & \mathrm{e}^{-\mathrm{i} \alpha}
\end{array}\right)
$$

so here is another two-dimensional representation. But is it different? In fact, no, because the transformation $\mathbf{S}^{-1} \mathbf{D}^{(1)} \mathbf{S}$, with the single matrix $\mathbf{S}=\left(\begin{array}{cc}1 & i \\ i & 1\end{array}\right) / \sqrt{2}$, diagonalises $\mathbf{D}^{(1)}$ into $\mathbf{D}^{(2)}$ for any angle $\alpha$, ie. for all elements of the rotation group.
But there are more: each linearly independent function $\mathrm{e}^{\mathrm{i} \alpha}$ and $\mathrm{e}^{-\mathrm{i} \alpha}$ is also a perfectly acceptable one-dimensional representation of $S O(2)$ ! Both $\mathbf{D}^{(1)}$ and $\mathbf{D}^{(2)}$ can be viewed as a joining of these one-dimensional representations, which we shall call $\mathbf{D}^{(3)}$ and $\mathbf{D}^{(4)}$ Since $\mathrm{e}^{\mathrm{i} \alpha}$ is a representation of $U(1)$, this establishes the isomorphism of $S O(2)$ and $U(1)$ via the mapping $\mathrm{e}^{\mathrm{i} \alpha}$. Obviously, there is something special about $\mathrm{e}^{ \pm \mathrm{i} \alpha}$. Before we discover what it is, let us look at another instructive example.
Example 2.6. Let us work out a three-dimensional representation of the left action of $S_{3}, \pi^{-1} \mathbf{x}$, on $\mathbb{R}^{3}$. Since $S_{n}$ merely shuffles the components of $\mathbf{x}$ it preserves its length, which is the definition of orthogonal matrices, ie., those whose transpose is their inverse. In fact, $S_{n} \subset O(n)$ ! Then, from eq.
(2.4), $\pi_{k}^{-1} \mathbf{x}=x^{i} \mathbf{e}_{j} D^{j}{ }_{i}\left(\pi_{k}^{-1}\right)=\left(x^{i} D_{i}^{j}\left(\pi_{k}\right)\right) \mathbf{e}_{j}$ so as to view the permutations as a shuffling of the components of $\mathbf{x}$ (written as row vectors-see the Appendix at the end of the Module!), and we have:
$\mathbf{D}^{(1)}\left(\pi_{1}\right)=\left(\begin{array}{lll}1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1\end{array}\right)$,
$\mathbf{D}^{(1)}\left(\pi_{2}\right)=\left(\begin{array}{lll}0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 1\end{array}\right)$
$\mathbf{D}^{(1)}\left(\pi_{3}\right)=\left(\begin{array}{lll}0 & 0 & 1 \\ 0 & 1 & 0 \\ 1 & 0 & 0\end{array}\right)$
$\mathbf{D}^{(1)}\left(\pi_{4}\right)=\left(\begin{array}{lll}1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0\end{array}\right)$,
$\mathbf{D}^{(1)}\left(\pi_{5}\right)=\left(\begin{array}{lll}0 & 1 & 0 \\ 0 & 0 & 1 \\ 1 & 0 & 0\end{array}\right)$,
$\mathbf{D}^{(1)}\left(\pi_{6}\right)=\left(\begin{array}{lll}0 & 0 & 1 \\ 1 & 0 & 0 \\ 0 & 1 & 0\end{array}\right)$

This so-called left defining (fundamental) left representation of $S_{3}$ is faithful: it is one-to-one with the $\pi_{i} \in S_{3}$. An analogous $n$-dim defining representation can be constructed for any $S_{n}$.
Now, I claim that there exists another (two-dimensional!) representation of $S_{3}$ :

$$
\begin{aligned}
& \mathbf{D}^{(2)}\left(\pi_{1}\right)=\mathbf{D}^{(2)}\left(\pi_{5}\right)=\mathbf{D}^{(2)}\left(\pi_{6}\right)=\left(\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right) \\
& \mathbf{D}^{(2)}\left(\pi_{2}\right)=\mathbf{D}^{(2)}\left(\pi_{3}\right)=\mathbf{D}^{(2)}\left(\pi_{4}\right)=\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right)
\end{aligned}
$$

One immediately checks that products of these matrices are consistent with the group products of $S_{3}$ as given earlier in its Cayley table. So they do form a representation of $S_{3}$, albeit not a faithful one.
Even less faithful, but no less acceptable, is the one-dimensional representation obtained by mapping the $\pi_{i}$ to their parity values:

$$
\begin{gathered}
\mathbf{D}^{(3)}\left(\pi_{1}\right)=\mathbf{D}^{(3)}\left(\pi_{5}\right)=\mathbf{D}^{(3)}\left(\pi_{6}\right)=1 \\
\mathbf{D}^{(3)}\left(\pi_{2}\right)=\mathbf{D}^{(3)}\left(\pi_{3}\right)=\mathbf{D}^{(3)}\left(\pi_{4}\right)=-1
\end{gathered}
$$

And, of course, we can always map all the $\pi_{i}$ to 1 and get another (trivial) representation!
On the other hand, we could join $\mathbf{D}^{(1)}$ and $\mathbf{D}^{(2)}$ into a $\mathbf{D}^{(4)}=\mathbf{D}^{(1)} \oplus \mathbf{D}^{(2)}$ (direct sum) representation whose six matrices are 5-dimensional and block-diagonal, each with the submatrices on the diagonal taken, one from $\mathbf{D}^{(2)}$ (the upper one, say), and the other from $\mathbf{D}^{(1)}$, for a given permutation $\pi_{i}$.

A less trivial object, the direct product $\mathbf{D}^{\alpha \otimes \beta}=\mathbf{D}^{(\alpha)} \otimes \mathbf{D}^{(\beta)}$ of two representations $\mathbf{D}^{(\alpha)}$ and $\mathbf{D}^{(\beta)}$ is the matrix formed from all the products of their elements:

$$
\begin{equation*}
\left(\mathbf{D}^{\alpha \otimes \beta}\right)_{j l}^{i k}=\left(\mathbf{D}^{(\alpha)}\right)_{j}^{i}\left(\mathbf{D}^{(\beta)}\right)_{l}^{k} \tag{2.6}
\end{equation*}
$$

### 2.4.4 The regular representation of finite groups

Definition 2.28. The regular representation of the left action of a finite group is the set of matrices $\mathbf{D}^{\mathrm{L}}(\mathrm{g})$, derived from the group product, with $\mathrm{g} \in G$, such that:

$$
\mathbf{D}^{\mathrm{L}}(\mathrm{~g}) \mathrm{g}_{i}=\mathrm{g} \mathrm{~g}_{i}=\mathrm{g}_{j} D^{j}{ }_{i}(\mathrm{~g}) \quad \forall \mathrm{g} \in G \quad D_{i}^{j}(\mathrm{~g})= \begin{cases}1 & \mathrm{~g} \mathrm{~g}_{i}=\mathrm{g}_{j} \\ 0 & \mathrm{~g}_{i} \neq \mathrm{g}_{j}\end{cases}
$$

The regular representation is seen to be closely related to the Cayley table of the group. Its dimension is equal to $N_{\mathrm{G}}$, the order of the group. We can also see that $D^{j}{ }_{i}(e)=\delta^{j}{ }_{i}$, ie. $\mathbf{D}^{\mathrm{L}}(e)=\mathbf{I}$. Also, the other matrices in the representation must have a 1 as their $(j i)^{\text {th }}$ element and 0 for all other elements in row $j$ and column $i$; by inspection, this 1 is never on the diagonal.
Similarly, we can define a regular representation, $\mathbf{D}^{R}(\mathrm{~g})$, for the right action of a group:

$$
\mathbf{D}^{\mathrm{R}}(\mathrm{~g}) \mathrm{g}_{i}=\mathrm{g}_{i} \mathrm{~g}^{-1}=D_{i}^{j}(\mathrm{~g}) \mathrm{g}_{j} \quad \quad D_{i}^{j}(\mathrm{~g})=\left\{\begin{array}{lll}
1 & \mathrm{~g}_{i} \mathrm{~g}^{-1}=\mathrm{g}_{j} & \text { or } \quad \mathrm{g}_{j} \mathrm{~g}=\mathrm{g}_{i} \\
0 & \mathrm{~g}_{i} \mathrm{~g}^{-1} \neq \mathrm{g}_{j} & \text { or } \\
\mathrm{g}_{j} \mathrm{~g} \neq \mathrm{g}_{i}
\end{array}\right.
$$

A note of caution: do not confuse the dimension of a representation, ie. of its carrier space (the space of functions on which group operators act), with the dimension of the coordinate space on which these functions act.

### 2.4.5 Invariant Spaces and Kronecker sum

To understand what relationship may exist between representations, it is time to bring in a very useful concept:
Definition 2.29. Let $\left\{f^{(i)}\right\}$ be a subspace $\mathcal{H}^{(i)}$ of the space $\mathcal{H}$ of functions on which the linear transformations $\left\{T_{\mathrm{g}}\right\}$ associated with a group $G$ act. If, $\forall f^{(i)} \in \mathcal{H}^{(i)}$ and $\forall \mathrm{g} \in G, T_{\mathrm{g}} f^{(i)} \in \mathcal{H}^{(i)}$, the subspace is invariant under $G$.

Definition 2.30. Let $\mathcal{H}^{(1)}$ and $\mathcal{H}^{(2)}$ be subspaces of a Hilbert space $\mathcal{H}$ such that $\mathcal{H}$ is the sum of the two subspaces with the zero function as the only overlap between the two. Then, if any function in $\mathcal{H}$ can be written uniquely as the sum of a function in $\mathcal{H}^{(1)}$ and another in $\mathcal{H}^{(2)}$, $\mathcal{H}$ is called the Kronecker (or direct) sum of $\mathcal{H}^{(1)}$ and $\mathcal{H}^{(2)}$, written $\mathcal{H}=\mathcal{H}^{(1)} \oplus \mathcal{H}^{(2)}$. The dimension of $\mathcal{H}$ is the sum of the dimensions of $\mathcal{H}^{(1)}$ and $\mathcal{H}^{(2)}$.

### 2.4.6 Reducible and irreducible representations (BF10.5)

Definition 2.31. If some function space $\mathcal{H}$ has a proper subspace invariant under $G$, in the sense that the action of $\mathbf{D}(\mathrm{g}), \forall \mathrm{g}$ and on any function in the subspace, always yields a function in the same subspace, then the representation consisting of the matrices $\mathbf{D}(\mathrm{g})(\forall \mathrm{g} \in G)$ is said to be reducible.

Definition 2.32. If a space $\mathcal{H}$ has no proper subspace invariant under a representation of a group $G$, that representation is said to be irreducible. Its dimension is often called degree by mathematicians.

Definition 2.33. If, $\forall \mathrm{g} \in G$, the matrices in a representation $\mathbf{D}(\mathrm{g})$ can be brought into diagonal block form by the same similarity transformation, then the representation is reducible to lower-dimensional representations whose elements are the block matrices.
If there is another level of invariant subspaces, so that any or all of these block matrices can themselves be written in diagonal block form, and so on, until we are left with only irreducible representations, then $\mathbf{D}(\mathrm{g})$ is fully reducible, in the sense that:

$$
\begin{equation*}
\mathbf{D}(\mathrm{g})=a_{1} \mathbf{D}^{(1)}(\mathrm{g}) \oplus a_{2} \mathbf{D}^{(2)}(\mathrm{g}) \oplus \cdots \oplus a_{N} \mathbf{D}^{(N)}(\mathrm{g}) \tag{2.7}
\end{equation*}
$$

where $a_{i}$ is the number of times the irreducible representation $\mathbf{D}^{(i)}(\mathrm{g})$ occurs in the direct sum.
It can be shown that every representation of a finite group is fully reducible.

When the $n$-dimensional function space $\mathcal{H}$ has proper invariant subspaces, it means that while $\mathcal{H}$ may have a set of linearly independent functions, there are at least two subspaces in $\mathcal{H}$, each of which has its own smaller set of linearly independent functions which transform among themselves. Indeed, let $\mathcal{H}^{A}$ be an invariant subspace of dimension $d$, and let $\left\{\mathbf{e}_{1}, \ldots, \mathbf{e}_{d}, \ldots\right\}$ be a basis of $\mathcal{H}$ with $\left\{\mathbf{e}_{1}, \ldots, \mathbf{e}_{d}\right\}$ a basis of $\mathcal{H}^{A}$. We write vectors of functions in $\mathcal{H}$ in block form $\binom{A}{B}$, where $A \in \mathcal{H}^{A}$ has dimension $d$, and $B$ belongs to the complement subspace $\mathcal{H}^{B}$, of dimension $n-d$. When the complement subspace is also invariant, as it always is in cases of interest to physics (see section 2.4.7 just below), then a representation matrix:

$$
\mathbf{D}(\mathrm{g})=\left(\begin{array}{cc}
\mathbf{D}^{A}(\mathrm{~g}) & 0  \tag{2.8}\\
0 & \mathbf{D}^{B}(\mathrm{~g})
\end{array}\right)
$$

maps vectors $\binom{A}{B}$ into other vectors $\binom{A^{\prime}}{B^{\prime}}$ where $A^{\prime} \in \mathcal{H}^{A}$ and $B^{\prime} \in \mathcal{H}^{B}$. The submatrix $\mathbf{D}^{A}(\mathrm{~g})$ has dimension $d$, and $\mathbf{D}^{B}(\mathrm{~g})$ has dimension $n-d$. Also, since:

$$
\left(\begin{array}{cc}
\mathbf{D}^{A}(\mathrm{~g}) & 0 \\
0 & \mathbf{D}^{B}(\mathrm{~g})
\end{array}\right)\left(\begin{array}{cc}
\mathbf{D}^{A}\left(\mathrm{~g}^{\prime}\right) & 0 \\
0 & \mathbf{D}^{B}\left(\mathrm{~g}^{\prime}\right)
\end{array}\right)=\left(\begin{array}{cc}
\mathbf{D}^{A}(\mathrm{~g}) \mathbf{D}^{A}\left(\mathrm{~g}^{\prime}\right) & 0 \\
0 & \mathbf{D}^{B}(\mathrm{~g}) \mathbf{D}^{B}\left(\mathrm{~g}^{\prime}\right)
\end{array}\right)
$$

$\mathbf{D}^{A}(\mathrm{~g})$ and $\mathbf{D}^{B}(\mathrm{~g})$ do preserve the group product, eg: $\mathbf{D}^{A}(\mathrm{~g}) \mathbf{D}^{A}\left(\mathrm{~g}^{\prime}\right)=\mathbf{D}^{A}(\mathrm{~g} \mathrm{~g})$, as they should.
Going back to $S O(2)$, the $\mathbf{D}^{(2)}$ representation we have obtained is clearly fully reducible as it is written in block-diagonal form, with $1 \times 1$ submatrices $\mathbf{D}^{(3)}=e^{\mathrm{i} \alpha}$ and $\mathbf{D}^{(4)}=\mathrm{e}^{-\mathrm{i} \alpha}$ (ie. it is diagonal), and we can write it as $\mathbf{D}^{(2)}=\mathbf{D}^{(3)} \oplus \mathbf{D}^{(4)}$, where both $\mathbf{D}^{(3)}$ and $\mathbf{D}^{(4)}$ are one-dimensional and, therefore, irreducible.

Example 2.7. The 5-dimensional representation, $\mathbf{D}^{(4)}$, we have constructed for $S_{3}$ in Example 2.6 is (by construction) reducible since it is in block-diagonal form, so $\mathbf{D}^{(4)}=\mathbf{D}^{(1)} \oplus \mathbf{D}^{(2)}$. What about that last two-dimensional representation $\left(\begin{array}{ll}1 & 0 \\ 0 & 1\end{array}\right)$ and $\left(\begin{array}{ll}0 & 1 \\ 1 & 0\end{array}\right)$ ? The first is already in block-diagonal form and the second can be diagonalised to $\left(\begin{array}{rr}1 & 0 \\ 0 & -1\end{array}\right)$, Therefore, we obtain two 1-dim irreducible representations, one identical to the identity representation $\mathbf{D}^{(5)}=1$, and the other the "parity" representation $\mathbf{D}^{(3)}$. Then we can write: $\mathbf{D}^{(4)}=\mathbf{D}^{(1)} \oplus \mathbf{D}^{(3)} \oplus \mathbf{D}^{(5)}$. As for $\mathbf{D}^{(1)}$, it is itself reducible (EXERCISE) to two irreducible representations, $\mathbf{D}^{(5)}$, and a set of six 2-dim matrices. Therefore, the 5-dim $\mathbf{D}^{(4)}$ is fully reducible, since it can be written as the direct sum of four representations, one 2 -dim, one 1-dim, and two 1-dim (the parity of the elements), all irreducible. Can you see why these irreducible representations could not all be one-dimensional?
So this reduction algorithm certainly works, but it would be nice not to have to rely on looking for invariant subspaces and similarity transformations, which can get quite involved.

### 2.4.7 Unitary representations (BF10.6)

A representation $\mathbf{D}(\mathrm{g})$ is unitary if $\mathbf{D}^{\dagger}(\mathrm{g})=\mathbf{D}\left(\mathrm{g}^{-1}\right), \forall \mathrm{g} \in G$. In terms of matrix elements, $D_{i j}\left(\mathrm{~g}^{-1}\right)=D_{j i}^{*}(\mathrm{~g})$.
Notice that $\mathbf{D}^{(3)}$ and $\mathbf{D}^{(4)}$ for $S O(2)$ are unitary,. So are the left and right regular representations of a group.
Now, if $\mathbf{D}(\mathrm{g})$ is not already unitary, we can always find a similarity transformation matrix $\mathbf{S}$, the Hermitian square root of $\mathbf{S}^{2}=\sum_{g} \mathbf{D}^{\dagger}(g) \mathbf{D}(g)$, such that $\mathbf{D}^{\prime}(\mathrm{g})=\mathbf{S} \mathbf{D}(\mathrm{g}) \mathbf{S}^{-1}$ is unitary (EXERCISE-first, show that $\left.\mathbf{D}^{\dagger}\left(g^{\prime}\right) \mathbf{S}^{2} \mathbf{D}\left(g^{\prime}\right)=\mathbf{S}^{2}\right)$. Any representation of a finite group is equivalent to a unitary representation, whose matrices are unitary. This is also true for certain infinite (continuous) groups, such as compact Lie groups. Also, it can be shown that if a subspace is invariant under a unitary representation, its complement must also be invariant.

### 2.5 Schur's Lemmas and the Classification of Group Representations (BF10.6)

We now present two fundamental results of group theory which provide useful criteria for the irreducibility of representations and from which can be derived relations that are of tremendous help in classifying representations.

### 2.5.1 Schur's Lemmas

Schur's First Lemma: The only complex matrix $\mathbf{M}$ that commutes with the matrices of a given irreducible representation $\mathbf{D}(\mathrm{g}), \forall \mathrm{g} \in G$, is a multiple of the identity matrix.

Indeed, since any $\mathbf{M} \in G L(n, \mathbb{C})$ has at least one (perhaps complex) non-zero eigenvector, let $\mathbf{M} \mathbf{A}=\lambda \mathbf{A}$. If $\mathbf{M}$ and $\mathbf{D}$ commute, we have $\mathbf{M}(\mathbf{D} \mathbf{A})=\mathbf{D} \mathbf{M} \mathbf{A}=\lambda(\mathbf{D} \mathbf{A})$. This means that the $\mathbf{D}\left(\mathrm{g}_{p}\right) \mathbf{A}$ are also eigenvectors of $\mathbf{M} \forall \mathrm{g}_{p} \in G$; since they span the same vector space as the eigenvectors $\mathbf{A}$ of $\mathbf{M}$, that space is invariant under $G$.

By definition 2.32 of irreducibility, however, that space $\mathcal{H}$ on which $\mathbf{D}$ acts has no proper subspace itself invariant under $G$. Therefore, the subspace spanned by the eigenvectors associated with this one eigenvalue is the whole of $\mathcal{H}$. In other words, $\forall \psi \in \mathcal{H}, \psi=a_{p} \mathbf{A}^{(p)}$, where $\mathbf{A}^{(p)}=\mathbf{D}\left(\mathrm{g}_{p}\right) \mathbf{A}$, and:

$$
\mathbf{M} \psi=a_{p} \mathbf{M} \mathbf{A}^{(p)}=\lambda a_{p} \mathbf{A}^{(p)}=\lambda \psi
$$

so that all vectors in $\mathcal{H}$ are eigenvectors of $\mathbf{M}$, with the same eigenvalue $\lambda$. This can happen only if $\mathbf{M}=\lambda \mathbf{I}$.
As a corollary, if a matrix can be found which is not a multiple of $\mathbf{I}$ and yet commutes with all matrices in a representation, that representation must be reducible.

From this Lemma follows an immediate consequence for Abelian groups, whose elements commute, so that any matrix $\mathbf{D}(\mathrm{g})$ in a given representation commutes with the matrices for all other group elements in this representation. Assuming a $n$-dim irreducible representation, the the Lemma requires that $\mathbf{D}(\mathrm{g})=\lambda_{\mathrm{g}} \mathbf{I}, \forall \mathrm{g} \in G$. But the $n \times n$ identity matrix, which is diagonal, cannot be irreducible if it represents all group elements, contradicting our assumption. We conclude that all irreducible representations of an Abelian group are one-dimensional.

It is not so hard to find matrices that satisfy the conditions of Schur's First Lemma. For instance, let $\mathcal{C}$ be a class of $G$, and let $\mathbf{D}\left(\mathrm{g}_{p}\right)$ be a matrix in a given irreducible representation of $G$. Construct a matrix by summing all the matrices $\mathbf{D}\left(\mathrm{g}_{p}\right)$ in the representation that correspond to elements $\mathrm{g}_{p}$ of $\mathcal{C}$ :

$$
\begin{equation*}
\mathbf{M}=\sum_{p} \mathbf{D}\left(\mathrm{~g}_{p}\right) \quad \forall \mathrm{g}_{p} \in \mathcal{C} \tag{2.9}
\end{equation*}
$$

Now, if $\mathrm{g}^{\prime}$ is some arbitrary element of $G$, we have:

$$
\mathbf{D}\left(\mathrm{g}^{\prime}\right) \mathbf{M} \mathbf{D}\left(\mathrm{g}^{\prime-1}\right)=\sum_{p} \mathbf{D}\left(\mathrm{~g}^{\prime}\right) \mathbf{D}\left(\mathrm{g}_{p}\right) \mathbf{D}\left(\mathrm{g}^{\prime-1}\right)=\sum_{p} \mathbf{D}\left(\mathrm{~g}^{\prime} \mathrm{g}_{p} \mathrm{~g}^{\prime-1}\right)=\mathbf{M}
$$

where the next-to-last equality expresses the fact that the matrices preserve the group product, and the last equality results from the fact that, since $\mathrm{g}^{\prime} \mathrm{g}_{p} \mathrm{~g}^{\prime-1} \in \mathcal{C}$, the left-hand side of the last equality is just a rearrangement of the previous equation. Thus, $\mathbf{D}(\mathrm{g}) \mathbf{M}=\mathbf{M} \mathbf{D}(\mathrm{g}) \forall \mathrm{g} \in G$, and, from Schur's First Lemma, $\mathbf{M}=\lambda \mathbf{I}$, with $\lambda$ a constant that depends on the class and on the representation. Thus, the trace of $\mathbf{M}, \operatorname{Tr} \mathbf{M}$, is $n \lambda$, with $n$ the dimension of the representation.

Before proceeding to find $\lambda$, we establish an interesting fact: In a given representation, all matrices associated with elements of the same class have the same trace. Recall that the class to which $g$ belongs is made of $\left\{\mathrm{g}^{\prime} \mathrm{g} \mathrm{g}^{\prime-1}\right\} \forall \mathrm{g}^{\prime} \in G$. Then the trace of $\mathbf{D}\left(\mathrm{g}^{\prime} \mathrm{g} \mathrm{g}^{\prime-1}\right)$ is equal ${ }^{\dagger}$ to the trace of $\mathbf{D}(\mathrm{g})$, which we denote by $\chi$.

Using this fact and the definition of $\mathbf{M}$, we have: $\mathbf{T r} \mathbf{M}=n_{\mathrm{c}} \chi$, where $n_{\mathrm{c}}$ is the number of elements in the class. Since that trace is also $n \lambda$, we find:

$$
\begin{equation*}
\mathbf{M}=\frac{n_{\mathrm{c}}}{n} \chi \mathbf{I} \tag{2.10}
\end{equation*}
$$

Schur's Second Lemma: If a non-zero matrix $\mathbf{M}$ exists such that $\mathbf{D}^{(\alpha)}(\mathrm{g}) \mathbf{M}=\mathbf{M} \mathbf{D}^{(\beta)}(\mathrm{g}) \forall \mathrm{g} \in G$, then $\mathbf{D}^{(\alpha)}$ and $\mathbf{D}^{(\beta)}$ must be equivalent irreducible representations. If $\mathbf{D}^{(\alpha)}$ and $\mathbf{D}^{(\beta)}$ are inequivalent, $\mathbf{M}=0$.

This lemma can be proved (pp. BF615-617) by assuming unitary representations. This makes for no loss of generality for finite or compact Lie groups, since these (eg. $O(n)$ ) have finite-dimensional representations.

### 2.5.2 An orthogonality relation (BF10.6)

Another important consequence of Schur's Lemmas is the fact that the matrix elements of all the inequivalent irreducible representations of a finite group, or those for infinite groups that have finite-dimensional representations, form a set of orthogonal functions of the elements of the group. More specifically, if $\left\{\mathbf{D}\left(\mathrm{g}_{p}\right)\right\}$ is the set of all matrices $\mathbf{D}_{p}$ in an irreducible representation $\mathbf{D}$, then, for two such representations labelled by $\alpha$ and $\beta::$

$$
\begin{equation*}
\sum_{p=1}^{N_{\mathrm{G}}}\left(D_{p}^{(\alpha)}\right)^{i}{ }_{k}\left(D_{\mathrm{g}_{p}^{-1}}^{(\beta)}\right)^{l}{ }_{j}=\frac{N_{\mathrm{G}}}{n_{\alpha}} \delta^{i}{ }_{j} \delta^{l}{ }_{k} \delta_{\alpha \beta} \tag{2.11}
\end{equation*}
$$

where $N_{\mathrm{G}}$ is the order of the group and $n_{\alpha}$ is the dimension of $\mathbf{D}^{(\alpha)}$. The sum is not matrix multiplication! Each term is the product of some $i k$ entry of $D_{p}^{(\alpha)}$ and $l k$ entry of $D_{\mathrm{g}^{-1}}^{(\beta)}$, with $i k$ and $l k$ the same for each term.

[^11]In the case of unitary representations, this relation simplifies to:

$$
\begin{equation*}
\sum_{p=1}^{N_{\mathrm{G}}}\left(D_{p}^{(\alpha)}\right)^{i}{ }_{k}\left(D_{p}^{(\beta) *}\right)_{j}^{l}=\frac{N_{\mathrm{G}}}{n_{\alpha}} \delta^{i}{ }_{j} \delta^{l}{ }_{k} \delta_{\alpha \beta} \tag{2.12}
\end{equation*}
$$

These relations set powerful constraints on the matrix elements of representations.
Eq. (2.11) is so important that it deserves a proof. Fortunately, this proof is fairly easy. Construct a matrix:

$$
\begin{equation*}
\mathbf{M}=\sum_{p=1}^{N_{\mathrm{G}}} \mathbf{D}_{p}^{(\alpha)} \mathbf{X}\left[\mathbf{D}_{p}^{(\beta)}\right]^{-1} \tag{2.13}
\end{equation*}
$$

where $\mathbf{D}^{(\alpha)}$ and $\mathbf{D}^{(\beta)}$ are $m$-dim and $n$-dim inequivalent irreducible matrix representations of $G$, and $X$ is any arbitrary operator represented by a $m \times n$ matrix $\mathbf{X}$. Note that the sum runs over all the group elements, each one labelled by a value of $p$. Then, for some $\mathrm{g}_{p^{\prime}} \in G$,

$$
\mathbf{D}_{p^{\prime}}^{(\alpha)} \mathbf{M}\left[\mathbf{D}_{p^{\prime}}^{(\beta)}\right]^{-1}=\sum_{p=1}^{N_{\mathrm{G}}} \mathbf{D}_{p^{\prime} p}^{(\alpha)} \mathbf{X}\left[\mathbf{D}_{p^{\prime} p}^{(\beta)}\right]^{-1}
$$

The sum on the right-hand side is just a different rearrangement of the sum that defines $\mathbf{M}$, so that:

$$
\mathbf{M}=\mathbf{D}_{p^{\prime}}^{(\alpha)} \mathbf{M}\left[\mathbf{D}_{p^{\prime}}^{(\beta)}\right]^{-1}
$$

Thus, $\mathbf{M}$ meets the condition for Schur's Lemmas: $\mathbf{D}_{p}^{(\alpha)} \mathbf{M}=\mathbf{M} \mathbf{D}_{p}^{(\beta)} \forall \mathrm{g} \in G$. When $\alpha \neq \beta \mathbf{M}=0$ from the Second Lemma. When $\alpha=\beta$, let us choose $\mathbf{X}$ to be a matrix whose only non-zero element, 1 , is its $(k l)^{\text {th }}$ entry. We can write this formally as: $X^{m}{ }_{n}=\delta^{m}{ }_{k} \delta_{n}{ }^{l}$ to obtain:

$$
M^{i}{ }_{j}=\sum_{p=1}^{N_{\mathrm{G}}}\left(\mathbf{D}_{p}^{(\alpha)}\right)^{i}{ }_{m} X^{m}{ }_{n}\left(\mathbf{D}_{\mathrm{g}_{p}^{-1}}^{(\alpha)}\right)^{n}{ }_{j}=\sum_{p=1}^{N_{\mathrm{G}}}\left(\mathbf{D}_{p}^{(\alpha)}\right)^{i}{ }_{k}\left(\mathbf{D}_{\mathrm{g}_{p}^{-1}}^{(\alpha)}\right)^{l}{ }_{j}
$$

Because $\mathbf{D}_{p}^{(\alpha)} \mathbf{M}=\mathbf{M} \mathbf{D}_{p}^{(\alpha)}$, Schur's First Lemma requires that $\mathbf{M}=\lambda_{\mathbf{x}} \mathbf{I}$, ie.:

$$
M^{i}{ }_{j}=\sum_{p=1}^{N_{\mathrm{G}}}\left(\mathbf{D}_{p}^{(\alpha)}\right)^{i}{ }_{k}\left(\mathbf{D}_{\mathrm{g}_{p}^{-1}}^{(\alpha)}{ }_{j}^{l}=\lambda_{k}{ }^{l} \delta^{i}{ }_{j}\right.
$$

Taking the trace (setting $i=j$ with summation over repeated indices) and interchanging the $\mathbf{D}$ factors to get a matrix product, there comes:

$$
\sum_{p=1}^{N_{\mathrm{G}}}\left(\mathbf{D}_{\mathrm{g}_{p}^{-1}}^{(\alpha)}\right)_{j}^{l}\left(\mathbf{D}_{p}^{(\alpha)}\right)^{j}{ }_{k}=\lambda_{k}{ }^{l} n_{\alpha}
$$

The product inside the sum now implies matrix multiplication and, because the $\mathbf{D}$ matrices preserve the group product of $G$, we have:

$$
\sum_{p=1}^{N_{\mathrm{G}}}\left(\mathbf{D}_{\mathrm{g}_{p}^{-1}}^{(\alpha)}{ }_{j}^{l}\left(\mathbf{D}_{p}^{(\alpha)}\right)^{j}{ }_{k}=\sum_{p=1}^{N_{\mathrm{G}}}\left(\mathbf{D}_{\mathrm{g}_{p} \mathrm{~g}_{p}^{-1}}^{(\alpha)}\right)_{k}^{l}=N_{\mathrm{G}} \delta^{l}{ }_{k}\right.
$$

Comparing, we get: $\lambda^{l}{ }_{k}=N_{\mathrm{G}} \delta^{l}{ }_{k} / n_{\alpha}$. Then there only remains to collect the results for $\alpha=\beta$ and $\alpha \neq \beta$ into the compact forms (2.11) or (2.12).

This means that the matrix elements $\sqrt{\frac{n_{\alpha}}{N_{\mathrm{G}}}} D_{i j}^{(\alpha)}$ of a unitary irreducible representation must be orthonormal functions of the group elements g , and therefore linearly independent over the space spanned by the group elements. Furthermore, they form a complete set.

EXERCISE: Show that for a finite group the sum over all elements $g_{p}$ of the matrix elements $\left(D_{p}\right)^{i}{ }_{j}(i$ and $j$ fixed) of an irreducible representation is zero. This property can provde a useful.check.

### 2.5.3 Characters of a representation (BF10.7); orthogonality of the characters

Definition 2.34. The character of a representation $\mathbf{D}(\mathrm{g})$ of a group $G$ is defined as a map from $G$ to $\mathbb{C}$ :

$$
\chi(\mathrm{g})=\operatorname{Tr} \mathbf{D}(\mathrm{g})
$$

Therefore the character of a representation for $g \in G$ is just the trace of its representing matrix $\mathbf{D}(\mathrm{g})$. Characters of reducible representations are compound; those of irreducible representations are called simple. Language alert: Mathematicians speak of the "character" of a representation as the set of traces of the matrices for the classes in the representation.

Matrices for equivalent representations all have the same character; in other words, any statement about characters is basis-independent! Also, if $g$ and $g^{\prime}$ are conjugate to one one another, there must exist an element $g^{\prime \prime}$ of $G$ such that

$$
\mathbf{D}\left(\mathrm{g}^{\prime}\right)=\mathbf{D}\left(\mathrm{g}^{\prime \prime}\right) \mathbf{D}(\mathrm{g}) \mathbf{D}^{-1}\left(\mathrm{~g}^{\prime \prime}\right)
$$

Therefore, all the matrices representing group elements in the same class and representation have the same character. If they are to be inequivalent, two representations must have at least one different character for some class,

Now we can set $k=i$ and $l=j$ in eq. (2.12):

$$
\sum_{p=1}^{N_{\mathrm{G}}}\left(D_{p}^{(\alpha)}\right)^{i}{ }_{i}\left(D_{p}^{(\beta) *}\right)^{j}{ }_{j}=\frac{N_{\mathrm{G}}}{n_{\alpha}} \delta^{i}{ }_{j} \delta_{i}{ }^{j} \delta_{\alpha \beta}=\frac{N_{\mathrm{G}}}{n_{\alpha}} \delta^{i}{ }_{i} \delta_{\alpha \beta}
$$

where repeated indices are summed over. Since $\delta^{i}{ }_{i}=n_{\alpha}$, this can be rewritten as:

$$
\begin{equation*}
\sum_{p=1}^{N_{\mathrm{G}}} \chi_{p}^{(\alpha)} \chi_{p}^{*(\beta)}=N_{\mathrm{G}} \delta_{\alpha \beta} \tag{2.14}
\end{equation*}
$$

This is our first orthogonality relation between the characters of irreducible representations.
Some of the terms in this sum will be identical since they correspond to group elements in the same class. So, instead of thinking of this sum as over all the elements of the group, we can collect all terms belonging to the same class, which we label with $k$, and sum over the classes, ie. over $k$ :

$$
\begin{equation*}
\sum_{k=1}^{N_{\mathrm{c}}} n_{k} \chi_{k}^{(\alpha)} \chi_{k}^{*(\beta)}=N_{\mathrm{G}} \delta_{\alpha \beta} \tag{2.15}
\end{equation*}
$$

where $n_{k}$ is the number of elements in class $k$, and $N_{\mathrm{c}}$ is the number of classes in the group. This looks for all the world like an orthogonality relation between two vectors, $\sqrt{n_{k} / N_{\mathrm{G}}} \chi^{(\alpha)}$ and $\sqrt{n_{k} / N_{\mathrm{G}}} \chi^{(\beta)}$, each of dimension $N_{\mathrm{c}}$.

For a given irreducible representation, eq. (2.15) becomes:

$$
\begin{equation*}
\sum_{k=1}^{N_{\mathrm{c}}} n_{k}\left|\chi_{k}^{(\alpha)}\right|^{2}=N_{\mathrm{G}} \tag{2.16}
\end{equation*}
$$

This is a necessary and sufficient condition for the representation to be irreducible!
Example 2.8. Take for instance the $3 \times 3$ representation of $S_{3}$ found in section 2.4.3. There is one element, the identity, with trace 3 , in its own class, three elements (the transpositions) in another class with trace 1 , and two elements (the cyclic permutations) with trace 0. Eq. (2.16) gives: $n_{1} \chi_{1}^{2}+n_{2} \chi_{2}^{2}+$ $n_{3} \chi_{3}^{2}=1(3)^{2}+3(1)^{2}+2(0)^{2}=12$. This is not equal to 6 , the number of elements in $S_{3}$, therefore the representation must be reducible.

According to eq. (2.15), the character "vectors" of the $N_{\mathrm{r}}$ different irreducible representations are orthogonal. There are $N_{\mathrm{r}}$ such orthogonal vectors, and their number may not exceed the dimensionality of the space, $N_{\mathrm{c}}$, so that $N_{\mathrm{r}} \leq N_{\mathrm{c}}$. We will need this result a little later.

### 2.5.4 Multiplicity of irreducible representations and a sum rule for their dimension

Now consider the decomposition of a fully reducible representation into a direct sum of irreducible ones, given in eq. (2.7). If we take its trace, we get an equation for the compound character $\chi(\mathrm{g}): \chi(\mathrm{g})=a_{\alpha} \chi^{(\alpha)}(\mathrm{g})$, where the (direct) sum runs over the $N_{\mathrm{r}}$ irreducible representations of the group. The compound character is seen to be a linear combination of simple characters with positive coefficients (the number of times a given irreducible representation appears in the decomposition).

Multiplying this relation by $\chi^{\star(\beta)}(\mathrm{g})$ and summing over group elements now labelled by an index $p$, we find from eq. (2.14) which expresses the orthogonality of the characters:

$$
\sum_{p}^{N_{\mathrm{G}}} \chi_{p} \chi_{p}^{\star(\beta)}=a_{\alpha} \sum_{p}^{N_{\mathrm{G}}} \chi_{p}^{(\alpha)} \chi_{p}^{*(\beta)}=a_{\alpha} N_{\mathrm{G}} \delta_{\alpha \beta}=a_{\beta} N_{\mathrm{G}}
$$

Thus, the number of times-multiplicity-each irreducible representation of a group occurs in the direct-sum decomposition of a reducible representation is readily calculated in terms of characters, often making this direct sum easy to obtain:

$$
\begin{equation*}
a_{\alpha}=\frac{1}{N_{\mathrm{G}}} \sum_{p}^{N_{\mathrm{G}}} \chi_{p} \chi_{p}^{*(\alpha)} \tag{2.17}
\end{equation*}
$$

We can now exploit the regular representation to obtain other general results for irreducible representations.
As we have seen in section 2.4.4, the entries of the matrices of the regular representation can only be 1 or 0 . Since only the identity will map a group element to itself, the only matrix with 1 anywhere on the diagonal is the identity matrix. Therefore, in the regular representation, the characters $\chi(\mathrm{g})$ all vanish except for $\chi(e)=N_{\mathrm{G}}$.

Now, when $\mathrm{g}_{p}=e$, the multiplicity relation (2.17) gives:

$$
a_{\alpha}=\frac{1}{N_{\mathrm{G}}} \chi(e) \chi^{*(\alpha)}(e)=\chi^{*(\alpha)}(e)=n_{\alpha}
$$

where, as before, $n_{\alpha}$ is the dimension of the $\alpha^{\text {th }}$ irreducible representation and we have used the fact that only $\chi(e)$ can contribute to the sum since the characters of all other group elements in the regular representation vanish.

Therefore, the multiplicity of an irreducible representation in the decomposition of the regular representation is the dimension of that irreducible representation, and it is never zero. Next, taking the trace of the Kronecker decomposition (2.7) for the identity group element of the regular representation yields a relation between its compound and simple characters: $N_{\mathrm{G}}=\sum_{\alpha} a_{\alpha} n_{\alpha}$. Combining those results, there comes an important sum rule:

$$
\begin{equation*}
N_{\mathrm{G}}=\sum_{\alpha} n_{\alpha}^{2} \tag{2.18}
\end{equation*}
$$

This powerful sum rule tells us that the dimension of any irreducible representation must be smaller than the square root of the order of the group. In fact, any representation of dimension larger than $\sqrt{N_{\mathrm{G}}}$ must be reducible. Thus, when $N_{\mathrm{G}}=2$ or 3 , all inequivalent irreducible representations are one-dimensional. When $N_{\mathrm{G}}=4$, we can have only four inequivalent 1-d irreducible representations; $n_{\alpha}=2$ is ruled out because there would be no room left for the identity 1-d representation. When $N_{\mathrm{G}}=5$, eq. (2.18) does allow the identity representation together with one 2 -d irreducible representation; but we know that this group, $Z_{5}$, is Abelian, and so admits only five inequivalent 1-d irreducible representations. For a group of order 6, the sum rule allows either six 1-d irreducible representations, or two $1-\mathrm{d}$ and one 2 -d irreducible representations.

We also see that all the irreducible representations of a group must appear in the Kronecker decomposition of its regular representation.

### 2.5.5 Another orthogonality relation

Let us go back to our first orthogonality relation, eq. (2.12). What it says is that the set $\left\{\sqrt{n_{\alpha} / n_{\mathrm{G}}}\left(D^{(\alpha)}\right)^{i}{ }_{j}\left(\mathrm{~g}_{\mathrm{p}}\right)\right\}$, with $i$ and $j$ fixed but $1 \leq p \leq N_{\mathrm{G}}$, of an irreducible representation $\alpha$, can be viewed as the $N_{\mathrm{G}}$ components of a
vector orthogonal to any other such vector corresponding to other matrix elements, whether or not they correspond to the same representation as that of the first vector. There are $N_{\mathrm{G}}$ such vectors and they form a complete set with completeness relation expressed as:

$$
\begin{equation*}
\sum_{\alpha}^{N_{\mathrm{r}}} \sum_{i, j}^{n_{\alpha}} \frac{n_{\alpha}}{N_{\mathrm{G}}}\left(D_{p}^{(\alpha)}\right)_{j}^{i}\left(D_{p^{\prime}}^{(\alpha) *}\right)_{j}^{i}=\delta_{p^{\prime} p} \tag{2.19}
\end{equation*}
$$

where $N_{r}$ is the number of irreducible representations. Again, the left-hand side is not matrix multiplication.
Well, what can we do with this result? We can take the equation for each $p$, where $p$ is an element of some class $k$, and sum over all elements of the class; we can also do this with $p^{\prime}$ over the elements of another class $k^{\prime}$. When $k \neq k^{\prime}$, the right-hand side of the double summation must vanish because classes are distinct; when $k=k^{\prime}$, the double sum collapses into one which adds up to $n_{k}$. On the left-hand side, we may recognise $\sum_{p}^{n_{k}}\left(D_{p}^{(\alpha)}\right)^{i}{ }_{j}$ as an element of the matrix $\mathbf{M}$ introduced in eq. (2.9). The completeness relation now reads:

$$
\sum_{\alpha}^{N_{\mathrm{r}}} \sum_{i, j}^{n_{\alpha}} \frac{n_{\alpha}}{N_{\mathrm{G}}}\left(M_{k}^{(\alpha)}\right)_{j}^{i}\left(M_{k^{\prime}}^{(\alpha) *}\right)_{j}^{i}=n_{k} \delta_{k^{\prime} k}
$$

But there we had also found that $\mathbf{M}^{(\alpha)}=\left(n_{k} / n_{\alpha}\right) \chi^{(\alpha)} \mathbf{I}$. Inserting and carrying out the sums over $i$ and $j$ gives another orthogonality relation:

$$
\begin{equation*}
\sum_{\alpha=1}^{N_{\mathrm{r}}} \frac{n_{k}}{N_{\mathrm{G}}} \chi_{k}^{(\alpha)} \chi_{k^{\prime}}^{*}(\alpha)=\delta_{k^{\prime} k} \tag{2.20}
\end{equation*}
$$

Thus, the characters in the $k^{\text {th }}$ class in a given irreducible representation can be considered as components of vectors forming a basis of a space whose dimension is $N_{\mathrm{r}}$, the number of irreducible representations. There are $N_{\mathrm{c}}$ such orthogonal vectors, one for each class. But, in a $N_{\mathrm{r}}$-dimensional space, there cannot be more than $N_{\mathrm{r}}$ orthogonal vectors, and $N_{\mathrm{c}} \leq N_{\mathrm{r}}$.

In section 2.5.3, however, we had argued that $N_{\mathrm{r}} \leq N_{\mathrm{c}}$. These results together lead to the important statement: The number of irreducible representations of a group is equal to the number of its classes: $N_{\mathrm{r}}=N_{\mathrm{c}}$.

### 2.5.6 Exploring representations with Young diagrams

We have already discussed how Young diagrams could be used to find and label classes of $S_{N}$. But, much more often, it is representations that they help to label. We will be looking at $S_{N}$ whose classes we have associated with partitions of $N$ and, noting that the number of irreducible representations is also the number of classes, we will construct their Young diagrams with the same partitions $\lambda_{i}$ of $N$, where the $\lambda_{i}$ sum up to $N$ and $\lambda_{1} \geq \ldots \geq \lambda_{N}$. So the Young diagrams for the irreducible respresentations of $S_{N}$ look exactly like those for its classes.

What will be different is the labelling of the Young diagrams: instead of taking the partitions as the number of boxes in columns from left to right, we take them as their number in rows from top to bottom. For $S_{3}$, this gives:


The sequence of representation labels is just the reverse of that for classes! If they are not cycles, what are they?
To discover the meaning of these Young diagrams we consider how the corresponding permutations act on functions in the carrier space of the $N$ !-dimensional regular representation. We start by giving ourselves a set of functions $\left\{\psi_{i}\right\}(1 \leq i \leq N)$, each of one variable, where the choice of the same symbol as for particle wavefunctions in quantum mechanics is intentional (some authors use the Dirac notation for them). Then with products of these we construct functions of $N$ variables $x_{j}$. For instance, the product $\psi_{(1 \ldots N)}=\psi_{1}\left(x_{1}\right) \cdots \psi_{1}\left(x_{N}\right)$ spans a one-dimensional subspace which contains functions which are obviously completely symmetric, ie. invariant,
under any of the $N$ ! possible permutations of the variables. Thus, our subspace qualifies as an invariant subspace for the regular representation, and it makes sense to associate it with the 1-d irreducible identity representation which has the same character, 1 , for all elements of $S_{N}$. We shall follow the usual convention by associating it with the single Young diagram with one row of $N$ boxes. Its label will therefore always be $(N)$.

With the same set $\left\{\psi_{i}\right\}$, we can also construct the completely antisymmetric function:

$$
\psi_{[1 \ldots N]}=\left|\begin{array}{ccc}
\psi_{1}\left(x_{1}\right) & \cdots & \psi_{1}\left(x_{N}\right) \\
\vdots & \cdots & \vdots \\
\psi_{N}\left(x_{1}\right) & \cdots & \psi_{N}\left(x_{N}\right)
\end{array}\right|
$$

This function changes sign under any transposition in its set of variables, and the 1-d subspace it spans is also invariant, because the function resulting from multiplying $\psi_{[1 \ldots N]}$ by $\pm 1$ is obviously in the same subspace. We associate this subspace with the 1-d irreducible representation which sends each element of $S_{N}$ to its parity, +1 or -1 . Again by convention, this in turn corresponds to the single one-column Young diagram with $N$ rows.

Other irreducible representations, and thus Young diagrams, have a mixed symmetry which can be used to find their dimension. This is even stronger than eq. (2.18) which is only a constraint on the possible dimensions. Here is one way to do this.

- Take the Young diagram for each irrep, and fill each of its $N$ boxes with numbers from 1 to $N$ in all possible permutations to generate $N$ ! Young tableaux. Then assign a function with $N$ subscripts, living in the carrier space of the regular representation of $S_{N}$, to each tableau. The order of the subscripts follows the order of numbers in the first row, then the second row, until the last row. These functions represent products of functions, each of one coordinate, but we no longer treat them explicitly as such. They form a basis for the carrier space of the regular representation.
- Symmetrise each function with respect to the numbers in each row of the tableau, and antisymmettrise the result with respect to the numbers in each column. This yields, for each diagram, a new, mixed-symmetry function, $\psi^{(i)}(1 \leq i \leq N)$, that is a linear combination of the previous $N$ ! basis functions for the carrier space of the regular representation.

Example 2.9. For the (21) irreducible representation of $S_{3}$, the Young tableaux and corresponding mixed-symmetry functions would be:

| $\begin{array}{\|l\|l\|} \hline 1 & 2 \\ \hline 3 & \\ \hline \end{array}$ | $\Psi^{(1)}=\psi_{123}+\psi_{213}-\psi_{321}-\psi_{231}$ | 1 3 <br> 2  | $\Psi^{(2)}=\psi_{132}+\psi_{312}-\psi_{231}-\psi_{321}$ |
| :---: | :---: | :---: | :---: |
| $\begin{array}{\|l\|l\|} \hline 2 & 1 \\ \hline 3 & \\ \hline & \\ \hline \end{array}$ | $\Psi^{(3)}=\psi_{213}+\psi_{123}-\psi_{312}-\psi_{132}$ | 2 3 <br> 1  | $\Psi^{(4)}=\psi_{231}+\psi_{321}-\psi_{132}-\psi_{312}$ |
| $\begin{array}{\|l\|l\|} \hline 3 & 1 \\ \hline 2 & \\ \hline \end{array}$ | $\Psi^{(5)}=\psi_{312}+\psi_{132}-\psi_{213}-\psi_{123}$ | 3 2 <br> 1  <br> 18  | $\Psi^{(6)}=\psi_{321}+\psi_{231}-\psi_{123}-\psi_{213}$ |

The question now is, are these mixed functions independent? Since we expect the regular representation to be reducible (fully reducible, in fact), there should exist a lower-dimensional invariant subspace, the carrier space of our irreducible representation of interest, and we should be able to show that there are only $n_{\alpha}<6$ (for $S_{3}$ ) independent combinations, where $n_{\alpha}$ will be the number of basis functions for the invariant subspace, and therefore the dimension of the irreducible representation of $S_{3}$ carried by that space.
We note immediately that linear combinations that differ by a transposition of numbers in a column of their tableaux cannot be independent: they are the negative of one another. So we have at most three linearly independent combinations. But we also see that $\Psi^{(1)}-\Psi^{(2)}-\Psi^{(3)}=0$, leaving only two independent combinations, which we take to be $\Psi^{(1)}$ and $\Psi^{(2)}$, and which are the basis functions for the carrier space of a 2 -dim irreducible representation.

This rather tedious procedure can be made much faster by filling the tableaux in all the possible ways subject to the following rules: the number 1 fills the uppermost, leftmost box; and the numbers must increase down any column and to the right along any row. The number of ways this can be done is the dimension of the representation. For instance, the (21) Young diagram of $S_{3}$ generates the two tableaux with so-called standard numbering:

each corresponding to one basis function in the 2-dimensional invariant subspace carrying the (21) irrep of $S_{3}$.
There is, however, a much more convenient method for calculating the dimension of the representation associated with a Young diagram if one does not wish to construct bases for the subspaces:

Definition 2.35. For any box in the Young diagram associated with an irreducible representation, draw a straight line down to the last box in its column and all the way to the right end of the box's row. The result is called a hook and the number of boxes traversed by the hook is the hook length of this box.

Then the dimension of the associated irreducible representation is the order of $S_{N}, N$ !, divided by the product of the $N$ hook lengths for this diagram..

Definition 2.36. Irreducible representations for which the Young diagrams are the transpose of each other, ie. for which the length of each row in one is equal to the length of the corresponding column in the other, are said to be conjugate. Their dimensions are the same.

A self-conjugate irreducible representation is one for which the Young diagram and its transpose are identical.

Now it can be shown that the product of an irreducible representation with a 1-d representation is itself an irreducible representation, which may be the same (when the 1-d representation is the trivial one). This goes for their characters also. When the completely antisymmetric $\left(1^{N}\right) 1$-d representation exists, as is the case for $S_{N}$, the characters of a conjugate irreducible representation can always be written, class by class, as the product of the characters of its conjugate representation and the characters in the $\left(1^{N}\right)$ representation. Therefore, characters for the same class in conjugate representations are either identical or differ only by their sign. Characters of a self-conjugate representation in a class that has negative parity must vanish.

### 2.5.7 Character tables

A character table contains all the distinct characters for all the irreducible representations of a group. Each row contains the characters of all the group elements in a given representation, and each column contains the characters of a class of the group in all its representations.

The first row corresponds to the identity 1 -dim irreducible representation, $(N)$; all entries in that row must be 1. Also, the first column corresponds to the identity class; each entry in that column must be the dimension of the representation (ie. the trace of the identity matrix for each representation) for the corresponding row.

If we are dealing with $S_{N}$, we always have another 1 -dim irreducible antisymmetric representation (called the sign representation by mathematicians), $\left(1^{N}\right)$, whose $1 \times 1$ matrices, and characters, are the parities $\pm 1$ of its classes. We choose to place the characters of this representation in the bottom row of the table.

What about the entries which are neither in the first/last row nor in the first column? Well, we can assign some algebraic symbol to the unknown characters and then spot conjugate representations. If there are any, the character in each column of one of representation in a conjugate pair must be the character of its conjugate multiplied by the character $( \pm 1)$ in the antisymmetric 1-d row. If there are self-conjugate irreducible representations, any character sitting in the same column as a -1 in the last 1 -dim row must be zero.

In the case of $S_{N}$, extra information can be extracted from the defining representation, of dimension $N$, of the group. This always reducible representation is in fact fully reducible to a 1 -d representation and a $(N-1)$ dimensional irreducible representation. Indeed, let $\left(x_{1}, \ldots, x_{N}\right)$ be a set of coordinates in the carrier space of the defining representation. It is easy to construct a fully symmetric combination of all those coordinates:

$$
X=\frac{x_{1}+\ldots+x_{N}}{N}
$$

This function spans the 1 -dim subspace of $\mathbb{R}^{N}$ invariant under any permutation of the coordinates and is the carrier space of the irreducible representation of $S_{N}$ labelled by $(N)$. Since the defining representation is unitary, the complementary subspace is itself invariant, and is the carrier space of another irreducible representation. Indeed, let this ( $N-1$ )-dim subspace be spanned by $N-1$ functions of the mixed-symmetry form:

$$
Y_{j-1}=\frac{x_{1}+\ldots+x_{j-1}-(j-1) x_{j}}{\sqrt{j(j-1)}} \quad 2 \leq j \leq N
$$

These $N-1$ Jacobi coordinates can be shown to be linearly independent, so that there is no proper invariant subspace, and the representation is irreducible. The functions are symmetrised with respect to $j-1$ coordinates and then antisymmetrised with respect to the $N^{\text {th }}$ one. This allows us to identify the representation with the ( $N-11$ ) irreducible representation of $S_{N}$.

The characters of this ( $N-11$ ) representation can be calculated as follows. First, we note that for a class labelled $\left(\ldots 2^{\beta} 1^{\alpha}\right)$, the characters of the defining representation are equal to the number of objects that the permutations in the class leave invariant, ie. $\alpha$ (see for instance the defining representation of $S_{3}$ in section 2.4.3). Since this character must be the sum of the characters of the $(N-11)$ and of the $(N)$ irreducible representations, we find that the characters of each class labelled by $\alpha$ in the ( $N-11$ ) irreducible representation are just $\alpha-1$.

If $S_{N}$ has an $N$-dim irreducible representation, other low-hanging fruit helping to find characters involves the $(N)$ class, whose permutations shuffle all $N$ objects, and whose $N$-dimensional matrices representing them must have diagonal entries 0 , resulting in a character that is 0 .

Next, we let eq. (2.15) and (2.20) provide constraints on the remaining unknowns:

- The first says that complete rows in the table (each for a different representations) are orthogonal, with the understanding that each term in the sum is weighted by the number of elements in the class (column).
- The second says that complete columns (each belonging to different classes) are othogonal.

Now, if $\beta$ refers to the identity representation, then, for any irrep $\alpha$ other than the identity, eq. (2.15) becomes:

$$
\begin{equation*}
\sum_{k=1}^{N_{\mathrm{c}}} n_{k} \chi_{k}^{(\alpha)}=0 \tag{2.21}
\end{equation*}
$$

This is usually the best way to use eq. (2.15), at least at the beginning, because the resulting constraints are linear. Unfortunately, many of these relations will be automatically satisfied and will not yield new information, because of the strong constraints on the characters imposed by conjugation and self-conjugation of the irreducible representations. When all possible information has been extracted from eq. (2.21) and (2.20), and there still remain unknowns, one can try to spot reasonably simple quadratic relations from eq. (2.15) as well as using the normalisation of rows and columns.

Two last but important remarks: the characters of the 1 -dim representations of any group (eg. those of an Abelian group) must preserve the group product. Also, although the characters of $S_{N}$ are real, characters of other groups (eg. $Z_{n}$ ) can be complex.

There exist even more sophisticated methods for determining the characters of a group (eg. by generating them from the characters of a subgroup, or of a factor group), but lack of time and space prevents us from discussing them here. In fact, character tables for well-known groups can be found in specialised books and on the web.

Let us use these rules to find the characters of $S_{3}$ as a $3 \times 3$ table, with classes corresponding to columns and irreducible representations to rows. The first and last row can be immediately written down from our knowledge of the parity of each class ( -1 for the transpositions and +1 for the cyclic permutations). Note also that the (21) representation is self-conjugate, so we can put 0 for the character in the (21) class, because the parity of that class (last character in the column) is -1 . The (21) representation is the ( $N-11$ ) representation discussed above, and its remaining character is determined by its belonging to a class with $\alpha=0$; thus, the character must be -1 . The linear constraint (2.21), as well as the other orthogonality rules, are automatically satisfied. Collecting yields:

|  | $\left(1^{3}\right)$ | $(21)$ | $(3)$ |
| :--- | ---: | ---: | ---: |
| $n_{k}$ | 1 | 3 | 2 |
| $(3)$ | 1 | 1 | 1 |
| $(21)$ | 2 | 0 | -1 |
| $\left(1^{3}\right)$ | 1 | -1 | 1 |

EXERCISE: work out the character table and irreducible representations of $Z_{4}$, the cyclic group of order 4. You may make the task easier by remembering that products of characters belonging to a 1-d irreducible representation, which are the actual representation matrices, must mimic the group product of the corresponding elements.

## Example 2.10. Lifting of a degeneracy by a weak interaction

Consider a physical system in a rotationally-invariant potential that depends only on distance to the origin. This often occurs in quantum mechanics, and the result is that the eigenstates labelled by the integers that characterise eigenvalues of $L^{2}$ and $L_{z}, l$ and $m$, with $-l \leq m \leq l$, exhibit a $2 l+1$ fold degeneracy, in the sense that they all have the same energy. This is also manifested by the way spherical harmonics, which are eigenfunctions of $L^{2}$ and $L_{z}$ for a given value of $l$, as well as of the Hamiltonian, transform under a rotation by some angle $\alpha$. Using eq. (2.3), we have:

$$
\left[\mathcal{R}_{\alpha} Y_{l m}\right](\theta, \phi)=Y_{l m^{\prime}}(\theta, \phi)\left(D^{(l)}\right)_{m}^{m^{\prime}}(\alpha)
$$

where the $\mathbf{D}^{(l)}$ matrix is an irreducible representation of the rotation group $S O(3)$ which acts on the invariant space spanned by the $2 l+1 Y_{l m}$ for that $l$. Also, a sum over $m^{\prime}$ from $-l$ to $l$ is implied.
We can simplify things by noting that rotations by an angle $\alpha$ about any axis are all equivalent to (in the same class as) a rotation by that angle around any axis. Since we are only interested in characters, take this axis to be the $z$-axis. Then $\left[\mathcal{R}_{\alpha} Y_{l m}\right](\theta, \phi)=\mathrm{e}^{\mathrm{i} m \alpha} Y_{l m}(\theta, \phi)=Y_{l m}(\theta, \phi+\alpha)$ because the dependence of the spherical harmonics on $\phi$ is $\mathrm{e}^{\mathrm{i} m \phi}$. The $\mathbf{D}^{(l)}(\alpha)$ matrix is diagonal, with entries running from $\mathrm{e}^{-\mathrm{i} l \alpha}$ to $\mathrm{e}^{\mathrm{i} l \alpha}$, and its character is not hard to compute:

$$
\begin{equation*}
\chi^{(l)}(\alpha \neq 0)=\sum_{m=-l}^{l}\left(\mathrm{e}^{\mathrm{i} \alpha}\right)^{m}=\mathrm{e}^{-\mathrm{i} l \alpha} \sum_{n=0}^{2 l}\left(\mathrm{e}^{\mathrm{i} \alpha}\right)^{n}=\mathrm{e}^{-\mathrm{i} l \alpha}\left(\frac{1-\mathrm{e}^{\mathrm{i}(2 l+1) \alpha}}{1-\mathrm{e}^{\mathrm{i} \alpha}}\right)=\frac{\sin (l+1 / 2) \alpha}{\sin (\alpha / 2)} \tag{2.22}
\end{equation*}
$$

where we have recast the sum as a geometric series by redefining the index as $m=n-l$.
Now let us turn on a weak interaction whose corresponding potential is no longer rotationally-invariant, but still retains that invariance for rotations by a restricted, finite set of angles, which we collectively denote by $\beta$. This would happen, for instance, if we embed our spherically-symmetric atom in a crystal lattice. Suppose this restricted set of rotations actually is a group, or more precisely, a subgroup of $S O(3)$. Then the matrix $\mathbf{D}^{(l)}(\beta)$ should be a representation of that subgroup, but that representation may no longer be irreducible. This will certainly happen for any $\mathbf{D}^{(l)}$ whose dimension is too large to satisfy the sum rule (2.18) that applies to the finite subgroup.
The set of $Y_{l m}$ transform as: $\mathcal{R}_{\beta} Y_{l m}=Y_{l m^{\prime}}\left(D^{(l)}\right)^{m^{\prime}}{ }_{m}(\beta)$, with summation over repeated indices implied. If the induced representation $\mathbf{D}$ of the restricted-symmetry subgroup is reducible, there exists a matrix $\mathbf{S}$ independent of $\beta$ which transforms all its matrices into block-diagonal matrices $\mathbf{D}^{\prime}=\mathbf{S}^{-1} \mathbf{D} \mathbf{S}$, something which was impossible when there was no restriction on the angles.

But we do not have to know $\mathbf{S}$ to extract useful information. Indeed, because $\mathbf{D}$ and $\mathbf{D}^{\prime}$ have the same trace, we can calculate the characters of $\mathbf{D}^{(l)}(\beta)$ for all elements of the restricted-angle subset in $S O(3)$. Then we find the character table of the restricted-symmetry group, which is finite. If there is a row in the table that exactly matches the $S O(3)$ characters of $\mathbf{D}^{(l)}(\beta)$, then $\mathbf{D}^{(l)}(\beta)$ is not only an irreducible representation of $S O(3)$, it is also an irrep of its subgroup defined by the angles allowed by the restricted symmetry. The corresponding invariant subspaces are identical, and the original $2 l+1$-fold degeneracy for that value of $l$ is still present after the perturbation has been turned on. As $l$ increases, however, the dimension $2 l+1$ of $\mathbf{D}^{(l)}(0)$, which always appears as the first character corresponding to the identity class of $S O(3)$, will eventually exceed the fixed dimension of any irreducible representation of the subgroup. Then all the corresponding $\mathbf{D}^{(l)}(\beta)$ will be reducible to a direct sum of the irreducible representations of the subgroup, given by eq. (2.7), with the multiplicity of each irrep calculable from eq. (2.17).
For instance, suppose that the perturbation has cubic symmetry. A cube is invariant under ${ }^{1}$ :

- 6 rotations by $\pm \pi / 2$ around the three axes through its centre that intersect faces through their centre;
- 3 rotations by $\pi$ around these same axes;
- 8 rotations by $\pm 2 \pi / 3$ around the four axes through diagonally opposed corners (vertices).
- 6 rotations by $\pi$ around the six axes intersecting the centre of two diagonally opposed edges;

With the identity rotation, these add up to 24 elements forming a subgroup of $S O(3)$ isomorphic to $S_{4}$. The correspondence between rotations and permutations is obtained by considering each rotation as a shuffling of the four pairs of diagonally opposed vertices, each pair labelled 1 to 4 . The five classes of $S_{4}$ are $\left(1^{4}\right)$ ( the identity), $\left(21^{2}\right)$ ( 6 rotations by $\pi$ ), $\left(2^{2}\right)(3$ rotations by $\pi),(31)$ (rotations by $\pm 2 \pi / 3$ ), and (4) (rotations by $\pm \pi / 2$ ). The character table of $S_{4}$ is:

|  | $\left(1^{4}\right)$ | $\left(21^{2}\right)$ | $\left(2^{2}\right)$ | $(31)$ | $(4)$ |
| :--- | ---: | ---: | ---: | ---: | ---: |
|  | $n_{k}$ | 1 | 6 | 3 | 8 |

Here, the irreps of $S_{4}$ (or of the group of symmetries of the cube) are ordered by increasing dimension instead of their mixed-symmetry structure. With eq. (2.22), we calculate the characters of the representations of $S_{4}$ induced by $\mathbf{D}^{(l=1)}(\beta)$ and $\mathbf{D}^{(l=2)}(\beta)$, with angles $\beta$ running through the values corresponding to the five classes of $S_{4}$ :

|  | $\left(1^{4}\right)$ | $\left(21^{2}\right)$ | $\left(2^{2}\right)$ | $(31)$ | $(4)$ |
| :--- | ---: | ---: | ---: | ---: | ---: |
| $\mathbf{D}^{(l=1)}$ | 3 | 1 | -1 | 0 | -1 |


|  | $\left(1^{4}\right)$ | $\left(21^{2}\right)$ | $\left(2^{2}\right)$ | $(31)$ | $(4)$ |
| :--- | ---: | ---: | ---: | ---: | ---: |
| $\mathbf{D}^{(l=2)}$ | 5 | -1 | 1 | -1 | 1 |

The $l=1$ irrep of $S O(3)$ restricted to the angles allowed by the cubic-symmetry subgroup has the same dimension and the same characters as the representation (31) of $S_{4}$ in the above character table for $S_{4}$. The invariant spaces are the same and there is no lifting of the unperturbed 3 -fold degeneracy. The $l=2$ irrrep of $S O(3)$, however, has no identical row in the $S_{4}$ character table, and must correspond to areducible representation of $S_{4}$. With eq. (2.17), we calculate the following

[^12]multiplicy for each irrep of $S_{4}$ that can appear in the decomposition of $\mathbf{D}^{(l=2)}(\beta): a_{1}=a_{2}=a_{4}=0$, and $a_{3}=a_{5}=1$. Then we have the decomposition:
$$
\mathbf{D}^{((=2)}(\beta)=D_{(3)}(\beta) \oplus D_{(5)}(\beta)
$$

The unperturbed 5 -fold degeneracy of the $l=2$ states is partially lifted to become two "levels", one 3 -fold and one 2 -fold degenerate.

### 2.6 Other Examples of Symmetry Helping to Solve Problems

In what follows, I wish to illustrate with a couple of other simple examples how powerful the use of symmetry can be in the solution of many problems.
Example 2.11. First, consider something ultra-simple: two identical bodies of mass $m$ separately undergoing simple (undamped) harmonic motion due to the same external force. Let $\mathbf{X}=\binom{x_{1}}{x_{2}}$ be their displacement vector away from their respective equilibrium position. (We have assumed that both objects lie some distance away from one another on the $x$ axis.) This vector is a solution of Newton's $2^{\text {nd }}$ Law for the system, $\ddot{\mathbf{X}}=-\mathbf{M}^{-1} \mathbf{K X}$, where:

$$
\mathbf{M}=\left(\begin{array}{cc}
m & 0 \\
0 & m
\end{array}\right), \quad \mathbf{K}=\left(\begin{array}{cc}
k & 0 \\
0 & k
\end{array}\right)
$$

with $k$ the restoring constant associated with the motion. We call $\mathbf{M}^{-1} \mathbf{K}$ the dynamical matrix of the system. Of course, as we all know, both bodies oscillate at the same frequency $\omega_{0}=\sqrt{k / m}$. The space of solutions is spanned by the two eigenvectors $\binom{1}{1}$ and $\binom{1}{-1}$, both with the same eigenvalue $\omega_{0}$. This is a two-fold degeneracy.

Now suppose that we introduce an internal interaction which is linear in the bodies' displacement (ie., usually, it is weak enough). We then say that the two bodies are weakly coupled.

The dynamical matrix becomes:

$$
\mathbf{M}^{-1} \mathbf{K}=\left(\begin{array}{cc}
\omega_{0}^{2}+\omega_{1}^{2} & -\omega_{1}^{2} \\
-\omega_{1}^{2} & \omega_{0}^{2}+\omega_{1}^{2}
\end{array}\right)
$$

where $\omega_{1}$ parametrises the coupling.
This coupled system retains a "mirror" symmetry, in the sense that, under reflection about a perpendicular plane halfway between the rest positions of the bodies, $\mathbf{X} \rightarrow \mathbf{X}^{\prime}=-\binom{x_{2}}{x_{1}}$. Indeed, we write in matrix form:

$$
\mathbf{X}^{\prime}=\mathbf{S} \mathbf{X} \quad \mathbf{S}=\left(\begin{array}{rr}
0 & -1 \\
-1 & 0
\end{array}\right)
$$

Apply $S$ to Newton's $2^{\text {nd }}$ Law. In matrix form:

$$
\ddot{\mathbf{X}}^{\prime}=\mathbf{S} \ddot{\mathbf{X}}=-\mathbf{S} \mathbf{M}^{-1} \mathbf{K} \mathbf{X}=-\left(\mathbf{S ~ M}^{-1} \mathbf{K} \mathbf{S}^{-1}\right)(\mathbf{S X})
$$

that is: $\ddot{\mathbf{X}}^{\prime}=-\left(\mathbf{S} \mathbf{M}^{-1} \mathbf{K} \mathbf{S}^{-1}\right) \mathbf{X}^{\prime}$. But, if Newton's $2^{\text {nd }}$ Law is to be invariant under reflection, $\ddot{\mathbf{X}}^{\prime}=-\mathbf{M}^{-1} \mathbf{K} \mathbf{X}^{\prime}$. This forces $\mathbf{S M}^{-1} \mathbf{K}=\mathbf{M}^{-1} \mathbf{K S}$. In operator language, because $\mathbf{S}$ and $\mathbf{M}^{-1} \mathbf{K}$ commute, they must have a common set of eigenvectors.

Call these eigenvectors $\mathbf{A}$. Finding the eigenvalues and eigenvectors of $\mathbf{M}^{-1} \mathbf{K}$ may be hard; finding the eigenvalues and eigenvectors of $\mathbf{S}$ is not. We know that $\mathbf{S}^{2}=\mathbf{I}$, so that if $\mathbf{S A}=\lambda \mathbf{A}, \lambda^{2}=1$. Then $\mathbf{A}_{+}=\binom{1}{-1}$, with eigenvalue $\lambda_{+}=1$, and $\mathbf{A}_{-}=\binom{1}{1}$, with eigenvalue $\lambda_{-}=-1$. Now it is a simple matter, knowing the eigenvectors of $\mathbf{M}^{-1} \mathbf{K}$, to find the normal frequencies of the system, ie. the eigenvalues $\omega_{ \pm}$of the $\mathbf{M}^{-1} \mathbf{K}$ operator, which satisfy $\mathbf{M}^{-1} \mathbf{K} \mathbf{A}_{ \pm}=\omega_{ \pm}^{2} \mathbf{A}_{ \pm}$. These eigenvectors are called normal modes; the entries of a normal-mode vector are the relative amplitudes of the motion for each body when the system oscillates in that particular mode.

Their importance lies in the fact that any displacement vector $\mathbf{X}$ of the system must be a linear combination of its normal modes.

Using group-theoretic language, $\mathbf{S}$ forms the regular 2-dim representation of the group of reflections of two objects about the plane, which has the same group-composition table as $S_{2}$. This can be reduced to 2 1-dim (irreducible) representations: 1 and -1 . Notice how the coupling has broken the initial degeneracy of the uncoupled system: now $\omega_{-}=\omega_{0}$ and $\omega_{+}^{2}=\omega_{0}^{2}+2 \omega_{1}^{2}$, with one eigenvector for each.
Example 2.12. This time, we take six identical bodies arranged on a circle $60^{\circ}$ apart. Each is subject to an identical external linear restoring force giving rise to small oscillations about their equilibrium position and tangent to the circle. Each are coupled by an identical weak interaction to their nearest neighbours $\pm 60^{\circ}$ away; similarly, they are coupled to their second next neighbours, $\pm 120^{\circ}$ away, by another (even weaker) interaction that is identical for both these neighbours; finally, a third (weakest) interaction couples each one to its opposite counterpart, $180^{\circ}$ away. We wish to study the effect of the coupling on the motion of the bodies tangent to the circle.

Because of the symmetry of the interactions and of the system, the dynamical $\mathbf{M}^{-1} \mathbf{K}$ matrix must have the form:

$$
\mathbf{M}^{-1} \mathbf{K}=\left(\begin{array}{rrrrrr}
\omega_{0}^{2} & -\omega_{1}^{2} & -\omega_{2}^{2} & -\omega_{3}^{2} & -\omega_{2}^{2} & -\omega_{1}^{2} \\
-\omega_{1}^{2} & \omega_{0}^{2} & -\omega_{1}^{2} & -\omega_{2}^{2} & -\omega_{3}^{2} & -\omega_{2}^{2} \\
-\omega_{2}^{2} & -\omega_{1}^{2} & \omega_{0}^{2} & -\omega_{1}^{2} & -\omega_{2}^{2} & -\omega_{3}^{2} \\
-\omega_{3}^{2} & -\omega_{2}^{2} & -\omega_{1}^{2} & \omega_{0}^{2} & -\omega_{1}^{2} & -\omega_{2}^{2} \\
-\omega_{2}^{2} & -\omega_{3}^{2} & -\omega_{2}^{2} & -\omega_{1}^{2} & \omega_{0}^{2} & -\omega_{1}^{2} \\
-\omega_{1}^{2} & -\omega_{2}^{2} & -\omega_{3}^{2} & -\omega_{2}^{2} & -\omega_{1}^{2} & \omega_{0}^{2}
\end{array}\right)
$$

How can we use the symmetry to find the normal modes of the system? By recognising that the system must be invariant under $60^{\circ}$ rotations. This operation is isomorphic to a cyclic permutation: $(123456) \in Z_{6}$. The regular representation matrix for this element of $Z_{6}$ looks like:

$$
\mathbf{S}=\left(\begin{array}{llllll}
0 & 1 & 0 & 0 & 0 & 0 \\
0 & 0 & 1 & 0 & 0 & 0 \\
0 & 0 & 0 & 1 & 0 & 0 \\
0 & 0 & 0 & 0 & 1 & 0 \\
0 & 0 & 0 & 0 & 0 & 1 \\
1 & 0 & 0 & 0 & 0 & 0
\end{array}\right)
$$

Invariance under $\mathbf{S}$ means that $\mathbf{M}^{-1} \mathbf{K}$ and $\mathbf{S}$ commute. In fact, this last statement can be used to obtain the form of the $\mathbf{M}^{-1} \mathbf{K}$ matrix given above.

Now, as in the last example, the eigenvectors of $\mathbf{S}$ satisfy $\mathbf{S A}=\lambda \mathbf{A}$. But since $\mathbf{S}^{6}=\mathbf{I}$, we immediately find that the eigenvalues are the sixth roots of 1 , as expected for the cyclic group. Therefore. $\lambda_{(m)}=\mathrm{e}^{\mathrm{i} m \pi / 3},(0 \leq m \leq$ 5). To each value of $m$ corresponds an eigenvector $\mathbf{A}_{(m)}$ with components $A_{(m)}^{j}=\lambda_{(m)}^{j-1}=\mathrm{e}^{\mathrm{i} m(j-1) \pi / 3}$.

As in Example 2.11, these eigenvectors are also the normal modes of the system. Inserting into the eigenvalue equation $\mathbf{M}^{-1} \mathbf{K} \mathbf{A}_{(m)}=\omega_{(m)}^{2} \mathbf{A}_{(m)}$ with the coupling parameters $\omega_{(5)}=\omega_{(1)}$ and $\omega_{(4)}=\omega_{(2)}$ yields the dispersion relation:

$$
\omega_{(m)}^{2}=\sum_{j=1}^{6} \omega_{j-1}^{2} \mathrm{e}^{\mathrm{i} m(j-1) \pi / 3}=\omega_{0}^{2}-2 \omega_{1}^{2} \cos m \pi / 3-2 \omega_{2}^{2} \cos 2 m \pi / 3-(-1)^{m} \omega_{3}^{2}
$$

We note that $\mathbf{A}_{(1)}^{*}=\mathbf{A}_{(5)}$, and $\mathbf{A}_{(2)}^{*}=\mathbf{A}_{(4)}$. These modes are complex, which is a problem if they are supposed to correspond to real relative amplitudes. But we also note that $\omega_{(1)}=\omega_{(5)}$, and $\omega_{(2)}=\omega_{(4)}$; therefore, the corresponding eigenvectors span two invariant 2-dim subspaces, which allows us to take appropriate linear combinations of the eigenvectors to turn them into real modes of the same frequency.

Again, the coupling has lifted the original 6-fold degeneracy of the uncoupled system, but there is still some degeneracy left because of the two 2-dim subspaces.

This is as far as we can go without knowing the interaction parameters themselves. But we have succeeded in nailing down the relative amplitudes of motion of the bodies in each normal mode without that explicit knowledge!

### 2.7 APPENDIX - The Right and Left Actions of a Group on a Vector, with $S_{n}$ as Example

First, we recall some important properties of linear transformations. For simplicity we shall take these transformations to act on vectors $\mathbf{x}=x^{i} \mathbf{e}_{\mathbf{i}} \in \mathbb{R}^{n}$, with the standard basis $\left\{\mathbf{e}_{\mathrm{i}}\right\}$. It is customary to distinguish between active transformations: $\mathbf{x} \longrightarrow \mathbf{x}^{\prime}=x^{\prime i} \mathbf{e}_{\mathbf{i}}=\mathbf{e}_{\mathbf{i}}\left(A^{i}{ }_{j} x^{j}\right)$, where the transformed coordinates $x^{\prime i}$ are those of a new vector; and passive transformations that leave $\mathbf{x}$ invariant but transform the basis to $\left\{\mathbf{e}_{\mathrm{i}}^{\prime}\right\}$ : $\mathbf{e}_{\mathbf{i}}^{\prime}=\mathbf{e}_{\mathbf{j}} P^{j}{ }_{i}$.

Paaive transformations, unlike active ones, require the transformed vector $\mathrm{x}^{\prime}$ to remain the same as the initial one, because all we have done is change the basis. To preserve $\mathbf{x}$, we must also transform its components with the inverse transformation:

$$
\mathbf{x}^{\prime}=\left(\mathbf{e}_{\mathbf{j}}^{\prime} x^{\prime j}\right)_{\mathrm{P}}=\mathbf{e}_{\mathbf{i}} P^{i}{ }_{j}\left(P^{-1}\right)^{j}{ }_{k} x^{k}=\mathbf{e}_{\mathbf{i}} \delta^{i}{ }_{k} x^{k}=\mathbf{e}_{\mathbf{i}} x^{i}=\mathbf{x}
$$

Now, although an active transformation is quite different from a passive one, the result should be mathematically equivalent to the result of transforming the basis with the inverse transformation. In other words, the components, $\left(x_{\mathrm{A}}^{\prime}\right)^{j}$, of the new vector produced by the active trnsformation should be the same as the those of the initial vector in the transformed basis, that is, $\left(x_{\mathrm{A}}^{\prime}\right)^{j}=\left(x_{\mathrm{P}}^{\prime}\right)^{j}=\left(P^{-1}\right)^{j}{ }_{k} x^{k}$. Comparing with $\left(x_{\mathrm{A}}^{\prime}\right)^{j}=A^{j}{ }_{k} x^{k}$, we see that that the passive transformation is indeed inverse to the active one.

### 2.7.1 Right action

This is often the easiest to work with. According to definition 2.25 directly applied as in eq. (2.5), $T_{\mathrm{g}}^{R}(\mathrm{x}):=\mathrm{xg}$. Now $\mathbf{x}=x^{i} \mathbf{e}_{i}$, and when it comes to representation matrices, we have the choice between acting on components $x^{i}$ and acting on basis vectors $\mathbf{e}_{i}$;

$$
\mathbf{x g}=\mathbf{e}_{i} x^{i} \mathrm{~g}=\mathbf{e}_{\mathrm{i}}\left[\left(D_{\mathrm{g}}^{\mathrm{R}}\right)_{j}^{i} x^{j}\right] \quad \begin{cases} & \Longleftrightarrow \\ \quad x^{i} \mathrm{~g}=\left(D_{\mathrm{g}}^{\mathrm{R}}\right)^{i}{ }_{j} x^{j} \\ & \Longleftrightarrow \\ \mathbf{e}_{\mathrm{i}} \mathrm{~g}=\mathbf{e}_{\mathrm{j}}\left(D_{\mathrm{g}-1}^{\mathrm{R}}\right)_{i}^{j}\end{cases}
$$

Notice that the action on the basis vectors involves the matrix for $\mathrm{g}^{-1}$, as expected for the passive transformation associated with the active one on the components.

## (a) Components

For $\mathrm{g}=\pi_{k} \in S_{n}$, the expression for $x^{i} \pi_{k}$ is equivalent to $x^{\pi_{k}(i)}=\left(D_{\mathrm{g}}^{\mathrm{R}}\right)^{i}{ }_{j} x^{j}$, so that $\left(D_{\mathrm{g}}^{\mathrm{R}}\right)^{i}{ }_{j}=\delta^{\pi_{k}(i)}{ }_{j}$, and the $i^{\text {th }}$ row of the $n$-dim $\mathbf{D}_{\mathrm{g}}^{\mathrm{R}}$ matrix is the standard-basis vector $\mathbf{e}_{\pi_{k}(i)}$. For $\pi_{k}=\pi_{5}=(132) \in S_{3}$, for instance, we find: $\left[\left(D_{\pi_{5}}^{\mathrm{R}}\right)_{1},\left(D_{\pi_{5}}^{\mathrm{R}}\right)^{1}{ }_{2},\left(D_{\pi_{5}}^{\mathrm{R}}\right)^{1}{ }_{3}\right]=(0,0,1)$, and for $\pi_{6}=(123):\left[\left(D_{\pi_{6}}^{\mathrm{R}}\right)_{1},\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{1}{ }_{2},\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{1}{ }_{3}\right]=$ $(0,1,0)$, etc. There comes:

$$
\begin{aligned}
& x^{i} \pi_{2}=\left(\begin{array}{lll}
0 & 1 & 0 \\
1 & 0 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{l}
x^{1} \\
x^{2} \\
x^{3}
\end{array}\right)=\left(\begin{array}{l}
x^{2} \\
x^{1} \\
x^{3}
\end{array}\right) \quad x^{i} \pi_{3}=\left(\begin{array}{lll}
0 & 0 & 1 \\
0 & 1 & 0 \\
1 & 0 & 0
\end{array}\right)\left(\begin{array}{l}
x^{1} \\
x^{2} \\
x^{3}
\end{array}\right)=\left(\begin{array}{lll}
x^{3} \\
x^{2} \\
x^{1}
\end{array}\right) \quad x^{i} \pi_{4}=\left(\begin{array}{lll}
1 & 0 & 0 \\
0 & 0 & 1 \\
0 & 1 & 0
\end{array}\right)\left(\begin{array}{l}
x^{1} \\
x^{2} \\
x^{3}
\end{array}\right)=\left(\begin{array}{l}
x^{1} \\
x^{3} \\
x^{2}
\end{array}\right) \\
& x^{i} \pi_{5}=\left(\begin{array}{lll}
0 & 0 & 1 \\
1 & 0 & 0 \\
0 & 1 & 0
\end{array}\right)\left(\begin{array}{l}
x^{1} \\
x^{2} \\
x^{3}
\end{array}\right)=\left(\begin{array}{l}
x^{3} \\
x^{1} \\
x^{2}
\end{array}\right) \quad x^{i} \pi_{6}=\left(\begin{array}{lll}
0 & 1 & 0 \\
0 & 0 & 1 \\
1 & 0 & 0
\end{array}\right)\left(\begin{array}{l}
x^{1} \\
x^{2} \\
x^{3}
\end{array}\right)=\left(\begin{array}{l}
x^{2} \\
x^{3} \\
x^{1}
\end{array}\right)
\end{aligned}
$$

where $\mathbf{D}_{\pi_{5}}^{\mathrm{R}}=\left(\mathbf{D}_{\pi_{6}}^{\mathrm{R}}\right)^{-1}$, and vice-versa, as expected from the Cayley table of $S_{3}$.

## (b) Basis vectors

Now the $\pi_{k}$ shuffle the basis vectors themselves (not components). The expression for $\mathbf{e}_{\mathrm{i}} \pi_{k}$ is not matrix multiplication! Indeed, $\mathbf{e}_{\mathrm{i}} \pi_{k}=\mathbf{e}_{\mathrm{j}}\left(D_{\pi_{k}^{-1}}^{\mathrm{R}}\right)^{j}$ gives the entries of the $i^{\text {th }}$ column of the $n$-dim $\mathbf{D}_{\mathrm{g}^{-1}}^{\mathrm{R}}$ matrix in the standard basis: $\mathbf{e}_{\mathrm{i}} \pi_{k}=\left[\left(D_{\pi_{k}^{-1}}^{\mathrm{R}}\right)^{1}, \ldots,\left(D_{\pi_{k}^{-1}}^{\mathrm{R}}\right)^{n}{ }_{i}\right]$. More succinctly, because $\mathbf{e}_{\pi_{k}(i)}=\mathbf{e}_{\mathbf{i}} \pi_{k}$, we get: $\left(D_{\pi_{k}^{-1}}^{\mathrm{R}}\right)_{i}^{j}=\delta^{j}{ }_{\pi_{k}^{-1}(i)}$, or the.$i^{\text {th }}$ column is the standard-basis ector $\mathbf{e}_{\pi_{k}(i)}$.

Using basis vectors to find $\mathbf{D}_{\pi_{5}}^{\mathrm{R}}$, for instance, we have: $\mathbf{e}_{1} \pi_{5}=\mathbf{e}_{3}=(0,0,1)=\left[\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{1}{ }_{1},\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{2}{ }_{1},\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{3}{ }_{1}\right]$, and $\mathbf{e}_{2} \pi_{5}=\mathbf{e}_{1}=(1,0,0)=\left[\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{1}{ }_{2},\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{2}{ }_{2},\left(D_{\pi_{6}}^{\mathrm{R}}\right)^{3}{ }_{2}\right]$, etc. The resulting matrix is the one that was obtained from the components for the right action of $\pi_{6}$, and its inverse will be the matrix for the right action of $\pi_{5}$, the same that we found somewhat more directly by acting on components.

### 2.7.2 Left action

Under the left action of g , we have: $T_{\mathrm{g}}^{\mathrm{L}}(\mathrm{x})=\mathrm{g}^{-1} \mathrm{x}$. Once again this can act on basis vectors or on components:

$$
\mathrm{g}^{-1} \mathbf{x}=x^{i} \mathrm{~g}^{-1} \mathbf{e}_{\mathbf{i}}=x^{i}\left[\mathbf{e}_{\mathrm{j}}\left(D_{\mathrm{g}^{-1}}^{\mathrm{L}}\right)^{j}{ }_{i}\right] \quad\left\{\begin{array}{l}
\Longleftrightarrow \\
\Longleftrightarrow \mathrm{g}^{-1} \mathbf{e}_{\mathbf{i}}=\mathbf{e}_{\mathrm{j}}\left(D_{\mathrm{g}^{-1}}^{\mathrm{L}}\right)^{j}{ }_{i}=\mathbf{e}_{\mathrm{j}}\left(D_{\mathrm{g}}^{\mathrm{L}}\right)_{i}{ }^{j} \\
\Longleftrightarrow \\
\mathrm{~g}^{-1} x^{i}=\left(D_{\mathrm{g}^{-1}}^{\mathrm{L}}\right)_{j}^{i} x^{j}=x^{j}\left(D_{\mathrm{g}}^{\mathrm{L}}\right)_{j}^{i}
\end{array}\right.
$$

where the last equalities on the right hold when the matrices are orthogonal (eg., rotations). In that case we can find the left-action matrix for g directly

## (a) Basis vectors

In the case of $\mathrm{g}=\pi_{k} \in S_{n}$, we have: $\pi_{k}^{-1} \mathbf{e}_{\mathrm{i}}=\mathbf{e}_{\pi_{k}^{-1}(i)}=\left[\left(D_{\pi_{k}}^{\mathrm{L}}\right)_{i}{ }^{1},\left(D_{\pi_{k}}^{\mathrm{L}}\right)_{i}{ }^{2}, \ldots,\left(D_{\pi_{k}}^{\mathrm{L}}\right)_{i}{ }^{n}\right]$ as the $i^{\text {th }}$ row of $\mathbf{D}_{\pi_{k}}^{\mathrm{L}}$. In short $\left(D_{\pi_{k}}^{\mathrm{L}}\right)_{i}{ }^{j}=\delta_{\pi_{k}^{-1}(i)}{ }^{j}$. For instance, taking as before $\pi_{k}=\left(\pi_{5}, \pi_{6}\right) \in S_{3}$ immediately leads to:

$$
\mathbf{D}_{\pi_{5}}^{\mathrm{L}}=\left(\begin{array}{ccc}
0 & 1 & 0 \\
0 & 0 & 1 \\
1 & 0 & 0
\end{array}\right) \quad \mathbf{D}_{\pi_{6}}^{\mathrm{L}}=\left(\begin{array}{ccc}
0 & 0 & 1 \\
1 & 0 & 0 \\
0 & 1 & 0
\end{array}\right)
$$

These are the $\mathbf{D}^{\mathrm{L}}$ matrices found in example 2.6. The left and right representations for each single transposition are identical.

## Components

In the case of $\mathrm{g}=\pi_{k} \in S_{n}$, we have: $x^{\pi_{k}(i)}=x^{j}\left(D_{\pi_{k}}^{\mathrm{L}}\right)_{j}{ }^{i}$, so that $\left(D_{\pi_{k}}^{\mathrm{L}}\right)_{j}{ }^{i}=\delta_{j}{ }^{\pi_{k}(i)}$. Working with our trusted friends $\left(\pi_{5}, \pi_{6}\right) \in S_{3}$, we obtain:

$$
\mathbf{D}_{\pi_{5}}^{\mathrm{L}}=\left(\begin{array}{ccc}
0 & 1 & 0 \\
0 & 0 & 1 \\
1 & 0 & 0
\end{array}\right) \quad \mathbf{D}_{\pi_{6}}^{\mathrm{L}}=\left(\begin{array}{ccc}
0 & 0 & 1 \\
1 & 0 & 0 \\
0 & 1 & 0
\end{array}\right)
$$

## 3 MODULE III - GROUP THEORY 2: Lie Groups

### 3.1 Definitions

In this module we focus on a class of groups with an infinite number of elements. As groups, they of course satisfy the algebraic properties of a group as set out in definition 2.1. But we shall put in an extra requirement: that each group element $\mathrm{g}_{p}$ be in correspondence with a point $P$ in some manifold. This means that the index " $p$ " can in fact be taken as a set of continuous, real variables, and we write a group element as $\mathrm{g}(P)$ and say that the manifold parametrises the group. More precisely:

Definition 3.1. Let $P$ be any point in a $n$-dim manifold $M^{n}$ which is obtained from two other points, $P_{1}$ and $P_{2}$ from a mapping $P=\phi\left(P_{1}, P_{2}\right)$. Let $\mathrm{g}\left(P_{1}\right) \star \mathrm{g}\left(P_{2}\right)=\mathrm{g}(P)$ be the group product of an infinite group $G$. If the map $\phi$ and its inverse are differentiable, then $G$ is a Lie group.
The important point to remember here is that since they correspond to points in a manifold, elements of a Lie group can be parametrised in terms of smooth coordinates on this manifold.
A Lie group is real if its manifold is real and complex if its manifold is complex.
The dimension of a Lie group is the dimension of its manifold.
Definition 3.2. A Lie group is connected if any pair of points on its manifold is connected by a continuous path.

The subset of all elements in a Lie group whose corresponding points in $M^{n}$ are connected by a continuous path to the identity must be a subgroup of the Lie group. Thus, even if it is itself not connected, a Lie group must contain a connected subgroup.
Example 3.1. An infinite line with a coordinate patch $-\infty<x<\infty(x \in \mathbb{R})$ is a 1-dim manifold. In section 2.1.1 we stated that $\mathbb{C}$ was a continuous group under addition. So is $\mathbb{R}$ itself, and if we write a group element as $\mathrm{g}(x)=\mathrm{e}^{x}$, we can easily deduce the function corresponding to the group product. Indeed, $\mathrm{g}(z)=\mathrm{g}(x) \star \mathrm{g}(y)=\mathrm{g}(x+y)$, and we are not surprised to find that: $\phi(x, y)=x+y$.
Example 3.2. What if we restrict $\theta=x \in \mathbb{R}$ so that $0 \leq \theta<2 \pi$ ? We can define group elements $\mathrm{g}(\theta)=\mathrm{e}^{\mathrm{i} \theta}$ with the group product:

$$
\mathrm{g}\left(\theta_{1}\right) \star \mathrm{g}\left(\theta_{2}\right)=\mathrm{g}\left(\theta_{1}+\theta_{2} \quad \bmod 2 \pi\right)
$$

The group manifold here is the unit circle, $S^{1}$, with each point on the circle parametrised by its polar angle $\theta$, and $\phi\left(\theta_{1}, \theta_{2}\right)=\theta_{1}+\theta_{2}$. Note that this manifold is real, so that the group is real even though its elements are complex! It is Abelian, and connected.
Example 3.3. Real invertible $2 \times 2$ matrices form a group whose elements can be written as $\mathbf{g}(x)=$ $\left(\begin{array}{ll}x_{1} & x_{2} \\ x_{3} & x_{4}\end{array}\right)$. If we impose the condition that the matrices have determinant 1 , we see that the number of parameters is lowered by 1 . We compute the set of three functions $z_{i}=\phi_{i}\left(x_{1}, x_{2}, x_{3}, y_{1}, y_{2}, y_{3}\right)$ consistent with the group product:

$$
\left(\begin{array}{cc}
x_{1} & x_{2} \\
x_{3} & \frac{1+x_{2} x_{3}}{x_{1}}
\end{array}\right)\left(\begin{array}{cc}
y_{1} & y_{2} \\
y_{3} & \frac{1+y_{2} y_{3}}{y_{1}}
\end{array}\right)=\left(\begin{array}{cc}
z_{1} & z_{2} \\
z_{3} & \frac{1+z_{2} z_{3}}{z_{1}}
\end{array}\right)
$$

We obtain:

$$
z_{1}=x_{1} y_{1}+x_{2} y_{3} \quad z_{2}=x_{1} y_{2}+x_{2} \frac{1+y_{2} y_{3}}{y_{1}} \quad z_{3}=x_{3} y_{1}+y_{3} \frac{1+x_{2} x_{3}}{x_{1}}
$$

In this particular parametrisation, the mappings $\phi_{i}$ are all differentiable only off the $x_{1}=0$ and $y_{1}=0$ planes. Whatever the manifold for the group is-see later-it cannot be covered with just this coordinate patch.
The inverse mapping corresponding to $\mathrm{g}(x)^{-1}$ can be read off the inverse matrix $\mathbf{g}^{-1}$.

Example 3.4. If we demand instead that the general complex $2 \times 2$ matrices be not only unimodular, but unitary as well, the treatment is simpler. Introduce the parametrisation:

$$
\left(\begin{array}{cc}
z & w \\
-w^{*} & z^{*}
\end{array}\right)=\left(\begin{array}{cc}
a_{0}+\mathrm{i} a_{3} & a_{2}+\mathrm{i} a_{1} \\
-\left(a_{2}-\mathrm{i} a_{1}\right) & a_{0}-\mathrm{i} a_{3}
\end{array}\right)
$$

with the condition $|z|^{2}+|w|^{2}=a_{0}^{2}+a_{1}^{2}+a_{2}^{2}+a_{3}^{2}=1$ which guarantees that the matrix is unitary with determinant equal to 1 . The group manifold is thus the unit 3 -sphere $S^{3}$ embedded in $\mathbb{R}^{4}$ with coordinates $\left(a_{0}, a_{1}, a_{2}, a_{3}\right) \in \mathbb{R}$; this is a real three-dimensional Lie group as was the last example.

Definition 3.3. A Lie group is compact when the volume of its manifold is finite.

### 3.2 Some Matrix Lie Groups

Amazingly enough, it turns out that almost all Lie groups of interest in physics, the so-called classical Lie groups, are either matrix groups or groups of transformations isomorphic to matrix groups. The only group product we ever have to consider is matrix multiplication, and inverse elements are just inverse matrices.

One useful and satisfying way of classifying Lie groups is to begin with $n \times n$ invertible matrices over some field $\mathbb{F}$ of numbers, the general linear group $G L(n, \mathbb{F})$, and then identify interesting subgroups by imposing constraints on the group elements of $G L(n, \mathbb{F})$. Some of these have already been mentioned in section 2.2.2. For a start, $G L(n, \mathbb{R}) \subset G L(n, \mathbb{C})$.

Following standard usage, we introduce the diagonal matrix $\mathbf{I}_{p}^{q}$ with $p$ elements equal to +1 and $q$ entries equal to -1 , where $p+q=n$. In this notation, $\mathbf{I}_{n}$ is the $n$-dim identity matrix, and also the orthonormal metric in Euclidean $n$-dim space.

### 3.2.1 Groups obtained from linear constraints

A host of matrix subgroups are obtained by imposing linear constraints on their elements so that they have some block structure (eg. upper or lower triangular block form), and perhaps further demanding that some of the diagonal elements be 1 . We shall not consider these here.

### 3.2.2 Bilinear or quadratic constraints: the metric (or distance)-preserving groups

Definition 3.4. The set of unitary transformations $\mathbf{T}$ of a complex matrix $\mathbf{M} \in G L(n, \mathbb{C})$ is defined by:

$$
\mathbf{M} \mapsto \mathbf{T M T}^{\dagger}
$$

where the subgroup of matrices $\mathbf{T}$ leaves $\mathbf{M}=\mathbf{I}_{n}$ (the Cartesian $n$-dim metric) invariant:

$$
\mathbf{T} \mathbf{I}_{n} \mathbf{T}^{\dagger}=\mathbf{T} \mathbf{T}^{\dagger}=\mathbf{I}_{n}
$$

that is, those matrices for which $\mathbf{T}^{-1}=\mathbf{T}^{\dagger}$, and we call that subgroup $U(n) \subset G L(n, \mathbb{C})$ : Both $U(n)$ and its matrices are unitary. Example 3.2 referred to $U(1)$.
Definition 3.5. The set of orthogonal transformations $\mathbf{T}$ of a real matrix $\mathbf{M} \in G L(n, \mathbb{R})$ is defined by:

$$
\mathbf{M} \mapsto \mathbf{T M}^{\mathrm{T}}
$$

( $\mathbf{T}^{\mathrm{T}}$ is the transpose of $\mathbf{T}$ ), such that $\mathbf{T}$ leave $\mathbf{I}_{n}$ invariant:

$$
\mathbf{T} \mathbf{I}_{n} \mathbf{T}^{\mathrm{T}}=\mathbf{T} \mathbf{T}^{\mathrm{T}}=\mathbf{I}_{n}
$$

that is, those matrices for which $\mathbf{T}^{-1}=\mathbf{T}^{\mathrm{T}}$. The group of such matrices is called $O(n)$ and is orthogonal.

Be aware that $n$ in $O(n)$ or $U(n)$ refers to the dimension of the matrices, not that of the group, which is the number of coordinates on its manifold! The $O(n)$ matrices have determinant $\pm 1$, whereas the absolute value of the complex determinant of $U(n)$ matrices is equal to 1 . Thus, (can you see why?) $O(n)$ is not connected. Neither, for that matter, is $U(n)$.

Finally, the symplectic group $S p(2 n, \mathbb{R}) \subset G L(2 n, \mathbb{R})$ leaves invariant the antisymmetric "metric" matrix $\left(\begin{array}{cc}0 & \mathbf{I}_{n} \\ -\mathbf{I}_{n} & 0\end{array}\right)$. The classical Hamiltonian equations of motion are form-invariant (covariant) under symplectic transformations.

The group manifolds (and thus these groups themselves) are compact because their matrices define closed, bounded subsets of the manifolds that parametrise $G L(n, \mathbb{C})$ and $G L(n, \mathbb{R}) . O(n)$ and $U(n)$ preserve the length (or norm) of $n$-vectors in Euclidean $\mathbb{R}^{n}$, and therefore also angles between those vectors (eg., the angles of any triangle are determined by the lengths of its sides).

We also have the non-compact groups which preserve the indefinite metric ${ }_{p}^{q}$, defined by the transformations:

$$
\begin{align*}
\mathbf{T} \mathbf{I}_{p}^{q} \mathbf{T}^{\mathrm{T}} & =\mathbf{I}_{p}^{q} & & O(p, q)  \tag{3.1}\\
\mathbf{T} \mathbf{I}_{p}^{q} \mathbf{T}^{\dagger} & =\mathbf{I}_{p}^{q} & & U(p, q) \tag{3.2}
\end{align*}
$$

A famous example is $O(3,1)$, aka the full Lorentz group, that leaves the pseudo-Euclidean Cartesian metric on $\mathbb{R}^{4}$ (or space-time distance) invariant; equivalently, the norm of a 4 -vector $\mathbf{x}$ is left invariant by 3 -dim rotations, Lorentz transformations (boosts), and space or time reflections. In principle, from the condition: $\mathbf{T} \mathbf{I}_{3}^{1} \mathbf{T}^{\mathrm{T}}=\mathbf{I}_{3}^{1}$, one could work out detailed constraints on the elements of the $O(3,1)$ matrices to find that there are six independent parameters, but this would be needlessly messy. There are far better ways of parametrising the group to extract all this information, and much more, as we shall see.

### 3.2.3 Multilinear constraints: the special linear groups

The special linear subgroups $S L(n, \mathbb{C}) \subset G L(n, \mathbb{C})$ and $S L(n, \mathbb{R}) \subset G L(n, \mathbb{R})$ contain matrices with determinant 1 . This last constraint is $n$-linear.

Example 3.3 actually referred to $S L(2, \mathbb{R})$ with the constraint $x_{1} x_{4}-x_{2} x_{3}=1$, a bilinear constraint. $S L(n, \mathbb{R})$ is often referred to as the volume-preserving group in $\mathbb{R}^{n}$. But it does not preserve all lengths!
Important subgroups of the special linear groups and the metric-preserving groups may be obtained as their intersections. Thus, $S O(n)=O(n) \cap S L(n, \mathbb{R})$ and $S U(n)=U(n) \cap S L(n, \mathbb{C})$. These groups are compact. The group of example 3.4 was $S U(2)$.

A little earlier, we stated that $O(n)$ and $U(n)$ are not connected, but we know that they must have connected subgroups, ie. groups with elements connected to the identity by a continuous path. These are $S O(n)$ and $S U(n)$.

### 3.2.4 Groups of transformations

Continuous transformations in physics act on vectors, or on functions of vectors. These transformations belong to groups which are usually isomorphic to Lie groups.

Here, instead of passive transformations as in Module II, we will use the more usual active transformations, which map a vector into a new vector in the same basis.

1. Translations Let $T_{\mathbf{a}}$ act on $\mathbf{x} \in \mathbb{R}^{3}$ so that $T_{\mathbf{a}} \mathbf{x}=\mathbf{x}+\mathbf{a}$, where $\mathbf{a} \in \mathbb{R}^{3}$. This can be generalised; indeed, let $f$ be an analytic function acting on $\mathbb{R} \times \mathbb{R}^{n}$. The left action on $f$ of the operator $\mathcal{T}_{\mathbf{a}}$ associated with $T_{\mathbf{a}}$ is:

$$
\left[\mathcal{T}_{\lambda \mathbf{a}} f\right](\mathbf{x})=f\left(T_{\lambda \mathbf{a}}^{-1} \mathbf{x}\right)=f(\mathbf{x}-\lambda \mathbf{a}) \quad \mathbf{a} \in \mathbb{R}^{n}, \lambda \in \mathbb{R}
$$

These are translations in the direction of $\mathbf{a}$; they form a group which is isomorphic to $G=\{\mathbb{R},+\}$, the addition being on the parameter $\lambda$.

## 2. Rotations

Parametrise 2-dim rotations in the $z=0$ plane of a Cartesian coordinate system by $R_{\alpha}$, with $R_{\alpha} \phi=\phi+\alpha$ (active rotation of a vector), with $\left[\mathcal{R}_{\alpha} f\right](\phi)=f(\phi-\alpha)$, and $-\pi<\phi \leq \pi$. In terms of the left action on $\mathbf{x} \in \mathbb{R}^{3}: \mathbf{x}^{\prime}=R_{\alpha}^{(z)} \mathbf{x}$ (ie. $\mathbf{x}^{\prime}$ obtained by rotating $\mathbf{x}$ by $+\alpha$ around the $z$ axis), the matrix associated with $R_{\alpha}^{(z)}$ now is:
$\left(\begin{array}{ccc}\cos \alpha & -\sin \alpha & 0 \\ \sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 1\end{array}\right) \quad$ Then : $\left[\mathcal{R}_{\alpha} f\right](\mathbf{x})=f\left(R_{\alpha}^{-1} \mathbf{x}\right)=f(x \cos \alpha+y \sin \alpha,-x \sin \alpha+y \cos \alpha, z)$.
At least formally, this generalises immediately to arbitrary rotations in 3-dim space: $[\mathcal{R} f](\mathbf{x})=f\left(R^{-1} \mathbf{x}\right)$, where $\mathcal{R}$ can be factorised as $\mathcal{R}_{\alpha, \beta, \gamma}=\mathcal{R}_{\alpha} \mathcal{R}_{\beta} \mathcal{R}_{\gamma}, \alpha, \beta$ and $\gamma$ being the famous Euler angles. We are not even going to try to write down the corresponding matrix; it is not so illuminating anyway. But it is not hard to see that it must have determinant 1 and (in Cartesian coordinates only) be orthogonal, which means it is an element of $S O(3)$.
3. We also have scale transformations $\mathbf{x}^{\prime}=a \mathbf{x}$, with $a \in \mathbb{R}$ a non-zero positive constant, and $\mathbf{x} \in \mathbb{R}^{n}$ in Cartesian coordinates. (Think of zooming in or out.) The group of these transformations is isomorphic to $G L_{+}(1, \mathbb{R})$. The restriction to Cartesian coordinates is important: in spherical coordinates over $\mathbb{R}^{3}$, only the radial coordinate would scale.

## 4. Lorentz and Poincaré transformations

Lorentz boosts are given in Jackson's Classical Electrodynamics, eq. (11.19), for the $\mathbb{R}^{4}$ coordinates $c t$ and 3-dim x :

$$
c t^{\prime}=\gamma(c t-\boldsymbol{\beta} \cdot \mathbf{x}) \quad \mathbf{x}^{\prime}=\mathbf{x}+\frac{\gamma-1}{\beta^{2}}(\boldsymbol{\beta} \cdot \mathbf{x}) \boldsymbol{\beta}-\gamma \boldsymbol{\beta}(c t)
$$

where $\boldsymbol{\beta}$ is the velocity of the primed frame in the unprimed frame, and $\gamma=1 / \sqrt{1-\beta^{2}}$. Jackson's eq. (11.98) expresses this transformation in matrix form. It is not pretty. To include 3 -dim rotations, just replace x by $R_{\alpha, \beta, \gamma}^{-1} \mathrm{x}$ in the second equation. It is not worth writing the resulting $4 \times 4$ matrix which will be an element of $S O(3,1)$ if we exclude time reversal and space reflection; otherwise the relevant group will be $O(3,1)$ (or $O(1,3)$ in a mostly negative metric), the full Lorentz group. The transformation is a homogeneous one, which in the 4 -vector formalism is written: $\mathbf{x}^{\prime}=\Lambda \mathbf{x}$, where $\mathbf{x}$ is a general 4 -vector (not necessarily position).
We can extend the full Lorentz transformations to include space-time translations $\mathbf{t}$; this makes them inhomogeneous:

$$
\mathrm{x}^{\prime}=\Lambda \mathrm{x}+\mathrm{t}
$$

Whereas the homogeneous transformations left the norm of of a 4 -vector invariant, these leave invariant the norm of the difference between two 4 -vectors.
If we call $\boldsymbol{\Lambda}$ the full Lorentz transformation matrix, we can construct the matrix for these transformations by adding to $\boldsymbol{\Lambda}$ a fifth row and column whose last element is a 1 that does not do anything, that is:

$$
\binom{\mathbf{x}^{\prime}}{1}=\left(\begin{array}{ll}
\mathbf{n} & \mathbf{t} \\
0 & 1
\end{array}\right)\binom{\mathbf{x}}{1}
$$

These $5 \times 5$ matrices are elements of the 10 -parameter inhomogeneous Lorentz group, or Poincaré group, $I S O(3,1)$.

These examples illustrate the isomorphism between physical transformations and matrix Lie groups. We can then identify, say, a rotation with a $S O(3)$ matrix. In fact, we go so far as calling $S O(3)$ the rotation group. The same applies to the other cases.

### 3.2.5 Operator realisation of groups of transformations: infinitesimal generators

Now we explore more deeply this isomorphism between groups of transformations of functions and Lie groups. We shall express the left action of a few transformations as differential operators, a far from gratuitous exercise.

## 1. Translations

We can first look just at functions $f(x)(x \in \mathbb{R})$. Then the result of a translation $T_{a} x=x+a(a \in \mathbb{R})$ on $f$, with $a \ll x$, can be Taylor-expanded about $x$ :

$$
\left[\mathcal{T}_{a} f\right](x)=f\left(T_{a}^{-1} x\right)=f(x-a)=\left(1-a \mathrm{~d}_{x}+\ldots\right) f(x)=\exp \left(-a \mathrm{~d}_{x}\right) f(x)
$$

In $\mathbb{R}^{3}$ this generalises to:

$$
\begin{equation*}
\left[\mathcal{T}_{\mathbf{a}} f\right](\mathbf{x})=f\left(T_{\mathbf{a}}^{-1} \mathbf{x}\right)=f(\mathbf{x}-\mathbf{a})=\sum_{n=0}^{\infty} \frac{1}{n!}\left(-a^{i} \partial_{i}\right)^{n} f(\mathbf{x})=\exp \left(-a^{i} \partial_{i}\right) f(\mathbf{x}) \tag{3.3}
\end{equation*}
$$

The operators $-\partial_{i}$ are called the infinitesimal generators of translations. In quantum mechanics, it is the Hermitian momentum operator $\mathbf{p}=-\mathrm{i} \hbar \boldsymbol{\partial}$ that is called the infinitesimal generator of translations. Then we would write the translation operator as: $T_{\mathbf{a}}=\mathrm{e}^{-\mathrm{i} \mathbf{a} \cdot \mathrm{p} / \hbar}$.

We note that the infinitesimal generators $-\partial_{i}$ commute amongst themselves.

## 2. Rotations

For rotations $R_{\alpha} \phi=\phi+\alpha$ in the (xy) plane by a small angle $\alpha$ :

$$
\left[\mathcal{R}_{\alpha} f\right](\phi)=f\left(R_{\alpha}^{-1} \phi\right)=f(\phi-\alpha)=\left(1-\alpha \mathrm{d}_{\phi}+\ldots\right) f(\phi)=\exp \left(-\alpha \mathrm{d}_{\phi}\right) f(\phi)
$$

As we have seen in the last section, in $\mathbb{R}^{3}$ with Cartesian coordinates, this gives for the left action of a rotation $R_{\alpha}^{(z)} \mathbf{x}=(x \cos \alpha-y \sin \alpha, x \sin \alpha+y \cos \alpha, z)$ :

$$
\left[\mathcal{R}_{\alpha}^{(z)} f\right](\mathbf{x})=f\left(R_{\alpha}^{-1} \mathbf{x}\right)=f(x \cos \alpha+y \sin \alpha,-x \sin \alpha+y \cos \alpha, z)
$$

If we Taylor-expand the right-hand side of the last equality, we obtain:

$$
\begin{equation*}
\left[\mathcal{R}_{\alpha}^{(z)} f\right](\mathbf{x})=\left[1+\alpha\left(y \partial_{x}-x \partial_{y}\right)+\ldots\right] f(\mathbf{x})=\exp \left(\alpha M_{z}\right) f(\mathbf{x}) \tag{3.4}
\end{equation*}
$$

where $M_{z}=y \partial_{x}-x \partial_{y}$. We proceed in exactly the same way for rotations about the $x$ and $y$ axes to arrive at the general rotation operator:

$$
\begin{equation*}
\mathcal{R}_{\alpha, \beta, \gamma}=\exp \left(\alpha M_{x}\right) \exp \left(\beta M_{y}\right) \exp \left(\gamma M_{z}\right) \tag{3.5}
\end{equation*}
$$

where $M_{x}=z \partial_{y}-y \partial_{z}, \quad M_{y}=x \partial_{z}-z \partial_{x}, \quad M_{z}=y \partial_{x}-x \partial_{y}$ or, more simply: $M_{i}=-\epsilon_{i j k} x^{k} \partial^{k}$. Again, in quantum mechanics, these infinitesimal generators of rotations are redefined as $\mathbf{L}=\mathrm{i} \hbar \mathbf{M}$ and interpreted as the (Hermitian) angular momentum operators in Cartesian coordinates.
The infinitesimal generators do not commute. Indeed: $\left[M_{i}, M_{j}\right]=\epsilon_{i j}^{k} M_{k}, \quad$ or $\quad\left[L_{i}, L_{j}\right]=\mathrm{i} \hbar \epsilon_{i j}^{k} L_{k}$.
(Note: we could have written - some do! - the infinitesimal generators of translations to define $\mathbf{M}$ as the negative of the above. The cost, however, would be an extra minus sign in the commutation relations.)

## 3. Dilatations or scale transformations

In a $n$-dim space with Cartesian coordinates $x^{\mu}$, consider the scale transformation: $T_{\kappa} x^{\mu}=(1+\kappa) x^{\mu}$. In the limit of small $\kappa, T_{\kappa}^{-1} x^{\mu} \approx(1-\kappa) x^{\mu}$. As before we Taylor-expand a function $f\left(T_{\kappa}^{-1} x^{\mu}\right)$ in the small parameter $\kappa$ :

$$
\begin{equation*}
f\left(T_{\kappa}^{-1} x^{\mu}\right)=\left[1-\kappa x^{\mu} \partial_{\mu}+\cdots\right] f\left(x^{\mu}\right)=\exp \left(-\kappa x^{\mu} \partial_{\mu}\right) f\left(x^{\mu}\right) \tag{3.6}
\end{equation*}
$$

We identifiy $D=-x^{\mu} \partial_{\mu}$ as the infinitesimal generator of dilatations.

We can now find the infinitesimal generators of an arbitrary group of transformations with $m$ parameters $a^{i}$ in the neighbourhood of the identity, such that $a^{i}=0 \forall i$ for the identity group element. These transformations map a point in a manifold $M^{n}$ (not the group manifold!) to another not far from the first that can certainly be described by the same coordinate patch.

Let the transformations act (left action!) on a space (aka carrier space) of differentiable functions $f$ defined on $M^{n}$ :

$$
\left[T_{\mathbf{a}} f\right](\mathbf{x})=f\left(T_{\mathbf{a}}^{-1} \mathbf{x}\right)
$$

Focus on $T_{\mathbf{a}} f$, and take $f$ as a function of the parameters $a^{i}$. As before, Taylor-expand the right-hand side to first order, this time around the identity parametrised by $\mathbf{a}=0$ :

$$
\left[T_{\mathbf{a}} f\right](\mathbf{x})=\left(1+\left.a^{i} \partial_{a^{i}}\left(T_{\mathbf{a}}^{-1} x\right)^{j}\right|_{\mathbf{a}=0} \partial_{j}+\ldots\right) f(\mathbf{x})
$$

where $i$ runs over the number $(m)$ of parameters, ie. the dimension of the group, and $j$ from 1 to the dimension $n$ of the space on which the functions $f$ act.

Definition 3.6. The operators:

$$
\begin{equation*}
X_{i}=\left.\partial_{a^{i}}\left(T_{\mathbf{a}}^{-1} x\right)^{j}\right|_{\mathbf{a}=0} \partial_{j} \tag{3.7}
\end{equation*}
$$

are called the infinitesimal generators of the group of transformations. In some references, the right-hand side is multiplied by $-i$ (with appropriate adjustment to the expansion) so as to ensure hermiticity.

For example, considering rotations around the $z$ axis in $\mathbb{R}^{3}$, there is just one parameter (angle) $a^{1}=\alpha$, and only $x$ and $y$ derivatives can occur since $z$ does not depend on $\alpha$, and the second term in the square bracket of eq. (3.4) is recovered.

### 3.2.6 Infinitesimal generators of matrix Lie groups

Now we show how to linearise matrix groups and find their infinitesimal generators. This is not hard at all if we know the matrices. In general, the matrix elements will be analytic functions of some (non-unique!) set of group parameters $a^{i}$, and all we have to do is Taylor-expand the matrix to first order in the group parameters around the identity element $\mathbf{I}_{n}$, for which the $a^{i}$ all vanish:

$$
\begin{equation*}
\mathbf{M}_{\mathbf{a}}=\mathbf{I}_{n}+a^{i} \mathbf{X}_{i} \quad \mathbf{X}_{i}=\left.\partial_{a^{i}} \mathbf{M}_{\mathbf{a}}\right|_{\mathbf{a}=0} \tag{3.8}
\end{equation*}
$$

where we understand that differentiating a matrix means differentiating each of its elements. The matrices $\mathbf{X}_{i}$ are the infinitesimal generators of the group. Again, some prefer the definition $\mathbf{X}_{i}=-\left.i \partial_{a^{i}} \mathbf{M}_{\mathbf{a}}\right|_{\mathbf{a}=0}$.

Example 3.5. $G L(n, \mathbb{R})$
Let the matrix elements $M^{i}{ }_{j}$ be written as $M^{i}{ }_{j}=\delta^{i}{ }_{j}+a^{i}{ }_{j}$. In matrix form, this is

$$
\mathbf{M}_{\mathbf{a}}=\mathbf{I}_{n}+a_{j}{ }_{j} \mathbf{E}_{i}^{j} \quad\left(\mathbf{E}_{i}^{j}\right)^{k}{ }_{l}=\delta^{j}{ }_{l} \delta^{k}{ }_{i}
$$

The $\mathbf{E}_{i}^{j}$ matrices are the infinitesimal generators of $G L(n, \mathbb{R})$; they have zero elements except for the $(i j)^{\text {th }}$ element which is 1.

Example 3.6. A slightly more exciting example is $\mathbf{M}_{\theta} \in S O(2):\left(\begin{array}{cc}\cos \theta & -\sin \theta \\ \sin \theta & \cos \theta\end{array}\right)$, for $0 \leq \theta<2 \pi$, that effects rotations in a plane. Taylor-expand to first order:

$$
\mathbf{M}_{\theta} \approx\left(\begin{array}{rr}
1 & -\theta \\
\theta & 1
\end{array}\right)=\mathbf{I}_{2}+\theta \mathbf{X}
$$

Then the infinitesimal generator of $S O(2)$ is a matrix fully consistent with the constraints on $S O(n)$ generators as we shall discover in section 3.3.4:

$$
\mathbf{X}=\left.\partial_{\theta} \mathbf{M}_{\theta}\right|_{\theta=0}=\left(\begin{array}{rr}
0 & -1 \\
1 & 0
\end{array}\right)
$$

We shall write the space it spans as:

$$
\mathfrak{s o}(2)=\left(\begin{array}{rr}
0 & -\theta \\
\theta & 0
\end{array}\right)
$$

Another example (EXERCISE) that is quite easy to work out is $S L(2, \mathbb{R})$; it will have three infinitesimal generators. Similarly, using the parametrisation of example 3.4, we see that an element of $S U(2)$ may be written as $a^{0} \mathbf{I}_{2}+a^{i} \mathbf{X}_{i}$, where $\mathbf{X}_{i}=\mathrm{i} \boldsymbol{\sigma}_{i}$ are the generators of $S U(2)$, with $\boldsymbol{\sigma}_{i}$ the Pauli matrices.

How the generators may be found when the group matrices are not known will be discussed after we formalise this linearisation procedure.

A couple of remarks are now in order.

- The term "infinitesimal generator" means exactly that: an operator that effects an infinitesimal transformation away from the identity. We want to reconstruct a finite transformation out of a succession of infinitesimal transformations involving only the first-order contribution in the expansion of a transformation or a matrix. That is:

$$
\mathbf{M}_{\mathbf{a}}=\lim _{n \rightarrow \infty}\left(M_{\mathbf{a} / n}\right)^{n}=\lim _{n \rightarrow \infty}\left(1+\left(a^{i} / n\right) \mathbf{X}_{i}\right)^{n}
$$

Now the following relations on any (well-behaved) linear operator $A$ hold:

$$
\begin{equation*}
\exp (A) \equiv \lim _{n \rightarrow \infty}\left(1+\frac{1}{n} A\right)^{n} \quad \mathrm{e}^{A}=\left.\sum_{n=0}^{\infty} \frac{1}{n!} A^{n} \mathrm{~d}_{A}^{n} \mathrm{e}^{A}\right|_{A=0}=\sum_{n=0}^{\infty} \frac{A^{n}}{n!} \tag{3.9}
\end{equation*}
$$

since the right-hand side of the first relation is equal to its derivative when $n \rightarrow \infty$. Therefore, $\mathbf{M}_{\mathbf{a}}=$ $\exp \left(a^{i} \mathbf{X}_{i}\right)$. This exponential map, useful though it is, must still be handled with some care as we shall discover later.

- To understand the importance of infinitesimal generators, consider that taking linear combinations of group elements, such as, for instance, the matrices of $S O(3)$ does not give an element of $S O(3)$. In other words, linear combinations of matrix representations of the rotation group are not rotations. In general, group products are non-linear in the group parameters, so linear combinations cannot be expected to preserve them.

Linear combinations of infinitesimal generators of rotations, however, are generators of rotations! Indeed, there is a set $\left\{X_{i}\right\}$ of infinitesimal generators of a Lie group that forms a basis of a linear vector space. An arbitrary vector in the space can always be written as $b^{i} X_{i}$, with $b_{i}$ the group parameters.

### 3.3 Lie Algebras

### 3.3.1 Linearisation of a Lie group product

The study of a Lie group $G$ can be considerably simplified by linearising its group product around the identity. To linearise the group product, we first write $\mathrm{g}, \mathrm{g}^{\prime} \in G$ in the neighbourhood of the identity as $\mathrm{g}=e+a \epsilon X$ and $\mathrm{g}^{\prime}=e+b \epsilon Y$, where $\epsilon$ is an arbitrarily small real number and $a$ and $b$ are real, but arbitrary. $X$ and $Y$ are infinitesimal generators of $g$ and $g^{\prime}$, respectively. Now expand $g g^{\prime}$ to first order in $\epsilon$ :

$$
\mathrm{gg}^{\prime} \approx e+\epsilon(a X+b Y)+\ldots
$$

Since $a X+b Y$ is a generator for the group product $\mathrm{g}^{\prime}$, we see that the generators indeed form a linear vector space, as asserted at the end of the last section.

Now consider the combined group product $h=\mathrm{g} \mathrm{g}^{\prime}\left(\mathrm{g}^{\prime} \mathrm{g}\right)^{-1} \in G . h=e$ if the group is Abelian. Let us expand it to first non-vanishing order, this time writing $\mathrm{g}=e+\epsilon_{1} X+\epsilon_{1}^{2} X^{2} / 2, \mathrm{~g}^{\prime}=e+\epsilon_{2} Y+\epsilon_{2}^{2} Y^{2} / 2$, where $\epsilon_{1}, \epsilon_{2}$ are arbitrarily small:

$$
\begin{aligned}
\mathrm{gg}^{\prime}\left(\mathrm{g}^{\prime} \mathrm{g}\right)^{-1} & \approx\left(e+\epsilon_{1} X+\frac{1}{2} \epsilon_{1}^{2} X^{2}\right)\left(e+\epsilon_{2} Y+\frac{1}{2} \epsilon_{2}^{2} Y^{2}\right)\left(e-\epsilon_{1} X+\frac{1}{2} \epsilon_{1}^{2} X^{2}\right)\left(e-\epsilon_{2} Y+\frac{1}{2} \epsilon_{2}^{2} Y^{2}\right)+\ldots \\
& \approx e+\epsilon_{1} \epsilon_{2}(X Y-Y X)+\ldots
\end{aligned}
$$

All other order- $\epsilon_{1} \epsilon_{2}$ contributions cancel out. We write $[X, Y]=X Y-Y X$, the commutator of the infinitesimal generators $X$ and $Y$. [ $X, Y]$ is the generator for $\mathrm{gg}^{\prime}\left(\mathrm{g}^{\prime} \mathrm{g}\right)^{-1}$ and must thus be an element of the same vector space as $X$ and $Y$. When the group product commutes, the corresponding commutator of the generators vanishes. Mathematicians often define $g g^{\prime}\left(g^{\prime} g\right)^{-1}$ as the "commutator" for the group product, but we shall reserve the term for $[X, Y]$.

It is straightforward if somewhat tedious to show that the Jacobi identity holds, just by expanding it:

$$
\begin{equation*}
[X,[Y, Z]]+[Y,[Z, X]]+[Z,[X, Y]]=0 \tag{3.10}
\end{equation*}
$$

Because $[X,[Y, Z]]-[[X, Y], Z]] \neq 0$, Lie algebras are non-associative.

### 3.3.2 Definition of a Lie algebra

Now we are ready for an important definition that collects and generalises our findings:
Definition 3.7. A Lie algebra $\mathfrak{g}$ (no relation to the metric of Module $I$ ) is a finite-dimensional vector space equipped with a bilinear product, the Lie bracket $[$,$] , which:$

- is linear: $[a X+b Y, Z]=a[X, Z]+b[Y, Z] \quad \forall a, b \in \mathbb{R}$ or $\mathbb{C} ;$
- is antisymmetric: $[X, Y]=-[Y, X]$;
- satisfies the Jacobi identity: $[X,[Y, Z]]+[Y,[Z, X]]+[Z,[X, Y]]=0$.

In physics, we usually take the Lie bracket to be the commutator $X Y-Y X$. Also, many physicists, because they always deal with the algebra, not the group, use $G$ to denote $\mathfrak{g}$; this is confusing.

It is important to keep in mind that in a Lie algebra the action of one element $X$ on another one, $Y$, is not given by some linear combination of them, nor by the straightforward composition of the two, $X Y$ (which may not be in the algebra); it is taken to be their commutator. The closure property of a Lie group translates into the closure of its algebra, in the sense that the commutator of any two of its elements is also an element of the algebra.

The algebra $\pm \mathrm{i} \mathfrak{g}$ is said to be an essentially real Lie Algebra. We have already seen that the linear and orbital angular momentum operators of quantum mechanics were related to real infinitesimal generators in that way.

Sometimes, however, it proves very convenient to construct a complex extension of a real or essentially real algebra, by allowing the coefficients of the expansion of its elements over a basis to be complex. For instance, we might wish to construct $J_{ \pm}=J_{x} \pm \mathrm{i} J_{y}$. This gives us more flexibility in constructing useful bases.

The dimension $n$ of a Lie algebra is that of the manifold of the group with which it is associated.

### 3.3.3 Structure constants of a Lie algebra

The commutators of its $n$ infinitesimal generators which form a basis of a Lie algebra are themselves elements of the algebra, so they must be written as linear combinations of those basis generators:

$$
\begin{equation*}
\left[X_{i}, X_{j}\right]=C_{i j}^{k} X_{k} \tag{3.11}
\end{equation*}
$$

The coefficients $C_{i j}^{k}$ are called the structure constants of the Lie algebra, whose structure they are said to specify. In fact, with some rarely relevant caveats, they pretty much tell us everything about the group itself.

The structure constants inherit the antisymmetry of the commutators: $C_{j i}{ }^{k}=-C_{i j}{ }^{k}$. When the structure constants all vanish, we say that the algebra is Abelian; then $[X, Y]=0 \forall X, Y \in \mathfrak{g}$.

The Jacobi identity on elements of an algebra induces (EXERCISE) a relation between the structure constants:

$$
\begin{equation*}
C_{i j}^{l} C_{k l}^{m}+C_{j k}^{l} C_{i l}^{m}+C_{k i}^{l} C_{j l}^{m}=0 \tag{3.12}
\end{equation*}
$$

Defining a matrix $\left(\mathbf{D}_{i}\right)_{j}^{k}=-C_{i j}{ }^{k}$, we find (EXERCISE) that $\mathbf{D}$ satisfies the commutation relation (3.11). If we can take the group's representations to be unitary, as for compact groups such as $S U(n)$ and $S O(n)$, the corresponding representations of the algebra are anti-Hermitian and we immediately find (EXEERCISE), since they must satisfy the commutation relattions, that the structure constants are real.

The structure constants for the essentially real algebra $\pm \mathrm{i} \mathfrak{g}$ are just (exercise) $\pm \mathrm{i} C_{i j}{ }^{k}$. Very often, in the case of essentially real algebras, people will call the $C_{i j}{ }^{k}$ themselves the structure constants instead of $\pm \mathrm{i} C_{i j}{ }^{k}$.

Two Lie algebras are said to be isomorphic when they have the same dimension and structure constants (up to a redefinition of their generators, eg. rescaling).

### 3.3.4 A direct way of finding Lie algebras

Suppose we do not have an explicit form for the matrix rrealisation of a Lie group in terms of the group parameters. All we know is the constraint(s) on the group elements. This is sufficient to find the Lie algebra; then the group matrix can be reconstructed from the algebra.

The first thing to do is to linearise the constraints. We listed some important ones at the beginning of section 3.2. For Cartesian metric-preserving compact groups, $\mathbf{M} \mathbf{I}_{n} \mathbf{M}^{\dagger}=\mathbf{I}_{n}$; for non-compact metric-preserving groups (when the metric is indefinite), we have $\mathbf{M} \mathbf{I}_{p}^{q} \mathbf{M}^{\dagger}=\mathbf{I}_{p}^{q}$, with $p+q=n$. In both cases, the real version just entails replacing the complex conjugate transpose by the transpose.

Linearising for the compact groups, we get:

$$
\left(\mathbf{I}_{n}+\epsilon \mathbf{A}\right)\left(\mathbf{I}_{n}+\epsilon \mathbf{A}\right)^{\dagger} \approx \mathbf{I}_{n}+\epsilon\left(\mathbf{A}^{\dagger}+\mathbf{A}\right)=\mathbf{I}_{n}
$$

Therefore the matrices representing the algebra are antihermitian: $\mathbf{A}^{\dagger}=-\mathbf{A}$. Their diagonal matrix elements are pure imaginary for unitary group algebras $\mathfrak{u}(n)$; for orthogonal group algebras $\mathfrak{o}(n), \mathbf{A}$ is real skew-symmetric, with $n(n-1) / 2$ independent parameters. Thus, $\mathfrak{o}(n)$ is the set of all real skew-symmetric matrices of rank $n$.

If we choose to use essentially real algebras instead (eg. $L$ as generators of $\mathfrak{s o}(3)$ instead of $M$ in section 3.2.5), then $\mathbf{M}=I_{n}+\mathrm{i} \in \mathbf{A}$, and we find that the $\mathbf{A}$ matrices must be Hermitian: $\mathbf{A}^{\dagger}=\mathbf{A}$.

If the group is an indefinite orthogonal group, which is non-compact, the same process yields: $\mathbf{A}^{\dagger} I_{p}^{q}=-I_{p}^{q} \mathbf{A}$. This is a bit messier, but we can simplify it by breaking $\mathbf{A}$ into block matrices. If $\mathbf{S}$ is a $q \times q$ matrix, $\mathbf{T}$ a $q \times p$ matrix, $\mathbf{U}$ a $p \times q$ matrix, and $\mathbf{V}$ a $p \times p$ matrix, then:

$$
\left(\begin{array}{ll}
\mathbf{S}^{\dagger} & \mathbf{U}^{\dagger} \\
\mathbf{T}^{\dagger} & \mathbf{V}^{\dagger}
\end{array}\right)\left(\begin{array}{cc}
-\mathbf{I}_{q} & \mathbf{0} \\
\mathbf{0} & \mathbf{I}_{p}
\end{array}\right)+\left(\begin{array}{cc}
-\mathbf{I}_{q} & \mathbf{0} \\
\mathbf{0} & \mathbf{I}_{p}
\end{array}\right)\left(\begin{array}{cc}
\mathbf{S} & \mathbf{T} \\
\mathbf{U} & \mathbf{V}
\end{array}\right)=0
$$

Expanding, we arrive (exercise) at three conditions on the block matrices:

$$
\mathbf{S}^{\dagger}=-\mathbf{S}, \quad \mathbf{V}^{\dagger}=-\mathbf{V}, \quad \mathbf{T}^{\dagger}=\mathbf{U}
$$

Both the $\mathbf{S}$ and $\mathbf{V}$ diagonal blocks are antihermitian. The off-diagonal blocks are each other's adjoint. Over $\mathbb{R}$, this means that $\mathbf{A}$ has two antisymmetric diagonal block matrices, one $q \times q$ and one $p \times p$; the off-diagonal blocks are the transpose of one another. The number of parameters of the infefinite orthogonal group is then $p(p-1) / 2+q(q-1) / 2+p q=n(n-1) / 2 .$, the same as for the compact orthogonal group $O(n)$.

There only remains to notice that the non-zero elements of the infinitesimal generators can only be $\pm 1$ (over $\mathbb{R}$ ) and also $\pm i$ (over $\mathbb{C}$ ) because of the linearisation.

As we saw, another important constraint can be imposed on a group matrix $\mathbf{M}: \operatorname{det} \mathbf{M}=1$, which defines $S L(n, \mathbb{R})$. Anticipating a little, its algebra can be written as a sum of matrices $\mathbf{C}_{i}$, and we have $\mathbf{M}=$ $\exp \mathbf{C}_{1} \exp \mathbf{C}_{2} \exp \mathbf{C}_{3} \cdots$.

Now, since the determinant of a product of matrices is equal to the product of the determinants of the matrix factors, and because-when a matrix $\mathbf{A}$ is diagonalisable- there exists a similarity transformation $\mathbf{S A S}^{-1}$ which takes $\mathbf{A}$ to $\mathbf{A}^{\prime}=\operatorname{diag}\left(\lambda_{1}, \ldots, \lambda_{i}, \ldots\right)$, we conclude that $\operatorname{det} \mathbf{A}$ is equal to the product of the eigenvalues of $\mathbf{A}$.

Also:

$$
\mathbf{S} \mathrm{e}^{\mathbf{A}} \mathbf{S}^{-1}=\mathbf{S} \mathbf{I} \mathbf{S}^{-1}+\mathbf{S} \mathbf{A} \mathbf{S}^{-1}+\frac{1}{2!} \mathbf{S} \mathbf{A} \mathbf{S}^{-1} \mathbf{S} \mathbf{A} \mathbf{S}^{-1}+\ldots=\mathbf{I}+\mathbf{A}^{\prime}+\frac{1}{2!}\left(\mathbf{A}^{\prime}\right)^{2}+\ldots=\mathrm{e}^{\mathbf{A}^{\prime}}
$$

where $\mathrm{e}^{\mathbf{A}^{\prime}}$ is a diagonal matrix with $\mathrm{e}^{\lambda_{i}}$ as entries. In other words, the eigenvalues of $\mathrm{e}^{\mathbf{A}}$ are just $\mathrm{e}^{\lambda_{i}}$. Then:

$$
\operatorname{det} \mathrm{e}^{\mathbf{A}}=\prod_{i} \mathrm{e}^{\lambda_{i}}=\exp \sum_{i} \lambda_{i}=\mathrm{e}^{\operatorname{Tr} \mathbf{A}^{\prime}}
$$

But $\operatorname{Tr} \mathbf{A}^{\prime}=\operatorname{Tr}\left(\mathbf{S A S}{ }^{-1}\right)=\operatorname{Tr} \mathbf{A}$. We obtain via this elegant (but limited to diagonalisable matrices!) derivation an important basis-independent relation:

$$
\begin{equation*}
\operatorname{det}\left(\mathrm{e}^{\mathbf{A}}\right)=\mathrm{e}^{\operatorname{Tr} \mathbf{A}} \tag{3.13}
\end{equation*}
$$

valid for any square matrix. This is immediately extended to $\operatorname{det}\left(e^{\mathbf{A}} e^{\mathbf{B}} \ldots\right)=e^{\operatorname{Tr}(\mathbf{A}+\mathbf{B}+\ldots)}$, and since all $S L(n, \mathbb{R})$ matrices can be written as a prodiuct $\mathrm{e}^{\mathbf{A}} \mathrm{e}^{\mathbf{B}}$ (to be shown later), we immediately deduce that all matrices in the algebra $\mathfrak{s l}(n, \mathbb{R})$ must have vanishing trace, including the generators that live in $\mathfrak{s u}(n)$ and $\mathfrak{s o}(n)$. This is why it can be said that $\mathfrak{s l}(n, \mathbb{R})$ consists of the set of all traceless matrices of rank $n$.

Note the interesting fact that since antisymmetric real matrices must be traceless, $\mathfrak{o}(n)$ and $\mathfrak{s o}(n)$ are identical! This is very much related to the absence of a continuous path from the $O(n)$ identity (which is unimodular) to orthogonal matrices with determinant $-1: O(n)$ is not connected. For instance, spatial inversions cannot be linearised; one cannot invert axes by a "small" amount! It does not make sense to speak of the infinitesimal generators of spatial inversions if we cannot take a system gradually, in steps as small as we wish, from its initial state to its final transformed state. In three dimensions parametrised with spherical coordinates, spatial inversion is performed by $f(r, \theta, \phi) \rightarrow f(r, \pi-\theta, \pi+\phi)$, but without intermediate steps. So the infinitesimal generators of $O(3)$ are those corresponding to its connected $S O(3)$ subgroup, the group of rotations.

While we are at it, and since the exponential representation of group matrices in terms of their algebra matrices can allow us to reconstruct the group elements from the algebra, we note an important but difficult to prove expression which says that the familiar rule $\mathrm{e}^{a} \mathrm{e}^{b}=\mathrm{e}^{a+b}$ does not hold for matrices unless they commute! This is the so-called Baker-Campbell-Hausdorff $(\mathrm{BCH})$ formula:

$$
\begin{equation*}
\mathrm{e}^{\mathbf{A}} \mathrm{e}^{\mathbf{B}}=\mathrm{e}^{\mathbf{C}} \quad \mathbf{C}=\mathbf{A}+\mathbf{B}+\frac{1}{2}[\mathbf{A}, \mathbf{B}]+\frac{1}{12}([\mathbf{A},[\mathbf{A}, \mathbf{B}]]+[[\mathbf{A}, \mathbf{B}], \mathbf{B}])+\ldots \tag{3.14}
\end{equation*}
$$

By comparison, it requires little effort to show (EXERCISE) that:

$$
\mathrm{d}_{x} \mathbf{A}^{-1}=-\mathbf{A}^{-1}\left(\mathrm{~d}_{x} \mathbf{A}\right) \mathbf{A}^{-1}
$$

Also, if $\mathbf{A}$ and $\mathbf{B}$ are square matrices, and if the explicit dependence of another matrix $\mathbf{C}$ on a real parameter $\lambda$ is given by $\mathbf{C}(\lambda)=\mathrm{e}^{\lambda \mathbf{B}} \mathbf{A} \mathrm{e}^{-\lambda \mathbf{B}}$, expanding plus repeated differentiation with respect to $\lambda$ yields (EXERCISE):

$$
\mathrm{e}^{\mathbf{B}} \mathbf{A} \mathrm{e}^{-\mathbf{B}}=\mathbf{A}+[\mathbf{B}, \mathbf{A}]+\frac{1}{2}[\mathbf{B},[\mathbf{B}, \mathbf{A}]]+\ldots+\frac{1}{n!}[\mathbf{B},[\mathbf{B}, \ldots[\mathbf{B}, \mathbf{A}]]+\ldots
$$

This relation also holds for linear operators.

Example 3.7. The generators of $S O(3)$ live in a three-parameter algebra:

$$
\mathfrak{s o}(3)=\left(\begin{array}{ccc}
0 & -\theta_{z} & \theta_{y}  \tag{3.15}\\
\theta_{z} & 0 & -\theta_{x} \\
-\theta_{y} & \theta_{x} & 0
\end{array}\right)=\theta^{i} \mathbf{M}_{i}
$$

But how did we pick the signs above the diagonal? Consider counterclockwise rotations by a small angle $\theta$ around an axis whose direction is specified by the vector $\hat{\mathbf{n}}$. An active transformation rotates a vector x by adding a small vector that is perpendicular to both the axis and to x . By geometry, we find that, to first-order, the transformed vector is $\mathbf{x}^{\prime}=\mathbf{x}+\theta \hat{\mathbf{n}} \times \mathbf{x}$. Expanding gives:

$$
\begin{array}{rlrl}
x^{\prime} \approx x+\theta\left(n_{y} z-n_{z} y\right) & y^{\prime} & \approx y+\theta\left(n_{z} x-n_{x} z\right) & z^{\prime} \\
& \Longleftrightarrow & \mathbf{x}^{\prime}=\mathbf{x}+\left(\begin{array}{ccc}
0 & -\theta_{z} & \theta_{y} \\
\theta_{z} & 0 & -\theta_{x} \\
-\theta_{y} & \theta_{x} & 0
\end{array}\right)\left(\begin{array}{l}
x \\
y \\
z
\end{array}\right)
\end{array}
$$

where $\boldsymbol{\theta}=\theta \hat{\mathbf{n}}$. The matrix is indeed the $\mathfrak{s o ( 3 )}$ algebra of eq. (3.15). How does this compare to the operator algebra as laid out in eq. (3.4) and the couple of equations that follow it? By identifying $\alpha=\theta n_{z}$, etc., we can write the first order in the expansion of the general rotation operator as:

$$
\left(\begin{array}{lll}
x & y & z
\end{array}\right)\left(\begin{array}{ccc}
0 & -\theta_{z} & \theta_{y} \\
\theta_{z} & 0 & -\theta_{x} \\
-\theta_{y} & \theta_{x} & 0
\end{array}\right)\left(\begin{array}{l}
\partial_{x} \\
\partial_{y} \\
\partial_{z}
\end{array}\right)
$$

The matrix is the $\mathfrak{s o}(3)$ algebra matrix. The operator algebra has the same commutator structure, $\left[M_{i}, M_{j}\right]=\epsilon_{i j}{ }^{k} M_{k}$, as the matrix algebra $\mathfrak{s o}(3)$ (and also as the $\mathfrak{s u}(2)$ matrix algebra!), which establishes the isomorphism of the algebras (but not of the groups...).
A rotation by a finite angle $\theta$ around an axis $\hat{\mathbf{n}}$ can now be written as:

$$
R(\boldsymbol{\theta})=\mathrm{e}^{-\theta^{k} \mathbf{M}_{k}}
$$

with the generator matrices:

$$
\mathbf{M}_{x}=\left(\begin{array}{ccc}
0 & 0 & 0 \\
0 & 0 & -1 \\
0 & 1 & 0
\end{array}\right), \quad \mathbf{M}_{y}=\left(\begin{array}{ccc}
0 & 0 & 1 \\
0 & 0 & 0 \\
-1 & 0 & 0
\end{array}\right), \quad \mathbf{M}_{z}=\left(\begin{array}{ccc}
0 & -1 & 0 \\
1 & 0 & 0 \\
0 & 0 & 0
\end{array}\right)
$$

Often, $S O(3)$ generators are written as $J_{i j}=\epsilon_{i j k} M^{k}$. Since $\left(\mathbf{M}_{i}\right)_{j k}=-\epsilon_{i j k}$, the matrix elements are: $\left(J_{i j}\right)^{l m}=-\epsilon_{i j k} \epsilon^{k l m}=-\left(\delta_{i}^{l} \delta_{j}^{m}-\delta_{i}^{m} \delta_{j}^{l}\right)$. The labels $(i j), i<j$ for $J$ refer to the plane in which thoe rotation is generated. This allows a straightforward generalisation to rotations in $n>3$ dimensions, where an axis does not uniquely define a plane of rotation. Their commutator algebra can now be obtained by computing (EXERCISE): $\left[J_{m n}, J_{p q}\right]^{i}{ }_{j}=\left(J_{m n}\right)^{i}{ }_{k}\left(J_{p q}\right)^{k}{ }_{j}-\left(J_{p q}\right)^{i}{ }_{l}\left(J_{m n}\right)^{l}{ }_{j}$ and rearranging the eight resulting terms, yielding a result ${ }^{\dagger}$ for the commutator that is important because it appplies to rotations in dimensions $N>3$ :

$$
\begin{equation*}
\left[J_{m n}, J_{p q}\right]=\delta^{m}{ }_{p} J_{n q}+\delta^{q}{ }_{n} J_{m p}-\delta^{q}{ }_{m} J_{n p}-\delta^{p}{ }_{n} J_{m q} \quad 1 \leq m<n \leq N, 1 \leq p<q \leq N \tag{3.16}
\end{equation*}
$$

Only one term on the right can contribute (EXERCISE), and the commutator vanishes unless one (and only one) number in the pair ( $m n$ ) is equal to one (and only one) number in the pair ( $p q$ ).
Example 3.8. The $\mathfrak{s o ( 4 )}$ Lie algebra is the set of all antisymmetric $4 \times 4$ real matrices. It is now appropriate to use the $4(4-1) / 2=6 J_{i j}$ generators, introduced in example 3.7, that generate rotations in the $(i j)$-plane. With eq. (3.16) it is easy to compute the nine non-trivial $\mathfrak{s o}(4)$ commutators. Taking $J_{i 4}=N_{i}$ and $J_{i j}=\epsilon_{i j k} M^{k}(1 \leq i, j<k \leq 3)$, they can be written as:

$$
\begin{equation*}
\left[M_{i}, M_{j}\right]=\epsilon_{i j}^{k} M_{k}, \quad\left[M_{i}, N_{j}\right]=\epsilon_{i j}^{k} N_{k}, \quad\left[N_{i}, N_{j}\right]=\epsilon_{i j}^{k} M_{k} \tag{3.17}
\end{equation*}
$$

[^13]The commutators can be decoupled by transforming to the basis:

$$
Y_{i}=\frac{1}{2}\left(M_{i}+N_{i}\right) \quad Z_{i}=\frac{1}{2}\left(M_{i}-N_{i}\right)
$$

from which we immediately obtain the decoupled relations:

$$
\begin{equation*}
\left[Y_{i}, Y_{j}\right]=\epsilon_{i j}^{k} Y_{k}, \quad\left[Y_{i}, Z_{j}\right]=0, \quad\left[Z_{i}, Z_{j}\right]=\epsilon_{i j}^{k} Z_{k} \tag{3.18}
\end{equation*}
$$

By inspection, the $Y_{i}$ and $Z_{i}$ separate into two $\mathfrak{s u}(2)$ (or $\left.\mathfrak{s o}(3)\right)$ algebras, and $\mathfrak{s o}(4)=\mathfrak{s u}(2) \oplus \mathfrak{s u}(2)$. In terms of dimensions, $\mathbf{6}=\mathbf{3} \oplus \mathbf{3}$. At group level, we say that $S O(4)$ is locally isomorphic to the direct product $S U(2) \times S U(2)$; it is globally isomorphic to $(S U(2) \times S U(2)) / Z_{2}$ since a pair of elements of $S U(2)$ and their negatives correspond to the same $S O(4)$ rotation.
The algebra now reads: $\mathfrak{s o}(4)=a^{i} Y_{i}+b^{i} Z_{i}$, and since $\left[Y_{i}, Z_{j}\right]=0$, an element of $S O(4)$ takes the form: $\mathrm{e}^{a^{i} Y_{i}} \mathrm{e}^{b^{i} Z_{i}}$.
Example 3.9. The $\mathfrak{s o}(3,1)$ algebra of the group $S O(3,1)$ derived from the metric-preserving constraint is:

$$
\mathfrak{s o}(3,1)=\left(\begin{array}{cccc}
0 & \zeta_{x} & \zeta_{y} & \zeta_{z}  \tag{3.19}\\
\zeta_{x} & 0 & -\theta_{z} & \theta_{y} \\
\zeta_{y} & \theta_{z} & 0 & -\theta_{x} \\
\zeta_{z} & -\theta_{y} & \theta_{x} & 0
\end{array}\right)=\theta^{\mu} \mathbf{M}_{\mu}+\zeta^{\nu} \mathbf{K}_{\nu}
$$

where the infinitesimal generators can be read off:

$$
\mathbf{M}_{x}=\left(\begin{array}{cccc}
0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 \\
0 & 0 & 0 & -1 \\
0 & 0 & 1 & 0
\end{array}\right) \quad \mathbf{M}_{y}=\left(\begin{array}{cccc}
0 & 0 & 0 & 0 \\
0 & 0 & 0 & 1 \\
0 & 0 & 0 & 0 \\
0 & -1 & 0 & 0
\end{array}\right) \quad \mathbf{M}_{z}=\left(\begin{array}{cccc}
0 & 0 & 0 & 0 \\
0 & 0 & -1 & 0 \\
0 & 1 & 0 & 0 \\
0 & 0 & 0 & 0
\end{array}\right)
$$

and

$$
\mathbf{K}_{x}=\left(\begin{array}{cccc}
0 & 1 & 0 & 0 \\
1 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0
\end{array}\right) \quad \mathbf{K}_{y}=\left(\begin{array}{cccc}
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 0 \\
1 & 0 & 0 & 0 \\
0 & 0 & 0 & 0
\end{array}\right) \quad \mathbf{K}_{z}=\left(\begin{array}{cccc}
0 & 0 & 0 & 1 \\
0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 \\
1 & 0 & 0 & 0
\end{array}\right)
$$

One shows (EXERCISE) that the commutators of the infinitesimal generators are:

$$
\begin{equation*}
\left[M_{i}, M_{j}\right]=\epsilon_{i j}{ }^{k} M_{k} \quad\left[M_{i}, K_{j}\right]=\epsilon_{i j}{ }^{k} K_{k} \quad\left[K_{i}, K_{j}\right]=-\epsilon_{i j}^{k} M_{k} \tag{3.20}
\end{equation*}
$$

Although the number of generators is identical to $\mathfrak{s o}(4)$, there is an important difference between these relations and the ones derived in example 3.8: the minus sign in the relation for the $K$. It follows if we take $N \rightarrow \mathrm{i} K$. Then the complex basis in which the commutators decouple is: $L_{i}^{ \pm}=\left(M_{i} \pm \mathrm{i} K_{i}\right) / 2$, yielding (EXERCISE) : $\left[L_{i}^{ \pm}, L_{j}^{ \pm}\right]=\epsilon_{i j}{ }^{k} L_{k}^{ \pm}$and $\left[L_{i}^{ \pm}, L_{j}^{\mp}\right]=0$.
As in example 3.7, by defining $J_{i j}=\epsilon_{i j}{ }^{k} M_{k}$ and $J_{0 i}=K_{i}, 1 \leq i \leq 3$, one rewrites the commutator relations (3.20) as a relation valid for any $\mathfrak{s o}(p, q)$ algebra:

$$
\begin{equation*}
\left[J_{\mu \nu}, J_{\alpha \beta}\right]=\eta_{\mu \alpha} J_{\nu \beta}+\eta_{\nu \beta} J_{\mu \alpha}-\eta_{\mu \beta} J_{\nu \alpha}-\eta_{\nu \alpha} J_{\mu \beta} \quad 0 \leq(\mu, \nu) \leq 3 \tag{3.21}
\end{equation*}
$$

where $J_{\nu \mu}=-J_{\mu \nu}$ (indices are labels of the generators $J$, not components!), and $\eta_{\mu \nu}$ is a diagonal matrix with 3 entries $\pm 1$ and one entry $\mp 1$, depending on the sign convention, or in general, the diagonal matrix with $p$ entries $\pm 1$ and $q$ entries $\mp 1$, with $p+q=N$.
One very important realisation of this algebra interprets $\theta_{i}$ as the three angles rotating around Cartesian axis $1 \leq i \leq 3$, and $\zeta_{i}=\hat{\beta}_{i} \tanh ^{-1} \beta$ the rapidity parameters for pure Lorentz boosts along the $x$, $y$ and $z$ axes. Then $\mathfrak{s o}(3,1)$ is called the Lorentz algebra for Minkowski spacetime. In this context, $\eta_{\mu \nu}$ in eq. (3.21) are the components of the Cartesian metric tensor. Then the relation can also be derived (EXERCISE) in the operator realisation: $J_{\mu \nu}=x_{\nu} \partial_{\mu}-x_{\mu} \partial_{\nu}$. This algebra is also that of the infinitesimal transformations $x^{\prime \mu}=\omega^{\mu}{ }_{\nu} x^{\nu}$, with $\omega^{\mu}{ }_{\nu}$ given in eq. (3.19), and $\omega_{\nu \mu}=-\omega_{\mu \nu}$.

### 3.3.5 Hard-nosed questions about the exponential map - the fine print

Three theorems by Lie, which we have implicitly made use of, show that for any Lie group a Lie algebra can be found which is characterised by the structure constants. But can we go back from the algebra to the group? Up to now, we have relied on the exponential map to do this, but it is not always possible, at least with just one map.

We already know that we cannot recover orthogonal matrices of determinant -1 in $O(3)$ from the $\mathfrak{o}(3)$ algebra. Do we fare better with special linear transformations whose matrices $S L(n, \mathbb{R})$ all have determinant 1 ? No, according to the following counter-example (provided, I believe, by Cartan). One form of the $\mathfrak{s l}(2, \mathbb{R})$ algebra, which is the set of all traceless matrix operators, is: $\mathbf{X}=\left(\begin{array}{cc}x_{1} & x_{2}-x_{3} \\ x_{2}+x_{3} & -x_{1}\end{array}\right)$. Exponentiating gives the basisindependent result (EXERCISE):

$$
\mathrm{e}^{\mathbf{X}}=\sum_{n} \frac{1}{n!} \mathbf{X}^{n}= \begin{cases}\mathbf{I}_{2} \cosh r+\mathbf{X} \frac{\sinh r}{r} & r^{2}>0 \\ \mathbf{I}_{2}+\mathbf{X} & r^{2}=0 \\ \mathbf{I}_{2} \cos r+\mathbf{X} \frac{\sin r}{r} & r^{2}<0\end{cases}
$$

where $r^{2}=x_{1}^{2}+x_{2}^{2}-x_{3}^{2}=-\operatorname{det} \mathbf{X}$, which makes the results basis-independent. The structure is reminiscent of the light-cone structure obtained by endowing the parameter space $\mathbb{R}^{3}$ with an indefinite metric invariant under $S O(2,1)$. Inside the light-cone, for any value of $x_{3}$, the values of the other two parameters are confined inside a circle of radius smaller than $x_{3}$. The corresponding generators map to compact group elements. Outside the light-cone, however, the generators can grow without restriction; they map to non-compact elements of $S L(2, \mathbb{R})$.

So far, so good. But a glance at the above expressions shows that $\operatorname{Tr} \mathrm{e}^{\mathbf{X}} \geq-2$ always. Yet $S L(2, \mathbb{R})$ has a large subset of elements with trace smaller than -2 : matrices of the type $\left(\begin{array}{cc}-\lambda & 0 \\ 0 & -1 / \lambda\end{array}\right)(\lambda>1)$, for instance. These cannot be reached with the above exponential map.

Cartan argued that all the group elements could nevertheless be reached by writing:

$$
\mathbf{X}=\mathbf{X}_{a}+\mathbf{X}_{b}=\left(\begin{array}{cc}
x_{1} & x_{2} \\
x_{2} & -x_{1}
\end{array}\right)+\left(\begin{array}{cc}
0 & -x_{3} \\
x_{3} & 0
\end{array}\right)
$$

and taking the product of the exponentials of $\mathbf{X}_{a}$ and $\mathbf{X}_{b}$, which is not $\mathrm{e}^{\mathbf{X}}$ since $\left[\mathbf{X}_{a}, \mathbf{X}_{b}\right] \neq 0$. Then (EXERCISE):

$$
\mathrm{e}^{\mathbf{x}_{a}} \mathrm{e}^{\mathbf{x}_{b}}=\left(\begin{array}{cc}
z+y & x \\
x & z-y
\end{array}\right)\left(\begin{array}{cc}
\cos x_{3} & -\sin x_{3} \\
\sin x_{3} & \cos x_{3}
\end{array}\right)
$$

where $z \equiv \cosh r^{\prime} \geq 1, x \equiv \frac{x_{2}}{r^{\prime}} \sinh r^{\prime}$, and $y \equiv \frac{x_{1}}{r^{\prime}} \sinh r^{\prime}$, with $r^{\prime 2}=x_{1}^{2}+x_{2}^{2}$. Each matrix is unimodular, and the trace of the product is now $2 z \cos x_{3}=2 \cosh r^{\prime} \cos x_{3}$ which is unrestricted. One exponential, $\mathrm{e}^{\mathbf{X}_{b}}$, corresponds to a compact generator, and the other to non-compact generators.

In example 3.3 we noted that we needed more tools to tell us what the manifold of $S L(2, \mathbb{R})$ was. Now we know! The parameters of the non-compact matrix satisfy $z^{2}-\left(x^{2}+y^{2}\right)=1$ which is the positive- $z$ hyperboloid. Topologically, it is equivalent to $\mathbb{R}^{2}$. The parameter values $-\pi \leq x_{3} \leq \pi$ map the $\mathbf{X}_{b}$ subalgebra to $S O(2) \subset$ $S L(2, \mathbb{R})$, whose manifold is $S^{1}$. We conclude that $S L(2, \mathbb{R})$ is non-compact, and that its manifold is $\mathbb{R}^{2} \times S^{1}$.

These considerations do not detract from the usefulness of the exponential map as a link between a Lie algebra and a Lie group; they just mean that in the case of non-compact groups a little more care should be exercised.

### 3.4 Representations of Lie Groups and Algebras

### 3.4.1 Representations of Lie Groups

Definition 3.8. As for finite groups, a representation $\mathbf{T}_{\mathrm{g}}$ of a Lie group $G(\mathrm{~g} \in G)$ is a homomorphism of $G$ to the group of general linear matrix transformations $G L(\mathcal{V})$ acting on a vector space $\mathcal{V}$ called its carrier space.

For compact Lie groups, $\mathcal{V}$ is a finite-dimensional Hilbert space $\mathcal{H}$, ie. a vector space over $\mathbb{C}$ equipped with an inner product. For non-compact groups, however, it may well happen that the Hilbert space is infinite-dimensional.

Of particular interest are the irreducible representations of $G$. They satisfy Schur's lemma in its version for continuous groups:

A unitary representation $\mathbf{T}_{\mathrm{g}}$ of a group G is irreducible if, and only if, the only operator $A$ on $\mathcal{H}$ that satisfies $A \mathbf{T}_{\mathrm{g}}=\mathbf{T}_{\mathrm{g}} A \forall \mathrm{~g} \in G$ is a multiple of the identity.

The following statements, which we quote without proof, apply to compact Lie groups:

- An irreducible representation of a compact Lie group must be equivalent to an unitary representation. All unitary representations of a compact Lie group are finite-dimensional. Thus, all irreducible representations are finite-dimensional.
- Every representation of a compact Lie group that is not already irreducible is fully reducible, in the sense that it can be written as the direct sum of irreducible unitary representations.
- An important theorem by Peter and Weyl states that:

Let $\left\{\mathbf{T}^{(\alpha)}\right\}$ be the set of all irreducible non-equivalent unitary representations of a compact Lie group. Then the functions $\sqrt{n_{\alpha}} T_{i j}^{(\alpha)}$, where $n_{\alpha}$ is the dimension of the representation and $1 \leq i, j \leq n_{\alpha}$, form an orthonormal basis for $\left(\mathcal{L}_{2}\right)_{G}$, the Hilbert space of square-integrable functions on the manifold of $G$; that is, if $f_{\mathrm{g}} \in\left(\mathcal{L}_{2}\right)_{G}$, we have that:

$$
f_{\mathrm{g}}=\sum_{\alpha, i, j} b_{i j}^{(\alpha)}\left(T_{\mathrm{g}}^{(\alpha)}\right)_{i j}, \quad b_{i j}^{(\alpha)}=n_{\alpha} \int_{G} f_{\mathrm{g}}\left(T_{\mathrm{g}}^{(\alpha)}\right)_{i j}^{*} \mathrm{~d} \mu_{\mathrm{g}}
$$

where $\mathrm{d} \mu_{\mathrm{g}}$ is the volume element of the group manifold.
Example 3.10. Consider the Abelian group $U(1)$ with $S^{1}$ as its manifold. Its unitary irreducible representations are of course 1-dim: $T^{(m)}(\theta)=\mathrm{e}^{\mathrm{i} m \theta},\left(m \in \mathbb{Z}\right.$ because $\left.T^{(m)}(\theta+2 \pi)=T^{(m)}(\theta)\right)$. Then the Peter-Weyl theorem asserts that $\left\{T^{(m)}\right\}$ is a basis for $\mathcal{L}_{2}\left(S^{1}\right)$, the space of differentiable square-integrable functions on $S^{1}$, which must be periodic:

$$
f(\theta)=\sum_{m=-\infty}^{\infty} c_{m} \mathrm{e}^{\mathrm{i} m \theta} \quad c_{m}=\int_{0}^{2 \pi} f(\theta) \mathrm{e}^{-\mathrm{i} m \theta} \mathrm{~d} \theta
$$

which we recognise as a Fourier expansion.

### 3.4.2 Representations of Lie algebras

Lie algebras, as we have seen, can be realised as (differential) operators, or also as $\operatorname{gl}(\mathcal{V})$, the set of all linear transformations on some vector space $\mathcal{V}$. $\mathcal{V}$ will be a Hilbert space over $\mathbb{R}$ or $\mathbb{C}$, and we will have $\mathfrak{g l}(n, \mathbb{R})$ or $\mathfrak{g l}(n, \mathbb{C})$ realised as $n \times n$ real or complex matrices. In fact, a finite-dimensional algebra will always be isomorphic to some matrix algebra.

Definition 3.9. Let $\mathfrak{g}$ be a Lie algebra. A representation $\mathbf{T}$ of $\mathfrak{g}$ maps elements of the algebra to elements of the general linear invertible matrix transformations on its carrier space (or module) $\mathcal{V}$. The mapping is a homomorphism. The dimension of a representation is that of its carrier space.
$\mathfrak{g}$ has a Lie bracket, the commutator, and its representations must satisfy this product. Thus, if $\mathbf{T}$ is a representation of $\mathfrak{g}$, we must have $\forall X, Y \in \mathfrak{g}$ :

$$
\mathbf{T}_{[X, Y]}=\left[\mathbf{T}_{X}, \mathbf{T}_{Y}\right]
$$

The defining (fundamental) representation of $\mathfrak{g}$ is the matrix obtained directly from the constraints on the algebra.

### 3.4.3 The regular (adjoint) representation and the classification of Lie algebras

We have already noted how eq. (3.12) for the structure constants could be written as the commutator of matrices whcih we now recognise as providing a new representation of the algebra:

Definition 3.10. The regular (adjoint) representation of a Lie algebra associates with each element $Z$ of the algebra a matrix $\mathbf{R}_{Z}\left(\right.$ or $\left.\mathrm{ad}_{Z}\right)$ such that $\left[Z, X_{i}\right]=\left(R_{Z}\right)_{i}^{j} X_{j}$, where the $X_{i}$ are the infinitesimal generators that form a basis for the algebra. (Warning! Some authors use the definition $\left.\left[Z, X_{i}\right]=\left(R_{Z}\right)_{i}^{j} X_{j}.\right)$
Clearly, then, the regular representation of a basis generator is just the structure constants: $\left[X_{i}, X_{j}\right]=$ $\left(R_{X_{i}}\right)_{j}^{k} X_{k}=C_{i j}^{k} X_{k}$. And its dimension is that of the algebra, ie. the number of generators (or parameters).

We confirm that $\mathbf{R}$ is a representation (EXERCISE, with the Jacobi identity): $\left[\mathbf{R}_{X_{i}}, \mathbf{R}_{X_{j}}\right] X_{k}=\mathbf{R}_{\left[X_{i}, X_{j}\right]} X_{k}$.
Example 3.11. Take the defining, two-dimensional representation of the essentially real version of the $\mathfrak{s u}(2)$ algebra with as basis set the three Hermitian Pauli matrices, $\boldsymbol{\sigma}_{i}$, with commutators: $\left[\boldsymbol{\sigma}_{i}, \boldsymbol{\sigma}_{j}\right]=$ $2 \mathrm{i} \epsilon_{i j}{ }^{k} \boldsymbol{\sigma}_{k}$. Then $\operatorname{ad}_{\boldsymbol{\sigma}_{i}}\left(\boldsymbol{\sigma}_{j}\right)=2 \mathrm{i} \epsilon_{i j}{ }^{k} \boldsymbol{\sigma}_{k}$, and $\left(\operatorname{ad}_{\boldsymbol{\sigma}_{i}}\right)_{j}^{k}=2 \mathrm{i} \epsilon_{i j}{ }^{k}$, and we have:

$$
\operatorname{ad}_{\boldsymbol{\sigma}_{1}}=\left(\begin{array}{ccc}
0 & 0 & 0 \\
0 & 0 & 2 \mathrm{i} \\
0 & -2 \mathrm{i} & 0
\end{array}\right) \quad \operatorname{ad}_{\boldsymbol{\sigma}_{2}}=\left(\begin{array}{ccc}
0 & 0 & -2 \mathrm{i} \\
0 & 0 & 0 \\
2 \mathrm{i} & 0 & 0
\end{array}\right) \quad \operatorname{ad}_{\boldsymbol{\sigma}_{3}}=\left(\begin{array}{ccc}
0 & 2 \mathrm{i} & 0 \\
-2 \mathrm{i} & 0 & 0 \\
0 & 0 & 0
\end{array}\right)
$$

If we write the regular representation of a generic element of this algebra as $\mathbf{R}=a^{i} \operatorname{ad}_{\boldsymbol{\sigma}_{i}}$, we obtain the Hermitian matrix:

$$
\mathbf{R}=\left(\begin{array}{ccc}
0 & 2 \mathrm{i} a_{3} & -2 \mathrm{i} a_{2} \\
-2 \mathrm{i} a_{3} & 0 & 2 \mathrm{i} a_{1} \\
2 \mathrm{i} a_{2} & -2 \mathrm{i} a_{1} & 0
\end{array}\right)
$$

Note that the commutation relations $\left[\boldsymbol{\sigma}_{i}, \boldsymbol{\sigma}_{j}\right]=2 \mathrm{i} \epsilon_{i j}{ }^{k} \boldsymbol{\sigma}_{k}$ can be written equivalently as: $\left[S_{i}, S_{j}\right]=$ $\mathrm{i} \epsilon_{i j}{ }^{k} S_{k}$, simply by the redefinition $\mathbf{S}=\boldsymbol{\sigma} / 2$. This particular form is often preferred in quantum mechanics, particle physics and related fields.

Definition 3.11. A subalgebra of an algebra is just a subspace of the algebra $\mathfrak{g}$ which closes under commutation. A subalgebra $\mathfrak{g}_{\text {sub }}$ is invariant if, for each $X \in \mathfrak{g}_{\text {sub }}$ and $\forall Y \in \mathfrak{g},[X, Y] \in \mathfrak{g}_{\text {sub }}$. An invariant subalgebra is sometimes called an ideal, but we shall not be using this term.
The centre $\mathfrak{z}$ of an algebra is the largest subalgebra whose elements commute with all elements of the algebra. This means that, for a commutative (Abelian) algebra, the centre is the algebra itself. By definition, $\mathfrak{z}$ is an Abelian invariant subalgebra, but not necessarily the only one.

An invariant subalgebra generates an invariant subgroup via the exponential map. To prove this, we must show that if $\mathrm{g}=\mathrm{e}^{Y}$ and $\mathrm{g}^{\prime}=\mathrm{e}^{Y^{\prime}}$, where $Y$ is a generator in an invariant subalgebra and $Y^{\prime} \in \mathfrak{g}$, then $\mathrm{g}^{\prime-1} \mathrm{gg}^{\prime}=$ $\mathrm{e}^{Y^{\prime-1}} \mathrm{e}^{Y} \mathrm{e}^{Y^{\prime}}=\mathrm{e}^{Z}$ is also in the subgroup generated by the subalgebra. But this is the case by virtue of the Baker-Campbell-Hausdorff formula, eq. (3.14), since $Z$ is calculated from commutators of $Y$ with other generators and so must be in the invariant subalgebra.

Definition 3.12. If a sequence of transformations exists that puts the regular representation of a nonAbelian Lie algebra into block-diagonal form, with the blocks irreducible non-zero subrepresentations, the representation is said to be fully reducible. In this case, the regular representation can be written as a direct sum of irreducible representations. Of course, these irreducible representations cannot all be one-dimensional. In this basis, the block submatrices commute with one another.

Like the structure constants, the regular representation summarises the structure of the Lie algebra. This algebra is a vector space spanned by a basis of generators. But we can decide to transform to another basis via a similarity transformation. The question is: can we transform the regular representation to a basis where it takes a form that might help classify the algebra?

Definition 3.13. If an algebra has no non-trivial invariant subalgebra, its regular representation is irreducible and we say that the algebra is simple. In that case, the representation leaves no proper subspace of its carrier space invariant.

Definition 3.14. A Lie algebra that contains no Abelian invariant subalgebra is said to be semisimple, ie. it has zero centre (no non-zero element commutes with all other elements). A semisimple Lie algebra is either simple or the sum of simple Lie algebras (that may occur more than once in the sum), and its representations are the direct sums of irreducible representations. A semisimple algebra always has at least two complementary invariant subalgebras, with all the generators of one commuting with all the generators of the other(s), but not amongst themselves.

From these two definitions it follows that all simple algebras are semisimple since they are already in (single) block form. Non-simple semisimple algebras must contain a proper, non-Abelian, invariant subalgebra.

Abelian Lie algebras (eg. $\mathfrak{u}(1), \mathfrak{s o}(2)$ ) are not semisimple, and therefore not simple. Apart from $\mathfrak{s o}(4)$ (see below), the non-Abelian $\mathfrak{s o}(n)$ algebras are all simple, and so are the $\mathfrak{s u}(n)$ and $\mathfrak{s l}(n, \mathbb{R})$ algebras.

Example 3.12. From eq. (3.17), no basis generator of $\mathfrak{s o ( 4 )}$ commutes with all others: The algebra has no non-zero centre! It is therefore ${ }^{\dagger}$ semisimple. Its structure constants determine the 6 -dim regular representation of a generic element of $\mathfrak{s o}(4)$ in block-diagonal form:

$$
\mathbf{R}=\left(\begin{array}{cccccc}
0 & a_{3} & -a_{2} & 0 & 0 & 0 \\
-a_{3} & 0 & a_{1} & 0 & 0 & 0 \\
a_{2} & -a_{1} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & b_{3} & -b_{2} \\
0 & 0 & 0 & -b_{3} & 0 & b_{1} \\
0 & 0 & 0 & b_{2} & -b_{1} & 0
\end{array}\right)
$$

The blocks cannot be further reduced, $\mathfrak{s o}(3)$ being simple; $\mathfrak{s o}(4)$ is semisimple, but not simple.

### 3.4.4 The Cartan-Killing form

Again, we recall that a Lie algebra is a vector space. As such, not only does it have a basis which can be chosen at our convenience, it can also be equipped with a (non-unique!) inner product. One such inner product ${ }^{\ddagger}$ is:

$$
(Y, Z)=\operatorname{Tr} \mathbf{Y} \mathbf{Z}
$$

Definition 3.15. The Cartan-Killing form (CK-form) of a Lie algebra is a symmetric bilinear form whose components are defined as the inner product of all pairs of elements of the algebra in their adjoint (regular) representation:

$$
\begin{equation*}
(Y, Z)=\operatorname{Tr}\left(\mathbf{R}_{Y} \mathbf{R}_{Z}\right)=\left(\mathbf{R}_{Y}\right)_{l}^{k}\left(\mathbf{R}_{Z}\right)_{k}^{l} \tag{3.22}
\end{equation*}
$$

The CK-form for basis generators $X_{i}$ is easily calculated: $\left(X_{i}, X_{j}\right)=C_{i l}{ }^{k} C_{j k}{ }^{l}$. If the algebra has $n$ parameters, the CK-form has $n(n+1) / 2$ components.

[^14]An important property of the CK-form is its invariance under the action of any element g in the Lie group associated with a Lie algebra. Let $X$ and $Y$ be elements of a Lie Algebra. Then:

$$
\left(\mathrm{g} X \mathrm{~g}^{-1}, \mathrm{~g} Y \mathrm{~g}^{-1}\right)=\operatorname{Tr}\left(\mathbf{R}_{\mathrm{g}} \mathbf{R}_{X} \mathbf{R}_{\mathrm{g}}{ }^{-1} \mathbf{R}_{\mathrm{g}} \mathbf{R}_{Y} \mathbf{R}_{\mathrm{g}}{ }^{-1}\right)=\operatorname{Tr}\left(\mathbf{R}_{X} \mathbf{R}_{Y}\right)=(X, Y)
$$

where we have used the property $\operatorname{Tr} \mathbf{A} \mathbf{B}=\operatorname{Tr} \mathbf{B} \mathbf{A}$. Linearising this after writing $g=\mathrm{e}^{\epsilon Z}$, we obtain (EXERCISE):

$$
\begin{equation*}
([Z, X], Y)+(X,[Z, Y])=0 \tag{3.23}
\end{equation*}
$$

This could be shown directly from $(Y, Z)=\operatorname{Tr} \mathbf{Y} \mathbf{Z}$, but it would not be so easy to interpret the result. Because it can be derived from the invariance of the CK-form under the group of which the algebra is the linearisation, we can say that eq. (3.23) is a manifestation of the invariance of the CK-form under the algebra as well as the group.

Definition 3.16. A CK-form is degenerate (or singular) if there exists at least one element $Z$ in the algebra for which $(Z, Y)=0 \forall Y \in \mathfrak{g}$. This can happen only if the matrix with elements $\left(X_{i}, X_{j}\right)$ has a row and column entirely populated with zeros, which forces its determinant to vanish. Otherwise, the CK-form is non-degenerate.

Alternatively, a CK-form is non-degenerate if there exists a basis in which all its diagonal elements are non-zero (so that its determinant does not vanish). Then we say that it induces a Cartan metric $\mathbf{g}$ on a Lie algebra, with components $g_{\mu \nu}=\left(X_{\mu}, X_{\nu}\right)$, where $\left\{X_{\mu}\right\}$ is that basis. If the algebra is compact, we can transform to an orthonormal Cartan metric $\mathbf{g}=k \mathbf{I}_{n}$; if the algebra is non-compact, we can transform to an indefinite metric $k \mathbf{I}_{p}^{q}$, with $p+q=n$, the dimension of the algebra (and rank of the adjoint rep). In these two cases, it is habitual to call $\mathbf{I}_{n}$ and $\mathbf{I}_{p}^{q}$ themselves the metric, which is then manifestly orthonormal.

Like all metrics, an orthonormal Cartan metric can be used to raise and lower indices. In particular, introduce $f_{\mu \nu \lambda}=C_{\mu \nu}^{\rho} g_{\rho \lambda}$. Inserting $g_{\rho \lambda}=\left(X_{\rho}, X_{\lambda}\right)$, and using eq. (3.23), it is straightforward to show (EXERCISE) that $f_{\mu \nu \lambda}$ is antisymmetric in all its indices.

Now, if a Lie algebra has a non-zero centre (ie. a subalgebra whose elements commute with all the elements of the algebra), its CK-form must be degenerate because the adjoint representation of any element of the centre vanishes trivially. We have also said that the centre of an algebra is an Abelian invariant subalgebra. Cartan's criterion asserts that a CK-form is degenerate if, and only if, an algebra has an Abelian invariant subalgebra that is, if there exists at least one generator which has vanishing commutator with all other generators.

This leads to the useful alternate definition:

Definition 3.17. A Lie algebra is semisimple if, and only if, its CK-form is non-degenerate,
Example 3.13. $x^{i} \partial_{j}$ is a basis of the operator realisation of $\mathfrak{g l}(3, \mathbb{R})$. Then $x^{i} \partial_{i}$ commutes with every other element of the algebra, and $\mathfrak{g l}(3, \mathbb{R})$ has a non-zero centre; its CK-form is degenerate, therefore, it is not semisimple.

Example 3.14. In example 3.11, we have already obtained the adjoint representation for the generators of $\mathfrak{s u}(2)$. If we use the more standard representation corresponding to $\mathbf{S}=\boldsymbol{\sigma} / 2$, then the adjoint representation matrix is merely divided by 2 , and is also the one for $\mathfrak{s o}(3)$ because the structure constants for the two algebras are now identical. With $\mathbf{S}$ in the adjoint rep, eq. (3.22) then gives:

$$
\left(S_{i}, S_{i}\right)=\operatorname{Tr}\left(\mathbf{S}_{1} \mathbf{S}_{1}\right)=\operatorname{Tr}\left(\mathbf{S}_{2} \mathbf{S}_{2}\right)=\operatorname{Tr}\left(\mathbf{S}_{3} \mathbf{S}_{3}\right)=2
$$

with all other traces 0 . The CK-form is then $2 \mathbf{I}$. This confirms that the CK-form for $\mathfrak{s u}(2)$ induces an invertible definite (Euclidean) orthonormal metric, $\mathbf{g}=\mathbf{I}$. Therefore, the group is compact as well as semisimple, and we can write the structure constants as the skew-symmetric $f_{i j k}=\mathrm{i} \epsilon_{i j k}$ (or in some sources just $\epsilon_{i j k}$ ).

The CK-form as defined above in terms of the regular representation for each generator can be tedious to calculate if we need all its components. That representation generates huge matrices for any but the smallest groups. If it is invertible (non-degenerate), however, we can extract useful information from it with much less work by working in terms of the parameters $a^{\mu}$ such that $\mathbf{R}=a^{\mu} \mathbf{X}_{\mu}$ :

$$
\begin{equation*}
(R, R)=a^{\mu} a^{\nu} \operatorname{Tr}\left(\mathbf{X}_{\mu} \mathbf{X}_{\nu}\right)=a^{\mu} a^{n} u g_{\mu \nu}=a^{\mu} a_{\mu} \tag{3.24}
\end{equation*}
$$

where we have used the Cartan metric to lower the parameter index. Inspection of $a^{\mu} a_{\mu}$ will reveal whether the group is compact or non-compact.

Example 3.15. Going back to the defining representation that we wrote for $Z \in \mathfrak{s l}(2, \mathbb{R})$ in section (3.3.5):

$$
\mathbf{Z}=\left(\begin{array}{cc}
x_{1} & x_{2}+x_{3} \\
x_{2}-x_{3} & -x_{1}
\end{array}\right)=x_{1}\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)+x_{2}\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right)+x_{3}\left(\begin{array}{rr}
0 & 1 \\
-1 & 0
\end{array}\right)
$$

The corresponding independent non-zero structure constants are: $C_{12}{ }^{3}=2, C_{31}{ }^{2}=-2$, and $C_{23}{ }^{1}=$ -2 . From these we build the regular-representation matrix:

$$
\mathbf{R}=\left(\begin{array}{ccc}
0 & -2 x_{3} & -2 x_{2} \\
2 x_{3} & 0 & 2 x_{1} \\
-2 x_{2} & 2 x_{1} & 0
\end{array}\right)
$$

Now, we only need to calculate the diagonal elements of $\mathbf{R}^{2}$ and sum them to get: $(R, R)=8\left(x_{1}^{2}+\right.$ $x_{2}^{2}-x_{3}^{2}$ ). We deduce that the algebra is non-compact: $X_{1}$ and $X_{2}$ are non-compact, while $X_{3}$ is compact. This is consistent with what we found earlier about the $S L(2, \mathbb{R})$ manifold.
Interestingly enough, using the defining representation directly, we would find (EXERCISE) $2\left(x_{1}^{2}+\right.$ $x_{2}^{2}-x_{3}^{2}$ ). Far from being a coincidence, this opens up the possibility of calculating $a^{\mu} a_{\mu}$ in eq. (3.24) directly from the defining representation instead of the more unwieldy regular representation. This is because for semisimple algebras the defining and regular representations are both faithful, and thus contain the same amount of information.

### 3.4.5 Cartan subalgebra of a semismple algebra

Now we would very much like to find whether some elements $H_{i}$ of a semisimple algebra have a diagonalisable adjoint-representation matrix, and satisfy the eigenvalue equation:

$$
\begin{equation*}
R_{H_{i}}(Y)=\left[H_{i}, Y\right]=\lambda_{Y} Y \tag{3.25}
\end{equation*}
$$

for some $Y \in \mathfrak{g}$, which makes $Y$ an eigengenerator of $H_{i}$. In fact, we would like to know the maximal subset of $r$ simultaneously diagonalisable operators in a given Lie algebra. These operators, $H_{i}$, must all commute between themselves, ie. they form an Abelian (but non-invariant!) subalgebra $\mathfrak{H}$.

Definition 3.18. A maximal Abelian subalgebra of a semisimple Lie algebra is called a Cartan subalgebra $\mathfrak{h}$. Its dimension $r<n$ defines the rank of the algebra. It is not unique, although its dimension is. The elements of a Cartan subalgebra are called the Cartan generators of the algebra.

Example 3.16. Let $\left\{J_{+}, J_{-}, J_{0}\right\}=\left\{J_{1}+\mathrm{i} J_{2}, J_{1}-\mathrm{i} J_{2}, J_{3}\right\}$ be a basis of a Lie algebra, with $\left[J_{i}, J_{j}\right]=\mathrm{i} \epsilon_{i j}{ }^{k} J_{k}$. These commutation relations translate into: $\left[J_{0}, J_{ \pm}\right]= \pm J_{ \pm},\left[J_{+}, J_{-}\right]=2 J_{0}$. The corresponding adjoint-representation matrices of $J_{0}$ and $J_{+}$are immediately found:

$$
\operatorname{ad}_{J_{0}}=\left(\begin{array}{ccc}
1 & 0 & 0 \\
0 & -1 & 0 \\
0 & 0 & 0
\end{array}\right) \quad \operatorname{ad}_{J_{+}}=\left(\begin{array}{rrr}
0 & 0 & 0 \\
0 & 0 & 2 \\
-1 & 0 & 0
\end{array}\right)
$$

Because ad $J_{0}$ is diagonal, $J_{0}$ is a Cartan generator; comparing with eq. (3.25), ad $J_{J_{0}}$ has a complete set $\left\{J_{+}, J_{-}, J_{0}\right\}$ of eigengenerators for the corresponding eigenvalues $\{1,-1,0\}$. But ad $J_{+}$has only one eigenvalue, 0 , and therefore one eigengenerator which cannot form a basis of the algebra. It is not diagonalisable and not a Cartan generator; the same holds for $J_{-}$. Thus, the algebra contains only one Cartan generator and is of rank 1.

Another important thing we learn from this is that the structure constants in the complex extension of an algebra can be quite different from those of the algebra itself, even in its essentially real version. Indeed, the adjoint representation of $J_{3}$ found - up to an overall constant - in example 3.11 is not diagonal, and has only zeros on its diagonal, in contrast with with $\mathrm{ad}_{J_{0}}$ in the complex extension. Of course, this does not affect the CK-form which, being a trace, is basis-independent. We also notice that the diagonal elements of ad $_{J_{0}}$ are precisely the structure constants involving $J_{0}$. Is this a coincidence?
We arrive at the same results by working out the secular equation for eq. (3.25), $\left\|\mathbf{R}-\lambda \mathbf{I}_{3}\right\|=0$, $\mathbf{R}$ being the regular presentation of example 3.11 for the basis $\left\{J_{i}\right\}=\left\{\boldsymbol{\sigma}_{i} / 2\right\}$. Then $\left\|\mathbf{R}-\lambda \mathbf{I}_{3}\right\|=$ $-\lambda\left(\lambda^{2}+a^{2}\right)=0$, where $a^{2}=a_{1}^{2}+a_{2}^{2}+a_{3}^{2}$. The fact that the eigenvalues $\{0, \pm \mathrm{i} a\}$ are complex tells us that we should use the complex extension of the algebra (which indeed we did when using the basis $\left\{J_{+}, J_{-}, J_{0}\right\}!$, and the number of independent non-zero eigenvalues gives the rank of the algebra, here 1.
The secular equation itself also yields the rank of the algebra: the number of independent functions of the parameters (just one, $a$, in our example) in the secular equation is the rank of the algebra. It can be shown that the rank of a $\mathfrak{s u}(n)$ algebra is $n-1$; also, $\mathfrak{s o}(2 n)$ and $\mathfrak{s o}(2 n+1)$ algebras have rank $n$.

An important point is that the same analysis could have been carried through using the three $2 \times 2$ matrices of the defining representation of the algebra $(\mathfrak{s u}(2)$ ), with the same conclusions. Indeed, in the usual Pauli matrix basis:

$$
\left\|\mathfrak{s u}(2)-\lambda \mathbf{I}_{2}\right\|=\left|\begin{array}{cc}
a_{3}-\lambda & a_{1}-\mathrm{i} a_{2} \\
a_{1}+\mathrm{i} a_{2} & -a_{3}-\lambda
\end{array}\right|=\lambda^{2}-a^{2}=0 \quad a^{2}=a_{1}^{2}+a_{2}^{2}+a_{3}^{2}
$$

gives $\lambda= \pm a$. Again, the secular equation has one independent function of the parameters.
As noted previously, the advantage of the defining representation is that the dimension of its matrices is most often much smaller than those of the adjoint representation.

### 3.5 Weights and Roots of a Representation of a Compact Semisimple Algebra

Definition 3.19. Let $|\mu\rangle$ be an eigenvector common to all Cartan basis generators $H_{i}$, living in the carrier space of some representation $\mathbf{D}$ of the generators. Then $H_{i}|\mu\rangle=\mu_{i}|\mu\rangle$. The set $\left\{\mu_{i}\right\}(1 \leq$ $i \leq r$ ) corresponding to each eigenvector can be viewed as the components of a $r$-dimensional vector called a weight $\mu$ of the representation. The number of these weights, ie. the number of eigenvectors, is equal to the dimension of $\mathbf{D}$.

Definition 3.20. Following standard notation, let $\left\{E_{\boldsymbol{\alpha}} \in \mathfrak{g}\right\}$ denote the set of eigengenerators of an element $H_{i}$ of the Cartan subalgebra of the complex extension of a semisimple Lie algebra, such that $\left[H_{i}, E_{\boldsymbol{\alpha}}\right]=\alpha_{i} E_{\boldsymbol{\alpha}}, 1 \leq i \leq r$. Then the set of eigenvalues, $\left\{\alpha_{i}\right\}$, of the $\left[H_{i}, \quad\right]$ operator associated with each eigengenerator $E_{\alpha}$ can be viewed as the components of a $r$-dimensional vector called the $\boldsymbol{r o o t} \boldsymbol{\alpha}$. We can also write $\left[\mathbf{H}, E_{\boldsymbol{\alpha}}\right]=\boldsymbol{\alpha} E_{\boldsymbol{\alpha}}$.

Do keep in mind the crucial distinction between the eigengenerators, whose associated eigenvalues are the root components, and the eigenvectors that live in the carrier space and whose eigenvalues are the weights.

Clearly, any Cartan generator $H_{i}$ has the root $\boldsymbol{\alpha}=0$, because $\left[H_{i}, H_{j}\right]=0 \forall j$. Thus, we can write an algebra $\mathfrak{g}$ as the sum of its Cartan subalgebra, with roots zero, and the non-Cartan generators with non-zero roots. The set
of all non-zero roots is called the root system of the algebra. The Cartan and non-Cartan generators form a basis of the algebra, called its Cartan-Weyl basis.

As we are soon to discover, all the information about a semisimple algebra is encoded in its system of $r$ dimensional roots. Like the weight vectors, these roots can be seen as living in a $r$-dimensional space. A metric is induced on this root space by the metric of the Cartan subalgebra, so that we can represent it as having $r$ Cartesian axes, each associated with a Cartan generator $H_{i}$. The root vectors can then be represented in a root diagram. The $i^{\text {th }}$ component of each root is the projection of the root along the $H_{i}$ axis. This root space is almost always much easier to work with than the algebra itself, simply because it is of smaller dimension.

Also, each $n$-dimensional representation (often called a multiplet) of a semisimple Lie algebra is associated with $n$ weights which are $r$-dim "vectors" with as $i^{\text {th }}$ component the corresponding eigenvalue of the simultaneously diagonalisable Cartan generators, $H_{i}$, of the algebra, acting on their carrier function space, as in definition 3.19. Just like roots, these weights can be plotted on a diagram, the weight diagram, or lattice, with each eigenvalue component of a weight being its projection along the associated $H_{i}$ axis.

### 3.5.1 Properties of eigengenerators in the Cartan-Weyl basis

Those eigengenerators of $H_{i}, E_{\boldsymbol{\alpha}} \in \mathfrak{g}$ (all of them generators in the complex extension!), which are not Cartan generators are quite interesting. An important fact, which we shall not prove, is that they are uniquely labelled by their roots. To each non-zero root corresponds one and only one such generator, which spans a 1-dim subalgebra.

Now let $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ be two non-zero roots. Then, from the Jacobi identity and definition 3.20 of the roots, there comes:

$$
\left[H_{i},\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]\right]=\left[\left[H_{i}, E_{\boldsymbol{\alpha}}\right], E_{\boldsymbol{\beta}}\right]+\left[E_{\boldsymbol{\alpha}},\left[H_{i}, E_{\boldsymbol{\beta}}\right]\right]=\left(\alpha_{i}+\beta_{i}\right)\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]
$$

Therefore, either $\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]=0$, which happens when $\boldsymbol{\alpha}+\boldsymbol{\beta}$ is not a root, or it is proportional to the eigengenerator $E_{\boldsymbol{\alpha}+\boldsymbol{\beta}}$ of $H_{i}$, with eigenvalue $\alpha_{i}+\beta_{i}$. and we can write:

$$
\begin{equation*}
\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]=C_{\boldsymbol{\alpha} \boldsymbol{\beta}} E_{\boldsymbol{\alpha}+\boldsymbol{\beta}} \tag{3.26}
\end{equation*}
$$

Note how only one generator contributes to the commutator in the Cartan-Weyl basis when $\boldsymbol{\beta} \neq-\boldsymbol{\alpha}$. When either $\boldsymbol{\alpha}$ or $\boldsymbol{\beta}$ vanishes, $C_{\boldsymbol{\alpha} \boldsymbol{\beta}}$ is equal to the other root, and we recover the defining equation for the roots in def. 3.20.

Going to definition 3.10 of the adjoint represepresentation, one should now be able to see that, in the CartanWeyl basis, the adjoint representation of a Cartan generator is a diagonal matrix, with $r$ 0s and $\alpha_{i}$ as the rest of the $n-r$ diagonal entries. And all diagonal entries of the adjoint representation of any other generator $E_{\boldsymbol{\alpha}}$ must be zero. From this, the following statements about the CK-form of an algebra can be derived (EXERCISE):

$$
h_{i j}=\left(H_{i}, H_{j}\right)=k_{\mathrm{D}} \delta_{i j}, \quad\left(H_{i}, E_{\boldsymbol{\alpha}}\right)=0, \quad\left(E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right)=0 \quad(\boldsymbol{\alpha}+\boldsymbol{\beta} \neq 0 \text { and not a root })
$$

where $h_{i j}$ are the components of the metric of the Cartan subalgebra.
To go further, work with Hermitian Cartan generators: $H_{i}^{\dagger}=H_{i}$ of the essentially real algebra. Then, if $\left[H_{i}, E_{\boldsymbol{\alpha}}\right]=\alpha_{i} E_{\boldsymbol{\alpha}}$, we immediately find by taking the adjoint that $\left[H_{i}, E_{\boldsymbol{\alpha}}^{\dagger}\right]=-\alpha_{i} E_{\boldsymbol{\alpha}}^{\dagger}$, which we can write as $\left[H_{i}, E_{-\boldsymbol{\alpha}}\right]=-\alpha_{i} E_{-\boldsymbol{\alpha}}$. Thus, non-Cartan generators and non-zero roots always come in pairs, $\left\{E_{\boldsymbol{\alpha}}, E_{-\boldsymbol{\alpha}}=E_{\boldsymbol{\alpha}}^{\dagger}\right\}$ and $\pm \boldsymbol{\alpha}$. In fact, $-\boldsymbol{\alpha}$ is the only possible multiple of $\boldsymbol{\alpha}$ that is a root. The generators $E_{ \pm \alpha}$ of the complex extension are computed from the pairs $X_{k}$ and $X_{l}$ of non-Cartan generators of the algebra as: $\left(X_{k} \pm \mathrm{i} X_{l}\right) / \sqrt{2}$.

When $\boldsymbol{\beta}=-\boldsymbol{\alpha}$, eq. (3.26) would map $\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]$ to a generator with zero root, ie. one that lives in the Cartan subalgebra. Therefore, $\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]$ must belong to $\mathfrak{H}$ as a linear combination of the $r$ Cartan basis generators:

$$
\left[E_{\boldsymbol{\alpha}}, E_{-\boldsymbol{\alpha}}\right]=\lambda^{i} H_{i} \quad 1 \leq i \leq r
$$

Now take the inner product of this equation with $H_{j}$. For semisimple algebras, we know that $\left(X_{i}, X_{j}\right)=g_{i j}$. Then the right-hand side gives $\lambda^{j} h_{i j}$. For the left-hand side, we have:

$$
\left(H_{i},\left[E_{\boldsymbol{\alpha}}, E_{-\boldsymbol{\alpha}}\right]\right)=\left(E_{-\boldsymbol{\alpha}},\left[H_{i}, E_{\boldsymbol{\alpha}}\right]\right)=\alpha_{i}\left(E_{-\boldsymbol{\alpha}}, E_{\boldsymbol{\alpha}}\right)
$$

where we have used eq. (3.23) (or, alternatively, the cyclic property of the trace of a product). Therefore, $\lambda^{i}=$ $h^{i j} \alpha_{j}\left(E_{-\boldsymbol{\alpha}}, E_{\boldsymbol{\alpha}}\right)$. Now the $E$ generators live in the complex extension, so that $\left(E_{-\boldsymbol{\alpha}}, E_{\boldsymbol{\alpha}}\right)=\left(X_{k}, X_{k}\right)=g_{k k}$. We also recall that for a compact semisimple algebra, $\mathbf{g}=k \mathbf{l}$ : the metric on $\mathfrak{g}$, including its restriction on $\mathfrak{H}$, is diagonal with identical entries. Then $h^{i i}=1 / g_{k k}$, and $\lambda^{i}=\alpha_{i}$. We arrive at yet another useful result for the commutator of two non-Cartan generators with opposite root vectors:

$$
\begin{equation*}
\left[E_{\boldsymbol{\alpha}}, E_{-\boldsymbol{\alpha}}\right]=\alpha_{i} H_{i} \equiv \boldsymbol{\alpha} \cdot \mathbf{H} \tag{3.27}
\end{equation*}
$$

Here, $\boldsymbol{\alpha} \cdot \mathbf{H}$ should be seen as just another way to write $\alpha_{i} H_{i}$, and there is a summation over $i$ implied.
Now is a good time to discover what those non-Cartan generators of the complex extension do for a living. We have:

$$
H_{i} E_{ \pm \boldsymbol{\alpha}}=\left[H_{i}, E_{ \pm \boldsymbol{\alpha}}\right]+E_{ \pm \boldsymbol{\alpha}} H_{i}= \pm \alpha_{i} E_{ \pm \boldsymbol{\alpha}}+E_{ \pm \boldsymbol{\alpha}} H_{i}
$$

When this acts on a common eigenvector of the $H_{i}$ in the carrier space, there comes:

$$
\begin{equation*}
H_{i} E_{ \pm \boldsymbol{\alpha}}|\mu\rangle= \pm \alpha_{i} E_{ \pm \boldsymbol{\alpha}}|\mu\rangle+E_{ \pm \boldsymbol{\alpha}} H_{i}|\mu\rangle=\left(\mu_{i} \pm \alpha_{i}\right) E_{ \pm \boldsymbol{\alpha}}|\mu\rangle \tag{3.28}
\end{equation*}
$$

We see that $E_{ \pm \boldsymbol{\alpha}}|\mu\rangle$ is an eigenvector of $H_{i}$ with eigenvalue $\mu_{i} \pm \alpha_{i}$, from which we conclude that the $E_{ \pm \boldsymbol{\alpha}}$ act as raising and lowering operators on the carrier space of the Cartan generators, changing weights by $\pm \boldsymbol{\alpha}$. This result holds for any irreducible representation. Thus, if we can find the weight vectors, the root vectors must be all the possible differences between them; often, this is the quickest way to obtain the roots.

Example 3.17. Start with the defining representation of the 3-dim $\mathfrak{s u}(2)$ algebra:

$$
J_{1}=\frac{1}{2}\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right) \quad J_{2}=\frac{1}{2}\left(\begin{array}{rr}
0 & -\mathrm{i} \\
\mathrm{i} & 0
\end{array}\right) \quad J_{0}=\frac{1}{2}\left(\begin{array}{rr}
1 & 0 \\
0 & -1
\end{array}\right)
$$

$\mathfrak{s u}(2)$ is semisimple, and one Cartan and one pair of non-Cartan generators can fit in it, thus one independent non-zero root vector. This is an algebra of rank 1, and the roots are one-dimensional. The diagonal generator $J_{0}$ is identified with the sole Cartan generator. The weights of $J_{0}$ in the defining representation are $1 / 2$ and $-1 / 2$ (corresponding to the doublet of eigenvectors $\binom{1}{0}$ and $\binom{0}{1}$ ). Once the weights have been found, the roots must be among all the possible differences between the weights, ie. $\pm 1$. These roots raise or lower the weights by 1.
Then, without any direct computation, we can state from eq. (3.27) that $\left[E_{1}, E_{-1}\right]=J_{0}$. Also, from the definition of roots, $\left[J_{0}, E_{ \pm 1}\right]= \pm E_{ \pm 1}$. The structure of the algebra then determines the non-Cartan generators. $\left[J_{0}, E_{ \pm 1}\right]= \pm E_{ \pm 1}$ gives, up to a normalisation constant $A$ :

$$
E_{1}=A\left(\begin{array}{ll}
0 & 1 \\
0 & 0
\end{array}\right)=A\left(J_{1}+\mathrm{i} J_{2}\right), \quad \quad E_{-1}=A\left(\begin{array}{ll}
0 & 0 \\
1 & 0
\end{array}\right)=A\left(J_{1}-\mathrm{i} J_{2}\right)
$$

$A$ is determined by the other commutation relation, yielding finally: $E_{ \pm 1}=\left(J_{1} \pm \mathrm{i} J_{2}\right) / \sqrt{2}$, as we should have expected. We recognise the set $\left\{J_{0}, E_{ \pm 1}\right\}$ as the basis for the complex extension of $\mathfrak{s u}(2)$.
Because $\mathfrak{s u}(2)$ has rank 1 , this is a somewhat trivial application of the machinery we are developing, but it comes into its own with algebras of higher rank.

### 3.5.2 Structure constants and the root system

We now concern ourselves with finding the (real!) structure constants $C_{\boldsymbol{\alpha} \boldsymbol{\beta}}$ in eq. (3.26), without explicit calculation of commutators. They satisfy several symmetry relations. For instance:

$$
\begin{equation*}
C_{\boldsymbol{\beta} \boldsymbol{\alpha}}=-C_{\boldsymbol{\alpha} \boldsymbol{\beta}} \quad C_{-\boldsymbol{\alpha},-\boldsymbol{\beta}}=-C_{\boldsymbol{\alpha} \boldsymbol{\beta}}^{*}=-C_{\boldsymbol{\alpha} \boldsymbol{\beta}} \tag{3.29}
\end{equation*}
$$

The first is read off the equation, while taking the adjoint of the equation gives the second relation.

Also, let $\boldsymbol{\alpha}, \boldsymbol{\beta}$, and $\boldsymbol{\alpha}+\boldsymbol{\beta}$ be non-zero roots; then $\boldsymbol{\gamma}=-(\boldsymbol{\alpha}+\boldsymbol{\beta})$ is also a non-zero root, Using the Jacobi identity on $E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}$, and $E_{\gamma}$, plus eq. (3.26) and (3.27), leads (EXERCISE) to:

$$
\left(\boldsymbol{\alpha} C_{\boldsymbol{\beta} \gamma}+\boldsymbol{\beta} C_{\gamma \boldsymbol{\alpha}}+\gamma C_{\boldsymbol{\alpha} \boldsymbol{\beta}}\right) \cdot \mathbf{H}=0
$$

since the $H_{i}$ are linearly independent, this equation can only be satisfied if:

$$
\boldsymbol{\alpha} C_{\boldsymbol{\beta} \gamma}+\boldsymbol{\beta} C_{\gamma \boldsymbol{\alpha}}+\gamma C_{\boldsymbol{\alpha} \boldsymbol{\beta}}=\boldsymbol{\alpha}\left(C_{\boldsymbol{\beta} \boldsymbol{\gamma}}-C_{\boldsymbol{\alpha} \boldsymbol{\beta}}\right)+\boldsymbol{\beta}\left(C_{\boldsymbol{\gamma} \boldsymbol{\alpha}}-C_{\boldsymbol{\alpha} \boldsymbol{\beta}}\right)=0
$$

which yields additional symmetries on the structure constants of a semisimple algebra:

$$
\begin{equation*}
C_{\boldsymbol{\beta},-\boldsymbol{\alpha}-\boldsymbol{\beta}}=C_{-\boldsymbol{\alpha}-\boldsymbol{\beta}, \boldsymbol{\alpha}}=C_{\boldsymbol{\alpha} \boldsymbol{\beta}} \tag{3.30}
\end{equation*}
$$

A very important result can now be derived.
Going back again to eq. (3.26), we can write: $\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}+\boldsymbol{\alpha}}\right]=C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+\boldsymbol{\alpha}} E_{\boldsymbol{\beta}+2 \alpha}, \ldots,\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}+k \boldsymbol{\alpha}}\right]=$ $C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+k \boldsymbol{\alpha}} E_{\boldsymbol{\beta}+(k+1) \boldsymbol{\alpha}}$. But there must exist a value $k=p \geq 0$ such that $\boldsymbol{\beta}+(p+1) \boldsymbol{\alpha}$ is not a root, so that $C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+p \boldsymbol{\alpha}}=0$. Similarly, if we start from: $\left[E_{-\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}}\right]=C_{-\boldsymbol{\alpha}, \boldsymbol{\beta}} E_{\boldsymbol{\beta}-\boldsymbol{\alpha}}$, there must exist a value $k=-q \leq 0$ such that $\boldsymbol{\beta}-(q+1) \boldsymbol{\alpha}$ is not a root, and $C_{-\boldsymbol{\alpha}, \boldsymbol{\beta}-q \boldsymbol{\alpha}}=0$.

Next, start from the always useful Jacobi identity and evaluate the commutators using eq. (3.26) and (3.27):

$$
\begin{array}{r}
{\left[E_{\boldsymbol{\alpha}},\left[E_{\boldsymbol{\beta}+k \boldsymbol{\alpha}}, E_{-\boldsymbol{\alpha}}\right]\right]+\left[E_{\boldsymbol{\beta}+k \boldsymbol{\alpha}},\left[E_{-\boldsymbol{\alpha}}, E_{\boldsymbol{\alpha}}\right]\right]+\left[E_{-\boldsymbol{\alpha}},\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}+k \boldsymbol{\alpha}}\right]\right]=0} \\
{\left[E_{\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}+(k-1) \boldsymbol{\alpha}}\right] C_{\boldsymbol{\beta}+k \boldsymbol{\alpha},-\boldsymbol{\alpha}}-\left[E_{\boldsymbol{\beta}+k \boldsymbol{\alpha}}, \boldsymbol{\alpha} \cdot \mathbf{H}\right]+\left[E_{-\boldsymbol{\alpha}}, E_{\boldsymbol{\beta}+(k+1) \boldsymbol{\alpha}}\right] C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+k \boldsymbol{\alpha}}=0} \\
C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+(k-1) \boldsymbol{\alpha}} C_{\boldsymbol{\beta}+k \boldsymbol{\alpha},-\boldsymbol{\alpha}}+\boldsymbol{\alpha} \cdot(\boldsymbol{\beta}+k \boldsymbol{\alpha})+C_{-\boldsymbol{\alpha}, \boldsymbol{\beta}+(k+1) \boldsymbol{\alpha}} C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+k \boldsymbol{\alpha}}=0
\end{array}
$$

Applying the symmetry relations (3.29) and then (3.30) to the first and last term on the left yields the recursion relation:

$$
C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+(k-1) \boldsymbol{\alpha}}^{2}=C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+k \boldsymbol{\alpha}}^{2}+\boldsymbol{\alpha} \cdot(\boldsymbol{\beta}+k \boldsymbol{\alpha})
$$

We already know that, by definition of $p, C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+p \boldsymbol{\alpha}}=0$. Then, from our recursion relation, $C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+(p-1) \boldsymbol{\alpha}}^{2}=$ $\boldsymbol{\alpha} \cdot \boldsymbol{\beta}+p|\boldsymbol{\alpha}|^{2}, C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+(p-2) \boldsymbol{\alpha}}^{2}=C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+(p-1) \boldsymbol{\alpha}}^{2}+\boldsymbol{\alpha} \cdot \boldsymbol{\beta}+(p-1)|\boldsymbol{\alpha}|^{2}=2 \boldsymbol{\alpha} \cdot \boldsymbol{\beta}+(p-2)|\boldsymbol{\alpha}|^{2}$, etc. Generically:

$$
C_{\boldsymbol{\alpha}, \boldsymbol{\beta}+(k-1) \boldsymbol{\alpha}}^{2}=(p-k+1)\left[\boldsymbol{\alpha} \cdot \boldsymbol{\beta}+\frac{p+k}{2}|\boldsymbol{\alpha}|^{2}\right]
$$

The recursion stops when $k=-q$, ie. when $C_{-\boldsymbol{\alpha}, \boldsymbol{\beta}-q \boldsymbol{\alpha}}=-C_{\boldsymbol{\alpha},-(\boldsymbol{\beta}-q \boldsymbol{\alpha})}=-C_{\boldsymbol{\beta}-(q+1) \boldsymbol{\alpha}, \boldsymbol{\alpha}}=0$ :

$$
0=C_{\boldsymbol{\alpha}, \boldsymbol{\beta}-(q+1) \boldsymbol{\alpha}}^{2}=(p+q+1)\left[\boldsymbol{\alpha} \cdot \boldsymbol{\beta}+\frac{p-q}{2}|\boldsymbol{\alpha}|^{2}\right]
$$

or:

$$
\begin{equation*}
2 \frac{\boldsymbol{\alpha} \cdot \boldsymbol{\beta}}{|\boldsymbol{\alpha}|^{2}}=-(p-q) \tag{3.31}
\end{equation*}
$$

the so-called " master formula". We will use it to prove a crucial fact a little later.
Each pair $F_{ \pm \alpha} \equiv \sqrt{2} E_{ \pm \alpha} /|\alpha|$ of normalised non-Cartan generators of a semisimple algebra, together with the combination: $\tilde{H}_{\boldsymbol{\alpha}}=2 \boldsymbol{\alpha} \cdot \mathbf{H} /|\boldsymbol{\alpha}|^{2}$, forms a $\mathfrak{s u}(2)$ subalgebra. There is a distinct $\mathfrak{s u}(2)$ subalgebra for each pair of non-zero roots. Indeed:

$$
\left[\frac{2 \boldsymbol{\alpha} \cdot \mathbf{H}}{|\boldsymbol{\alpha}|^{2}}, \frac{\sqrt{2} E_{ \pm \boldsymbol{\alpha}}}{|\boldsymbol{\alpha}|}\right]=2 \sqrt{2} \frac{\boldsymbol{\alpha}}{|\boldsymbol{\alpha}|^{3}} \cdot\left[\mathbf{H}, E_{ \pm \boldsymbol{\alpha}}\right]= \pm 2 \sqrt{2} \frac{|\boldsymbol{\alpha}|^{2}}{|\boldsymbol{\alpha}|^{3}} E_{ \pm \boldsymbol{\alpha}}= \pm 2\left(\frac{\sqrt{2} E_{ \pm \boldsymbol{\alpha}}}{|\boldsymbol{\alpha}|}\right)
$$

Together with eq. (3.27), we recover the $\mathfrak{s u}(2)$ structure constants in the basis of example 3.16, with $\tilde{H}_{\boldsymbol{\alpha}}=2 J_{0}$ and $F_{ \pm \alpha}=J_{ \pm}$. The non-immediately zero commutators with elements of the other $\mathfrak{s u}(2)$ subalgebras can also be found. For instance, one easily shows (EXERCISE) that $\left[\tilde{H}_{\boldsymbol{\alpha}}, F_{ \pm \boldsymbol{\beta}}\right]= \pm(q-p) F_{ \pm \boldsymbol{\beta}}$. And, provided one
accepts the following formula: $|\boldsymbol{\alpha}+\boldsymbol{\beta}| /|\boldsymbol{\beta}|=\sqrt{q+1 / p}$ (check with all possible angles allowed by eq. (3.32) below), then it is not hard to obtain: $\left[F_{ \pm \boldsymbol{\alpha}}, F_{\boldsymbol{\beta}}\right]= \pm(q+1) F_{\boldsymbol{\alpha}+\boldsymbol{\beta}}$. In this basis, all the structure constants are real integers. This result, due I think to Chevalley ( $\sim 1955$ ), means that any semisimple algebra, not just $\mathfrak{s u}(2)$, has a real version.

Thus, if a semisimple algebra has dimension $n$ and rank $r$, it contains $(n-r) / 2$ distinct $\mathfrak{s u}(2)$ subalgebras, each having as Cartan generator a different element of the Cartan subalgebra, plus two non-Cartan generators corresponding to the root that labels the $\mathfrak{s u}(2)$ subalgebra.

Once we have put all the roots on a root diagram, we notice how symmetrical the diagram looks. This is because, of course, not all the $n-r$ roots can be linearly independent; only $r$ of them can be.

Definition 3.21. A positive root is one whose first non-zero component is positive; otherwise, it is negative. The $r$ (rank of the algebra) positive roots which cannot be obtained from a linear combination of other positive roots are called simple, or independent. The other positive roots can be obtained as linear combinations of the simple roots, with positive coefficients.
Roots may be obtained from other roots via a Weyl reflection with respect to a Weyl hyperplane in the root diagram. There is a Weyl hyperplane perpendicular to each root through the origin of the root space. For instance, the reflection of a root with respect to its own Weyl hyperplane gives $\boldsymbol{\alpha} \rightarrow-\boldsymbol{\alpha}$.

Roots are in fact extremely constrained. If, when we derived eq. (3.31), we had started instead with $\left[E_{\boldsymbol{\beta}}, E_{\boldsymbol{\alpha}}\right]=$ $C_{\boldsymbol{\beta} \boldsymbol{\alpha}} E_{\boldsymbol{\alpha}+\boldsymbol{\beta}}$, and added/subtracted integer multiples of $\boldsymbol{\beta}$ to $\boldsymbol{\alpha}$, we would have found that $2 \boldsymbol{\beta} \cdot \boldsymbol{\alpha} /|\boldsymbol{\beta}|^{2}=-\left(p^{\prime}-q^{\prime}\right)$. Multiplying the two expressions, we arrive at the important expression:

$$
\begin{equation*}
\frac{(\boldsymbol{\alpha} \cdot \boldsymbol{\beta})^{2}}{|\boldsymbol{\alpha}|^{2}|\boldsymbol{\beta}|^{2}}=\cos ^{2} \theta_{\boldsymbol{\alpha} \boldsymbol{\beta}}=\frac{1}{4}(p-q)\left(p^{\prime}-q^{\prime}\right) \tag{3.32}
\end{equation*}
$$

The relative length of the roots is seen to be constrained to $|\boldsymbol{\alpha}| /|\boldsymbol{\beta}|=\sqrt{\left(p^{\prime}-q^{\prime}\right) /(p-q)}$. Also, if $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ are simple roots, $\boldsymbol{\alpha}-\boldsymbol{\beta}$ cannot be a root; otherwise, either it or $\boldsymbol{\beta}-\boldsymbol{\alpha}$ must be positive, and a simple root could be constructed out of two different positive roots: eg., $\boldsymbol{\beta}=(\boldsymbol{\beta}-\boldsymbol{\alpha})+\boldsymbol{\alpha}$. This means that $C_{\boldsymbol{\alpha},-\boldsymbol{\beta}}=0$. Comparing with $C_{\boldsymbol{\alpha},-\boldsymbol{\beta}+q \boldsymbol{\alpha}}=0$, derived from eq. (3.29) applied to $C_{-\boldsymbol{\alpha}, \boldsymbol{\beta}-q \boldsymbol{\alpha}}=0$, we see that $q=0$ for simple roots. Therefore, from the master formular (3.31), the angle between two simple roots satisfies $\cos \theta_{\boldsymbol{\alpha} \boldsymbol{\beta}} \leq 0$, so that $\pi / 2 \leq \theta_{\alpha \beta} \leq \pi$.

Since $(p-q)\left(p^{\prime}-q^{\prime}\right)$ must be an integer, There are only five possible values allowed for $\cos ^{2} \theta_{\boldsymbol{\alpha} \boldsymbol{\beta}}$ in eq. (3.32), and this, for any two roots of any semisimple algebra: $0 \Rightarrow \theta_{\alpha \beta}= \pm 90^{\circ} ; 1 / 4 \Rightarrow \theta_{\alpha \boldsymbol{\beta}}=60^{\circ}, 120^{\circ}$; $1 / 2 \Rightarrow \theta_{\boldsymbol{\alpha} \boldsymbol{\beta}}=45^{\circ}, 135^{\circ} ; 3 / 4 \Rightarrow \theta_{\boldsymbol{\alpha} \boldsymbol{\beta}}=30^{\circ}, 150^{\circ}$; and $1 \Rightarrow \theta_{\boldsymbol{\alpha} \boldsymbol{\beta}}=0^{\circ}, 180^{\circ}$.

Thanks to all these constraints, a systematic and exhaustive procedure exists to construct the root space for all four families of classical semisimple groups, and for the five so-called exceptional groups. With the subscript denoting the rank of the algebra, the four families are:

- $A_{n-1}(n>1)$, corresponding to $S U(n), S L(n, \mathbb{R}), S U(p, q)$, with $p+q=n$ (no relation to the $p$ and $q$ above!)
- $B_{n}$, corresponding to $S O(2 n+1)$ and $S O(p, q)$, with $p+q=2 n+1$.
- $C_{n}$, corresponding to $S p(n)$ and $S p(p, q)$, with $p+q=2 n$.
- $D_{n}$, corresponding to $S O(2 n)$ and $S O(p, q)$, with $p+q=n$.
$S U(2), S L(2, \mathbb{R})$, both $A_{1}, S O(2)\left(D_{1}\right)$, and $S O(3)\left(B_{1}\right)$, all have the same one-dim root space with the two roots $\pm 1$. Only five two-dimensional root spaces (four classical and one exceptional) can satisfy all our constraints; but $B_{2}$ and $C_{2}$ are rotated from each other by $45^{\circ}$, so are taken to be the same. And there are only four threedimensional root spaces. Beyond three dimensions, root spaces can no longer be represented on root diagrams.

Instead, one uses Dynkin diagrams, which are planar and represent only the simple roots and the angle between them. They are equivalent to a root diagram..

Finally, a few words about weight diagrams. One of the Cartan generators, say $H_{1}$, will always be the Cartan generator of a $\mathfrak{s u}(2)$ (and $\mathfrak{s o}(3)$ - see section 3.6 .1 below) subalgebra. Then weight points are arranged on lines parallel to the $H_{1}$ axis, with each line corresponding to an irreducible representation (multiplet) of $\mathfrak{s u}(2)$ labelled with $j$, an integer multiple of $1 / 2$, and containing $2 j+1$ weights. These weights can be generated by applying the lowering non-Cartan generator of $\mathfrak{s u}(2)$ to the weights in each $\mathfrak{s u}(2)$ multiplet, starting with the highest one, ie. by repeated addition of the $r$ - $\operatorname{dim}$ root, $(-1,0, \ldots, 0)$, to that highest weight. This root, as well as $(1,0, \ldots, 0)$ (which moves you up from the lowest to the highest weight) , is always a root of the semisimple algebra. Needless to say, as you move parallel to the $H_{1}$ axis, all other eigenvalues in the weights remain the same. since you are moving perpendicular to all other $H_{i}$ axes. One can use other roots to move from one line of weights to its neighbouring lines.

The number of weights for these different $\mathfrak{s u}(2)$ multiplets must add up to the dimension of the multiplet of the semisimple algebra. The $\mathfrak{s u}(2)$ multiplets must fit snugly inside this multiplet. For instance, take the 10dimensional representation (decuplet) of $\mathfrak{s u}(3)$. The rank of this algebra is 2, and thus the weights are 2-component "vectors". There are four eigenvalues for the other Cartan generator, $H_{2}$, in this representation, and the weights are organised on an inverted-triangle lattice made of one horizontal $\mathfrak{s u}(2)$ quadruplet, triplet, doublet and singlet, in the direction of decreasing $H_{2}$ eigenvalues.

### 3.5.3 Casimir invariant operators

Each irreducible representation of a Lie algebra can be labelled with the eigenvalues of some function of the basis generators of the algebra.

Definition 3.22. A Casimir invariant operator $C$ for a representation of a Lie algebra is an operator that commutes with all the generators of the representation.

When the representation is irreducible, $C$ has to be a multiple of the identity by Schur's lemma. All elements of an invariant subspace of the carrier space of the representation will be eigenvectors of $C$ with the same eigenvalue. When the algebra is semisimple, a theorem by Chevalley guarantees the existence of a set of Casimir operators as polynomials in the generators, whose eigenvalues may be used to characterise the irreducible representations of the algebra. More precisely, each invariant subspace of the carrier space has a set of basis vectors, each labeled by an eigenvalue of each Casimir operator. The number of Casimir operators is the rank of the algebra.

In other words, if $f(\mathbf{x})$ is in an invariant subspace of the Hilbert space which is the carrier space of the algebra, for each Casimir operator $C_{i}, C_{i} f(\mathbf{x})=g(\mathbf{x})$ is a set of equations invariant under the group transformations, in the sense that under the action of the group, any $f(\mathbf{x})$ in the invariant subspace is sent into another function $g(\mathbf{x})$ in that same invariant subspace.

Because a Cartan metric can always be defined for a semisimple algebra, I claim that $g^{\mu \nu} X_{\mu} X_{\nu}$ is a Casimir operator, where the $X_{\mu}$ are basis generators of such an algebra. Indeed:

$$
\begin{aligned}
{\left[g^{\mu \nu} X_{\mu} X_{\nu}, X_{\rho}\right] } & =g^{\mu \nu}\left(X_{\mu}\left[X_{\nu}, X_{\rho}\right]+\left[X_{\mu}, X_{\rho}\right] X_{\nu}\right) \\
& =g^{\mu \nu} C_{\mu \rho}{ }^{\lambda}\left(X_{\nu} X_{\lambda}+X_{\lambda} X_{\nu}\right) \\
& =g^{\mu \nu} g^{\alpha \lambda} f_{\mu \rho \alpha}\left(X_{\nu} X_{\lambda}+X_{\lambda} X_{\nu}\right) \\
& =0
\end{aligned}
$$

because $g^{\mu \nu} g^{\alpha \lambda} f_{\mu \rho \alpha}$ is antisymmetric in, and the term in round brackets is symmetric in, $\nu$ and $\lambda$. For instance, we found in example 3.14 that for $\mathfrak{s o}(3)$, the Cartan metric is $g_{\mu \nu}=\delta^{\mu}{ }_{\nu}$. We immediately get:

$$
C=X^{\mu} X_{\mu}=J_{x}^{2}+J_{y}^{2}+J_{z}^{2}=J^{2}
$$

where $\mathbf{J}$ is the angular momentum operator of quantum mechanics. Because $\mathfrak{s o}(3)$ has rank 1 , this is the only Casimir invariant in this case. Then the eigenvalues of $J^{2}$ each label an irreducible representation of $\mathfrak{s o}(3)$.

### 3.6 More on finding irreducible representations

### 3.6.1 Irreducible representations of so(3)

To find the irreducible representations of $\mathfrak{s o}(3)$, our approach is to find the eigenvalues $\lambda$ of the Casimir operator $J^{2} . \mathfrak{s o}(3)$ has one Cartan generator, $J_{z}$, whose eigenvalues $m$ are the weights of the representations. The other generators, $J_{x}$ and $J_{y}$, do not commute with $J_{z}$ and therefore are non-Cartan. We choose a Hermitian representation for the generators.

First, go to the complex extension of $\mathfrak{s o}(3)$ via the change of basis: $J_{ \pm}=\left(J_{x} \pm \mathrm{i} J_{y}\right) / \sqrt{2}, J_{0}=J_{z} . J_{x}$ and $J_{y}$ are Hermitian, but $J_{ \pm}$are not: we have $J_{+}^{\dagger}=J_{-}$, and $J_{-}^{\dagger}=J_{+} . J_{0}$ and $J_{ \pm}$commute with $J^{2}$, by definition of a Casimir operator. The commutation relations between $J_{x}, J_{y}$ and $J_{z}$ give $\left[J_{0}, J_{ \pm}\right]= \pm J_{ \pm}$, and we identify from definition 3.20 the roots of $J_{ \pm}$as $\pm 1$, with +1 the simple root. Eq. (3.27) then leads directly to: $\left[J_{+}, J_{-}\right]=J_{0}$.

Also, eq. (3.28) becomes for $\mathfrak{s o}(3)$ :

$$
J_{0}\left(J_{ \pm} f\right)=(m \pm 1)\left(J_{ \pm} f\right)
$$

and, since $J^{2}$ must commute with $J_{ \pm}$:

$$
J^{2}\left(J_{ \pm} f\right)=J_{ \pm}\left(J^{2} f\right)=\lambda\left(J_{ \pm} f\right)
$$

True to their nature as ladder operators, $J_{+}$raises, and $J_{-}$lowers, the weights $m$ by 1 , but they cannot transform $f$ into an eigenfunction of $J^{2}$ with a different eigenvalue. Consequently, all the eigenfunctions of $J_{0}$ reachable with the ladder operators in a given invariant subspace are also eigenfunctions of $J^{2}$ with the same eigenvalue $\lambda$.

Another expression will come in handy, relating $J^{2}$ to the generators. Use the definition of $J_{ \pm}$to write:

$$
J_{ \pm} J_{\mp}=\frac{1}{2}\left(J_{x}^{2}+J_{y}^{2} \mp \mathrm{i}\left[J_{x}, J_{y}\right]\right)=\frac{1}{2}\left(J_{x}^{2}+J_{y}^{2} \pm J_{0}\right)=\frac{1}{2}\left(J^{2}-J_{0}^{2} \pm J_{0}\right)
$$

so that:

$$
\begin{equation*}
J^{2}=2 J_{ \pm} J_{\mp}+J_{0}^{2} \mp J_{0} \tag{3.33}
\end{equation*}
$$

For a given value of $\lambda$, we expect that there should exist a highest weight, $m_{\max } \equiv j$, as well as a lowest weight, $m_{\min } \equiv j^{\prime}$, since the corresponding irreducible representation must be finite-dimensional.

Now act with $J_{+}$on the highest-weight eigenfunction of $J_{0}, f_{j}$. Then $J_{+} f_{j}=0$ and, from the identity (3.33), we find:

$$
J^{2} f_{j}=j^{2} f_{j}+j f_{j}=j(j+1) f_{j}=\lambda f_{j}
$$

Similarly, act with $J^{2}$ on the lowest-weight eigenfunction of $J_{0}, f_{j^{\prime}}$, keeping in mind that $f_{j^{\prime}}$ is also an eigenfunction of $J^{2}$ with the same eigenvalue as $f_{j}$ :

$$
J^{2} f_{j^{\prime}}=\left(j^{\prime}\right)^{2} f_{j^{\prime}}-j^{\prime} f_{j^{\prime}}=j^{\prime}\left(j^{\prime}-1\right) f_{j^{\prime}}=\lambda f_{j^{\prime}}
$$

Comparing yields $\lambda=j(j+1)=j^{\prime}\left(j^{\prime}-1\right)$, and thus $j^{\prime}=-j$. It follows that the weights $m$ go from $-j$ to $j$ in $N$ integer steps, ie, $j=-j+N$, so $j=N / 2$.

We conclude that:

- The eigenvalues of the Casimir operator $J^{2}$ are $j(j+1)$, where $j$ is a positive integer or a half-integer.
- For a given value of $j$, the weights $m$ can take $2 j+1$ values, from $-j$ to $j$. Therefore, odd-dimensional irreducible representations correspond to integer $j$ and even-dimensional ones to half-integer $j$.

With the help of eq. (3.33), we can now exhibit the full action of $J_{-}$on an eigenstate $f_{j m}$ of $J^{2}$ and $J_{0}$. Let $J_{-} f_{j m}=c_{-} f_{j, m-1}$. Then, if the $f_{j m}$ are normalised:

$$
\begin{aligned}
\left\langle f_{j m} \mid J_{+} J_{-} f_{j m}\right\rangle & =\int f_{j m}^{*} J_{+} J_{-} f_{j m} \mathrm{~d}^{3} x \\
& =\left\langle J_{-} f_{j m} \mid J_{-} f_{j m}\right\rangle \\
& =c_{-}^{*} c_{-}=\left|c_{-}\right|^{2}
\end{aligned}
$$

But since $2 J_{ \pm} J_{\mp}=J^{2}-J_{0}^{2} \pm J_{0}$, we also have that:

$$
\left\langle f_{j m} \mid J_{+} J_{-} f_{j m}\right\rangle=\frac{1}{2}\left\langle f_{j m} \mid\left(J^{2}-J_{0}^{2}+J_{0}\right) f_{j m}\right\rangle=\frac{1}{2}\left(j(j+1)-m^{2}+m\right)
$$

Comparing yields $c_{-}$up to an unimportant phase factor which we put equal to 1 . We find the coefficient in $J_{+} f_{j m}=c_{+} f_{j, m+1}$ in a strictly analogous way. The results for both ladder operators are:

$$
\begin{equation*}
J_{ \pm} f_{j m}=\frac{1}{\sqrt{2}} \sqrt{j(j+1)-m(m \pm 1)} f_{j, m \pm 1} \tag{3.34}
\end{equation*}
$$

Each value of $j$ labels an invariant $2 j+1$-dimensional subspace of the carrier space of $\mathfrak{s o}(3)$ of which the $2 j+1$ eigenfunctions $f_{j m}$ form a basis.

The entries of the three representation matrices $\mathcal{D}^{j}\left(J_{0}\right)=\left\langle f_{j m^{\prime}} \mid J_{0} f_{j m}\right\rangle, \mathcal{D}^{j}\left(J_{ \pm}\right)=\left\langle f_{j m^{\prime}} \mid J_{ \pm} f_{j m}\right\rangle$ are:

$$
\begin{equation*}
\mathcal{D}_{m^{\prime} m}^{j}\left(J_{0}\right)=m \delta_{m^{\prime} m} \quad \mathcal{D}_{m^{\prime} m}^{j}\left(J_{ \pm}\right)=\frac{\delta_{m^{\prime}, m \pm 1}}{\sqrt{2}} \sqrt{(j \mp m)(j \pm m+1)} \quad|m| \leq j \tag{3.35}
\end{equation*}
$$

This form for the coefficients is often quoted, but the equivalent form in eq. (3.34) is often easier to use since only the second factor in the root changes. The representation matrices for $J_{x}=\left(J_{+}+J_{-}\right) / \sqrt{2}, J_{y}=\left(J_{+}-J_{-}\right) /(\mathrm{i} \sqrt{2})$ and $J_{z}=J_{0}$ are easily recovered if needed. Keeping in mind that the rows and columns are labelled by the values of $m$ from $j$ to $-j$, we have for, say, $j=1$ :

$$
\mathcal{D}^{1}\left(J_{0}\right)=\left(\begin{array}{ccc}
1 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & -1
\end{array}\right) \quad \mathcal{D}^{1}\left(J_{+}\right)=\left(\begin{array}{ccc}
0 & 1 & 0 \\
0 & 0 & 1 \\
0 & 0 & 0
\end{array}\right) \quad \mathcal{D}^{1}\left(J_{-}\right)=\left(\begin{array}{ccc}
0 & 0 & 0 \\
1 & 0 & 0 \\
0 & 1 & 0
\end{array}\right)
$$

### 3.6.2 Representations of $\operatorname{su}(2), \operatorname{SU}(2)$, and $\operatorname{SO}(3)$

The $j=1 / 2$, defining irreducible representation of $\mathfrak{s u}(2)$ can be written in terms of the three matrices $\mathbf{s}_{0}$ and $\mathbf{s}_{ \pm}$:

$$
\mathbf{s}_{0}=\left(\begin{array}{cc}
\frac{1}{2} & 0 \\
0 & -\frac{1}{2}
\end{array}\right) \quad \mathbf{s}_{+}=\left(\begin{array}{cc}
0 & 1 \\
0 & 0
\end{array}\right) \quad \mathbf{s}_{-}=\left(\begin{array}{cc}
0 & 0 \\
1 & 0
\end{array}\right)
$$

Transform to the basis $\mathbf{s}_{x}=\left(\mathbf{s}_{+}+\mathbf{s}_{-}\right) / 2, \mathbf{s}_{y}=\left(\mathbf{s}_{+}-\mathbf{s}_{-}\right) / 2 \mathrm{i}, \mathbf{s}_{z}=\mathbf{s}_{0}$. Then the structure constants for this representation are identical to those of the $j=1 / 2$ irreducible representation of $\mathfrak{s o}(3)$. Indeed,

$$
\left[s_{i}, s_{j}\right]=\mathrm{i} \epsilon_{i j}^{k} s_{k}
$$

The s matrices are Hermitian; they are also traceless, as expected for $\mathfrak{s u}(2)$. As before, a finite $S U(2)$ transformation can be reconstructed with the exponential map, corresponding to a rotation parametrised by $\boldsymbol{\theta}=\theta \hat{\mathbf{n}}$ : $S(\boldsymbol{\theta})=\mathrm{e}^{\mathrm{i} \theta \hat{\mathbf{n}} \cdot \mathbf{s}}$, where the direction of $\hat{\mathbf{n}}$ is the axis of rotation.

But the isomorphism between $\mathfrak{s u}(2)$ and $\mathfrak{s o}(3)$ does not translate into an isomorphism between $S U(2)$ and $S O(3)$ ! Indeed, whereas a $S O(3)$ rotation by $2 \pi$ is identical to the identity, a $S U(2)$ rotation by $2 \pi$ is equivalent to minus the identity, because of the factor $1 / 2$ lurking in the $\mathbf{s}$ matrices.

There is a $2 \rightarrow 1$ homomorphism that maps $S U(2)$ to $S O(3): \pm S(\boldsymbol{\theta}) \rightarrow R(\boldsymbol{\theta})$, and because of this $S U(2)$ can be represented by $S O(3)$ matrices. But the map is not uniquely invertible, and therefore $S U(2)$ matrices are not stricto sensu representations of $S O(3)$. Only the ones that correspond to integer $S U(2) j$ are; those with half-integer $S U(2) j$ are called spinor representations, and we say that integer and half-integer representations of $S U(2)$ together form projective representations $R_{\mathrm{g}}$ of $S O(3)$, in the sense that $R_{\mathrm{g}_{1}} \cdot R_{\mathrm{g}_{2}}=\alpha_{\mathrm{g}_{1}, \mathrm{~g}_{2}} R_{\mathrm{g}_{1} \mathrm{~g}_{2}}$, with $\alpha \in \mathbb{C}$

Wigner matrices $\mathcal{D}_{\theta}^{j}=\mathrm{e}^{\mathrm{i} \theta \hat{\mathbf{n}}^{j} \cdot \mathbf{s}_{j}}$ (with $\mathbf{s}_{j}$ the triplet of $\mathfrak{s u}(2)$ basis generators of the irreducible representation labelled by $j$ ), is the name given to the irreducible representations of $S U(2)$, and the matrix elements are called Wigner functions. They can be rather complicated, except when $\hat{\mathbf{n}}=\hat{\mathbf{z}}$ and $s_{z}=s_{0}$ is diagonal, in which case $\left(\mathcal{D}_{\theta}^{j}\right)^{m^{\prime}}{ }_{m}=\mathrm{e}^{\mathrm{i} m \theta} \delta^{m^{\prime}}{ }_{m}(|m| \leq j)$. Fortunately, they are tabulated in many places for small values of $j$ and are easily calculated by computer.

### 3.6.3 Tensor product representations

Definition 3.23. Let $f_{j_{1} m_{1}}$ and $f_{j_{2} m_{2}}$ be two basis functions associated, respectively, with irreducible representations $\mathcal{D}_{\mathrm{g}}^{j_{1}}$ and $\mathcal{D}_{\mathrm{g}}^{j_{2}}$ of $\mathrm{g} \in S U(2)$ or $S O(3)$, such that:

$$
S_{\mathrm{g}} f_{j_{1} m_{1}}=f_{j_{1} m_{1}^{\prime}}\left(\mathcal{D}_{\mathrm{g}}^{j_{1}}\right)^{m_{1}^{\prime}}{ }_{m_{1}}, \quad S_{\mathrm{g}} f_{j_{2} m_{2}}=f_{j_{2} m_{2}^{\prime}}\left(\mathcal{D}_{\mathrm{g}}^{j_{2}}\right)^{m_{2}^{\prime}}{ }_{m_{2}}
$$

where $S_{\mathrm{g}}$ is the transformation associated with g . Then we form the tensor product representation $\mathcal{D}_{\mathrm{g}}^{j_{1}} \otimes \mathcal{D}_{\mathrm{g}}^{j_{2}}:$

$$
\begin{equation*}
S_{\mathrm{g}} f_{j_{1} m_{1}} f_{j_{2} m_{2}}=f_{j_{1} m_{1}^{\prime}} f_{j_{2} m_{2}^{\prime}}\left(\mathcal{D}^{j_{1}}\right)^{m_{1}^{\prime}}{ }_{m_{1}}\left(\mathcal{D}^{j_{2}}\right)^{m_{2}^{\prime}}{ }_{m_{2}} \tag{3.36}
\end{equation*}
$$

In Dirac notation, the product of the basis functions would read: $\left|j_{1} m_{1}, j_{2} m_{2}\right\rangle=\left|j_{1} m_{1}\right\rangle\left|j_{2} m_{2}\right\rangle$.

Such a product is needed when a system responds to transformations in more than one way, either because of the coupling of two separate systems (eg. particles) or because two distinct dynamical variables of one system get coupled. A common transformation on the whole system is to be written as a direct product of transformations on each of its parts in its own subspace.

Linearise eq. (3.36) using the generic expansion $\mathcal{D}=\mathbf{I}+a^{i} \mathbf{X}_{i}$, where $\mathbf{X}$ stands for a generator of $S U(2)$ or $S O(3)$ in that representation. We find that the generators of the composite representation are the sums of the generators of the distinct terms in the tensor product, so that:

$$
\begin{equation*}
\mathbf{X}^{(1 \otimes 2)}\left(f_{j_{1} m_{1}} f_{j_{2} m_{2}}\right)=\left(X^{(1)} f_{j_{1} m_{1}}\right) f_{j_{2} m_{2}}+f_{j_{1} m_{1}}\left(X^{(2)} f_{j_{2} m_{2}}\right) \tag{3.37}
\end{equation*}
$$

that is: $\mathbf{X}^{(1 \otimes 2)}=X^{(1)} \otimes \mathbf{I}+\mathbf{I} \otimes X^{(2)}$ or, more sloppily, $\mathbf{X}=\mathbf{X}^{(1)}+\mathbf{X}^{(2)}$. When the generators have diagonal representations, as happens with $J_{0}(S O(3))$ or $s_{0}(S U(2))$, we find, eg.:

$$
J_{0}\left(f_{j_{1} m_{1}} f_{j_{2} m_{2}}\right)=\left(m_{1}+m_{2}\right) f_{j_{1} m_{1}} f_{j_{2} m_{2}}
$$

Note that $\left[X^{(1)}, X^{(2)}\right]=0$, because they act on distinct subspaces.
As before, we expect the product representation to be reducible, ie. there should exist linear combinations $\phi_{j m}$ (or $|j m\rangle$ ) of the product basis functions $f_{j_{1} m_{1}} f_{j_{1} m_{2}}$ which transform among themselves. In other words, we are looking for invariant subspaces of the Hilbert product space. Those linear combinations take the form of the invertible transformation:

$$
\begin{equation*}
\phi_{j m}=\sum_{m_{1}, m_{2}}\left(j_{1} m_{1}, j_{2}, m_{2} \mid j m\right) f_{j_{1} m_{1}} f_{j_{2} m_{2}} \tag{3.38}
\end{equation*}
$$

where $m=m_{1}+m_{2}$, and $\left|j_{1}-j_{2}\right| \leq j \leq j_{1}+j_{2}$. The real coefficients $\left(j_{1} m_{1}, j_{2}, m_{2} \mid j m\right)$ are known as Clebsch-Gordan or Wigner coefficients. They are unique up to a phase convention and can be calculated or looked up in tables.

One easy way to obtain the $\phi_{j m}$ in terms of the $f_{j_{1} m_{1}} f_{j_{2} m_{2}}$ is to start with the highest weight component, $m=j_{1}+j_{2}$, of the highest $j$ irreducible representation: $j=j_{1}+j_{2}$. Of course, $\phi_{j_{1}+j_{2}, j_{1}+j_{2}}=f_{j_{1} j_{1}} f_{j_{2} j_{2}}$. Next, apply $J_{-}$on the left and on the right, using eq. (3.37), until the lowest weight component of the $j$ irreducible representation, $\phi_{j,-j}$, is reached. Now obtain the linear combination for the highest weight of the $j-1$ representation, $\phi_{j-1, j-1}$, by demanding that it be orthogonal to $\phi_{j, j-1}$, and repeat with $J_{-}$. Continue until all values of $j$ allowed by $\left|j_{1}-j_{2}\right| \leq j \leq j_{1}+j_{2}$ have been reached.

### 3.6.4 Irreducible (Spherical) tensors

Suppose that a set of functions $f_{j m}$ in the carrier space of $S U(2)$ or $S O(3)$ transforms under a group element parametrised by $\boldsymbol{\theta}=\theta \mathbf{n}$ as: $R_{\boldsymbol{\theta}} f_{j m}=f_{j m^{\prime}}\left(\mathcal{D}_{\boldsymbol{\theta}}^{j}\right)^{m^{\prime}}{ }_{m}$. Then the set $\left\{f_{j m}\right\}$ form a basis for an irreducible representation of $S U(2)$ or $S O(3)$ labelled by $j$.

Definition 3.24. The components $T_{j m}$ of an irreducible spherical tensor $\mathbf{T}_{j}$ of rank $j$ are operators that transform as:

$$
\begin{equation*}
R_{\boldsymbol{\theta}} T_{j m} R_{\boldsymbol{\theta}}^{-1}=T_{j m^{\prime}}\left(\mathcal{D}_{\boldsymbol{\theta}}^{j}\right)_{m}^{m^{\prime}} \tag{3.39}
\end{equation*}
$$

If we linearise this equation, we obtain (EXERCISE) an alternative definition of spherical tensors in terms of generators $\mathbf{J}^{(j)}$ of an irreducible representations of the algebra:

$$
\begin{equation*}
\left[J^{(j)}, T_{j m}\right]=T_{j m^{\prime}}\left(J^{(j)}\right)_{m}^{m^{\prime}} \quad \text { (no summation on } j \text { ) } \tag{3.40}
\end{equation*}
$$

where $j$ is the label of the irreducible representation.
For $\mathrm{SU}(2)$ or $\mathrm{SO}(3)$ :

$$
\begin{equation*}
\left[J_{0}^{(j)}, T_{j m}\right]=m T_{j m}, \quad\left[J_{ \pm}^{(j)}, T_{j m}\right]=\sqrt{(j \mp m)(j \pm m+1)} T_{j, m \pm 1} \tag{3.41}
\end{equation*}
$$

As a direct consequence of these commutation relations, the matrix element of $T_{j m},\left\langle j_{2} m_{2}\right| T_{j m}\left|j_{1} m_{1}\right\rangle$, vanishes unless $m_{2}=m_{1}+m$ and $\left|j_{1}-j\right| \leq j_{2} \leq j_{1}+j$. These are the famous vector addition rules as applied to spherical tensors.

### 3.6.5 The Wigner-Eckart theorem

The Wigner-Eckart theorem says that if $\mathbf{T}_{j}$ is a spherical tensor under $S U(2)$, then its matrix elements, written in bra-ket notation, $\left\langle j_{2} m_{2}\right| T_{j m}\left|j_{1} m_{1}\right\rangle$, can be factored as:

$$
\begin{equation*}
\left\langle j_{2} m_{2}\right| T_{j m}\left|j_{1} m_{1}\right\rangle=\frac{\left(j_{1} m_{1}, j m \mid j_{2} m_{2}\right)}{\sqrt{2 j_{2}+1}}\left\langle j_{2}\left\|\mathbf{T}_{j}\right\| j_{1}\right\rangle \tag{3.42}
\end{equation*}
$$

where $\left\langle j_{2}\left\|\mathbf{T}_{j}\right\| j_{1}\right\rangle$ is called the reduced matrix element and does not depend on $m, m_{1}$ or $m_{2}$. So the dependence of the matrix element on these numbers is carried entirely by the Clebsch-Gordan coefficient!

The Wigner-Eckart theorem applies to unitary representations of Lie groups, not only to those of $S U(2)$. The Clebsch-Gordan coefficients and the labelling with eigenvalues of Casimir operators will be appropriate to the Lie group.

As a result, ratios of matrix elements for a given $j$ but different values of $m$ are just ratios of Clebsch-Gordan coefficients.

Example 3.18. Suppose that $\mathbf{T}$ transforms as a scalar under some Lie group. Then the relevant representation matrix of the group is just the identity matrix. If the Lie group is $S U(2), j=m=0$, and the vector-addition rules collapse the Wigner-Eckart theorem to:

$$
\left\langle j_{2} m_{2}\right| \mathbf{T}\left|j_{1} m_{1}\right\rangle=\frac{\left\langle j_{2}\right|\left|\mathbf{T} \| j_{1}\right\rangle}{\sqrt{2 j_{2}+1}} \delta^{j_{1}}{ }_{j_{2}} \delta^{m_{1}}{ }_{m_{2}}
$$

From this we see that matrix elements of scalar operators between weights associated with different irreducible representations of a group vanish.

Essentially, the importance of the Wigner-Eckart theorem resides in its separating symmetry-related ("geometrical") aspects of matrix elements from other ("dynamical") aspects which may be difficult to calculate and which the theorem shoves into the reduced matrix element.

### 3.6.6 Decomposing product representations

The problem of decomposing representations of a semisimple group into their irreducible representations can often be treated in a fairly intuitive way. Consider $S O(3)$ again, and its 3 -dim carrier space of functions $f(\mathbf{x})$ and $g(\mathbf{y})$ (eg. in quantum mechanics, the wave-functions of two particles), each of which transforms in some known way under 3-dim rotations. We can form tensor products, $f(\mathbf{x}) \otimes g(\mathbf{y})$, of such functions, whose transformation properties are derived from those of the functions.

For instance, if our functions were 3 -dim vectors, we would have a 9 -dim product representation (ie. one with nine weights, or basis vectors for its carrier space), with components $T^{i j}$, which under rotations $R$ would tranform as:

$$
\begin{equation*}
T^{\prime i j}=R^{i}{ }_{k} R^{j}{ }_{l} T^{k l} \tag{3.43}
\end{equation*}
$$

We know that the $T^{i j}$ can be decomposed into a symmetric and an antisymmetric part, each of which transforms separately under rotations, in the sense that the 6 -dim symmetric part rotates into a symmetric object, and the 3 -dim antisymmetric part into an antisymmetric one. Thus, we have easily found invariant subspaces. But we can go even further. The trace of $T^{i j}, T_{i}^{i}$, is invariant under rotations, forming a 1-dim invariant subspace which should be separated out from the symmetric part.

Note that the trace is obtained by contracting $T^{i j}$ with the metric of the carrier space, with components $g_{i j}$, which here is just the identity matrix invariant under rotations. Similarly, the antisymmetric part can be obtained with the Levi-Civita symbol which is also invariant under rotation. Thus, we can write:

$$
\begin{equation*}
T^{i j}=\frac{1}{2}\left(T^{i j}+T^{j i}\right)+\frac{1}{2} \epsilon^{i j k} \epsilon_{k l m} T^{l m}=\frac{1}{2}\left(T^{i j}+T^{j i}-\frac{2}{3} g^{i j} T_{k}{ }_{k}\right)+\frac{1}{2}\left(T^{i j}-T^{j i}\right)+\frac{1}{3} g^{i j} T^{k}{ }_{k} \tag{3.44}
\end{equation*}
$$

The numerical coefficient of the trace term has been chosen so as to make the symmetric term traceless.
But we can also think of eq. (3.43) as a $3 \otimes 3$ exterior direct product a rotation with itself, so a $9 \times 9$ matrix, with each row labelled by a pair $\{i j\}$ and each column labelled by a pair $\{k l\}$, acting on a $9 \times 1$ matrix with entries $T^{k l}$ labelled by the pairs $\{k l\}$. The direct-product matrix is a representation of $S O(3)$. Indeed, under a rotation $R_{1}$ followed by $R_{2}, T^{i j} \longrightarrow\left(R_{2} R_{1}\right)^{i}{ }_{m}\left(R_{2} R_{1}\right)^{j}{ }_{n} T^{m n}$, where now the $9 \times 9$ matrix is formed from the matrix product $R_{2} R_{1}$. The representation is reducible, that is, it can be transformed via an angle-independent similarity matrix to a block-diagonal matrix with a symmetric traceless $6 \times 6$ block (which acts only on the symmetric traceless part of $\mathbf{T}$ ), an amtisymmetric $3 \times 3$ block acting only on the antisymmetric part of $\mathbf{T}$, and a 1 acting only on the trace of $\mathbf{T}$.

We obtain the following decomposition into irreducible representations: $\mathbf{9}=\mathbf{5} \oplus \mathbf{3} \oplus \mathbf{1}$
As expected, the total dimensions on the left and right match. The result is also consistent with what we would find by decomnposing a $j_{1} \otimes j_{2}=1 \otimes 1 S O(3)$ product representation with the method of section 3.6.3 to obtain a direct sum of three irreducible representations labelled by $j=2$ (of dimension 5 ), $j=1$ (of dimension 3 ), and $j=0($ of dimension 1 ).

## 4 MODULE IV - Elementary Theory of Analytic Functions

### 4.1 Complex Numbers

We construct an extended algebra over the real numbers with the following structure: $z=x+\epsilon y$, where $x$ (the real part) and $y$ (the imaginary part) are real numbers, $\epsilon$ is an independent, non-real quantity endowed with a rule that characterises the algebra. One can think of it as a device that allows us to extend arithmetical operations from the real numbers with essentially no change.

Addition and multiplication are straightforward:

$$
\begin{aligned}
& z_{1}+z_{2}=x_{1}+\epsilon y_{1}+x_{2}+\epsilon y_{2}=x_{1}+x_{2}+\epsilon\left(y_{1}+y_{2}\right) \\
& z_{1} z_{2}=\left(x_{1}+\epsilon y_{1}\right)\left(x_{2}+\epsilon y_{2}\right)=x_{1} x_{2}+\epsilon^{2} y_{1} y_{2}+\epsilon\left(x_{1} y_{2}+y_{1} x_{2}\right)
\end{aligned}
$$

If the extended structure is to be at least a ring of numbers, it should close under multiplication. This is achieved if $\epsilon^{2}$ itself is of the form $a+\epsilon b$, with $a, b \in \mathbb{R}$. It is usual to choose $b=0$ and $a= \pm 1$.

To any extended number $z$ one associates its conjugate $z^{*} \equiv \bar{z}=x-\epsilon y . z=z^{*}$ if and only if $z$ is real. Since $\epsilon^{2}$ is real, so is $|z|^{2}=z z^{*}=x^{2}-\epsilon^{2} y^{2}$. Also, to perform a division, one must first make the denominator real: $z_{1} / z_{2}=z_{1} z_{2}^{*} /\left|z_{2}\right|^{2}$. This of course requires that $\left|z_{2}\right| \neq 0$. Multiplication is both commutative and associative.

One extended algebra, that of the hyperbolic-complex numbers, is obtained by choosing $\epsilon^{2}=+1$. Another algebra, that of the complex numbers, is characterised by $\epsilon \equiv \mathrm{i}$, where $\mathrm{i}^{2}=-1$.

When multiplying or dividing such numbers it is usually more convenient to use the polar representation:

$$
z=r \mathrm{e}^{\epsilon \theta}
$$

To find $r \equiv|z|$ (the modulus, or absolute value) and $\theta$ (the phase, or argument), we invoke the beautiful generalised Euler relation:

$$
\mathrm{e}^{\epsilon \theta}= \begin{cases}\cosh \theta+\epsilon \sinh \theta & \epsilon^{2}=+1 \\ \cos \theta+\epsilon \sin \theta & \epsilon^{2}=-1\end{cases}
$$

which can be proved by expanding the exponential in a Taylor series.
If we write a spacetime event at position $x$ and time $t$ as $z=c t+\epsilon x$, hyperbolic-complex numbers provide ${ }^{\dagger}$ a natural language for special relativity. For instance, consider the "hyperbolic rotation": $z^{\prime}=z \mathrm{e}^{-\epsilon \phi}$, with $\tanh \phi=\beta \equiv v / c$. Then $\cosh \phi=1 / \sqrt{1-\beta^{2}} \equiv \gamma$, and $\sinh \phi=\gamma \beta$. Expanding with Euler's relation yields:

$$
z^{\prime}=c t^{\prime}+\epsilon x^{\prime}=(\cosh \phi-\epsilon \sinh \phi)(c t+\epsilon x)=\gamma(c t-\beta x)+\epsilon[\gamma(x-\beta c t)]
$$

which we recognise as Lorentz transformations. From now on, however, we focus on complex numbers.
The connection between Cartesian and polar representations of complex numbers is provided by:

$$
z=r \cos \theta+\mathrm{i} r \sin \theta=x+\mathrm{i} y \quad \begin{cases}x=r \cos \theta & r=\sqrt{x^{2}+y^{2}} \\ y=r \sin \theta & \theta=\tan ^{-1}(y / x)\end{cases}
$$

We recognise these expressions as the transformations between Cartesian and polar coordinates for vectors in two-dimensional Euclidean space. Thus, complex numbers can be mapped to points in a complex plane with coordinates $(x, y)$ or $(r, \theta)$. The last expression containing the tangent cannot distinguish between angles in the first and third quadrant and between the second and fourth quadrant of the complex plane. Which angle to choose must be decided from the signs of $x$ and $y$.

The field $\mathbb{C}$ cannot be ordered. Even saying that $\mathrm{i}>0$ or $\mathrm{i}<0$ leads to $-1>0$, as can be seen from $\mathrm{i}^{2}>0$ and $0=-\mathrm{i}(\mathrm{i}-\mathrm{i})<-\mathrm{i}(-\mathrm{i})=-1$, using the muliplication rules for a field.

Two-dim rotations and complex numbers are representations of one another. To perform a rotation by angle $\theta$ in the complex plane, multiply by $\mathrm{e}^{\mathrm{i} \theta}$.

[^15]
### 4.2 Holomorphic Functions of a Complex Variable (BF 6.1)

### 4.2.1 Differentiability of complex functions

Introduce a one-to-one mapping $f: \mathbb{C} \longrightarrow \mathbb{C}$ from one complex plane to another $w=u+\mathrm{i} v$ : Of course, we are interested in mappings, or functions, that are continuous and differentiable. But $\mathrm{d}_{z} f(z)$ is not quite like the gradient of a two-dimensional vector which is function of $x$ and $y$. With the gradient, we ask how the vector changes under $x+\mathrm{d} x$ or $y+\mathrm{d} y$, keeping the other variable constant; in the case of $f(z)=f\left(r \mathrm{e}^{\mathrm{i} \theta}\right)$, we think that the derivative $\mathrm{d}_{z} f(z)$ is well-defined only if it is invariant under $\theta+\mathrm{d} \theta$. In other words, it should not matter in which direction we vary $z$ : the derivative should be isotropic. More formally:

$$
\begin{equation*}
\mathrm{d}_{z} f=\lim _{\Delta z \rightarrow 0} \frac{f(z+\Delta z)-f(z)}{\Delta z} \tag{4.1}
\end{equation*}
$$

where the limit exists, and therefore must be independent of the phase of $\Delta z$.
It is difficult to overemphasise how strong a constraint this is on functions of a complex variable! It imposes stronger restrictions than any that appply to functions of a real variable. All the properties and applications we are going to review follow from it.

### 4.2.2 Cauchy-Riemann Relations

We need a more workable criterion than eq. (4.1) for complex differentiability. This can be readily derived by writing $f(z)$ in the often useful representation:

$$
\begin{equation*}
f(z)=u(x, y)+\mathrm{i} v(x, y) \tag{4.2}
\end{equation*}
$$

where $u$ and $v$ are real functions of the real and imaginary parts of $z, x$ and $y$. Then $f$ maps regions in the $z$ plane to regions in the $w=u+\mathrm{i} v$ plane.

If we write $\mathrm{d}_{z} f(z)=p(x, y)+\mathrm{i} q(x, y)$, then:

$$
\delta f=\mathrm{d}_{z} f(z) \delta z=[p(x, y)+\mathrm{i} q(x, y)](\delta x+\mathrm{i} \delta y)=\delta u+\mathrm{i} \delta v
$$

Equating real and imaginary parts gives:

$$
\begin{aligned}
\delta u & =p \delta x-q \delta y=\partial_{x} u \delta x+\partial_{y} u \delta y \\
\delta v & =q \delta x+p \delta y=\partial_{x} v \delta x+\partial_{y} v \delta y
\end{aligned}
$$

By omparing these equations, we find that complex differentiability yields a system of coupled differential equations for $u$ and $v$, the Cauchy-Riemann equations:

$$
\begin{equation*}
\partial_{x} u(x, y)=\partial_{y} v(x, y), \quad \partial_{y} u(x, y)=-\partial_{x} v(x, y) \tag{4.3}
\end{equation*}
$$

We see that the real and imaginary parts of a complex-differentiable function are not independent: if one is known, the other can be calculated up to a constant. For instance, given $u(x, y), \partial_{y} v(x, y)$ is calculable from one CauchyRiemann condition, and if it can be integrated with respect to $y, v$ is known up to some function $g(x)$. The other Cauchy-Riemann equation provides the derivative of $g$ which can then be retrieved up to a constant.

Definition 4.1. Let the function $f(z)$ satisfy the Cauchy-Riemann equations (let it be complexdifferentiable) at a point $z_{0}$; in addition, let it be complex-differentiable over a finite, open neighbourhood of $z_{0}$, (or, alternatively, continuous over that neighbourhood). Then $f(z)$ is said to be holomorphic, or regular, at $z_{0}$.
A function $f$ is said to be analytic at $z_{0}$ if there exists a series $\sum a_{n}\left(z-z_{0}\right)^{n}(n \geq 0)$, and also a $r>0$ such that the series converges to $f(z)$ for $\left|z-z_{0}\right|<r$. We see that an analytic function is holomorphic; the reverse will be shown to hold (but not for functions of a real variable!) later.
Points at which a function is not holomorphic are called singularities (or singular points); if a function is nowhere singular in the complex plane, we say that it is an entire function.

By differentiating one Cauchy-Riemann relation with respect to $x$ and the other with respect to $y$ (and viceversa), we show that the real and imaginary parts of a holomorphic function must satisfy the Laplace equation: $\nabla^{2} u=\nabla^{2} v=0$, and are thus both harmonic. This intimate link between the Laplace equation and holomorphic functions is crucial. Because, as we shall see in a later chapter, boundary conditions uniquely determine the solutions of the Laplace equation, we expect functions holomorphic over a region in the complex plane to be determined by their values on the boundary of the region. This, as we will soon discover, provides the justification for the important process of analytic continuation of a function.

In general, a function $f=u(x, y)+\mathrm{i} v(x, y)$ might be expected to depend on both $z$ and $z^{*}$ since $x=\left(z+z^{*}\right) / 2$ and $y=\left(z-z^{*}\right) / 2$ i. If so, then:

$$
\begin{gather*}
\partial_{z} f=\partial_{x} f \partial_{z} x+\partial_{y} f \partial_{z} y=\frac{1}{2}\left(\partial_{x} u+\partial_{y} v\right)+\frac{\mathrm{i}}{2}\left(\partial_{x} v-\partial_{y} u\right)  \tag{4.4}\\
\partial_{z^{*}} f=\partial_{x} f \partial_{z^{*}} x+\partial_{y} f \partial_{z^{*}} y=\frac{1}{2}\left(\partial_{x} u-\partial_{y} v\right)+\frac{\mathrm{i}}{2}\left(\partial_{x} v+\partial_{y} u\right)
\end{gather*}
$$

The Cauchy-Riemann conditions are seen to be equivalent to imposing $\partial_{z^{*}} f=0$, so that $f$ is a function of a complex variable, not just a complex function of two real variables. The message is clear: a holomorphic function of a complex variable cannot depend on the complex conjugate of that variable.

Holomorphic functions can always be constructed out of harmonic functions $u$ of two real variables. When this is done, one should be able to write the functions $f=u+\mathrm{i} v$ in terms of $z$ only, with no $z^{*}$ dependence.

From eq. (4.4), it should also be clear that $\mathrm{d}_{z} f$ can be written in terms of derivatives of either $u$ or $v$. For instance, $\partial_{z} f=\partial_{x} u-\mathrm{i} \partial_{y} u$, and the complex derivative of $f$ can be calculated with the sole knowledge of the real part of $f$ !

### 4.2.3 Single- and Multi-valued Functions (BF 6.2)

It is not hard to come up with holomorphic functions. We have already seen how they could be generated from harmonic functions of two real variables. Useful examples are the exponential function $\mathrm{e}^{z}$ and the trigonometric functions $(\sin z, \cos z, \tan z)$, which share with their real cousins all the usual relations and properties. They are obtained by a process called analytic continuation from the real axis into the complex plane. Apart from $\tan z$, which is singular at $z=(n+1 / 2) \pi$, they are entire functions.

Simply by replacing $z=x+\mathrm{i} y$, one readily obtains the $u+\mathrm{i} v$ forms:

$$
\begin{align*}
\mathrm{e}^{z} & =\mathrm{e}^{x}(\cos y+\mathrm{i} \sin y) \\
\cos z & =\cos x \cosh y-\mathrm{i} \sin x \sinh y  \tag{4.5}\\
\sin z & =\sin x \cosh y+\mathrm{i} \cos x \sinh y
\end{align*}
$$

Unlike $\cos x$ and $\sin x$, however, neither $|\cos z|$ nor $|\sin z|$ is bounded:

$$
|\cos z|^{2}=\cos ^{2} x+\sinh ^{2} y, \quad|\sin z|^{2}=\sin ^{2} x+\sinh ^{2} y
$$

Another, and perhaps the most important, example is the power function $z^{n}$, with $n \in \mathbb{Z}^{+}$. Indeed, most common functions can be written as power series, which means that they can be approximated by polynomials. The closed unit disk centered on the origin of the complex plane is defined by $|z| \leq 1$. Then it should be clear that the function $z^{n}$ maps the disk into itself.

Roots and logarithms require a bit more care. Let us look at roots first. Here we must use the polar representation of $z, r \mathrm{e}^{\mathrm{i} \theta}$, and we have:

$$
z^{1 / n}=r^{1 / n} \mathrm{e}^{\mathrm{i}(\theta+2 \pi k) / n} \quad k=0,1, \ldots, n-1
$$

While $r^{1 / n}$ is unique, the whole function is multi-valued: if we start anywhere on the positive real axis $(\theta=0)$, every time we go around by $2 \pi$, we arrive at a different value; only after $n$ loops do we recover the value $z(\theta=$
$0)=r^{1 / n}$. This means that, starting from the singular point at the origin (where the root has no finite derivative), the function is not continuous, and therefore not holomorphic anywhere on the positive real axis!

Whenever there exists a point $z_{0}$ around which we must circle by an angle larger than $2 \pi$ in order to come back to the same value $f\left(z_{1}\right)$ that we had at $z_{1}$ when we started, we say that $z_{0}$ is a branch point. We can recover single-valuedness of a function by using the Riemann construction: we imagine that along the line of singularities starting at the branch point, the complex plane defined by $0 \leq \theta \leq 2 \pi$ is joined smoothly to another complex plane $2 \pi \leq \theta \leq 4 \pi$, and so on until the last plane $2(n-1) \pi \leq \theta \leq 2 n \pi$ is joined smoothly to the first complex plane. The different planes are known as Riemann sheets on each of which a different branch of the function is defined, with the line on which the sheets are joined a branch cut.

More simply, we can think of the root as a bunch of $n$ single-valued functions, each defined on one sheet $2(k-1) \pi \leq \theta \leq 2 k \pi$, with $1 \leq k \leq n$.

Another well-known example of a multi-valued function is $\ln z$ :

$$
\ln z=\ln r+\mathrm{i}(\theta+2 \pi n)
$$

where any integer value of $n$ corresponds to the same point $z$. Thus, for a given $z$ there are infinitely many values of the $\log$, each separated from its neighbours by $\pm 2 \pi$ i. This time, each Riemann sheet $2(k-1) \pi \leq \theta \leq 2 k \pi$ may be joined on the real axis to make $\ln z$ holomorphic everywhere except at $z=0$, where it goes to $-\infty$. Or we can think of each sheet as providing one branch which is non-holomorphic on the real axis. In the range $0 \leq \theta \leq 2 \pi, \ln z$ is often called the principal branch, or principal value, and denoted by $\log z$ in tables of integrals or mathematical handbooks.

Useful representations of inverse trigonometric functions in terms of logarithms can be obtained, for instance, to discover the principal branch of those functions:

$$
z=\tan w=\frac{1}{\mathrm{i}} \frac{\mathrm{e}^{\mathrm{i} w}-\mathrm{e}^{-\mathrm{i} w}}{\mathrm{e}^{\mathrm{i} w}+\mathrm{e}^{-\mathrm{i} w}}
$$

is easily inverted to yield:

$$
w=\tan ^{-1} z=\frac{\mathrm{i}}{2} \ln \left(\frac{1-\mathrm{i} z}{1+\mathrm{i} z}\right)
$$

with principal branch $\mathrm{i}[\log (1-\mathrm{i} z)-\log (1+\mathrm{i} z)] / 2$. Similar expressions can be found (EXERCISE) for $\cos ^{-1} z$ and $\sin ^{-1} z$.

### 4.2.4 Conformal transformations

When $f=u+\mathrm{i} v$ is holomorphic, the functions $u(x, y)$ and $v(x, y)$ define a transformation whose Jacobian is:

$$
\frac{\partial(u, v)}{\partial(x, y)}=\partial_{x} u \partial_{y} v-\partial_{y} u \partial_{x} v=\left(\partial_{x} u\right)^{2}+\left(\partial_{y} u\right)^{2}=\left|\partial_{x} u+\mathrm{i} \partial_{y} u\right|^{2}=\left|\mathrm{d}_{z} f\right|^{2}
$$

because of the Cauchy-Riemann conditions and eq. (4.4). The transformation is one-to-one invertible at points where the Jacobian, and therefore $\mathrm{d}_{z} f$, does not vanish. Points where this is not true are said to be critical.

Example 4.1. Consider the transformation $w=\ln z$, with $z=r \mathrm{e}^{\mathrm{i} \phi}$. Then $u=\ln r$ and $v=\phi$. Focus on the principal branch $0 \leq \phi \leq 2 \pi$. Circles of radius $r$ centered on the origin are mapped into vertical lines $u=\ln r$. Annular regions $r_{1} \leq r \leq r_{2}$ are mapped into vertical strips $\ln r_{1} \leq u \leq \ln r_{2}$ and height $0 \leq v \leq 2 \pi$. Finally, rays centered on the origin $\phi=\phi_{0}$ are mapped into horizontal lines $v=\phi_{0}$.

Holomorphic functions considered as transformations turn out to be the two-dimensional version of a very important class of transformations!

In order to find what such transformations are, consider first a $n$-dimensional space with line element $\mathrm{d} s^{2}=$ $g_{\mu \nu}(x) \mathrm{d} x^{\mu} \mathrm{d} x^{\nu}$, where $g_{\mu \nu}$ are the covariant components of the metric tensor for this space (see module I for more details). Now we know that $\mathrm{d} x^{\mu}=\partial_{\alpha^{\prime}} x^{\mu} \mathrm{d} x^{\prime \alpha}$. Since $\mathrm{d} s^{2}$ is independent of coordinates, we obtain the usual general law for the transformation of covariant tensor components:

$$
\mathrm{d} s^{2}=g_{\alpha \beta}^{\prime}\left(x^{\prime}\right) \mathrm{d} x^{\prime \alpha} \mathrm{d} x^{\prime \beta}=g_{\mu \nu}(x) \partial_{\alpha^{\prime}} x^{\mu} \partial_{\beta^{\prime}} x^{\nu} \mathrm{d} x^{\prime \alpha} \mathrm{d} x^{\prime \beta}
$$

Comparing without imposing restrictions on the $\mathrm{d} x$ gives the usual transformations of tensor components (implicit summation over repeated indices):

$$
\begin{equation*}
g_{\alpha \beta}^{\prime}\left(x^{\prime}\right)=g_{\mu \nu}(x) \partial_{\alpha^{\prime}} x^{\mu} \partial_{\beta^{\prime}} x^{\nu} \quad \text { or } \quad g^{\prime \mu \nu}\left(x^{\prime}\right)=g^{\alpha \beta}(x) \partial_{\alpha} x^{\prime \mu} \partial_{\beta} x^{\prime \nu} \tag{4.6}
\end{equation*}
$$

Definition 4.2. Let $\Omega(x)$ be a positive function. Then the product $g^{\mu \nu}(x)=\Omega(x) g^{\mu \nu}(x)$ is a conformal transformation. It can also be interpreted as mapping a metric tensor in $x$ coordinates to a metric tensor $g^{\mu \nu}\left(x^{\prime}\right)$ in $x^{\prime}$ coordinates.

A conformal transformation preserves the magnitude and sense of the angle between two vectors $\mathbf{A}$ and B, eg., the tangent vectors of two curves intersecting at some point.

Specialise to a two-dimensional Euclidean space with $g^{\mu \nu}=\delta_{\mu \nu}$. Call $x^{1}=x, x^{2}=y$, and the transformation functions: $x^{1}=u(x, y)$ and $x^{\prime 2}=v(x, y)$. Using: $\delta^{\alpha \beta}(x) \partial_{\alpha} x^{\mu} \partial_{\beta} x^{\nu}=\Omega \delta^{\mu \nu}$, we quickly find (EXERCISE):

$$
\begin{aligned}
\left(\partial_{x} u\right)^{2}+\left(\partial_{y} u\right)^{2} & =\left(\partial_{x} v\right)^{2}+\left(\partial_{y} v\right)^{2} & & \text { from the } 11 \text { and } 22 \text { components } \\
\partial_{x} u \partial_{x} v & =-\partial_{y} u \partial_{y} v & & \text { from the } 12 \text { component }
\end{aligned}
$$

The first line can be rearranged as:

$$
\left(\partial_{x} u\right)^{2}\left[1+\left(\frac{\partial_{y} u}{\partial_{x} u}\right)^{2}\right]=\left(\partial_{y} v\right)^{2}\left[1+\left(\frac{\partial_{x} v}{\partial_{y} v}\right)^{2}\right]
$$

Inserting the second line yields:

$$
\begin{equation*}
\partial_{x} u= \pm \partial_{y} v \quad \partial_{x} v=\mp \partial_{y} u \tag{4.7}
\end{equation*}
$$

which, when they hold over an open neighbourhood, we recognise as the Cauchy-Riemann equations guaranteeing that $u+\mathrm{i} v$ is holomorphic. Since, as we shall see later, holomorphic functions are analytic, it means that conformal transformations in two dimensions have an infinite number of parameters (the coefficients of the series). In higher dimensions, the number of parameters is very much finite (eg. 15 in four dimensions).

Alternatively, if less generally, we can show that holomorphic functions rotate the tangents to all curves at that point by the same amount. Indeed, consider the tangent to some curve at $z_{0}$ and its image $w$ via $f$. If the curve is parametrised by $t, \mathrm{~d}_{t} z$ and $\mathrm{d}_{t} w$ are the tangent vectors at $z_{0}$ and at its image $w_{0}$, respectively. Then:

$$
\left.\mathrm{d}_{t} w\right|_{w_{0}}=\left.\mathrm{d}_{z} w \mathrm{~d}_{t} z\right|_{z_{0}}=\left.\mathrm{d}_{z} f \mathrm{~d}_{t} z\right|_{z_{0}}
$$

Writing $\mathrm{d}_{z} f=R \mathrm{e}^{\mathrm{i} \alpha}$, we conclude that $\left.\mathrm{d}_{t} z\right|_{z_{0}}$ is rotated by an angle $\alpha$ to obtain $\left.\mathrm{d}_{t} w\right|_{w_{0}}$. Since this applies to the tangent of any curve through $z_{0}$, the angle between the tangents of any two curves at $z_{0}$ is left unchanged by the transformation, and the latter is conformal.

If $f=u+\mathrm{i} v$ is holomorphic, one can show (exercise) by applying the Cauchy-Riemann conditions to the scalar product of the gradients of $u$ and $v$ that curves of constant $u$ and $v$ are perpendicular to each other.

One of the most useful conformal transformations is the so-called fractional (aka bilinear, homographic, Möbius) transformation:

$$
\begin{equation*}
f(z)=\frac{a z+b}{c z+d} \quad a d-b c \neq 0 \tag{4.8}
\end{equation*}
$$

where the condition $a d-b c \neq 0$ ensures that the transformation has a unique inverse. It is usual to scale the four parameters (leaving the transformation invariant) so that $a d-b c=1$. The transformation can be rewritten:

$$
f(z)=\lambda+\frac{\mu}{z+\nu}
$$

where $\lambda, \mu$ and $\nu$ are constants. Since $z+\nu$ represents a translation, $1 /(z+\nu)$ an inversion, and multiplication by $\mu$ a dilatation combined with a rotation, the bilinear transformation is a combination of all these.

Example 4.2. If $z_{0}$ lies in the upper half-plane, the bilinear transformation:

$$
w=\mathrm{e}^{\mathrm{i} \phi_{0}} \frac{z-z_{0}}{z-\bar{z}_{0}}
$$

maps the upper half-plane into the interior of the unit circle, $|w| \leq 1$. Indeed, $|w|=\left|z-z_{0}\right| /\left|z-\bar{z}_{0}\right|$ which is smaller than 1 except when $z$ lies on the real axis, in which case $|w|=1$.

Example 4.3. Under transformation (4.8), circles are mapped into circles. To see this, we note first that the general equation of a circle in the $x y$ plane can be written: $A\left(x^{2}+y^{2}\right)+B x+C y+D=0$, with $A>0$ and $D>0$. Transforming to complex coordinates yields $\alpha z z^{*}+\beta z+\beta^{*} z^{*}+\gamma=0$, where $\alpha=A, \beta=\frac{1}{2}(B-\mathrm{i} C)$, and $\gamma=D$. When $\alpha=A=0$, the circle collapses to a straight line.
Now, under $w=1 / z$, this becomes $\gamma w w^{*}+\beta w^{*}+\beta^{*} w+\alpha=0$, a circle. Under $w=a z(a \in \mathbb{C})$, it becomes $\alpha w w^{*}+\beta a^{*} w+\beta^{*} a w^{*}+\gamma a a^{*}=0$, another circle. Translations also map circles to circles. Since a fractional transformation is a combination of inversion, dilatation plus rotation, and translation, then it indeed maps circles to circles.

Conformal transformations are often used in some fields of physics or engineering to map a complicated set of objects to a simpler one. Finding the right one can be a black art, but extensive catalogues do exist, eg. S. G. Krantz, Handbook of Complex Variables (Birkhäuser, 1999), pp. 163-184.

### 4.3 Complex Integrals (BF 6.3)

### 4.3.1 The Cauchy-Goursat theorem

The integral of $f(z)=u+\mathrm{i} v$ over some path $C$ between two points $A$ and $B$ in the complex plane is written:

$$
\begin{equation*}
\int_{C} f(z) \mathrm{d} z=\int_{C}(u \mathrm{~d} x-v \mathrm{~d} y)+\mathrm{i} \int_{C}(v \mathrm{~d} x+u \mathrm{~d} y)=\int_{C} \mathbf{F} \cdot \mathrm{~d} \mathbf{l}+\mathrm{i} \int_{C} \mathbf{G} \cdot \mathrm{~d} \mathbf{l} \tag{4.9}
\end{equation*}
$$

where $\mathbf{F}=u \hat{\mathbf{x}}-v \hat{\mathbf{y}}$ and $\mathbf{G}=v \hat{\mathbf{x}}+u \hat{\mathbf{y}}$. The integral from point $B$ to point $A$ on the same path is the negative of the integral from $A$ to $B$.

Now consider a closed path, also called a contour. If it encloses a simply-connected region $S$ (no "holes" with points not belonging to $S$ ) in the complex plane, and if the derivatives are continuous everywhere in $S$, Stokes theorem requires that:

$$
\oint_{C} f(z) \mathrm{d} z=\int_{S}(\boldsymbol{\nabla} \times \mathbf{F}) \cdot \mathrm{d} \mathbf{S}+\mathrm{i} \int_{S}(\boldsymbol{\nabla} \times \mathbf{G}) \cdot \mathrm{d} \mathbf{S}
$$

where $\mathrm{d} \mathbf{S}=\hat{\mathbf{z}} \mathrm{d} S$. Evaluating the " $z$ " component of the two curls gives $(\boldsymbol{\nabla} \times \mathbf{F})_{z}=-\left(\partial_{y} u+\partial_{x} v\right)$ and $(\boldsymbol{\nabla} \times \mathbf{G})_{z}=$ $\partial_{x} u-\partial_{y} v$.

Now, if $f(z)$ is holomorphic everywhere in $S$ and on its border C, the curls vanish because of the CauchyRiemann conditions, and the Cauchy-Goursat theorem ${ }^{\dagger}$ holds:

$$
\begin{equation*}
\oint_{C} f(z) \mathrm{d} z=0 \quad f(z) \text { holomorphic within and on } C \tag{4.10}
\end{equation*}
$$

[^16]The Cauchy-Goursat theorem is readily extended to multiply-connected regions. Indeed, suppose there is a hole in $S$. Consider a path that goes some distance along the exterior contour, then leaves it on a path that takes it to the boundary of the hole, which is then traversed in the direction opposite the exterior contour, coming back to the latter on the reverse interconnecting path, and then finishes the trip along the exterior contour back to the initial point. The net contribution to the integral of the two interconnecting paths vanishes. The complete contour encloses a simply connected region, so $\oint f(z) \mathrm{d} z=0$. But this is equivalent to an integral over the exterior boundary plus an integral over the interior boundary, along a path such that if you walk on this boundary, $S$ is always to your left.

The converse of the Cauchy-Goursat theorem, Morera's theorem, also holds: if $f(z)$ is continuous in a simply connected region and if $\oint f(z) \mathrm{d} z=0$ around every simple closed curve in that region, then $f(z)$ is holomorphic all over the region.

### 4.3.2 A few consequences of the Cauchy-Goursat Theorem

If $f(z)$ is holomorphic within and on a (closed) contour $C$ around a simply connected region, then we have the following:

- Obviously, the integral of $f(z)$ between any two points in that region is path-independent.
- Fundamental Theorem of Calculus:

$$
\begin{equation*}
\int_{a}^{b} f(z) \mathrm{d} z=F(b)-F(a) \quad \mathrm{d}_{z} F(z)=f(z) \tag{4.11}
\end{equation*}
$$

where $F$ is holomorphic.

## - Winding number:

On a circle $C$ of radius $r$ centered on $z_{0}$, we obtain by direct integration:

$$
\oint_{C} \frac{1}{z-z_{0}} \mathrm{~d} z=\int_{0}^{2 \pi} \frac{1}{r \mathrm{e}^{\mathrm{i} \theta}} \mathrm{i} r \mathrm{e}^{\mathrm{i} \theta} \mathrm{~d} \theta=2 \pi \mathrm{i}
$$

This result is totally independent of the radius of the circle, or indeed the shape of the contour, as well as of $z_{0}$ itself!
Definition 4.3. We define the winding number with respect to $z_{0}$ as:

$$
\begin{equation*}
W_{C}\left(z_{0}\right)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{\mathrm{~d} z}{z-z_{0}} \tag{4.12}
\end{equation*}
$$

where $W_{C}\left(z_{0}\right) \in \mathbb{Z}^{+}$and $C$ is any loop enclosing $z_{0}$. Since the circle can be deformed into a loop of any shape that encloses $z_{0}$, clearly the winding number counts the number of times any closed loop goes around $z_{0}$ in the counterclockwise direction.

## - Cauchy-Gauss Integral Formula:

Suppose that $f(z)$ is holomorphic on a closed region of the complex plane with border $C$. Consider a circle $C_{0}$, centered on a point $z_{0}$, with no point outside $C$. Then, from the Cauchy-Goursat theorrem:

$$
\oint_{C} \frac{f(z)}{z-z_{0}} \mathrm{~d} z=\oint_{C_{0}} \frac{f(z)}{z-z_{0}} \mathrm{~d} z
$$

We can also write:

$$
\oint_{C_{0}} \frac{f(z)}{z-z_{0}} \mathrm{~d} z=f\left(z_{0}\right) \oint_{C_{0}} \frac{\mathrm{~d} z}{z-z_{0}}+\oint_{C_{0}} \frac{f(z)-f\left(z_{0}\right)}{z-z_{0}} \mathrm{~d} z=2 \pi \mathrm{i} f\left(z_{0}\right)+\oint_{C_{0}} \frac{f(z)-f\left(z_{0}\right)}{z-z_{0}} \mathrm{~d} z
$$

The last term on the right vanishes. To see this, invoke Darboux's inequality:

$$
\begin{equation*}
\left|\int_{C} g(z) \mathrm{d} z\right| \leq \int_{C}|g(z) \mathrm{d} z|=\int_{C}|g(z)||\mathrm{d} z| \leq M L \tag{4.13}
\end{equation*}
$$

where $M$ is the maximum value of $g(z)$ on $C$ and $L$ is the length of $C$. The middle equality results from $\left|z_{1} z_{2}\right|=\left|z_{1}\right|\left|z_{2}\right|$, working in the complex representation, Now call $\delta=\left|z-z_{0}\right|$ the radius of $C_{0}$. Because $f(z)$, being holomorphic, is continuous at $z_{0}$, we know that for any $\epsilon>0$, no matter how small, we can choose $\delta$ small enough that $\left|f(z)-f\left(z_{0}\right)\right|<\epsilon$. From the Darboux inequality:

$$
\left|\oint_{C_{0}} \frac{f(z)-f\left(z_{0}\right)}{z-z_{0}} \mathrm{~d} z\right| \leq \frac{\epsilon}{\delta}(2 \pi \delta)=2 \pi \epsilon
$$

Since $\epsilon$ can be arbitrarily small, we can make the absolute value of the integral as small as we wish.
Therefore, for any $z_{0}$ inside a region, enclosed by a contour $C$, where $f$ is holomorphic, the Cauchy-Gauss Integral formula holds:

$$
\begin{equation*}
f\left(z_{0}\right)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f(z)}{z-z_{0}} \mathrm{~d} z \tag{4.14}
\end{equation*}
$$

Thus, $f(z)$ on the (closed) boundary of a region in the complex plane determines its values everywhere inside, provided only that $f$ is holomorphic over the whole region.

- Write the integration variable in the Cauchy-Gauss formula as $\zeta$, with $z_{0}=z$, and differentiate it with respect to $z$ :

$$
\mathrm{d}_{z} f(z)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{(\zeta-z)^{2}} \mathrm{~d} \zeta
$$

where the integral is well-defined so long as $z$ is not on the contour (see p. BF333 for a different proof). We can go on differentiating any number $n$ of times. Unlike for functions of a real variable, $f(z)$ being holomorphic implies the existence of not only its first derivative, but of all its derivatives, and the following rather cute relation holds:

$$
\begin{equation*}
\left.\mathrm{d}_{z}^{n} f(z)\right|_{z_{0}}=\frac{n!}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{\left(\zeta-z_{0}\right)^{n+1}} \mathrm{~d} \zeta \tag{4.15}
\end{equation*}
$$

Example 4.4. Evaluate, over the circle $|z|=3$ :

$$
\oint \frac{\mathrm{e}^{2 z}}{(z+2)^{4}} \mathrm{~d} z
$$

This might seem quite hard until we spot the resemblance of the integrand with that in eq. (4.15) with $z_{0}=-2$, and $n=3$. Then:

$$
\left.\mathrm{d}_{z}^{3} f(z)\right|_{z=-2}=\frac{3!}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{(\zeta-z)^{4}} \mathrm{~d} \zeta
$$

Here, $f(\zeta)=\mathrm{e}^{2 \zeta}$, and:

$$
\oint_{|z|=3} \frac{\mathrm{e}^{2 \zeta}}{(\zeta+2)^{4}} \mathrm{~d} \zeta=\left.\frac{\pi \mathrm{i}}{3} \mathrm{~d}_{z}^{3} \mathrm{e}^{2 z}\right|_{z=-2}=\frac{8 \pi \mathrm{i}}{3} \mathrm{e}^{-4}
$$

Note again that it matters not a jot that the contour is a circle and what its radius is, so long as the point $z=-2$ lies inside: the result is exactly the same.

## - Liouville's Theorem:

If $f(z)$ is an entire function whose modulus is bounded, ie., if $|f(z)| \leq M$, then $f(z)$ is a constant. To prove this, all we have to show is that $\mathrm{d}_{z} f(z)=0$. We have just seen (eq. (4.15)) that:

$$
\mathrm{d}_{z} f(z)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{(\zeta-z)^{2}} \mathrm{~d} \zeta
$$

where we can choose $C$ to be a circle of radius $r$ centered at $z$. Then, evaluating the absolute value in the polar representation and then using the Darboux inequality:

$$
\left|\mathrm{d}_{z} f(z)\right|=\frac{1}{2 \pi}\left|\int_{0}^{2 \pi} \frac{f(\zeta)}{r^{2} \mathrm{e}^{2 \mathrm{i} \theta}} \mathrm{i} r \mathrm{e}^{\mathrm{i} \theta} \mathrm{~d} \theta\right| \leq \frac{1}{2 \pi r} \int_{0}^{2 \pi}|f(\zeta)| \mathrm{d} \theta \leq \frac{M}{r}
$$

Since $f$ is entire, we can take $r$ to be arbitrarily large, and is derivative indeed goes to 0 , proving the theorem. This means that entire functions such as $\mathrm{e}^{z}$ and $\sin z$ are necessarily not bounded (as we have seen before in the case of the latter).

## - Fundamental Theorem of Algebra:

If $P(z)=\sum a_{n} z^{n}$ is a polynomial of degree $N \geq 1$, then it has as at least one root, ie. there exists $z=z_{1}$ such that $P\left(z_{1}\right)=0$.

To show this, it is sufficient to assume that $P(z) \neq 0$ everywhere, which leads to $1 /|P(z)|$ being everywhere holomorphic (entire). Moreover, $1 / P(z)$ is bounded, and thus, by Liouville's theorem, is a constant, which is impossible. Thus, there must exists at least one value $z_{1}$ of $z$ such that $P(z)=\left(z-z_{1}\right) Q(z)=0$, where $Q(z)$ is a poynomial of degree $N-1$.
We can go on to argue in the same way that $Q(z)$ must also have at least one root, ie. a value $z_{2}$ where it vanishes. At the end of this process, we find that $P(z)$ must have exactly $N$ roots.

## - Poisson's integral formulae

Let $f(z)$ be holomorphic on and inside a circle $C$ of radius $R$, and $z=r \mathrm{e}^{\mathrm{i} \theta}$ be any point inside the circle. Then, from the Cauchy-Gauss integral formula (4.14):

$$
f(z)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{\zeta-z} \mathrm{~d} \zeta
$$

Now the point $R^{2} / \bar{z}$ lies outside the circle, so that the function $f(\zeta) /\left(\zeta-R^{2} / \bar{z}\right)$ is holomorphic inside, and its integral along the circle vanishes from the Cauchy theorem. Therefore, we can write:

$$
\begin{aligned}
f(z) & =\frac{1}{2 \pi \mathrm{i}}\left(\oint_{C} \frac{f(\zeta)}{\zeta-z} \mathrm{~d} \zeta-\oint_{C} \frac{f(\zeta)}{\zeta-R^{2} / \bar{z}} \mathrm{~d} \zeta\right)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{z-R^{2} / \bar{z}}{(\zeta-z)\left(\zeta-R^{2} / \bar{z}\right)} f(\zeta) \mathrm{d} \zeta \\
& =\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{r^{2}-R^{2}}{\bar{z}(\zeta-z)(1-\bar{\zeta} / \bar{z})} f(\zeta) \frac{\mathrm{d} \zeta}{\zeta}=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{R^{2}-r^{2}}{|\zeta-z|^{2}} f(\zeta) \frac{\mathrm{d} \zeta}{\zeta} \\
& =\frac{1}{2 \pi} \int_{0}^{2 \pi} \frac{R^{2}-r^{2}}{R^{2}+r^{2}-2 R r \cos (\theta-\phi)} f\left(R \mathrm{e}^{\mathrm{i} \phi}\right) \mathrm{d} \phi=\oint_{C} P_{r}(\theta-\phi) f\left(R \mathrm{e}^{\mathrm{i} \phi}\right) \mathrm{d} l
\end{aligned}
$$

where, in the last line, we have gone to the polar representation $z=r \mathrm{e}^{\mathrm{i} \theta}$ and $\zeta=R \mathrm{e}^{\mathrm{i} \phi}$, and where:

$$
P_{r}(\theta-\phi):=\frac{1}{2 \pi R} \frac{R^{2}-r^{2}}{R^{2}+r^{2}-2 R r \cos (\theta-\phi)}
$$

is an example of a 2-dim Poisson kernel, here for a disc of radius $R$. This result is also an example of a convolution integral known as Poisson's integral formula on a circle.

There is a companion formula on the half-plane. Let $f(z)$ be holomorphic in the upper half-plane, and consider a semi-circle of radius $R$ in the upper half-plane, centered on the origin, and with its base on the real axis. Let $z=a+\mathrm{i} b$ be a point inside the semi-circle. Then the Cauchy-Gauss formula and Cauchy's theorem give, respectively:

$$
f(z)=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{\zeta-z} \mathrm{~d} \zeta, \quad 0=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f(\zeta)}{\zeta-\bar{z}} \mathrm{~d} \zeta
$$

Subtracting and combining yields:

$$
\begin{aligned}
f(z) & =\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{z-\bar{z}}{(\zeta-z)(\zeta-\bar{z})} f(\zeta) \mathrm{d} \zeta \\
& =\frac{1}{\pi} \int_{-R}^{R} \frac{b}{(x-a)^{2}+b^{2}} f(x) \mathrm{d} x+\frac{1}{\pi} \int_{\substack{|\zeta|=R \\
y>0}} \frac{b}{(\zeta-z)(\zeta-\bar{z})} f(\zeta) \mathrm{d} \zeta
\end{aligned}
$$

This time, however, we impose the condition $|f(\zeta) / \zeta| \rightarrow 0$ as $R \rightarrow \infty$. Then the second integral on the semi-circle vanishes as $R \rightarrow \infty$, and we are left with another Poisson integral formula:

$$
\begin{equation*}
f(z)=\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{b}{(x-a)^{2}+b^{2}} f(x) \mathrm{d} x \tag{4.16}
\end{equation*}
$$

What can we do with these Poisson formulae? Well, you can be the life of the party by easily proving that:

$$
\int_{0}^{2 \pi} \frac{R^{2}-r^{2}}{R^{2}+r^{2}-2 r R \cos (\theta-\phi)} \mathrm{d} \phi=2 \pi
$$

### 4.4 Power-Series Expansions of Analytic Functions — Laurent and Taylor series (BF 6.7)

Power-series expansions provide an important application of the complex differentiability of functions of a complex variables.

Let $f(z)$ be holomorphic between and on circles $C_{1}$ and $C_{2}$ centered on $z_{0}$ in the complex plane. Then, everywhere in that region, $f(z)$ can be expanded in a unique series, the Laurent series:

$$
\begin{equation*}
f(z)=\sum_{n=-\infty}^{\infty} A_{n}\left(z-z_{0}\right)^{n} \quad A_{n}=\frac{1}{2 \pi \mathrm{i}} \oint_{C} \frac{f\left(z^{\prime}\right)}{\left(z^{\prime}-z_{0}\right)^{n+1}} \mathrm{~d} z^{\prime} \quad(n \in \mathbb{Z}) \tag{4.17}
\end{equation*}
$$

where $C$ is any contour that encloses $z_{0}$ in the annular region, and we take $C_{1}$ to be the outer circle. The series converges everywhere inside the annular region between $C_{1}$ and $C_{2}$. Outside this region, the series generally does not converge, and a new one must be found that converges.

To prove this theorem, consider a contour that traverses $C_{1}$ counterclockwise, but leaves $C_{1}$ at some point to go to $C_{2}$ and traverse it clockwise, coming back to $C_{1}$ along the path connecting the two circles. Integrating $f\left(z^{\prime}\right) /\left(z^{\prime}-z\right)$ over this contour, it is clear that the interconnecting path makes no net contribution to the integral. And since $f$ is holomorphic everywhere on and inside the contour, the value of $f$ anywhere inside is given by the Cauchy-Gauss iItegral formula (4.14) which, applied to this contour, is:

$$
\begin{aligned}
f(z) & =\frac{1}{2 \pi \mathrm{i}} \oint_{C_{1}} \frac{f\left(z^{\prime}\right)}{z^{\prime}-z} \mathrm{~d} z^{\prime}-\frac{1}{2 \pi \mathrm{i}} \oint_{C_{2}} \frac{f\left(z^{\prime}\right)}{z^{\prime}-z} \mathrm{~d} z^{\prime} \\
& =\frac{1}{2 \pi \mathrm{i}} \oint_{C_{1}} \frac{f\left(z^{\prime}\right)}{\left[\left(z^{\prime}-z_{0}\right)-\left(z-z_{0}\right)\right]} \mathrm{d} z^{\prime}+\frac{1}{2 \pi \mathrm{i}} \oint_{C_{2}} \frac{f\left(z^{\prime}\right)}{\left[\left(z-z_{0}\right)-\left(z^{\prime}-z_{0}\right)\right]} \mathrm{d} z^{\prime}
\end{aligned}
$$

where the minus sign between the terms in the first line is due to $C_{2}$ being traversed clockwise.

Now we establish a useful identity:

$$
\frac{1}{z_{1}-z_{2}}=\frac{1}{z_{1}}+\frac{z_{2}}{z_{1}} \frac{1}{z_{1}-z_{2}}=\sum_{n=0}^{N} \frac{z_{2}^{n}}{z_{1}^{n+1}}+\left(\frac{z_{2}}{z_{1}}\right)^{N+1} \frac{1}{z_{1}-z_{2}}
$$

where the last equality is obtained by iterating the first equality (corresponding to $N=0$ ) $N$ times. With this identity, $f(z)$ becomes:

$$
\begin{array}{r}
f(z)=\frac{1}{2 \pi \mathrm{i}} \sum_{n=0}^{N} \oint_{C_{1}} \frac{\left(z-z_{0}\right)^{n}}{\left(z^{\prime}-z_{0}\right)^{n+1}} f\left(z^{\prime}\right) \mathrm{d} z^{\prime}+\frac{1}{2 \pi \mathrm{i}} \sum_{m=0}^{N} \oint_{C_{2}} \frac{\left(z^{\prime}-z_{0}\right)^{m}}{\left(z-z_{0}\right)^{m+1}} f\left(z^{\prime}\right) \mathrm{d} z^{\prime}+ \\
\frac{1}{2 \pi \mathrm{i}} \oint_{C_{1}}\left(\frac{z-z_{0}}{z^{\prime}-z_{0}}\right)^{N+1} \frac{f\left(z^{\prime}\right)}{z^{\prime}-z} \mathrm{~d} z^{\prime}+\frac{1}{2 \pi \mathrm{i}} \oint_{C_{2}}\left(\frac{z^{\prime}-z_{0}}{z-z_{0}}\right)^{N+1} \frac{f\left(z^{\prime}\right)}{z-z^{\prime}} \mathrm{d} z^{\prime}
\end{array}
$$

Define $M_{1,2}=|f(z)|_{\max }$ on $C_{1,2}, d_{1,2}=\left|z-z^{\prime}\right|_{\min }$ with $z^{\prime} \in C_{1,2}$, and $l_{1,2}=\left|z-z_{0}\right|_{\text {max,min }}$, We note that $l_{1}<r_{1}=\left|z^{\prime}-z_{0}\right|$ and $l_{2}>r_{2}=\left|z^{\prime}-z_{0}\right|$, where $r_{1,2}$ is the radius of $C_{1,2}$. Then the absolute value of the last two terms is bounded by:

$$
\frac{M_{1} r_{1}}{d_{1}}\left(\frac{l_{1}}{r_{1}}\right)^{N+1}+\frac{M_{2} r_{2}}{d_{2}}\left(\frac{r_{2}}{l_{2}}\right)^{N+1} \underset{N \rightarrow \infty}{\rightarrow} 0
$$

Now, from the Cauchy-Goursat theorem, each integral in the first and second term is equal to an integral over an arbitrary contour C that encloses $z_{0}$ between $C_{1}$ and $C_{2}$. All that is left is to redefine the index $m$ in the second integral as $m=-n-1$ and rewrite it as a sum over negative $n$. After letting $N \rightarrow \infty$, we merge the two sums and thus prove eq. (4.17). We can even deform $C_{1}$ and $C_{2}$ to other contours, so long as we do not cross any singular point on the way, ie., so long as $f(z)$ remains holomorphic within and on the deformed contours.

Definition 4.4. If $A_{n} \neq 0$ for some $n<0$, clearly $f(z)$ has a singularity at $z=z_{0}$. If $A_{n}=0 \forall n<$ $-N(N>0)$, then we say that $f(z)$ has a pole of order $N$ at $z_{0}$. If $N \rightarrow \infty$, ie., if the series of negative powers is infinite, and if it is isolated-in the sense that it is not a branch point-this singularity is said to be essential. $f(z)$ is called meromorphic in regions where it has no essential singularity.

Because of their uniform convergence in the annular region, one can manipulate Laurent series as if they were finite power series, ie., polynomials: they can be added, multiplied, and even divided to represent the addition, the multiplication, or the ratio of the functions which they represent.

Now suppose further that $f(z)$ is holomorphic everywhere inside $C_{1}$, including at $z_{0}$. Then $A_{n}=0 \forall n \leq-1$ so as to keep the Laurent series of $f(z)$ holomorphic at $z_{0}$. When only terms with $n \geq 0$ are present, we use eq. (4.15) to obtain directly Taylor's theorem for holomorphic functions:

$$
\begin{equation*}
f(z)=\left.\sum_{n=0}^{\infty} \frac{1}{n!} \mathrm{d}_{z}^{n} f(z)\right|_{z_{0}}\left(z-z_{0}\right)^{n} \tag{4.18}
\end{equation*}
$$

valid anywhere inside a circle centered on $z_{0}$ where $f(z)$ is holomorphic. It also establishes that holomorphic functions are analytic in the sense of Definition 4.1, justifying the assertion that the two properties are equivalent.

Beyond its well-known usefulness, the Taylor series is also an accurate representation of $f(z)$ within a radius of convergence which extends all the way to its closest singularity. But since the function and all its derivatives are known at any point inside the circle of convergence, we can choose another point inside the circle around which to Taylor-expand, extending the domain in which we can calculate $f(z)$ further, until another singularity is encountered. This so-called analytic continuation can be repeated any number of times.

We conclude that knowledge of $f(z)$ in some region of the complex plane is sufficient to determine it uniquely everywhere it is analytic! This explains why we are able to start from the definitions and relations applying to many functions on the real axis, such as $\mathrm{e}^{x}$ or the trig functions, and extend them with impunity to the whole complex plane, or at least to any region in which the same function, with $x \rightarrow z$, is analytic.

Example 4.5. The function $1 /(1-z)$ can be written as the geometric series $S_{1}(z)=\sum_{0}^{\infty} z^{n}$, provided that $|z|<1$. Similarly, if $|z+1 / 4|<5 / 4$ :

$$
S_{2}(z)=\frac{4}{5} \sum_{n=0}^{\infty}\left[\frac{4}{5}\left(z+\frac{1}{4}\right)\right]^{n}=\frac{4 / 5}{1-\frac{4}{5}(z+1 / 4)}=\frac{1}{1-z}
$$

The region in which $S_{2}$ converges is larger than the circle of convergence of $S_{1}$, and both represent the same function in the overlapping region. $S_{2}$ may be seen as an analytic continuation of $S_{1}$ beyond $|z|<1$. Both $S_{1}$ and $S_{2}$ are Laurent series in their respective region of validity.

Example 4.6. The fact that the $2 \pi \mathrm{i} A_{-1}$ coefficient of a Laurent series is equal to $\oint f(z) \mathrm{d} z$ can sometimes be used to evaluate the contour integral if $A_{-1}$ can be found in some other way. Take $f(z)=1 /\left[z^{2}(z-3)\right]$ which has singularities at $z=0$ and at $z=3$, but is analytic in the open annular region $0<|z|<3$. Now write $f(z)=-1 /\left[3 z^{2}(1-z / 3)\right]$, and focus on the factor:

$$
\frac{1}{1-z / 3}=\sum_{n=0}^{\infty}(z / 3)^{n}
$$

The right-hand side can be thought of either as a geometric series or as a Taylor expansion around $z=0$ (since there is no singularity there). Then we reconstruct $f(z)$ and find that it can be written as the Laurent series:

$$
f(z)=-\frac{1}{3 z^{2}}-\frac{1}{9 z}-\frac{1}{27}-\frac{z}{81}-\text { positive powers of } z \quad|z|<3
$$

Therefore, $A_{-1}=-1 / 9$, and we arrive at:

$$
\oint_{C} \frac{1}{z^{2}(z-3)} \mathrm{d} z=-\mathrm{i} \frac{2 \pi}{9}
$$

for any closed path that encloses $z=0$ in the region $0<|z|<3$.
For the region $|z|>3$, this Laurent series is not well defined because its positive-power terms make it diverge as $|z| \rightarrow \infty$. But we can write for that region only:

$$
f(z)=\frac{1}{z^{3}} \frac{1}{1-3 / z}=\frac{1}{z^{3}} \sum_{n=0}^{\infty}(3 / z)^{n}
$$

Example 4.7. As a highly non-trivial example of a Laurent series, consider:

$$
f(z)=\frac{1}{\sqrt{z^{2}-1}}=\frac{1}{\sqrt{r_{1} r_{2}}} \mathrm{e}^{-\mathrm{i} \theta_{1} / 2} \mathrm{e}^{-\mathrm{i} \theta_{2} / 2} \quad r_{1}=|z-1|, r_{2}=|z+1|
$$

If we start at some $z$ such that $|z-1|>1$, for instance at $\theta_{1}=0$ and $\theta_{2}=0$ (positive real axis), it is clear in this polar parametrisation that when we come back to the same point after traversing a circle that encloses both singularities, both angles go from 0 to $2 \pi$, and we come back to the same value of $f(z)$, so that the function is single-valued and analytic on the real axis for $|z|>1$. If, on the other hand, we start on the real axis between $z=1$ and $z=-1$, and go full circle round one only of the two singular points, only one angle goes from 0 to $2 \pi$, and we come back to the negative of the value of $f(z)$ that we started from. Thus, there is a branch cut running between $z=1$ and $z=-1$, and $f(z)$ is not analytic there.
Now $f(z)$ is analytic everywhere in the region (centered on the origin) $|z|>1$, and this is where we look for a Laurent series. If we take the inner boundary $C_{2}$ of the region of existence of the Laurent
series to be of radius $|z|=1+\epsilon$ around the origin, where $\epsilon$ can be as small as we wish but not zero, and if we further choose the outer boundary to have arbitrarily large radius, then $A_{n}=0 \forall n \geq 0$.
To find the $A_{n}$ for $n<0$, we choose a contour $C$ made of circles of arbitrarily small (but non-zero) radius around $z=1$ and $z=-1$, connected by a straight line hugging the real axis in the upper half-plane and another one very close to the real axis in the lower half-plane. Although this might seem to contradict the conditions $|z|>1$ for the existence of a Laurent series, it does not: we use the path-independence of contour integration valid anywhere in a region of analyticity to say that the result is the same as if we had actually integrated on a contour with $|z|>1$. The series itself, of course, is valid only for $|z|>1$.
The contributions from the two circles can be seen to vanish. Indeed, for the circle $C^{\prime}$ around $z=1$, we write:

$$
\oint_{C^{\prime}} \frac{\left(z^{\prime 2}-1\right)^{-1 / 2}}{\left(z^{\prime}\right)^{n+1}} \mathrm{~d} z^{\prime}=\oint_{C^{\prime}} \frac{\sqrt{z^{\prime}-1}}{\sqrt{z^{\prime}+1}\left(z^{\prime}\right)^{n+1}} \frac{1}{z^{\prime}-1} \mathrm{~d} z^{\prime}
$$

The integrand is of the form $f\left(z^{\prime}\right) /\left(z^{\prime}-1\right)$, where $f\left(z^{\prime}\right)$ is analytic on and within $C^{\prime}$. Therefore, from the Cauchy-Gauss integral formula (4.14), the integral is proportional to $f\left(z^{\prime}=1\right)=0$. Similar considerations apply to the circle around $z=-1$, and we are left with the contributions from the two straight lines above and below the real axis, arbitrarily close to it. Because of the discontinuity across the cut, these do not cancel one another. We write $\sqrt{x^{2}-1}= \pm \mathrm{e}^{\mathrm{i} \pi / 2} \sqrt{1-x^{2}}$, where the + sign applies to the path from 1 to -1 , in the upper half-plane, and the - sign to the path from -1 to 1, in the lower half-plane. So long as we keep the sign distinction, we can then integrate on the actual $x$ axis, and there comes for $A_{n} \neq 0$ :

$$
A_{n}=\frac{1}{2 \pi \mathrm{i}}\left[-\int_{-1}^{1} \frac{x^{-(n+1)} \mathrm{e}^{-\mathrm{i} \pi / 2}}{\sqrt{1-x^{2}}} \mathrm{~d} x+\int_{1}^{-1} \frac{x^{-(n+1)} \mathrm{e}^{-\mathrm{i} \pi / 2}}{\sqrt{1-x^{2}}} \mathrm{~d} x\right]=\frac{1}{\pi} \int_{-1}^{1} \frac{x^{-(n+1)}}{\sqrt{1-x^{2}}} \mathrm{~d} x \quad(n \text { odd })
$$

The integral can be evaluated by Maple/Mathematica or tables, and we finally get, with $\nu=-(n+$ $1) / 2$, the Laurent series:

$$
\frac{1}{\sqrt{z^{2}-1}}=\sum_{\nu=0}^{\infty} \frac{(2 \nu)!}{2^{2 \nu}(\nu!)^{2}} \frac{1}{z^{2 \nu+1}}=\frac{1}{z}+\frac{1}{2} \frac{1}{z^{3}}+\frac{3}{8} \frac{1}{z^{5}}+\ldots \quad|z|>1
$$

Although this series has an infinite number of negative powers, the singularities at $z= \pm 1$ are not essential because, being branch points at the ends of the cut between $z=1$ and $z=-1$, they are not isolated .

### 4.5 Cauchy Principal Value of an Integral (BF 6.5)

Consider the following integral over the real axis, with $-R<x_{0}<R$ :

$$
\lim _{\delta \rightarrow 0}\left[\int_{-R}^{x_{0}-\delta} \frac{f(x)}{x-x_{0}} \mathrm{~d} x+\int_{x_{0}+\delta}^{R} \frac{f(x)}{x-x_{0}} \mathrm{~d} x\right] \equiv \mathcal{P} \int_{-R}^{R} \frac{f(x)}{x-x_{0}} \mathrm{~d} x
$$

where the symbol $\mathcal{P}$ means the Cauchy principal value of $f(x) /\left(x-x_{0}\right)$. Note that $\delta$ must be the same in both integrals. If we can evaluate this, we can make sense of an integral over the real axis whose integrand has a singularity at some point.

The principal value may always be written as:

$$
\mathcal{P} \int_{-R}^{R} \frac{f(x)}{x-x_{0}} \mathrm{~d} x=f\left(x_{0}\right) \mathcal{P} \int_{-R}^{R} \frac{1}{x-x_{0}} \mathrm{~d} x+\mathcal{P} \int_{-R}^{R} \frac{f(x)-f\left(x_{0}\right)}{x-x_{0}} \mathrm{~d} x
$$

Now, if $-R<x_{0}<R$,

$$
\begin{aligned}
\mathcal{P} \int_{-R}^{R} \frac{1}{x-x_{0}} \mathrm{~d} x & =\lim _{\delta \rightarrow 0}\left[\int_{-R}^{x_{0}-\delta} \frac{1}{x-x_{0}} \mathrm{~d} x+\int_{x_{0}+\delta}^{R} \frac{1}{x-x_{0}} \mathrm{~d} x\right] \\
& =\lim _{\delta \rightarrow 0} \ln \left(\frac{R-x_{0}}{R+x_{0}}\right)=\ln \left(\frac{R-x_{0}}{R+x_{0}}\right)
\end{aligned}
$$

This Cauchy principal value vanishes when the limits of integration are taken to infinity $(R \rightarrow \infty)$.
Thus, we obtain the useful relation, valid when $R>\left|x_{0}\right|$ :

$$
\begin{equation*}
\mathcal{P} \int_{-R}^{R} \frac{f(x)}{x-x_{0}} \mathrm{~d} x=f\left(x_{0}\right) \ln \left(\frac{R-x_{0}}{R+x_{0}}\right)+\mathcal{P} \int_{-R}^{R} \frac{f(x)-f\left(x_{0}\right)}{x-x_{0}} \mathrm{~d} x \tag{4.19}
\end{equation*}
$$

The integrand of the integral on the right is not singular at $x=x_{0}$ when $f(x)$ is differentiable at $x_{0}$; then the $\mathcal{P}$ symbol can be dropped.

Example 4.8. To illustrate this, let $f(x)=x /\left(x^{2}+1\right)$, whose derivative exists everywhere. The integrand on the right-hand side of eq. (4.19) becomes:

$$
\frac{f(x)-f\left(x_{0}\right)}{x-x_{0}}=\left(\frac{x}{x^{2}+1}-\frac{x_{0}}{x_{0}^{2}+1}\right) \frac{1}{x-x_{0}}=\frac{1-x_{0} x}{\left(x^{2}+1\right)\left(x_{0}^{2}+1\right)}
$$

The integrand is manifestly not singular at $x_{0}$, and we can drop the principal value symbol. Then:

$$
\mathcal{P} \int_{-R}^{R} \frac{x}{\left(x^{2}+1\right)\left(x-x_{0}\right)} \mathrm{d} x=\frac{1}{x_{0}^{2}+1}\left[x_{0} \ln \left(\frac{R-x_{0}}{R+x_{0}}\right)+2 \tan ^{-1} R\right] \quad\left(R>\left|x_{0}\right|\right)
$$

The restriction to $-R<x_{0}<R$ becomes trivial if $R \rightarrow \infty$ :

$$
\mathcal{P} \int_{-\infty}^{\infty} \frac{x}{\left(x^{2}+1\right)\left(x-x_{0}\right)} \mathrm{d} x=\frac{\pi}{x_{0}^{2}+1}
$$

### 4.6 Hilbert Transforms (BF 6.5)

I have already mentioned soon after introducing holomorphic functions how, given the real part of some holomorphic function $f(z)$, its imaginary part could be found up to a constant, and vice-versa. Let $|f(z)| \rightarrow 0$ as $|z| \rightarrow \infty$ in either the upper or lower half-plane, but not both (since then $f$ would be entire and bounded, thus constant, ie. zero, by Liouville's theorem). We will show that with this condition a stronger result can be obtained for that same complex-valued function evaluated on the real axis: the real and imaginary parts of $f(x)$ are uniquely related via integrals over the real axis! Unlike when the Cauchy-Riemann conditions are used, we do not need detailed knowledge of $f(z)$ away from the real axis, just that it is analytic in, say, the upper half-plane, and that it vanishes at infinity in that region.

First, we would like to integrate the function $f(z) /\left(z-x_{0}\right)$ over a closed counter-clockwise path (contour) consisting of a half-circle of infinite radius in the upper half-plane. centered on the origin, with its base along the real axis. $f(z)$ is analytic on and within the contour, but $f(z) /\left(z-x_{0}\right)$ has a singularity at $x_{0}$ on the real axis, so it is not analytic on the contour.

A strategy to make sense of the integral is to "avoid" $x_{0}$ along a semi-circle $C_{\delta}$ of radius $\delta$ in the upper halfplane, centered on $x_{0}$, whose diameter extends from $x_{0}-\delta$ to $x_{0}+\delta$. Since now $f(z) /\left(z-x_{0}\right)$ is analytic everywhere within and on the closed path we have chosen, the contour integral vanishes by the Cauchy-Goursat theorem.

The contribution to the contour integral from the upper arc of radius $R \rightarrow \infty$ is easily evaluated:

$$
\lim _{R \rightarrow \infty} \int_{\substack{|z|=R \\ y>0}} \frac{f(z)}{z-x_{0}} \mathrm{~d} z=\lim _{R \rightarrow \infty} \mathrm{i} \int_{0}^{\pi} \frac{f(z)}{\mathrm{e}^{\mathrm{i} \theta}-x_{0} / R} \mathrm{e}^{\mathrm{i} \theta} \mathrm{~d} \theta
$$

Now, $\left|\mathrm{e}^{\mathrm{i} \theta}-x_{0} / R\right| \geq\left|1-x_{0} / R\right|$. Therefore, so long as $|f(z)| \rightarrow 0$ as $R \rightarrow \infty$, we have:

$$
\left|\int_{R \rightarrow \infty} \frac{f(z)}{z-x_{0}} \mathrm{~d} z\right| \leq \lim _{R \rightarrow \infty} \frac{1}{\left|1-x_{0} / R\right|} \int_{0}^{\pi}|f(z)| \mathrm{d} \theta \longrightarrow 0
$$

Another part of our closed path is the small semi-circular arc $C_{\delta}$ of radius $\delta$, traversed clockwise:

$$
\int_{-C_{\delta}} \frac{f(z)}{z-x_{0}} \mathrm{~d} z \equiv f\left(x_{0}\right) \int_{-C_{\delta}} \frac{\mathrm{d} z}{z-x_{0}}+\int_{-C_{\delta}} \frac{f(z)-f\left(x_{0}\right)}{z-x_{0}} \mathrm{~d} z=-\mathrm{i} \pi f\left(x_{0}\right)-\int_{C_{\delta}} \frac{f(z)-f\left(x_{0}\right)}{z-x_{0}} \mathrm{~d} z
$$

where $\left|z-x_{0}\right|=\delta$. The last integral vanishes as $\delta \rightarrow 0$. To see this, note that:

$$
\lim _{\delta \rightarrow 0}\left|\int_{C_{\delta}} \frac{f(z)-f\left(x_{0}\right)}{z-x_{0}} \mathrm{~d} z\right| \leq \lim _{\delta \rightarrow 0} \int_{0}^{\pi}\left|f(z)-f\left(x_{0}\right)\right| \mathrm{d} \theta
$$

But since $f$ is continuous (it is analytic), $\left|f(z)-f\left(x_{0}\right)\right|$ can always be made smaller than any positive number $\epsilon$, no matter how small that number may be, by taking $\delta$ small enough, which the limit implies anyway. Then the integral, which is smaller than $\pi \epsilon$ from eq. (4.13) (Darboux), can be made to vanish by choosing $\epsilon$ vanishingly small. There is a much niftier way of calculating this integral over $C_{\delta}$, but that is for a little later.

The remaining contribution is the principal value integral of $f(x) /\left(x-x_{0}\right)$. Thus, summing the non-zero contributions to the contour integral to zero, we have succeeded in evaluating the Cauchy principal value:

$$
\begin{equation*}
\mathcal{P} \int_{-\infty}^{\infty} \frac{f(x)}{x-x_{0}} \mathrm{~d} x=\mathrm{i} \pi f\left(x_{0}\right) \tag{4.20}
\end{equation*}
$$

(EXERCISE: Can this result be used to evaluate the Cauchy principal value of the function in example 4.8?) By taking the real and imaginary parts of this equation, we conclude that if $f(x)$ is a complex-valued function of a real variable such that $f(z)$ is analytic in the upper half-plane (or lower half-plane, by a similar argument) with $|f(z)| \rightarrow 0$ as $|z| \rightarrow \infty$, its real and imaginary parts are related by the Hilbert transforms:

$$
\begin{equation*}
\Re\left[f\left(x_{0}\right)\right]=\frac{1}{\pi} \mathcal{P} \int_{-\infty}^{\infty} \frac{\Im[f(x)]}{x-x_{0}} \mathrm{~d} x, \quad \Im\left[f\left(x_{0}\right)\right]=-\frac{1}{\pi} \mathcal{P} \int_{-\infty}^{\infty} \frac{\Re[f(x)]}{x-x_{0}} \mathrm{~d} x \tag{4.21}
\end{equation*}
$$

Also, with $R \rightarrow \infty$, the relation (4.19) we derived in the last section becomes:

$$
\mathcal{P} \int_{-\infty}^{\infty} \frac{f(x)}{x-x_{0}} \mathrm{~d} x=\mathcal{P} \int_{-\infty}^{\infty} \frac{f(x)-f\left(x_{0}\right)}{x-x_{0}} \mathrm{~d} x
$$

Now, because $f(z)$ is analytic on the real axis, it is differentiable at $x_{0}$, and we can drop the $\mathcal{P}$ symbol in the right-hand side. Then the Hilbert transforms become:

$$
\begin{equation*}
\Re\left[f\left(x_{0}\right]=\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{\Im[f(x)]-\Im\left[f\left(x_{0}\right)\right]}{x-x_{0}} \mathrm{~d} x, \quad \Im\left[f\left(x_{0}\right)\right]=-\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{\Re[f(x)]-\Re\left[f\left(x_{0}\right)\right]}{x-x_{0}} \mathrm{~d} x\right. \tag{4.22}
\end{equation*}
$$

### 4.7 Dispersion relations (BF 6.6)

### 4.7.1 Non-locality in time

The behaviour in time, $R(t)$, of a system in response to a stimulus $I(t)$ cannot be instantaneous. We shall now elucidate important consequences of this fact.

Suppressing any spatial dependence, the Fourier decompositions of $R(t)$ and $I(t)$ over the frequency domain are (note the choice of sign in the exponentials!):

$$
\begin{aligned}
R(t) & =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} r(\omega) \mathrm{e}^{-\mathrm{i} \omega t} \mathrm{~d} \omega & r(\omega) & =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} R\left(t^{\prime}\right) \mathrm{e}^{\mathrm{i} \omega t^{\prime}} \mathrm{d} t^{\prime} \\
I(t) & =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} i(\omega) \mathrm{e}^{-\mathrm{i} \omega t} \mathrm{~d} \omega & i(\omega) & =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} I\left(t^{\prime}\right) \mathrm{e}^{\mathrm{i} \omega t^{\prime}} \mathrm{d} t^{\prime}
\end{aligned}
$$

We are interested in situations where $r(\omega)=g(\omega) i(\omega)$, with $g(\omega)$ bounded $\forall \omega$. Then:

$$
\begin{aligned}
R(t) & =\frac{1}{2 \pi} \int_{-\infty}^{\infty} \mathrm{d} \omega g(\omega) \mathrm{e}^{-\mathrm{i} \omega t} \int_{-\infty}^{\infty} \mathrm{d} t^{\prime} \mathrm{e}^{\mathrm{i} \omega t^{\prime}} I\left(t^{\prime}\right)=\frac{1}{2 \pi} \int_{-\infty}^{\infty}\left[\int_{-\infty}^{\infty} \mathrm{d} \omega g(\omega) \mathrm{e}^{-\mathrm{i} \omega\left(t-t^{\prime}\right)}\right] I\left(t^{\prime}\right) \mathrm{d} t^{\prime} \\
& =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} G(\tau) I(t-\tau) \mathrm{d} \tau \quad\left(\tau=t-t^{\prime}\right)
\end{aligned}
$$

where $G(\tau)$ is the kernel, or response function:

$$
\begin{equation*}
G(\tau)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} g(\omega) \mathrm{e}^{-\mathrm{i} \omega \tau} \mathrm{~d} \omega \quad g(\omega)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} G(\tau) \mathrm{e}^{\mathrm{i} \omega \tau} \mathrm{~d} \tau \tag{4.23}
\end{equation*}
$$

We see that $R(t)$ depends on the whole history of $I(t)$, and that the Fourier formalism provides non-locality in time in a natural way, as the convolution integral (see section 5.5 of Module V ):

$$
R(t)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} G\left(t-t^{\prime}\right) I\left(t^{\prime}\right) \mathrm{d} t^{\prime}
$$

But this is a little strange: not only does $R(t)$ depend on $I\left(t^{\prime}<t\right)$, it also depends on what $I$ does at times later than $t\left(t^{\prime}>t\right)$ !

### 4.7.2 Causality and Analyticity of $g(\omega)$

If we impose causality, however, only values of the stimulus before $t$ can contribute to $R(t)$. Thus, $G(\tau \leq 0)=0$, and we can write a very general relation:

$$
\begin{equation*}
R(t)=\frac{1}{\sqrt{2 \pi}} \int_{0}^{\infty} G(\tau) I(t-\tau) \mathrm{d} \tau \tag{4.24}
\end{equation*}
$$

with the frequency dependence of the response function given by:

$$
\begin{equation*}
g(\omega)=\frac{1}{\sqrt{2 \pi}} \int_{0}^{\infty} G(\tau) \mathrm{e}^{\mathrm{i} \omega \tau} \mathrm{~d} \tau \tag{4.25}
\end{equation*}
$$

Now extend $\omega$ to complex values and, on a contour enclosing the upper-half plane, evaluate:

$$
\oint_{\mathrm{C}} g(\omega) \mathrm{e}^{-\mathrm{i} \omega \tau} \mathrm{~d} \omega=\int_{-\infty}^{\infty} g(\omega) \mathrm{e}^{-\mathrm{i} \omega \tau} \mathrm{~d} \omega+\int_{|\omega| \rightarrow \infty} g(\omega) \mathrm{e}^{-\mathrm{i} \omega \tau} \mathrm{~d} \omega
$$

Since we can choose any value of $\tau$, let $\tau<0$. Then the first term on the right, which is just $G(\tau)$ (up to $\sqrt{2 \pi}$ ), vanishes because of causality. The integrand of the second term is bounded by $|g| \mathrm{e}^{\Im(\omega) \tau}$; because here $\Im(\omega)>0$,
the integrand vanishes at infinity when $\tau<0$, and the contour integral is zero. This establishes ${ }^{\dagger}$ that $g(\omega)$ is analytic in the upper half-plane when causality is imposed. Thus, causality implies analyticity!

Other general information about the behaviour of the response function can be derived by further postulating that $G(\tau)$ goes to 0 at least as fast as $1 / \tau$ as $\tau \rightarrow \infty$. Then, if we repeatedly integrate by parts the integral in eq. (4.25), we get, up to an irrelevant factor of $1 / \sqrt{2 \pi}$ :

$$
g(\omega)=\frac{\mathrm{i}}{\omega} G(0)-\left.\frac{1}{\omega^{2}} \mathrm{~d}_{\tau} G\right|_{\tau=0}+\ldots
$$

The first term vanishes from causality, and we find that:

$$
\begin{equation*}
\Re[g(\omega)] \underset{\omega \rightarrow \infty}{\approx} \mathrm{O}\left(1 / \omega^{2}\right) \quad \Im[g(\omega)] \underset{\omega \rightarrow \infty}{\approx} \mathrm{O}\left(1 / \omega^{3}\right) \tag{4.26}
\end{equation*}
$$

Thus, $|g(\omega)| \rightarrow 0$ as $\omega \rightarrow \infty$, and we have shown that $g$ satisfies the conditions under which Hilbert transforms may be written:

$$
\begin{equation*}
\Re[g(\omega)]=\frac{1}{\pi} \mathcal{P} \int_{-\infty}^{\infty} \frac{\Im\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime}-\omega} \mathrm{d} \omega^{\prime}, \quad \Im[g(\omega)]=-\frac{1}{\pi} \mathcal{P} \int_{-\infty}^{\infty} \frac{\Re\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime}-\omega} \mathrm{d} \omega^{\prime} \tag{4.27}
\end{equation*}
$$

With a little extra manipulation, we can rewrite these as more useful integrals over positive frequencies. To do this, we must use the fact that reality of $I(t)$ and $R(t)$, and thus of $G(t)$, demands that $g(-\omega)=g^{*}(\omega)$. Indeed, $g(\omega)$ being a complex function of a real variable, take its complex conjugate. Because $R(t)$ and $I(t)$ are real, consistency demands (EXERCISE) that $g(-\omega)=g^{*}(\omega)$. This is a version of the Schwarz reflection principle.

In terms of real and imaginary parts, we have the crossing relations:

$$
\begin{equation*}
\Re[g(-\omega)]=\Re[g(\omega)] \quad \Im[g(-\omega)]=-\Im[g(\omega)] \tag{4.28}
\end{equation*}
$$

In other words, the real part of $g$ is even, and the imaginary part odd, in their argument. Then:

$$
\begin{aligned}
\Re[g(\omega)] & =\frac{1}{\pi} \mathcal{P}\left[\int_{0}^{\infty} \frac{\Im\left[g\left(-\omega^{\prime}\right)\right]}{-\omega^{\prime}-\omega} \mathrm{d} \omega^{\prime}+\int_{0}^{\infty} \frac{\Im\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime}-\omega} \mathrm{d} \omega^{\prime}\right] \\
& =\frac{1}{\pi} \mathcal{P}\left[\int_{0}^{\infty} \frac{\Im\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime}+\omega} \mathrm{d} \omega^{\prime}+\int_{0}^{\infty} \frac{\Im\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime}-\omega} \mathrm{d} \omega^{\prime}\right] \\
& =\frac{2}{\pi} \mathcal{P} \int_{0}^{\infty} \frac{\omega^{\prime} \Im\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime 2}-\omega^{2}} \mathrm{~d} \omega^{\prime}
\end{aligned}
$$

$\Im[g(\omega)]$ can be rewritten (exercise) in the same fashion, and we end up with the dispersion relations:

$$
\begin{align*}
& \Re[g(\omega)]=\frac{2}{\pi} \mathcal{P} \int_{0}^{\infty} \frac{\omega^{\prime} \Im\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime 2}-\omega^{2}} \mathrm{~d} \omega^{\prime}  \tag{4.29}\\
& \Im[g(\omega)]=-\frac{2 \omega}{\pi} \mathcal{P} \int_{0}^{\infty} \frac{\Re\left[g\left(\omega^{\prime}\right)\right]}{\omega^{\prime 2}-\omega^{2}} \mathrm{~d} \omega^{\prime}
\end{align*}
$$

These were originally derived in 1926-7 for electrical susceptibility $\chi(\omega)$ by Kramers and Kronig.
This type of relation, which can be established with a minimum of assumptions (causality), can be very useful and exists in other areas of physics (it was often used in particle physics in the sixties, for instance).

If $\Im[g(\omega)]$ is sharply peaked around some frequency, say $\omega_{\mathrm{p}}$, we approximate it by $\Im\left[g\left(\omega^{\prime}\right)\right]=g_{\mathrm{p}} \delta\left(\omega^{\prime}-\omega_{\mathrm{p}}\right)$, in which case the integrand in the first relation is no longer singular at $\omega$, and we can drop the $\mathcal{P}$ to get:

$$
\Re[g(\omega)]=\frac{2}{\pi} g_{\mathrm{p}} \frac{\omega_{\mathrm{p}}}{\omega_{\mathrm{p}}^{2}-\omega^{2}}
$$

[^17]Therefore, in a frequency range where there is a peak in $\Im[g(\omega)], \Re[g(\omega)]$ decreases. In optics, this is known as anomalous dispersion.

On the other hand, we can ask what happens if there is a frequency range over which $\Im[g(\omega)]$ is very small. Taking $\Im[g(\omega)] \approx 0$ for $\omega_{1}<\omega<\omega_{2}$ in the first dispersion relation of eq. (4.29), there should be little contribution from that frequency range. Then the rate of change of the real part of $g(\omega)$ is:

$$
\mathrm{d}_{\omega} \Re[g(\omega)]=\frac{4 \omega}{\pi} \int_{0}^{\omega_{1}} \frac{\omega^{\prime} \Im\left[g\left(\omega^{\prime}\right)\right]}{\left(\omega^{\prime 2}-\omega^{2}\right)^{2}} \mathrm{~d} \omega^{\prime}+\frac{4 \omega}{\pi} \int_{\omega_{2}}^{\infty} \frac{\omega^{\prime} \Im\left[g\left(\omega^{\prime}\right)\right]}{\left(\omega^{\prime 2}-\omega^{2}\right)^{2}} \mathrm{~d} \omega^{\prime}
$$

If $\Im[g(\omega)]>0$, as is the case for the index of refraction $n(\omega)$ in an absorptive medium, the two terms are positive. Therefore, because of causality, the index of refraction in a medium increases with frequency, except in regions where $\Im[n(\omega)]$ (absorption) is significant! This is called normal dispersion, because it is the usual case which holds away from the absorption peaks.

We can also say something about $\Re[g(\omega)]$ in the high-frequency regime. We remember that at high frequency, eq. (4.26) says that the imaginary part of $g(\omega)$, goes like $1 / \omega^{3}$. The integral in the dispersion relation for $\Re[g(\omega)]$ is then dominated by the low-frequency behaviour of $\Im[g(\omega)]$, in the sense that $\omega^{\prime} \ll \omega$, and we can set $\omega^{\prime 2}-\omega^{2} \approx-\omega^{2}$ in the dispersion relation. We obtain a sum rule that is fully consistent with eq. (4.26):

$$
\begin{equation*}
\lim _{\omega \rightarrow \infty} \omega^{2} \Re[g(\omega)]=-\frac{2}{\pi} \int_{0}^{\infty} \omega^{\prime} \Im\left[g\left(\omega^{\prime}\right)\right] \mathrm{d} \omega^{\prime} \tag{4.30}
\end{equation*}
$$

For much more detail about the application of the principle of causality to the response of a dispersive dielectric medium to electromagnetic waves, see sections 7.5 and 7.8 of Jackson's Classical Electrodynamics.

### 4.7.3 Dispersion relations with one subtraction

When deriving Hilbert transforms, it is possible to relax the condition that $|f(z)| \rightarrow 0$ at infinity, although there will be a (fairly modest) price to pay. Suppose that instead $|f(z)|$ goes to some constant which may not be known. The function $\left(f(z)-f\left(x_{1}\right)\right) /\left(z-x_{1}\right)$ does go to zero as $|z| \rightarrow \infty$ in the upper-half plane because of its $1 / z$ dependence; and if $f(z)$ is analytic, it is also analytic. Thus, when restricted to a real argument, it satisfies eq. (4.20):

$$
\mathcal{P} \int_{-\infty}^{\infty} \frac{f\left(x^{\prime}\right)-f\left(x_{1}\right)}{\left(x^{\prime}-x\right)\left(x^{\prime}-x_{1}\right)} \mathrm{d} x^{\prime}=\mathrm{i} \pi \frac{f(x)-f\left(x_{1}\right)}{x-x_{1}}
$$

or, rearranging:

$$
\begin{aligned}
\mathrm{i} \pi\left(f(x)-f\left(x_{1}\right)\right) & =\left(x-x_{1}\right) \mathcal{P} \int_{-\infty}^{\infty} \frac{f\left(x^{\prime}\right)}{\left(x^{\prime}-x\right)\left(\left(x^{\prime}-x_{1}\right)\right.} \mathrm{d} x^{\prime}-f\left(x_{1}\right) \mathcal{P} \int_{-\infty}^{\infty} \frac{x-x_{1}}{\left(x^{\prime}-x\right)\left(\left(x^{\prime}-x_{1}\right)\right.} \mathrm{d} x^{\prime} \\
& =\left(x-x_{1}\right) \mathcal{P} \int_{-\infty}^{\infty} \frac{f\left(x^{\prime}\right)}{\left(x^{\prime}-x\right)\left(\left(x^{\prime}-x_{1}\right)\right.} \mathrm{d} x^{\prime}-f\left(x_{1}\right)\left(\mathcal{P} \int_{-\infty}^{\infty} \frac{\mathrm{d} x^{\prime}}{x^{\prime}-x}-\mathcal{P} \int_{-\infty}^{\infty} \frac{\mathrm{d} x^{\prime}}{x^{\prime}-x_{1}}\right)
\end{aligned}
$$

In section 4.5 the last two principal-value integrals were evaluated and found to vanish when the limits of integration extend to infinity. Taking the real and imaginary parts as before, we arrive at the once-subtracted Hilbert transforms:

$$
\begin{align*}
& \Re[f(x)]=\Re\left[f\left(x_{1}\right)\right]+\frac{x-x_{1}}{\pi} \mathcal{P} \int_{-\infty}^{\infty} \frac{\Im\left[f\left(x^{\prime}\right)\right]}{\left(x^{\prime}-x\right)\left(\left(x^{\prime}-x_{1}\right)\right.} \mathrm{d} x^{\prime} \\
& \Im[f(x)]=\Im\left[f\left(x_{1}\right)\right]-\frac{x-x_{1}}{\pi} \mathcal{P} \int_{-\infty}^{\infty} \frac{\Re\left[f\left(x^{\prime}\right)\right]}{\left(x^{\prime}-x\right)\left(\left(x^{\prime}-x_{1}\right)\right.} \mathrm{d} x^{\prime} \tag{4.31}
\end{align*}
$$

The integrals can be split in two and once-subtracted dispersion relations involving only positive frequencies derived, analogous to eq. (4.29). We see that we need to know, eg. in the first relation, the real part of $f$ at some point on top of its imaginary part everywhere.

### 4.7.4 Arrival of a Signal After Propagation

As another interesting physical application of these ideas, consider a plane electromagnetic wave train at normal incidence from vacuum to a medium of refraction index $n(\omega)$. Then the Fourier representation of the wave amplitude inside the medium $(x>0)$ is:

$$
\psi(x, t)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty}\left(\frac{2}{1+n(\omega)}\right) A_{\mathrm{i}}(\omega) \mathrm{e}^{\mathrm{i} \omega[n(\omega) x / c-t]} \mathrm{d} \omega \quad x>0
$$

where:

$$
A_{\mathrm{i}}(\omega)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} \psi_{\mathrm{i}}\left(0, t^{\prime}\right) \mathrm{e}^{\mathrm{i} \omega t^{\prime}} \mathrm{d} t^{\prime} \quad \quad \quad(\operatorname{acts} \text { as } i(\omega))
$$

is the Fourier transform of the real incident wave amplitude as it reaches the medium. The frequency-domain kernel function would be $g(\omega)=2 \mathrm{e}^{\mathrm{i} \omega n(\omega) x / c} /(1+n(\omega))$.

The integral for $\psi$ can be evaluated by contour integration in the complex $\omega$ plane. The integral enclosing the whole upper half-plane vanishes: the integrand is analytic there, because electromagnetic considerations show that both $A(\omega)$ and $n(\omega)$ are analytic there. Also from electromagnetism, when $|\omega| \rightarrow \infty, n(\omega) \rightarrow 1$ and the argument of the exponential becomes $i \omega[x-c t] / c$, so that the contribution from the semi-circle at infinity also vanishes if $x>c t$. Then the contribution along the real $\omega$-axis must also vanish, and there is no wave amplitude for $x>c t$. This shows, without any detailed knowledge of $n(\omega)$, that no signal can propagate in any medium faster than $c$. We can say that analyticity implies causality (see also the more sophisticated argument at the end of BF 6.6).

### 4.8 Bessel Functions (BF 6.9)

The function $\mathrm{e}^{(w-1 / w) z / 2}$, which is analytic $\forall w \in \mathbb{C}$ except at $w=0$, can be expanded in a Laurent series for any $|w| \neq 0$ in a region centered on, but not including, $w=0$ :

$$
\mathrm{e}^{(w-1 / w) z / 2}=\sum_{n=-\infty}^{\infty} J_{n}(z) w^{n}
$$

where, choosing the unit $w$-circle as contour and making the change of variable $w=\mathrm{e}^{\mathrm{i} \theta}$ :

$$
\begin{align*}
J_{n}(z)=\frac{1}{2 \pi \mathrm{i}} \oint_{|w|=1} \frac{\mathrm{e}^{(w-1 / w) z / 2}}{w^{n+1}} \mathrm{~d} w & =\frac{1}{2 \pi} \int_{-\pi}^{\pi} \mathrm{e}^{\mathrm{i}(z \sin \theta-n \theta)} \mathrm{d} \theta \\
& =\frac{1}{\pi} \int_{0}^{\pi} \cos (n \theta-z \sin \theta) \mathrm{d} \theta \tag{4.32}
\end{align*}
$$

This is an integral representation for the $n^{\text {th }}$ order Bessel function of the first kind $J_{n}(z)(n \in \mathbb{Z})$, which satisfies (see pp. BF372-373) the Bessel equation:

$$
\begin{equation*}
z^{2} \mathrm{~d}_{z}^{2} J_{n}(z)+z \mathrm{~d}_{z} J_{n}(z)+\left(z^{2}-n^{2}\right) J_{n}(z)=0 \tag{4.33}
\end{equation*}
$$

$\mathrm{e}^{(w-1 / w) z / 2}$ is said to be a generating function for the $J_{n}(z)$; these are analytic at $z=0$.
Bessel functions arise in the theory of diffraction, in the description of electromagnetic modes inside cavities, and also as solutions of Laplace's equation in cylindrical coordinates.

### 4.9 Calculus of Residues and Some Applications (BF 6.8)

### 4.9.1 Theorem of Residues

We know from the Cauchy-Goursat theorem that the contour integral of a function analytic on and inside a contour $C$ vanishes. We now show how the integral can be evaluated even if there is a finite number of singularities within the contour.

The way to do this is to choose another contour within which the function is analytic, and therefore over which its integral still vanishes. This time the path leaves the original contour at a point closest to a singular point, goes straight to within a distance of this point, goes all the way around the point clockwise on a circle, and finally goes back to the original contour on the same interconnecting line, but in the opposite direction. The same procedure is repeated for each singularity. It should be obvious that the second path is the sum of the original contour, the incoming and outgoing paths to each singular point which generate cancelling contributions, and the sum of all the circular paths around each of the singular points. Since the original contour is traversed counterclockwise by convention, we have the Residue Theorem for $n$ singularities:

$$
\begin{equation*}
\oint_{C} f(z) \mathrm{d} z=2 \pi \mathrm{i} \sum_{i=1}^{n} \operatorname{Res}\left[f\left(z_{i}\right)\right] \quad \operatorname{Res}\left[f\left(z_{i}\right)\right]=\frac{1}{2 \pi \mathrm{i}} \oint_{C_{i}} f(z) \mathrm{d} z \tag{4.34}
\end{equation*}
$$

where each $C_{i}$ is the circle around one, and only one, singular point, which is now traversed counterclockwise, and $\operatorname{Res}\left[f\left(z_{i}\right)\right]$ is called the residue at the singular point $z_{i}$. It should be clear from eq. (4.17) that a residue is just the $A_{-1}$ coefficient of a Laurent series expansion of $f(z)$.

The Cauchy-Gauss integral formula (4.14) is often useful to evaluate residues, for instance, if the integrand in the residue integral only has simple (order 1) poles:

$$
f(z)=\frac{g(z)}{\left(z-z_{1}\right)\left(z-z_{2}\right) \ldots\left(z-z_{n}\right)}
$$

Then we simply write:

$$
\oint_{C_{i}} \frac{g(z) \mathrm{d} z}{\left(z-z_{1}\right)\left(z-z_{2}\right) \ldots\left(z-z_{n}\right)}=\oint_{C_{i}} \frac{g(z)}{\left(z-z_{1}\right) \cdots\left(z-z_{i-1}\right)\left(z-z_{i+1}\right) \cdots\left(z-z_{n}\right)} \frac{1}{z-z_{i}} \mathrm{~d} z
$$

The sum of all the residues is just:

$$
\sum \frac{1}{2 \pi \mathrm{i}} \oint_{C_{i}} f(z) \mathrm{d} z=\frac{g\left(z_{1}\right)}{\left(z_{1}-z_{2}\right) \ldots\left(z_{1}-z_{n}\right)}+\ldots+\frac{g\left(z_{n}\right)}{\left(z_{n}-z_{1}\right)\left(z_{n}-z_{2}\right) \ldots\left(z_{n}-z_{n-1}\right)}
$$

If there are poles of higher order, we can always use eq. (4.15) to express the residue in terms of derivatives, eg.:

$$
\frac{1}{2 \pi \mathrm{i}} \oint \frac{g(z)}{\left(z-z_{0}\right)^{n}} \mathrm{~d} z=\left.\frac{1}{(n-1)!} \mathrm{d}_{z}^{n-1} g(z)\right|_{z_{0}}
$$

Another technique, illustrated by example 4.6, is to expand the integrand in a Laurent series. The contour integral of the $n \geq 0$ terms vanishes because of analyticity, and the remaining negative powers integrate to:

$$
\frac{1}{2 \pi \mathrm{i}} \oint_{C_{0}} f(z) \mathrm{d} z=\frac{1}{2 \pi \mathrm{i}} \oint_{C_{0}} \sum_{0}^{\infty} \frac{A_{-n}}{\left(z-z_{0}\right)^{n}} \mathrm{~d} z=\left.\sum_{0}^{\infty} \frac{1}{(n-1)!} \mathrm{d}_{z}^{n-1} A_{-n}\right|_{z_{0}}=A_{-1}
$$

### 4.9.2 Real definite integrals

Many applications of the calculus of residues involve the evaluation of real definite integrals, which we would like to "complexify" so as to use contour integration. First, we establish a very useful result, generally referred to as Jordan's lemma:

Let $f(z)$ be a function whose absolute value is bounded in the upper half-plane, with that bound going uniformly to 0 as $|z| \rightarrow \infty$ for $0 \leq$ phase $(z) \leq \pi$; also, let $\alpha>0$. Then, on a semi-circle of radius $R$ centered on the origin, we have:

$$
\begin{equation*}
\lim _{R \rightarrow \infty} \int_{\substack{|z|=R \\ y>0}} f(z) \mathrm{e}^{\mathrm{i} \alpha z} \mathrm{~d} z=0 \tag{4.35}
\end{equation*}
$$

Indeed, in the polar representation with $z=R \mathrm{e}^{\mathrm{i} \theta}$, the absolute value of the integral is bounded:

$$
\left|\int_{\left\lvert\, \begin{array}{l}
|z|=R \\
y>0 \\
\hline
\end{array}\right.} f(z) \mathrm{e}^{\mathrm{i} \alpha z} \mathrm{~d} z\right| \leq 2|f(R)|_{\max } \int_{0}^{\pi / 2} R \mathrm{e}^{-\alpha R \sin \theta} \mathrm{~d} \theta \leq 2|f(R)|_{\max } \int_{0}^{\pi / 2} \mathrm{e}^{-2 \alpha R \theta / \pi} \mathrm{d}(R \theta)
$$

where the second inequality is justified by the fact that $\sin \theta \geq 2 \theta / \pi$ over the interval of integration ( $\sin \theta-2 \theta / \pi=$ 0 at the end-points, is positive at $\pi / 4$, and has no minimum in the interval, so is positive or zero everywhere). When $R \rightarrow \infty$, the integral on the right converges to a number and, because $|f(z)|_{\max } \rightarrow 0$ at infinity, the bound on the absolute value of the integral on the left-hand side also vanishes there. Actually, Maple easily evaluates the middle integral as $\alpha^{-1}$ in the limit $R \rightarrow \infty$. With $a=\alpha R$ :

$$
\begin{gathered}
\quad>\operatorname{limit}(a * i n t(\exp (-a * \sin (\text { theta) }), \text { theta=0..Pi/2), a=infinity)/alpha assuming } \\
a>0 ; \\
1 / \alpha
\end{gathered}
$$

Whichever way, we obtain the result sought.
Under the same condition on $|f(z)|$ in the lower-half plane, Jordan's lemma holds for $\alpha<0$, since $\sin \theta<0$ in that region, provided we use a clockwise contour. From the above proof, the lemma also holds when $\alpha=0$.
Integrals of the form: $\int_{-\infty}^{\infty} R(x) \mathrm{e}^{\mathrm{i} \alpha x} \mathrm{~d} x$

- We consider first the case where the rational function $R(x)$ has no pole on the real axis. The Residue theorem allows us to evaluate the contour integral in the upper half of the complex plane, $y>0$ :

$$
\oint_{C} R(z) \mathrm{e}^{\mathrm{i} \alpha z} \mathrm{~d} z=\int_{-\infty}^{\infty} R(x) \mathrm{e}^{\mathrm{i} \alpha x} \mathrm{~d} x+\lim _{\rho \rightarrow \infty} \int_{\substack{|z|=\rho \\ y>0}} R(z) \mathrm{e}^{\mathrm{i} \alpha z} \mathrm{~d} z=2 \pi \mathrm{i} \sum_{y>0} \operatorname{Res}\left[R(z) \mathrm{e}^{\mathrm{i} \alpha z}\right]
$$

Provided that $\alpha \geq 0$ and that $|R(z)| \rightarrow 0$ uniformly in $\theta$ as $\rho \rightarrow \infty$, the integral over the semi-circle at infinity in the upper half-plane must vanish because of Jordan's lemma.
[If $\alpha<0$, simply take instead a clockwise contour in the lower half-plane $(y<0)$, and integrate from $2 \pi$ to $\pi$ so that $\sin \theta<0$ in the exponential.]
The result is then the following:

$$
\int_{-\infty}^{\infty} R(x) \mathrm{e}^{\mathrm{i} \alpha x} \mathrm{~d} x= \begin{cases}2 \pi \mathrm{i} \sum_{y>0} \operatorname{Res}\left[R(z) \mathrm{e}^{\mathrm{i} \alpha z}\right] & \alpha>0  \tag{4.36}\\ -2 \pi \mathrm{i} \sum_{y<0} \operatorname{Res}\left[R(z) \mathrm{e}^{\mathrm{i} \alpha z}\right] & \alpha<0\end{cases}
$$

- Now we let $\alpha=0$ and take some function $Q(x)$ with simple poles and $|Q(z)| \rightarrow 0$ as $|z| \rightarrow \infty$. As we did in our treatment of Hilbert transforms, we take a contour consisting of a semi-circle of infinite radius based on the real axis. We avoid the simple poles on the real axis by moving around them clockwise on a small half-circle of radius $\rho$ in the upper half-plane $y>0$. Under our assumptions, the contribution from the semi-circle at infinity vanishes. If on the real axis $Q(z)$ has only one simple pole at $x=a$ :

$$
\mathcal{P} \int_{-\infty}^{\infty} Q(x) \mathrm{d} x+\lim _{\rho \rightarrow 0} \int_{\pi}^{0} \mathrm{i} Q\left(a+\rho \mathrm{e}^{\mathrm{i} \theta}\right) \rho \mathrm{e}^{\mathrm{i} \theta} \mathrm{~d} \theta=2 \pi \mathrm{i} \sum_{y>0} \operatorname{Res} Q(z)
$$

using the Residue theorem. To make further progress, we first establish another useful result:
If $f(z)$ has a simple pole at $z=a$, and $C_{\rho}$ is a counterclockwise arc of radius $\rho$ centered on $a$ and intercepting an angle $\alpha$, then:

$$
\begin{equation*}
\lim _{\rho \rightarrow 0} \int_{C_{\rho}} f(z) \mathrm{d} z=\left.\mathrm{i} \alpha \operatorname{Res} f(z)\right|_{z=a} \tag{4.37}
\end{equation*}
$$

To show this, note that because $f(z)$ must be of the form $g(z) /(z-a)$ :

$$
\int_{C_{\rho}} \frac{g(z)}{z-a} \mathrm{~d} z=\int_{0}^{\alpha} \mathrm{i} \frac{g\left(a+\rho \mathrm{e}^{\mathrm{i} \theta}\right)}{\rho \mathrm{e}^{\mathrm{i} \theta}} \rho \mathrm{e}^{\mathrm{i} \theta} \mathrm{~d} \theta=\int_{0}^{\alpha} \mathrm{i} g\left(a+\rho \mathrm{e}^{\mathrm{i} \theta}\right) \mathrm{d} \theta
$$

Since $g(z)$ is analytic at $a$, it can be Taylor-expanded around $a$, so that $g\left(a+\rho \mathrm{e}^{\mathrm{i} \theta}\right)=g(a)+$ terms proportional to positive powers of $\rho$ that do not contribute in the limit $\rho \rightarrow 0$. We are left with the result sought:

$$
\lim _{\rho \rightarrow 0} \int_{C_{\rho}} f(z) \mathrm{d} z=\mathrm{i} \alpha g(a)=\left.\mathrm{i} \alpha \operatorname{Res} f(z)\right|_{z=a}
$$

With $\alpha=-\pi$ (semi-circle traversed clockwise!), and for a finite number of simple poles on the real axis $y=0$,

$$
\begin{equation*}
\mathcal{P} \int_{-\infty}^{\infty} Q(x) \mathrm{d} x=2 \pi \mathrm{i} \sum_{y>0} \operatorname{Res} Q(z)+\mathrm{i} \pi \sum_{y=0} \operatorname{Res} Q(z) \tag{4.38}
\end{equation*}
$$

Integrals of the form: $\int_{0}^{\infty} x^{\lambda-1} R(x) \mathrm{d} x,(\lambda \notin \mathbb{Z})$
Here, $R(z)$ must be rational, ie. of the form $P_{n}(x) / Q_{m}(x), P$ and $Q$ being polynomials of degree $n$ and $m$, respectively. $R(z)$ must be analytic at the origin with no poles on the positive real axis. Then $Q_{m}(x)$ should have no zero or positive real root. As well, $\left|z^{\lambda} R(z)\right| \rightarrow 0$ both when $|z| \rightarrow \infty$ and $|z| \rightarrow 0$. These last conditions are satisfied if $\lambda>0$ and if $\lambda+n<m$, which means that $m \geq 1$. Then $R(z)$ is guranteed to have poles somewhere (other than zero) in the complex plane.

Since $\lambda$ is not an integer, we expect a branch cut in $z^{\lambda-1}$ that starts at the origin and which we take to be along the real axis. Then we only consider the branch:

$$
z^{\lambda-1}=|z|^{\lambda-1} \mathrm{e}^{\mathrm{i}(\lambda-1) \theta} \quad 0<\theta<2 \pi
$$

It means that, at an arbitrarily small but finite distance above the positive real axis, $z^{\lambda-1}=x^{\lambda-1}$, and $z^{\lambda-1}=$ $x^{\lambda-1} \mathrm{e}^{2 \pi \mathrm{i}(\lambda-1)}$ below (infinitesimally close to) the positive real axis.

We take a circular contour C at infinity centered on the origin, that runs back to the origin just below the positive real axis, encircles the origin at infinitesimally small radius, and goes back to infinity just above the positive real axis. The contributions to the contour integral from the two circles vanish, leaving those from the two sides of the cut. So, from the Residue theorem,

$$
\begin{aligned}
\int_{C} z^{\lambda-1} R(z) \mathrm{d} z & =\int_{\infty}^{0} x^{\lambda-1} \mathrm{e}^{2 \pi \mathrm{i}(\lambda-1)} R(x) \mathrm{d} x+\int_{0}^{\infty} x^{\lambda-1} R(x) \mathrm{d} x=-\frac{2 \mathrm{i} \sin \lambda \pi}{\mathrm{e}^{-\mathrm{i} \pi \lambda}} \int_{0}^{\infty} x^{\lambda-1} R(x) \mathrm{d} x \\
& =2 \pi \mathrm{i} \sum_{\text {inside } C} \operatorname{Res}\left[z^{\lambda-1} R(z)\right]
\end{aligned}
$$

which can be rearranged to:

$$
\int_{0}^{\infty} x^{\lambda-1} R(x) \mathrm{d} x=(-1)^{\lambda-1} \frac{\pi}{\sin \lambda \pi} \sum_{\text {inside } C} \operatorname{Res}\left[z^{\lambda-1} R(z)\right]
$$

## 5 MODULE V - Approximation of Arbitrary Functions with Special Functions

There are some things that we do all the time in physics without quite realising how non-trivial they are. For instance, we often expand functions over infinite series of known functions. We write solutions of the Schrödinger equation as a linear combination of an often infinite number of eigenstates; or we expand in Fourier series. When the series have a finite number of terms, there is normally no problem. When they are infinite, however, we must answer the following question: Given some function $h(x)$ defined over $[a, b]$, under what conditions is it possible to express $h(x)$ as a linear combination of some possibly infinite set of functions defined over the same interval. Since such linear combinations are in general infinite series, we want to know under what conditions on $h(x)$ and on the set $\left\{f_{n}(x)\right\}$ a series $\sum_{n}^{\infty} c_{n} f_{n}(x)$ :

- converges $\forall x \in[a, b]$;
- converges to $h(x)$ (almost) everywhere in $[a, b]$.

Once we have found what kind of functions admit this approximation, we shall have to worry about what happens when we take their derivative, and we will be led to a special kind of second-order differential operator with rich applications.

### 5.1 Hilbert Space Without Tears (BF 5.1)

### 5.1.1 Complete, normed vector spaces with inner product

Consider a vector space $\mathcal{V}$, ie. a set which is closed under addition and multiplication by a scalar:

$$
\begin{equation*}
\alpha f_{1}+f_{2} \in \mathcal{V} \quad \forall f_{1}, f_{2} \in \mathcal{V}, \quad \alpha \in \mathbb{C} \tag{5.1}
\end{equation*}
$$

Definition 5.1. $\mathcal{V}$ is said to be normed if, $\forall f \in \mathcal{V}$ there exists a real number, the norm $\|f\|$ of $f$, such that: $\|f\| \geq 0$, with $\|f\|=0$ only if $f=0$ "almost everywhere"; $\|\alpha f\|=|\alpha|\|f\| \forall \alpha \in \mathbb{C}$; and $\|f+g\| \leq\|f\|+\|g\|$.
Definition 5.2. Let $\left\{h_{n}\right\}$ be a sequence in a normed space. If, for any arbitrarily small positive number $\epsilon$, there is an integer $N(\epsilon)$ such that $\left\|h_{j}-h_{k}\right\|<\epsilon$ for $j>N, k>N$, we say that $\left\{h_{n}\right\}$ is a Cauchy sequence.
In an alternative and equivalent definition, the sequence is Cauchy when $\left\|h_{j}-h_{k}\right\| \rightarrow 0$ in the limit $j, k \rightarrow \infty$.

Definition 5.3. A vector space is said to be complete if there exists no Cauchy sequence of any of its elements which converges to an element which is not in the space.
The space of real numbers is complete in that sense, but the space of rational numbers is not, since, for instance, the (Cauchy) sequence $\left\{s_{N}\right\}$ of partial sums $s_{N}=\sum_{0}^{N} 1 / n$ ! converges to $e$, which is not rational.

Quite independently, we can also endow our vector space with a general inner product (or complex sesquilinear form, as mathematicians sometimes call it) that has the following properties:

1. $\left(f_{1}, f_{2}\right)$ is a complex number.
2. $(f, f)$ is a real number which is either zero or positive.
3. $\left(f_{1}, f_{2}\right)=\left(f_{2}, f_{1}\right)^{*}$.
4. $\left(f_{1}+f_{2}, \alpha f_{3}\right)=\alpha\left(f_{1}, f_{3}\right)+\alpha\left(f_{2}, f_{3}\right)$, where $\alpha$ is a complex scalar.
5. $\left(\alpha f_{1}, f_{2}\right)=\left(f_{2}, \alpha f_{1}\right)^{*}=\alpha^{*}\left(f_{2}, f_{1}\right)^{*}=\alpha^{*}\left(f_{1}, f_{2}\right)$.

Properties (4) and (5) follow the convention, widely used in physics but not in mathematics, that the inner product is linear in its second argument and antilinear (or conjugate-linear) in its first argument.

### 5.1.2 Banach and Hilbert spaces of square-integrable functions

Definition 5.4. A vector space which is normed and complete is called a Banach space. If, in addition, it is endowed with an inner product, it is also a Hilbert space $\mathcal{H}$.

Any finite-dimensional vector space is a Banach space. In general, this is not the case for the infinite-dimensional spaces which are inevitably involved in the infinite series that we hope to use to approximate functions.

Consider the set $\{f(x)\}$ of all complex-valued functions of a real variable defined on a closed interval $[a, b]$ and which are square-integrable, in the sense that $\int_{a}^{b}|f(x)|^{2} \mathrm{~d} x$ is finite. Then we take $(f, f)$ as the norm of $f$, and we choose as the inner product :

$$
\begin{equation*}
\left(f_{1}, f_{2}\right)=\int_{a}^{b} f_{1}^{*}(x) f_{2}(x) w(x) \mathrm{d} x \tag{5.2}
\end{equation*}
$$

where the real weight function $w(x) \geq 0$ everywhere in $[a, b]$.
A theorem by Riesz and Fischer guarantees that the space of square-integrable functions is complete, and therefore a Hilbert (and Banach) space.

There is a technicality which we should not gloss over. For eq. (5.2) to be an inner product such that $(f, f)$ is identified with the norm $\|f\|,(f, f)=0$ should imply that $f$ itself vanishes everywhere in $[a, b]$. But with our choice of inner product, $f$ is allowed not to vanish at a finite number of points (or, more generally, a countably infinite one), and $(f, f)$ will still vanish. If this happens, the set of points at which $f$ does not vanish is said to be a set of measure zero. Integrals involving sets of measure zero are usually handled by Lebesgue integration, as opposed to the usual Riemann kind. The long and the short is that if $f$ is zero "almost everywhere" in $[a, b]$ (but not everywhere), we can still have $(f, f)=0$.

With $f$ and $g$ square-integrable functions, it is straightforward to show the identity:

$$
|(f, g)|^{2} \equiv(f, f)(g, g)-\frac{1}{2} \iint|f(x) g(y)-f(y) g(x)|^{2} \mathrm{~d} x \mathrm{~d} y
$$

from which the Schwartz inequality immediately follows:

$$
\begin{equation*}
|(f, g)|^{2} \leq(f, f)(g, g) \tag{5.3}
\end{equation*}
$$

This makes it plain that square-integrability of two functions guarantees that their inner product is finite.

### 5.2 Orthonormality and Completeness of a Set of Functions (BF 5.2)

We will be especially concerned with the set of functions $f_{i} \in \mathcal{H}$ orthonormal with respect to a positive weight function $w(x)$ on $[a, b]$.

Definition 5.5. When

$$
\left(f_{i}, f_{j}\right)=\int_{a}^{b} f_{i}^{*}(x) f_{j}(x) w(x) \mathrm{d} x=\delta_{i j}
$$

we say that the $f_{i}$ are orthonormal with respect to $w(x)$ on $[a, b]$.
Now suppose there exists such an infinite orthonormal set $\left\{f_{j}(x)\right\}$, with $j=1,2, \ldots$ Pick some arbitrary $h(x) \in \mathcal{H}$ over $[a, b]$, and compute a finite number of inner products $c_{j}=\left(f_{j}, h\right)$. Then construct the sequence of partial sums:

$$
h_{n}(x)=\sum_{j}^{n} c_{j} f_{j}(x) \quad n=1,2, \ldots
$$

Then, for any given $h_{n}$, and using the orthonormality of the $f_{j}$, we can immediately write:

$$
\left(h, h_{n}\right)=\sum_{j}^{n}\left|c_{j}\right|^{2}, \quad\left(h_{n}, h_{n}\right)=\sum_{j}^{n}\left|c_{j}\right|^{2}
$$

while, from the Schwartz inequality, eq. (5.3), we have:

$$
\left(h_{n}, h_{n}\right)^{2}=\left(h_{n}, h\right)^{2} \leq(h, h)\left(h_{n}, h_{n}\right)
$$

from which follows the inequality: $s_{n}=\sum_{j}^{n}\left|c_{j}\right|^{2}=\left(h_{n}, h_{n}\right) \leq(h, h)$.
The sequence of $s_{n}$ increases monotonically with $n$, always remaining bounded, yielding Bessel's inequality:

$$
\begin{equation*}
(h, h) \geq \lim _{n \rightarrow \infty}\left(h_{n}, h_{n}\right)=\sum_{j=1}^{\infty}\left|\left(f_{j}, h\right)\right|^{2} \tag{5.4}
\end{equation*}
$$

That the sum on the right converges in the limit $n \rightarrow \infty$ is important, because it means that our seemingly ad hoc initial construction, the sequence of partial sums $h_{n}$, has finite norm and thus also converges in the limit $n \rightarrow \infty$. The question is: to what?

Ask a slightly different question: given a square-integrable function $h(x)$ over the interval $[a, b]$, we would like to "fit" it with a linear combination of our orthonormal $f_{j}$, which are also square-integrable. The best fit is obtained by minimising the mean quadratic error:

$$
\begin{equation*}
M=\lim _{n \rightarrow \infty}\left(h-\sum_{j}^{n} a_{j} f_{j}, h-\sum_{j}^{n} a_{j} f_{j}\right)=(h, h)-\sum_{j}^{\infty}\left|\left(f_{j}, h\right)\right|^{2}+\sum_{j}^{\infty}\left|a_{j}-\left(f_{j}, h\right)\right|^{2} \tag{5.5}
\end{equation*}
$$

By definition, $M \geq 0$, and we already know from Bessel's inequality that the sum of the first two terms on the right is also positive. We conclude that we can minimise $M$ by setting the coefficients $a_{j}$ equal to the inner products $c_{j}=\left(f_{j}, h\right)$ introduced previously.

Now, is it possible for the mean quadratic error to vanish? Before answering this, let us recall a couple of definitions for different types of convergence of series.

Consider the sequence of partial sums:

$$
h_{n}(x)=\sum_{i}^{n} k_{i}(x) \quad n=1,2, \ldots
$$

Definition 5.6. If, for any arbitrarily small positive number $\epsilon$, we can find an integer $n=N(\epsilon)$ for which $\left|h(x)-h_{n}(x)\right|<\epsilon$ anywhere in $[a, b]$ when $n>N$, then the sequence of partial sums $h_{n}(x)$ converges uniformly toward the value $h(x)$ :

$$
h(x)=\lim _{n \rightarrow \infty} h_{n}(x)
$$

If $N$ also depends on $x$, we have a weaker point-wise convergence.
It will be convenient to relax this criterion slightly by allowing it not to hold at a countable number of points in $[a, b]$. Then we say that we have convergence in the mean if:

$$
\begin{equation*}
\lim _{n \rightarrow \infty}\left(h-h_{n}, h-h_{n}\right)=0 \tag{5.6}
\end{equation*}
$$

Uniform convergence would force $h(x)=\lim h_{n}(x)$ everywhere in the interval. If this fails to occur at only a countable set of points, this set of points is of measure zero and contributes nothing to the integral in the inner product, so long as we treat the integral as a Lebesgue integral.

Definition 5.7. For a given set $\left\{f_{i}\right\}$ orthonormal on $[a, b]$, if it is possible by letting $n \rightarrow \infty$ to make the mean quadratic error vanish, ie. if the sequence $\left\{h_{n}\right\}$ with $h_{n}(x)=\sum c_{i} f_{i}(x)$ converges in the mean to $h(x)$, then the $f_{i}$ form a complete orthonormal set.

Note that we wrote the coefficients as $c_{i}$, not $c_{i n}$, because they do not change as we construct partial sums of higher $n$ to improve the approximation.

When the $f_{i}$ form a complete orthonormal set, the vanishing of the mean quadratic error (eq. (5.5)) as $n \rightarrow \infty$ transforms Bessel's inequality into the important completeness relation (Parseval equality):

$$
\begin{equation*}
(h, h)=\sum_{j}^{\infty}\left|c_{j}\right|^{2}=\sum_{j}^{\infty}\left|\left(f_{j}, h\right)\right|^{2} \tag{5.7}
\end{equation*}
$$

which holds for any square-integrable function $h(x)$.
A more intuitive-and certainly more useful than Definition 5.7-criterion for completeness is provided by the following theorem (theorem BF5.2 on p. 222) which we quote without proof:

A set of orthonormal functions in Hilbert space is complete if, and only if, there exists no function outside the set that is orthogonal to every function, in which case the set is said to be closed.

Essentially, what we are saying here is that a complete, orthonormal set of functions forms a basis for the Hilbert space, in the sense that any function in that space can be expressed as a (possibly infinite) sum over the basis functions.

Note that completeness of $\left\{f_{i}\right\}$ does not allow us to assert that an arbitrary function can be expressed as a linear combination of the $f_{i}$ everywhere in $[a, b]$; for this to be true, we need to establish uniform convergence. If only convergence in the mean can be proved, there will in general be a countable set of points in $[a, b]$ where we cannot say that $h=\sum c_{i} f_{i}$.

### 5.3 Approximation by Polynomials: the Theorem of Weierstrass (BF5.4)

According to the Theorem of Weierstrass (1885), for any function $f(q)$ continuous over $[a, b]$, there must exist a sequence of polynomials, $P_{N}(q)$, that converges uniformly to $f(q) \forall q \in[a, b]$ :

$$
\begin{equation*}
f(q)=\lim _{N \rightarrow \infty} P_{N}(q) \tag{5.8}
\end{equation*}
$$

where:

$$
P_{N}(q)=\sum_{i=0}^{N} a_{i}^{(N)} q^{i}
$$

Proofs of the theorem by "elementary" methods (see BF) are rather tedious, and more sophisticated proofs-the so-called Stone-Weierstrass Theorem (1937) - require concepts far beyond the scope of this course. One thing, however, should be clear from the outset about what the theorem asserts. Since the coefficients $a_{i}^{(N)}$ depend on $N$, we should not think of this representation of $f(q)$ as a power-series expansion, in which the coefficients would not change as more powers are included. As we improve the approximation with higher-order polynomials, coefficients of a given power of $q$ will change!

The theorem may be extended so as to approximate functions of $m$ variables :

$$
f\left(q_{1}, q_{2}, \ldots, q_{m}\right)=\lim _{N \rightarrow \infty} \sum_{i_{1}, \ldots, i_{m}=0}^{N} a_{i_{1}, \ldots i_{m}}^{(N)} q_{1}^{i_{1}} \ldots q_{m}^{i_{m}}
$$

Although the Theorem of Weierstrass is somewhat unwieldy in actual computations, it can help us discover various complete sets of polynomials over suitable closed 1 -dim intervals, as well as other sets of complete functions.

Let us illustrate this with one very important case of approximation in terms of trigonometric functions, the Fourier approximation.

### 5.4 Fourier Series (BF 5.6)

### 5.4.1 Fourier theorem

We look at continuous functions $g(x, y)$ defined in two dimensions, for which the Weierstrass theorem guarantees the existence of functions:

$$
g_{N}(x, y)=\sum_{j, k=0}^{N} a_{j k}^{(N)} x^{j} y^{k}
$$

which approximate uniformly $g(x, y)$ over a finite interval in the $N \rightarrow \infty$ limit. Since we are really interested in a 1-dim case, we restrict ourselves to functions $g(x, y)=f(\theta)$ which live on the unit circle and we go to polar coordinates with $x=\cos \theta, y=\sin \theta$. Expressing the cosine and sine in terms of exponentials, the theorem of Weierstrass then requires that $f(x)$ can be written in the form:

$$
\begin{aligned}
f(x) & =\lim _{N \rightarrow \infty} \frac{1}{\sqrt{2 \pi}} \sum_{-N}^{N} c_{n}^{(N)} \mathrm{e}^{\mathrm{i} n x} \\
& =\lim _{N \rightarrow \infty}\left[\frac{a_{0}^{(N)}}{2}+\sum_{n=1}^{N}\left(a_{n}^{(N)} \cos n x+b_{n}^{(N)} \sin n x\right)\right]
\end{aligned}
$$

where $f(x+2 \pi)=f(x)$. Actually, $f(x)$ need not be periodic so long as we are interested in approximating it only in the interval $[-\pi, \pi]$.

Though uniform, this approximation is not quite satisfactory since the coefficients depend on $N$. But it points to the set of functions $\mathrm{e}^{\mathrm{i} n x} / \sqrt{2 \pi}$ ( $n$ any integer) as being worthy of investigation. All that has to be done now is show that the set is orthonormal and complete. Indeed, one easily shows that the set is orthonormal on $[-\pi, \pi]$; moreover, it is complete (p. BF241). The equivalent set $\{1 / \sqrt{2 \pi},(\cos n x) / \sqrt{\pi},(\sin n x) / \sqrt{\pi}\}$ is also orthonormal and complete.

It follows immediately that an arbitrary function can be approximated in the mean over $[-\pi, \pi]$ by the expansions:

$$
\frac{1}{\sqrt{2 \pi}} \sum_{-\infty}^{\infty} c_{n} \mathrm{e}^{\mathrm{i} n x}=\frac{a_{0}}{2}+\sum_{n=1}^{\infty}\left(a_{n} \cos n x+b_{n} \sin n x\right)
$$

The convergence becomes uniform (theorems BF5.4 and $5.5 \mathrm{on} \mathrm{pp.242-245)} \mathrm{if} \mathrm{the} \mathrm{first} \mathrm{derivative} \mathrm{of} \mathrm{the} \mathrm{function}$ to be approximated is piecewise continuous over $[-\pi, \pi]$. Any point of discontinuity of the function requires special treatment.

A change of variable $x \rightarrow \pi x / l$ generalises BF's theorem 5.5 to any interval $[c, c+2 l]$, and we can now state the Fourier-series theorem in full:

Let $f(x)$ and its first derivative be piecewise continuous on a closed interval $[c, c+2 l]$, where $c \in \mathbb{R}$. Then in every closed subinterval over which $f(x)$ is continuous, we have uniform convergence of the sequence of partial sums to $f(x)$ :

$$
\begin{equation*}
f(x)=\frac{1}{\sqrt{2 l}} \sum_{-\infty}^{\infty} c_{n} \mathrm{e}^{\mathrm{i} n \pi x / l}=\frac{a_{0}}{2}+\sum_{n=1}^{\infty}\left(a_{n} \cos \frac{n \pi x}{l}+b_{n} \sin \frac{n \pi x}{l}\right) \tag{5.9}
\end{equation*}
$$

where $\left\{\mathrm{e}^{\mathrm{i} n \pi x / l} / \sqrt{2 l}\right\}$ and $\{1 / \sqrt{2 \pi},(\cos n x) / \sqrt{l},(\sin n x) / \sqrt{l}\}$ are complete, orthonormal sets over $[c, c+2 l]$. The Fourier coefficients are given by:

$$
\begin{align*}
& c_{n}=\frac{1}{\sqrt{2 l}} \int_{c}^{c+2 l} f(x) \mathrm{e}^{-\mathrm{i} n \pi x / l} \mathrm{~d} x \\
& a_{n}=\frac{1}{l} \int_{c}^{c+2 l} f(x) \cos \frac{n \pi x}{l} \mathrm{~d} x  \tag{5.10}\\
& b_{n}=\frac{1}{l} \int_{c}^{c+2 l} f(x) \sin \frac{n \pi x}{l} \mathrm{~d} x
\end{align*}
$$

At points of discontinuity of $f(x)$, its Fourier series converges to the average of the two limits, from the right and from the left, of the function.

If, and only if, $f(x+2 l)=f(x)$, ie. if $f$ is periodic, are these expressions for $f$ valid outside $[c, c+2 l]$. If $f$ is non-periodic, we can still use the Fourier theorem in an open, finite interval whose length we identify with $2 l$; outside this interval, we simply replace the actual $f$ by an infinite succession of replicas of $f$ over the interval so as to create a function of period $2 l$.

The Fourier coefficients $c_{n}$, and $\left\{a_{n}, b_{n}\right\}$ satisfy the completeness relation (5.7):

$$
(f, f)=\sum_{-\infty}^{\infty}\left|c_{n}\right|^{2}=l\left[\frac{a_{0}^{2}}{2}+\sum_{n=1}^{\infty}\left(a_{n}^{2}+b_{n}^{2}\right)\right]
$$

The Fourier theorem also applies to periodic analytic functions $f(z)$.

### 5.4.2 Gibbs phenomenon

Note the emphasis on a closed interval in the Fourier theorem. If the (sub)interval should be open because $f(x)$ is discontinuous at its end-points, the sequence of partial (truncated) Fourier series converges point-wise, but not uniformly, and truncated series may not approximate a function correctly at points that depend on how many terms are kept. To see this, rewrite the exponential Fourier series from eq. (5.9) as a truncated series:

$$
\begin{aligned}
& f_{N}(x)=\frac{1}{2 l} \int_{0}^{2 l} \mathrm{~d} x^{\prime} f\left(x^{\prime}\right) \sum_{-N}^{N} \mathrm{e}^{\mathrm{i} n \pi\left(x-x^{\prime}\right) / l}=\frac{1}{2 l} \int_{0}^{2 l} \mathrm{~d} x^{\prime} f\left(x^{\prime}\right) \frac{\sin \left[\left(N+\frac{1}{2}\right) \pi\left(x-x^{\prime}\right) / l\right]}{\sin \frac{\pi\left(x-x^{\prime}\right)}{2 l}} \\
& f_{N}(\alpha)=\frac{1}{2 \pi} \int_{-\alpha}^{2 \pi-\alpha} \mathrm{d} \theta f(l(\theta+\alpha) / \pi) \frac{\sin \left(N+\frac{1}{2}\right) \theta}{\sin \frac{\theta}{2}}
\end{aligned}
$$

where the last equality is obtained from the change of variable and redefinition: $\theta=\pi\left(x^{\prime}-x\right) / l=\pi x^{\prime} / l-\alpha$.
Now suppose there is a discontinuity in $f(x)$ at some point corresponding to $\alpha=\alpha_{0} \in(0,2 \pi)$. At the discontinuity, the jump in the truncated sum can be written as: $\Delta f_{N}=f_{N}\left(\alpha_{0}+\epsilon\right)-f_{N}\left(\alpha_{0}-\epsilon\right)$, with $\epsilon$ arbitrarily small but not zero. Calling the sine factor in the integrand $F(\theta)$ for brevity, we have:

$$
\begin{aligned}
f_{N}\left(\alpha_{0}+\epsilon\right) & =\frac{1}{2 \pi} \int_{-\alpha_{0}-\epsilon}^{2 \pi-\alpha_{0}-\epsilon} \mathrm{d} \theta f\left[l\left(\theta+\alpha_{0}+\epsilon\right) / \pi\right] F(\theta) \\
& =\frac{1}{2 \pi} \int_{-\alpha_{0}-\epsilon}^{-\alpha_{0}+\epsilon} \mathrm{d} \theta f\left[l\left(\theta+\alpha_{0}+\epsilon\right) / \pi\right] F(\theta)+\frac{1}{2 \pi} \int_{-\alpha_{0}+\epsilon}^{2 \pi-\alpha_{0}-\epsilon} \mathrm{d} \theta f\left[l\left(\theta+\alpha_{0}+\epsilon\right) / \pi\right] F(\theta)
\end{aligned}
$$

The first integral vanishes in the limit $\epsilon \rightarrow 0$ because the integrand is then continuous over the interval of integration. Similarly, we can split $f_{N}\left(\alpha_{0}-\epsilon\right)$ into an integral from $-\alpha_{0}+\epsilon$ to $2 \pi-\alpha_{0}-\epsilon$ plus another integral from $2 \pi-\alpha_{0}-\epsilon$ to $2 \pi-\alpha_{0}+\epsilon$. Again, continuity of the integrand makes the latter integral vanish. We are left with:

$$
\Delta f_{N}=\frac{1}{2 \pi} \int_{-\alpha_{0}+\epsilon}^{2 \pi-\alpha_{0}-\epsilon} \mathrm{d} \theta\left[f\left(l\left(\theta+\alpha_{0}+\epsilon\right) / \pi\right)-f\left(l\left(\theta+\alpha_{0}-\epsilon\right) / \pi\right)\right] F(\theta)
$$

Now, for large $N$, the function $F(\theta)$ more or less vanishes outside the interval $[-\pi /(N+1 / 2), \pi /(N+1 / 2)]$; inside this interval, $\theta$ is very small, and we can approximate the jump of the truncated sum with:

$$
\Delta f_{N} \approx 2 \frac{\Delta f}{\pi} \int_{0}^{\pi /(N+1 / 2)} \frac{\sin (N+1 / 2) \theta}{\theta} \mathrm{d} \theta=1.18 \Delta f
$$

from which we conclude that, very close $(\approx \pi / N)$ to a discontinuity, the Fourier approximation overshoots the function by about $18 \%$. This is often called the Gibbs phenomenon.

Here is an example of the Gibbs phenomenon applied to a square wave with a discontinuity of 1 at $x=0$. I have plotted the Fourier series:

$$
y(x)=\frac{1}{2}+\frac{2}{\pi} \sum_{k=1}^{N} \frac{1}{2 k+1} \sin \frac{(2 k+1) \pi x}{10}
$$

for $N=50$ (on the left) and $N=2000$ (on the right). As expected, the series overshoots 1 by about 0.09 , independent of $N$ so long as it is large enough. Note also the change of scale of the $x$-axis needed because the overshoot becomes narrower as $N$ increases.



### 5.4.3 Multidimensional Fourier series

The Fourier theorem is readily extended to more than one dimension. Let $F(\mathbf{x})$ be a $N$-dimensional separable function, in the sense that it is a product of $N 1$-dim periodic functions $f_{i}$, each depending on one variable $x_{i}$, each having the Fourier decomposition:

$$
f_{i}\left(x_{i}\right)=\frac{1}{\sqrt{2 l_{i}}} \sum_{n_{i}=-\infty}^{\infty} c_{n_{i}} \mathrm{e}^{\mathrm{i} k_{i} x_{i}}
$$

where $k_{i}=\pi n_{i} / l_{i}$. Then the Fourier expansion of $F$ is:

$$
\begin{equation*}
F(\mathbf{x})=\frac{1}{\sqrt{V}} \sum_{\mathbf{n}} c_{\mathbf{n}} \mathrm{e}^{\mathbf{i} \mathbf{k} \cdot \mathbf{x}} \tag{5.11}
\end{equation*}
$$

where $c_{\mathbf{n}}=c_{n_{1}} \cdots c_{n_{N}}$ and $V=2^{N}\left(l_{1} \cdots l_{N}\right)$. We shall take this expression as the definition of the Fourier expansion of a function of $N$ variables that is periodic in all the variables.

### 5.5 Fourier Integrals (BF 5.7) and Dirac Delta-Function (BF 5.3)

### 5.5.1 Fourier integrals

We would like to approximate a non-periodic function over the whole axis instead of just over some finite interval. Using the complex-exponential form of the Fourier theorem, we write:

$$
f(x)=\frac{1}{\sqrt{2 l}} \sum_{n=-\infty}^{\infty}\left(\frac{1}{\sqrt{2 l}} \int_{-l}^{l} f\left(x^{\prime}\right) \mathrm{e}^{-\mathrm{i} n \pi x^{\prime} / l} \mathrm{~d} x^{\prime}\right) \mathrm{e}^{\mathrm{i} n \pi x / l}
$$

Now, in the summation, we redefine the index $n$ so that $k_{n}=n \pi / l$, where the new running index jumps by values of $\Delta k=\pi / l$. Implementing this change in the summand gives:

$$
f(x)=\frac{1}{\sqrt{2 \pi}} \sum_{k_{n}=-\infty}^{\infty}\left(\frac{1}{\sqrt{2 \pi}} \int_{-l}^{l} f\left(x^{\prime}\right) \mathrm{e}^{-\mathrm{i} k_{n} x^{\prime}} \mathrm{d} x^{\prime}\right) \Delta k \mathrm{e}^{\mathrm{i} k_{n} x}
$$

Letting $l \rightarrow \infty$, or $\Delta k \rightarrow \mathrm{~d} k$, one transforms the summation into a Fourier integral for a piecewise smooth $f(x)$ :

$$
\begin{equation*}
f(x)=\sqrt{\frac{|a|}{2 \pi}} \int_{-\infty}^{\infty} g(k) \mathrm{e}^{\mathrm{i} a k x} \mathrm{~d} k \quad g(k)=\sqrt{\frac{1}{2 \pi|a|}} \int_{-\infty}^{\infty} f(x) \mathrm{e}^{-\mathrm{i} a k x} \mathrm{~d} x \tag{5.12}
\end{equation*}
$$

where $a$ is chosen by convention; some references choose $a=2 \pi$, others choose $a<0$. Here we shall opt for another popular choice: $a=1$. Also, we often write the Fourier transform of $f(x)$ as $f(k)$, or $\tilde{f}(k)$
$f(x)$ and $g(k)$ are called the Fourier transforms of each other ${ }^{\dagger}$ and provide equivalent representations of the same object. It is very easy to check (EXERCISE) that they inherit each other's symmetry (if any) in one dimension.

At points of discontinuity of $f(x)$, its Fourier transform converges to the average of the two limits, from the right and from the left, of the function.

It is tempting to see $f(x)$ in eq. (5.12) as an expansion over a continuous basis of functions $\mathrm{e}^{\mathrm{i} k x}$. But this is not quite like the discrete basis $\left\{\mathrm{e}^{\mathrm{i} n \pi x / l} / \sqrt{2 l}\right\}$ whose elements certainly belong to the Hilbert space defined over the finite interval $(-l, l)$. Here, the interval is infinite, and the functions $\mathrm{e}^{\mathrm{i} k x}$ are not square-integrable over that interval. Therefore, they do not belong to the corresponding Hilbert space. Clearly, there must be a restriction on $f(x)$ which propagates to its Fourier transform $g(k)$. This restriction is that $f(x)$ must be piecewise smooth (its derivative may not be discontinuous at more than a finite number of points) and that the integral of $|f(x)|$ over the whole $x$ axis must exist.

Example 5.1. A nice model for pulses is the Gaussian time-domain function $g(t)=\mathrm{e}^{-t^{2} / 2 \tau^{2}}$. This pulse is centered at $t=0$, and $\tau$ controls its width at any particular time. Transforming to the frequency domain gives:

$$
g(\omega)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} \mathrm{e}^{-t^{2} / 2 \tau^{2}-\mathrm{i} \omega t} \mathrm{~d} t=\frac{\mathrm{e}^{-\omega^{2} \tau^{2} / 2}}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} \mathrm{e}^{-\left(t+\mathrm{i} \omega \tau^{2}\right)^{2} / 2 \tau^{2}} \mathrm{~d} t
$$

where the last equality is obtained by completing the square in the argument of the exponential. The integral, of the form $\int \mathrm{e}^{-x^{2} / 2} \mathrm{~d} x$, is easily found by Maple/Mathematica or in tables, but it is instructive to see how we can evaluate it ourselves. Write in terms of two Cartesian variables and change to circular coordinates:

$$
\begin{aligned}
\iint_{-\infty}^{\infty} \mathrm{e}^{-\left(x^{2}+y^{2}\right) / 2} \mathrm{~d} x \mathrm{~d} y & =\int_{\phi=0}^{2 \pi} \mathrm{~d} \phi \int_{\rho=0}^{\infty} \mathrm{e}^{-\rho^{2} / 2} \rho \mathrm{~d} \rho=2 \pi \int_{u=0}^{\infty} \mathrm{e}^{-u} \mathrm{~d} u \\
& =2 \pi
\end{aligned}
$$

[^18]But this integral is just the square of $\int \mathrm{e}^{-x^{2} / 2} \mathrm{~d} x$. We arrive at:

$$
g(\omega)=\tau \mathrm{e}^{-\omega^{2} \tau^{2} / 2}
$$

Extending Fourier integrals to three dimensions is straightforward; with our conventions:

$$
\begin{align*}
& F(\mathbf{x})=\frac{1}{(2 \pi)^{3 / 2}} \int_{\text {all space }} G(\mathbf{k}) \mathrm{e}^{\mathrm{i} \mathbf{k} \cdot \mathbf{x}} \mathrm{~d}^{3} k \\
& G(\mathbf{k})=\frac{1}{(2 \pi)^{3 / 2}} \int_{\text {all space }} F(\mathbf{x}) \mathrm{e}^{-\mathrm{i} \mathbf{k} \cdot \mathbf{x}} \mathrm{~d}^{3} x \tag{5.13}
\end{align*}
$$

EXERCISE: Calculate the 3-dim Fourier transform, $f(\mathbf{k})$, of $f(r)=\mathrm{e}^{-\alpha r} / r$, where $\alpha>0$ and $r$ is the radial coordinate. Then take the limit $\alpha \rightarrow 0$. What happens if you had tried to find the Fourier transform of $1 / r$ directly?

### 5.5.2 Differentiating Fourier transforms

Differentiating Fourier transforms is not hard, and it can bring dividends:

$$
\partial_{n} f(\mathbf{x})=\frac{1}{(2 \pi)^{3 / 2}} \int \partial_{n}\left[f(\mathbf{k}) \mathrm{e}^{\mathrm{i} k_{j} x^{j}}\right] \mathrm{d}^{3} k=\frac{1}{(2 \pi)^{3 / 2}} \int f(\mathbf{k}) \mathrm{i} k_{n} \mathrm{e}^{\mathrm{i} k_{j} x^{j}} \mathrm{~d}^{3} k \quad \partial_{n}=\partial / \partial x^{n}
$$

Then it is easy to see that a second-order ordinary differential equation with constant coefficients: $c_{2} \mathrm{~d}_{x}^{2} f(x)+$ $c_{1} \mathrm{~d}_{x} f(x)+c_{0} f(x)=F(x)$, becomes:

$$
\frac{1}{\sqrt{2 \pi}} \int f(k)\left(-c_{2} k^{2}+\mathrm{i} c_{1} k+c_{0}\right) \mathrm{e}^{\mathrm{i} k x} \mathrm{~d} k=\frac{1}{\sqrt{2 \pi}} \int F(k) \mathrm{e}^{\mathrm{i} k x} \mathrm{~d} k
$$

Then, because the Fourier transform of the zero function vanishes everywhere, the differential equation is turned into the algebraic equation:

$$
\begin{equation*}
f(k)=\frac{F(k)}{-c_{2} k^{2}+\mathrm{i} c_{1} k+c_{0}} \tag{5.14}
\end{equation*}
$$

In principle, the solution $f(x)$ follows by taking the Fourier transform of $f(k)$, but that integral can be quite difficult to evaluate. We will look at some ways of doing this in a later module.

One important use for Fourier series and integrals is the decomposition of a suitable function into a superposition of oscillating functions in their trigonometric or complex exponential forms. This spectral decomposition is in terms of harmonics, or modes, of frequency $n \pi / l$ and amplitude $\left|c_{n}\right|$ [eq. (5.9)] when performed over a closed interval $[-l, l]$. Over an infinite interval, the spectrum is said to be continuous with spectral density given by $g(k)$ in eq. (5.13), where $k$ is the frequency.

The 3 -dim algebraic equation inherits (EXERCISE) the vector structure of the derivative: $\boldsymbol{\nabla} \rightarrow \mathrm{i} \mathbf{k}, \boldsymbol{\nabla} \cdot \rightarrow \mathrm{i} \mathbf{k}$, $\nabla \times \rightarrow \mathrm{ik} \times$.

### 5.5.3 Dirac delta function

A famous Fourier integral is of a finite wave train $f(x)=\mathrm{e}^{\mathrm{i} k_{0} x}$ over $[-l, l]$ and $f(x)=0$ elsewhere. Its spectral density is:

$$
g(k)=\frac{1}{\sqrt{2 \pi}} \int_{-l}^{l} \mathrm{e}^{-\mathrm{i}\left(k-k_{0}\right) x} \mathrm{~d} x=\sqrt{2 \pi} \frac{\sin \left(k-k_{0}\right) l}{\pi\left(k-k_{0}\right)}
$$

The longer the wave train, the more sharply peaked its spectral density around the single spatial frequency $k_{0}$. You can watch it as it happens by entering the following command in Maple:

$$
\text { plots [animate] }(p l o t, \sin ((k-k 0) * x) /(k-k 0), k=-a . a, x=1 \ldots N)
$$

where k 0 is a number and $a$ should be no more than ten percent of $\mathrm{k} 0 ; N$ can be very large but about twenty times k 0 should be sufficient. Right-click on the output plot and choose Animate $\rightarrow$ Play in the drop-down menu. You can vary the speed of the animation (FPS) at the top of the worksheet.

If we take the limit $l \rightarrow \infty$, we find that:

$$
\frac{1}{2 \pi} \int_{-\infty}^{\infty} \mathrm{e}^{-\mathrm{i}\left(k-k_{0}\right) x} \mathrm{~d} x=\lim _{u \rightarrow \infty} \frac{\sin \left(k-k_{0}\right) u}{\pi\left(k-k_{0}\right)}
$$

Another instructive manipulation consists in inserting $g(k)$ into $f(x)$, both in eq. (5.12):

$$
f(x)=\int_{-\infty}^{\infty} f\left(x^{\prime}\right)\left[\frac{1}{2 \pi} \int_{-\infty}^{\infty} \mathrm{e}^{\mathrm{i} k\left(x-x^{\prime}\right)} \mathrm{d} k\right] \mathrm{d} x^{\prime}
$$

Compare this to the defining equation for the Dirac $\delta$-function:

$$
\begin{equation*}
f(a)=\int_{-\infty}^{\infty} f(x) \delta(a-x) \mathrm{d} x \tag{5.15}
\end{equation*}
$$

to obtain the Fourier representation of the $\delta$-function, perhaps the most useful:

$$
\begin{equation*}
\delta\left(x-x^{\prime}\right)=\frac{1}{2 \pi} \int_{-\infty}^{\infty} \mathrm{e}^{\mathrm{i} k\left(x-x^{\prime}\right)} \mathrm{d} k=\lim _{u \rightarrow \infty} \frac{\sin \left(x-x^{\prime}\right) u}{\pi\left(x-x^{\prime}\right)} \tag{5.16}
\end{equation*}
$$

The last equality makes it plain that there must be an implicit limit in the integral since its integrand is not squareintegrable.

Other expressions (BF 5.3) for the $\delta$-function which can be useful are the Gaussian and Lorentzian representations:

$$
\begin{equation*}
\delta\left(x-x^{\prime}\right)=\lim _{\tau \rightarrow 0} \frac{1}{\sqrt{2 \pi} \tau} \mathrm{e}^{-\left(x-x^{\prime}\right)^{2} / 2 \tau^{2}}=\lim _{\epsilon \rightarrow 0} \frac{1}{\pi} \frac{\epsilon}{\left(x-x^{\prime}\right)^{2}+\epsilon^{2}} \tag{5.17}
\end{equation*}
$$

The $\delta$-function $\delta(x)$ vanishes everywhere except at $x=0$, where it is singular. Because its integral over any interval containing $x=0$ does not vanish, it is not a function in the usual sense. Although something like it was introduced by Weierstrass to prove his approximation theorem in 1885, only in 1947 was it rigorously interpreted by Laurent Schwartz (1915-2002) in his theory of distributions.

The $\delta$-function can be used to extend the concept of orthonormality to a set $\left\{f_{k}(x)\right\}$ of functions with a continuous index $k$. The functions are said to be orthonormal if:

$$
\begin{equation*}
\int_{-\infty}^{\infty} f_{k}(x) f_{k^{\prime}}(x) \mathrm{d} x=\delta\left(k-k^{\prime}\right) \tag{5.18}
\end{equation*}
$$

It also allows us to prove the Parseval-Plancherel theorem (or identity), for a Fourier integral:

$$
\begin{align*}
\int_{-\infty}^{\infty}|f(x)|^{2} \mathrm{~d} x & =\int \mathrm{d} k g^{*}(k) \int \mathrm{d} k^{\prime} g\left(k^{\prime}\right)\left[\frac{1}{2 \pi} \int \mathrm{e}^{\mathrm{i}\left(k^{\prime}-k\right) x} \mathrm{~d} x\right]=\int \mathrm{d} k g^{*}(k) \int \mathrm{d} k^{\prime} g\left(k^{\prime}\right) \delta\left(k^{\prime}-k\right) \\
& =\int_{-\infty}^{\infty}|g(k)|^{2} \mathrm{~d} k \tag{5.19}
\end{align*}
$$

This says that the norm of a function is invariant under a Fourier transform. See section 14.5 in Jackson's Classical Electrodynamics for an application to the calculation of the energy radiated per unit solid angle by a charge in arbitrary motion.

The generalisation of the $\delta$-function's Fourier representation to three dimensions is straightforward:

$$
\begin{equation*}
\delta^{3}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)=\frac{1}{(2 \pi)^{3}} \int_{-\infty}^{\infty} \mathrm{e}^{\mathrm{i} \cdot\left(\mathbf{x}-\mathrm{x}^{\prime}\right)} \mathrm{d}^{3} k \tag{5.20}
\end{equation*}
$$

Some care must be exercised, however, when relating 3-dim to 1 -dim $\delta$-functions in curvilinear coordinates:

$$
\delta^{3}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)=\delta\left(x-x^{\prime}\right) \delta\left(y-y^{\prime}\right) \delta\left(z-z^{\prime}\right)=\frac{1}{r^{2} \sin \theta} \delta\left(r-r^{\prime}\right) \delta\left(\theta-\theta^{\prime}\right) \delta\left(\phi-\phi^{\prime}\right)
$$

where the first expression is in Cartesian and the second line in spherical coordinates. This is so that, in any coordinate system, the integral of $\delta^{3}\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ over whole space is 1 . There is some awkwardness about $\delta(r)$ since $r \geq 0$ : normally, the interval of integration in $\int \delta(x) \mathrm{d} x=1$ contains 0 , where $\delta(x)$ is singular, but with $r$ the lower limit lies at the singularity.. An exception has to be allowed for the radial variable, and $\delta(r)$ defined so that $\int \delta(r) f(r) \mathrm{d} r=f(0)$

Moreover, the $\delta$-function can be expanded over any complete orthonormal set $\left\{f_{k}(x)\right\}$ (EXERCISE):

$$
\begin{equation*}
\delta\left(x-x^{\prime}\right)=\sum_{k}^{\infty} f_{k}(x) f_{k}^{*}\left(x^{\prime}\right) \tag{5.21}
\end{equation*}
$$

When $k$ is continuous, replace the summation by an integral over $k$. The $\delta$-function has lots of representations!

### 5.5.4 Convolution

Definition 5.8. The convolution $\left[f_{1} \star f_{2}\right](x)$ of two functions $f_{1}$ and $f_{2}$ is defined by:

$$
\begin{equation*}
\left[f_{1} \star f_{2}\right](x):=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} f_{1}\left(x^{\prime}\right) f_{2}\left(x-x^{\prime}\right) \mathrm{d} x^{\prime} \tag{5.22}
\end{equation*}
$$

This can be viewed as a weighted average of the input function $f_{1}$ at the point $x$ with weighting given by $f_{2}\left(x-x^{\prime}\right)$. As $x$ varies, the weight function emphasises different regions of the input function. The convolution is commutative, associative, distributive, and linear.

If we express the right-hand side of eq. (5.22) in terms of the Fourier transforms, $g_{1}$ and $g_{2}$, of $f_{1}$ and $f_{2}$ :

$$
\begin{align*}
{\left[f_{1} \star f_{2}\right](x) } & =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} f_{1}\left(x^{\prime}\right)\left[\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} g_{2}(k) \mathrm{e}^{\mathrm{i} k\left(x-x^{\prime}\right)} \mathrm{d} k\right] \mathrm{d} x^{\prime} \\
& =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty}\left[\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} f_{1}\left(x^{\prime}\right) \mathrm{e}^{-\mathrm{i} k x^{\prime}} \mathrm{d} x^{\prime}\right] g_{2}(k) \mathrm{e}^{\mathrm{i} k x} \mathrm{~d} k  \tag{5.23}\\
& =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} g_{1}(k) g_{2}(k) \mathrm{e}^{\mathrm{i} k x} \mathrm{~d} k \tag{5.24}
\end{align*}
$$

the convolution theorem then says that the Fourier transform $G(k)$ of $\left[f_{1} \star f_{2}\right](x)$ is the product $g_{1}(k) g_{2}(k)$ of the Fourier transforms of $f_{1}$ and $f_{2}$. In other words, in Fourier space, the convolution integral becomes multiplication.

Convolution integrals pop up everywhere. We have met them in the Poisson integral and in the response function (section 4.7) of module IV, and we will meet them again with Green functions in module VI.

### 5.5.5 Discrete and Fast Fourier Transforms

The Fourier transform is invaluable for the study of functions whose analytical form is known. In numerical calculations, however, all we have is a table with a set of points in, say, time, and values of the function at those discrete points. How can we adapt our Fourier formalism to deal with such cases?

We assume that $N$ values of this function $f\left(t_{j}\right),\left(t_{j}=j \Delta t\right)$, are given at equal intervals $\Delta t=T / N$ over a total period $T$. When we try to write the Fourier transform $f(\omega)$, we should not expect to be able to construct a continuous transform: it too will be discretised at frequencies $\omega_{k}=k \Delta \omega$, where the frequency interval $\Delta \omega$ is to be determined. Letting $\int \longrightarrow \sum$ and $\mathrm{d} t \longrightarrow \Delta t$, we write:

$$
f\left(\omega_{k}\right)=\frac{1}{\sqrt{2 \pi}} \frac{T}{N} \sum_{j=0}^{N-1} f\left(t_{j}\right) \mathrm{e}^{-\mathrm{i} \omega_{k} t_{j}}
$$

We want to reconstruct $f\left(t_{j}\right)$ from $f\left(\omega_{k}\right)$. Our final result will look prettier if we multiply this by $\sqrt{2 \pi N} / T$ and redefine $f\left(\omega_{k}\right)$ so that $\sqrt{2 \pi N} f\left(\omega_{k}\right) / T \longrightarrow f\left(\omega_{k}\right)$. Then, multiplying by $\mathrm{e}^{\mathrm{i} \omega_{k} t_{n}}$ and summing over $k$, there comes:

$$
\begin{equation*}
\sum_{k=0}^{N-1} f\left(\omega_{k}\right) \mathrm{e}^{\mathrm{i} \omega_{k} t_{n}}=\frac{1}{\sqrt{N}} \sum_{j=0}^{N-1} f\left(t_{j}\right) \sum_{k=0}^{N-1} \mathrm{e}^{\mathrm{i} \omega_{k}\left(t_{n}-t_{j}\right)} \tag{5.25}
\end{equation*}
$$

With $\omega_{k}=k \Delta \omega$, we recognise that the last sum over $k$ is a geometric series of the form: $\sum_{k=0}^{N-1} r^{k}=\frac{1-r^{N}}{1-r}$ with $r=\mathrm{e}^{\mathrm{i} \Delta \omega\left(t_{n}-t_{j}\right)}$, and we obtain:

$$
\sum_{k=0}^{N-1} \mathrm{e}^{\mathrm{i} k \Delta \omega\left(t_{n}-t_{j}\right)}=\frac{1-\mathrm{e}^{\mathrm{i} N \Delta \omega\left(t_{n}-t_{j}\right)}}{1-\mathrm{e}^{\mathrm{i} \Delta \omega\left(t_{n}-t_{j}\right)}}
$$

Now, if we can demand that this expression vanish when $n \neq j$, we will have an expression for $f\left(t_{j}\right)$ that looks usable. This will happen if $N \Delta \omega\left(t_{n}-t_{j}\right)=N \Delta \omega \Delta t(n-j)$ is equal to $2 \pi(n-j)$. Indeed, when $n \neq j$, the numerator vanishes, but the denominator does not. And when $n=j$, the sum is simply $N$. Also, since $\Delta t=T / N$, we find the frequency interval we had to determine: $\Delta \omega=2 \pi / T$. Eq. (5.25) now equals just $\sqrt{N} f\left(t_{n}\right)$, and we arrive at the discrete transforms:

$$
\begin{equation*}
f\left(\omega_{k}\right)=\frac{1}{\sqrt{N}} \sum_{j=0}^{N-1} f\left(t_{j}\right) \mathrm{e}^{-\mathrm{i} \omega_{k} t_{j}} \quad f\left(t_{j}\right)=\frac{1}{\sqrt{N}} \sum_{k=0}^{N-1} f\left(\omega_{k}\right) \mathrm{e}^{\mathrm{i} \omega_{k} t_{j}} \tag{5.26}
\end{equation*}
$$

where $t_{j}=j \Delta t=j T / N$ and $\omega_{k}=2 \pi k / T$. Other properties, such as the convolution theorem and Parseval's identity, are readily rewritten in their discretised form.

Discrete Fourier transforms can be used for the numerical solution of differential equations as well as for the spectrum analysis of signals, but they can lead to rather heavy computations. Because $k$ and $j$ each take $N$ values, our calculation contains a $N \times N$ matrix, and it will take a time of order $N^{2}$. Fortunately, there exists a ridiculously simple way of speeding up such calculations tremendously: the Fast Fourier Transform (FFT).

To get some idea of how it works, we shall require that $N=2^{l}$, with $l$ a positive integer. Now split the sum for $f\left(\omega_{k}\right)$ in eq. (5.26) into two sums, one over even, and one over odd, values of $j$. Since $\omega_{k} t_{j}=2 \pi k j / N$, we have:

$$
f\left(\omega_{k}\right)=\sum_{j^{\prime}=0}^{N^{\prime}-1} f\left(t_{2 j^{\prime}}\right) \mathrm{e}^{-\mathrm{i} 2 \pi k j^{\prime} / N^{\prime}}+\mathrm{e}^{-\mathrm{i} 2 \pi k / N} \sum_{j^{\prime}=0}^{N^{\prime}-1} f\left(t_{2 j^{\prime}+1}\right) \mathrm{e}^{-\mathrm{i} 2 \pi k j^{\prime} / N^{\prime}}
$$

Each sum looks exactly like the original sum, except that $N$ has been replaced by $N^{\prime}=N / 2$. Now, under $k \rightarrow k+N^{\prime}$, each sum is invariant and the exponential factor in front of the odd sum switches sign, so that only $N^{\prime}=N / 2$ values of $k$ need computing. Therefore, each sum corresponds to a $\frac{N}{2} \times \frac{N}{2}$ matrix, and here is the magic: the computation's time is now of order $2(N / 2)^{2}=N^{2} / 2$, half of what it was before.

Since $N / 2$ is an even number, we can repeat the process, dividing each sum into two, ending up with a computational time of order $4(N / 4)^{2}=N^{2} / 4$. After $l$ such steps, what remains is a sum of $N$ complex products! The computational time is now $l N$, ie. $N \log _{2} N$. For 1024 points, a FFT computation would be about 100 times faster than a direct calculation.

### 5.5.6 The Sampling Theorem (Whittaker,..., Shannon)

This important theorem asserts that a frequency band-limited function is completely determined by sampling at a frequency (the Nyquist frequency) that is at least twice the highest frequency occurring in the spectrum of the function. The sampling theorem also tells us how to reconstruct the function from its samples. Analog to digital conversion relies heavily on it.

With $\Delta t$ the period between samplings, we sample the continuous function $f(t)$ at $t_{n}=n \Delta t$ with an operator called the Dirac comb:

$$
\operatorname{comb}(t)=\sum_{n=-\infty}^{\infty} \delta\left(t-t_{n}\right)
$$

There is a corresponding sampling frequency $\omega_{\mathrm{s}}$. The result is a sampled function: $f_{\mathrm{s}}(t)=f(t)(\Delta t) \operatorname{comb}(t)$.
Now the comb-"function" admits a Fourier representation (EXERCISE):

$$
\operatorname{comb}(t)=\frac{1}{\Delta t} \sum_{k=-\infty}^{\infty} \mathrm{e}^{\mathrm{i} k \omega_{\mathrm{s}} t}
$$

Therefore:

$$
f_{\mathrm{s}}(t)=\sum_{k=-\infty}^{\infty} f(t) \mathrm{e}^{\mathrm{i} k \omega_{\mathrm{s}} t}
$$

What is the Fourier transform of $f_{\mathbf{s}}(t)$ ? By linearity, just the sum of the transforms of $f(t) \mathrm{e}^{\mathrm{i} k \omega_{\mathrm{s}} t}$, ie.:

$$
\tilde{f}_{\mathrm{s}}(\omega)=\sum_{k=-\infty}^{\infty} \frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} f(t) \mathrm{e}^{\mathrm{i}\left(k \omega_{\mathrm{s}}-\omega\right) t} \mathrm{~d} t=\sum_{k=-\infty}^{\infty} \tilde{f}\left(\omega-k \omega_{\mathrm{s}}\right)
$$

where we have used a tilde to make it clearer when we are talking about Fourier transforms.
What we have just found is that the sampled frequency spectrum is made of an infinite number of replicas of the spectrum of the original function around integer multiples of $\omega_{\mathrm{s}}$.

Now suppose the frequency spectrum of $f$ has limited bandwith, in the sense that $\tilde{f}(\omega)=0$ when $\omega>\omega_{\text {max }}$. The sampled spectrum still consists of copies of this original spectrum. In practice, this means that for some sampled value we cannot disentangle the contribution from $\tilde{f}\left(\omega<\omega_{\max }\right)$, which is the one that carries information, from contributions at $\omega+k \omega_{\mathrm{s}}$ which are spurious since there is no spectrum there! If, however, we sample at a frequency larger than the Nyquist frequency, $\omega_{\mathrm{N}}=2 \omega_{\max }$, then the copies all lie above $\omega_{\max }$, and we can pick out the actual spectrum by filtering $\tilde{f}_{\mathrm{s}}(\omega)$ :

$$
\tilde{f}(\omega)=\operatorname{rect}\left(\omega / \omega_{\mathrm{s}}\right) \tilde{f}_{\mathrm{s}}(\omega)
$$

where we have introduced the rectangular function:

$$
\operatorname{rect}(u)= \begin{cases}1 & |u|<\frac{1}{2}  \tag{5.27}\\ 0 & |u|>\frac{1}{2}\end{cases}
$$

According to the convolution theorem, $f(t)$ can be seen as the convolution of rect and $f_{\mathrm{s}}$ :

$$
f(t)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} \widetilde{\operatorname{rect}}\left(t^{\prime}\right) f_{\mathrm{s}}\left(t-t^{\prime}\right) \mathrm{d} t^{\prime}
$$

The Fourier transform of the rectangle function is easily found (EXERCISE): $\widetilde{\operatorname{rect}}(t)=\frac{\sqrt{2 \pi}}{\Delta t} \operatorname{sinc}(t / \Delta t)$, where:

$$
\begin{equation*}
\operatorname{sinc}(x) \equiv \frac{\sin \pi x}{\pi x} \tag{5.28}
\end{equation*}
$$

is the sine cardinal function ${ }^{\dagger}$. Inserting this in the integral expression for $f(t)$ yields:

$$
f(t)=\int_{-\infty}^{\infty} \frac{\operatorname{sinc}\left(t^{\prime} / \Delta t\right)}{\Delta t} f_{\mathrm{s}}\left(t-t^{\prime}\right) \mathrm{d} t^{\prime}=\sum_{n=-\infty}^{\infty} \int_{-\infty}^{\infty} \operatorname{sinc}\left(t^{\prime} / \Delta t\right) f\left(t-t^{\prime}\right) \delta\left(t-t_{n}-t^{\prime}\right) \mathrm{d} t^{\prime}
$$

Thus, we arrive at:

$$
\begin{equation*}
f(t)=\sum_{n=-\infty}^{\infty} f\left(t_{n}\right) \operatorname{sinc}\left(\frac{t-t_{n}}{\Delta t}\right) \tag{5.29}
\end{equation*}
$$

If the sampling frequency is above the Nyquist frequency, this exactly reconstructs $f$, even between samples!
Important as they may be, Fourier transforms fail to provide useful and accurate representations of many "realworld" functions. Also, even the FFT method may be too slow for very large sets of data. More recently, a new type of method, called wavelet analysis has been developed that obviates these problems. This topic lies well outside the scope of this course.

[^19]
### 5.6 Some Special Functions and their Defining Differential Equations

In the second half of this module we introduce a unified approach to the study of some functions which although called "special" are ubiquitous in mathematical physics. The reason for the name has more to do with the fact that they cannot be expressed in closed form in terms of elementary functions. Historically, they often appeared first as solutions to second-order differential equations. In fact, these equations, along with appropriate boundary conditions, contain all the information about the functions.

### 5.6.1 Self-adjoint differential operators (BF5.10)

Before we discuss the relevant ordinary differential equations, there are some things we should know about their main ingredient: the differential operator $\mathrm{d}_{x} \equiv \mathrm{~d} / \mathrm{d} x$. This will be acting on functions living in a Hilbert space endowed with the extended inner product introduced in Definition 5.2:

$$
\begin{equation*}
(f, g)=\int_{-\infty}^{\infty} f^{*}(x) g(x) w(x) \mathrm{d} x \tag{5.30}
\end{equation*}
$$

with the real weight function $w(x) \geq 0$. $w(x)$ can be used to restrict the interval to a finite one when appropriate.
In general, however, the outcome of differentiation will not be an element of the Hilbert space of which the function being differentiated is an element. It is easy to find examples of this, for instance the function $\sqrt{x-a}$ which is square-integrable over $[a, b]$ while its derivative is not. Because of this, we say that $\mathrm{d}_{x}$ is not bounded. This means that in a given interval we will have to restrict the set of functions on which the operator may act-its domain-by specifying boundary conditions.

The kind of differential operator we are interested in takes the general form:

$$
\begin{equation*}
L=\alpha(x) \mathrm{d}_{x}^{2}+\beta(x) \mathrm{d}_{x}+\gamma(x) \tag{5.31}
\end{equation*}
$$

where $\alpha(x), \beta(x)$ and $\gamma(x)$ are real. Assuming that $L$ has square-integrable eigenfunctions $f_{n}$, how can we guarantee that $L[f]$ is aquare-integrable if $f(x) \in \mathcal{H}$ for $x \in[a, b]$ ? A sufficient condition is that $\left\{f_{n}\right\}$ forms a basis of $\mathcal{H}$ on $[a, b]$. Indeed:

$$
L[f]=\sum a_{n} L\left[f_{n}\right]=\sum a_{n} \lambda_{n} f_{n}
$$

which does guarantee that $L[f] \in \mathcal{H}$. The question then becomes: what kind of operator can have eigenfunctions which form such a basis? Assume that $L(f) \in \mathcal{H}$ and introduce an adjoint operator, $L^{\dagger}$, which by definition satisfies:

$$
\begin{equation*}
\left(L^{\dagger}[f], g\right)=(f, L[g]) \tag{5.32}
\end{equation*}
$$

We shall demand that $L$ be self-adjoint (Hermitian), ie. satisfy the condition $(L[f], g)=(f, L[g])$ for any functions in Hilbert space on a given interval. Then the eigenfunctions $f_{n}$ that correspond to different eigenvalues are orthogonal and these eigenvalues are real. Indeed:

$$
\left(f_{n}, L\left[f_{m}\right]\right)-\left(L\left[f_{n}\right], f_{m}\right)=\left(\lambda_{m}-\lambda_{n}^{*}\right)\left(f_{n}, f_{m}\right)=0
$$

The case $m=n$ constrains $\lambda_{n}$ to be real, and $m \neq n$ yields $\left(f_{n}, f_{m}\right)=0$. This easily derived result is very important. When we can convince ourselves that the orthonormal $f_{n}$ form a complete set, we will have constructed a basis for any function $f \in \mathcal{H}$ functions in $[a, b]$, so that $L[f] \in \mathcal{H}$, as required.

What restrictions does all this put on the coefficients in $L$, and perhaps also on the functions on which it acts? Before we find out, let us bring in a very useful object:

Definition 5.9. Let $f_{1}$ and $f_{2}$ be differentiable functions, and $f^{\prime}(x) \equiv \mathrm{d}_{x} f(x)$. Their Wronskian is:

$$
\begin{equation*}
W_{f_{1} f_{2}}(x)=f_{1}(x) f_{2}^{\prime}(x)-f_{2}(x) f_{1}^{\prime}(x) \tag{5.33}
\end{equation*}
$$

$W(x)$ can be written as a determinant, which allows us to extend the definition to $n$ functions $f_{i}$ :

$$
W_{f_{1} \ldots f_{n}}(x)=\operatorname{det}\left(\begin{array}{cccc}
f_{1} & \mathrm{~d}_{x} f_{1} & \cdots & \mathrm{~d}_{x}^{n-1} f_{1}  \tag{5.34}\\
f_{2} & \mathrm{~d}_{x} f_{2} & \cdots & \mathrm{~d}_{x}^{n-1} f_{2} \\
\vdots & & \vdots & \\
f_{n} & \mathrm{~d}_{x} f_{n} & \cdots & \mathrm{~d}_{x}^{n-1} f_{n}
\end{array}\right)
$$

The Wronskian provides a very useful and easy test of linear independence: Let $f_{1}$ and $f_{2}$ be differentiable and non-vanishing everywhere in the interval $[a, b]$; their Wronskian, $W_{f_{1} f_{2}}(x)$, vanishes $\forall x \in[a, b]$ if, and only if, $f_{1}$ and $f_{2}$ are linearly dependent everywhere ( $f_{2}$ is the same multiple of $f_{1} \forall x$ ) in $[a, b]$ (EXERCISE).

Now we are ready to compute $(f, L] g])-(L[f], g)$. The result is:

$$
(f, L[g])-(L[f], g)=\left.w \alpha W_{f^{*} g}(x)\right|_{a} ^{b}-\int_{a}^{b}\left[(w \alpha)^{\prime}-w \beta\right] W_{f^{*} g}(x) \mathrm{d} x
$$

The right-hand side vanishes and $L$ is Hermitian (self-adjoint) if, an only if:

$$
\begin{align*}
\left.w \alpha W_{f^{*} g}\right|_{a} ^{b} & =0  \tag{5.35}\\
(w \alpha)^{\prime} & =w \beta \quad \Longleftrightarrow \quad w \alpha=C \exp \left[\int^{x} \frac{\beta\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right)} \mathrm{d} x^{\prime}\right] \tag{5.36}
\end{align*}
$$

where $C$ is an arbitrary constant.
It is clear from condition (5.36) that if $\beta=\alpha^{\prime}$, then $w$ is a constant which can be set to 1 , and $L$ is already in so-called self-adjoint form; it enjoys full self-adjoint status when, from eq. (5.35), either $\alpha$ or the Wronskian vanishes outside a finite interval (or fast enough at infinity), In general, however, $L$ is not self-adjoint. But we can always use the weight function $w(x)$ that satisfies eq. (5.36) to make it so! Indeed, simply multiply $L$ by $w$ and apply $w \beta=(w \alpha)^{\prime}$ :

$$
\begin{equation*}
w L=\mathrm{d}_{x}\left[(w \alpha) \mathrm{d}_{x}\right]+\gamma w \tag{5.37}
\end{equation*}
$$

Provided condition (5.35) is met, $(f, L[g])=(L[f], g)$, with the inner product now containing $w$ as determined by eq. (5.36) (or, equivalently, $w L$ as the operator and 1 as weight function), $w L$ is the self-adjoint form of $L$.

Example 5.2. One form of the equation satisfied by Bessel functions of order $n$ is:

$$
L\left[J_{n}(x)\right]=\left(\mathrm{d}_{x}^{2}+\frac{1}{x} \mathrm{~d}_{x}-\frac{n^{2}}{x^{2}}+1\right) J_{n}(x)=0
$$

A quick computation using the solution for $w(x)$ in eq. (5.36) yields $w(x)=x$, As $\beta=1 / x \neq \mathrm{d}_{x} \alpha$, $L$ is not in self-adjoint form. We can, however, multiply $L$ by $x$ to obtain a self-adjoint operator:

$$
L\left[J_{n}\right]=\mathrm{d}_{x}\left(x \mathrm{~d}_{x} J_{n}(x)\right)+\left(x-\frac{n^{2}}{x}\right) J_{n}(x)
$$

subject (since $w \alpha=x$ ) to appropriate boundary conditions on the $J_{n}(x)$.
Suppose now that we find that $w(x)$ vanishes on the boundaries of an interval of interest. Since the constant $C$ in eq. (5.36) is arbitrary, we can match $w(x)(x \in[a, b])$ at the boundaries to other functions $w(x)$ outside the interval for which the arbitrary constant $C$ in eq. (5.36) vanishes. $w(x)$ can then be defined as:

$$
w(x)= \begin{cases}\frac{1}{\alpha} \exp \left[\int^{x} \frac{\beta\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right)} \mathrm{d} x^{\prime}\right] & a \leq x \leq b \\ 0 & x>b, x<a\end{cases}
$$

This allows us to use the whole axis as the interval of integration in inner products if this makes the integral easier to calculate.

Definition 5.10. Over the interval $[a, b]$, the self-adjoint operator:

$$
\mathrm{d}_{x}\left[(w \alpha) \mathrm{d}_{x}\right]+\gamma w
$$

together with either the separated boundary conditions for the functions $f \in \mathcal{H}$ on which it acts:

$$
c_{1} f(a)+\left.c_{2} \mathrm{~d}_{x} f\right|_{a}=0, \quad \quad d_{1} f(b)+\left.d_{2} \mathrm{~d}_{x} f\right|_{b}=0
$$

(with at least one constant in each equation being non-zero) or the periodic boundary conditions:

$$
f(a)=f(b),\left.\quad \quad \mathrm{d}_{x} f\right|_{a}=\left.\mathrm{d}_{x} f\right|_{b}
$$

for the functions on which it acts, form ${ }^{\dagger}$ a Sturm-Liouville system.
With appropriate boundary conditions on its domain, the operator $L=\mathrm{d}_{x}^{2}+q(x)$ is already self-adjoint and in Sturm-Liouville form, since $\alpha=1$. It turns out that it is always possible to redefine the functions on which $L$ in eq. (5.31) acts so as to transform it into the form with only a second-order derivative with a constant coefficient. Some of you have done precisely this when you eliminated the first-order derivative in the radial Schrödinger equation for a spherically-symmetric potential by redefining the radial solutions: $R(r)=u(r) / r$.

Many differential equations in physics are eigenvalue equations of the form $L[f]=\lambda f$. If $L[f]$ is in its self-adjoint form, and if $f$ is subject to separated or periodic boundary conditions as defined above, we say that $L[f]=\lambda f$, where $\lambda$ is the eigenvalue associated with eigenfunction $f$, is a Sturm-Liouville problem.

We note that if there were two eigenfunctions, $f_{1}$ and $f_{2}$, differentiable everywhere and corresponding to the same eigenvalue, that satisfy the same eigenvalue equation and separated boundary condition, say at $x=a$, then the conditions could be trivially rewritten as:

$$
\left(\begin{array}{ll}
f_{1}(a) & f_{1}^{\prime}(a) \\
f_{2}(a) & f_{2}^{\prime}(a)
\end{array}\right)\binom{c_{1}}{c_{2}}=0
$$

But since $c_{1}$ and $c_{2}$ cannot both vanish, the Wronskian $W_{f_{1} f_{2}}(x)$ must vanish at the boundary and therefore (differentiate its definition and eliminate second-order derivatives using $L[f]=\lambda f$, then integrate to see that if $W_{f_{1} f_{2}}(x)=0$ anywhere in the interval, $W_{f_{1} f_{2}}(x)=0$ everywhere) and the functions are linearly dependent. This shows that the eigenvalues of a Sturm-Liouville operator with separated boundary conditions are non-degenerate.

It is important to remember that the eigenfunctions of a Sturm-Liouville operator do not form a basis for the whole of Hilbert space, ie. for all square-integrable functions, but only for the subspace of square-integrable functions that satisfy the appropriate boundary conditions for the corresponding Sturm-Liouville problem. Nevertheless, from Sturm-Liouville problems we can extract several bases composed of known functions over which to expand many functions of interest.

Example 5.3. In this example, we illustrate the results of choosing different intervals and boundary conditions for the same Sturm-Liouville operator, $\mathrm{d}_{x}^{2}$, with eigenvalue equation:

$$
\mathrm{d}_{x}^{2} f_{n}-\lambda_{n} f_{n}=0 \quad \lambda_{n}<0
$$

The eigenfunctions depend on the interval we choose. Over $[0, L]$ and with the separated boundary conditions $f_{n}(0)=0, f_{n}(L)=0$, the eigenfunctions are $\sin (n \pi x / L)$. Note that these conditions are not periodic! Indeed, the boundary conditions satisfied by the derivatives are $\left.\mathrm{d}_{x} f_{n}\right|_{x=0}=(n \pi / L)$ and $\left.\mathrm{d}_{x} f_{n}\right|_{x=L}=(-1)^{n}(n \pi / L)$ which are not identical for all $n$.
If instead we take $[-L, L]$ as our interval and periodic boundary conditions, $f_{n}(-L)=f_{n}(L)$ and $\left.\mathrm{d}_{x} f_{n}\right|_{x=-L}=\left.\mathrm{d}_{x} f_{n}\right|_{x=L}$, the allowed eigenfunctions are now $\sin (n \pi x / L)$ and $\cos (n \pi x / L)$.
We now have doubly degenerate eigenvalues, allowed since the boundary conditions are not separated.
The eigenvalues of the operator, however, are the same in all cases: $\lambda_{n}=-(n \pi / L)^{2}\left(n \in \mathbb{Z}^{+}\right)$.

[^20]
### 5.6.2 Orthonormal polynomial solutions of Sturm-Liouville problems (BF 5.10)

We now focus on a particular class of solutions to a Sturm-Liouville eigenvalue problem: orthonormal polynomials, $Q_{n}(x)$, of finite degree $n$. Weierstrass guarantees that functions can be approximated by polynomials only on finite, not semi-infinite or infinite, intervals, but if we can establish that a set of orthonormal polynomials is complete, we have a basis for our Hilbert space. There will in general also exist non-polynomial solutions which fall outside the scope of the following discussion.

We begin by establishing the completeness of the infinite set $\left\{Q_{n}\right\}$ of any polynomial eigenfunctions of a Sturm-Liouville problem over the whole real axis. This can be done by showing that $\left\{Q_{n}\right\}$ is closed, ie. that there exists no function orthogonal to all the $Q_{n}$. Using our extended inner product, and for any real $k$ :

$$
\left(\mathrm{e}^{\mathrm{i} k x}, f\right)=\int_{-\infty}^{\infty} \mathrm{e}^{-\mathrm{i} k x} f(x) w(x) \mathrm{d} x=\sum_{m=0}^{\infty} \frac{(-\mathrm{i} k)^{m}}{m!} \int_{-\infty}^{\infty} f(x) x^{m} w(x) \mathrm{d} x=\sum_{m=0}^{\infty} \frac{(-\mathrm{i} k)^{m}}{m!}\left(x^{m}, f\right)
$$

Now it is always possible to take the first polynomials of degree $\leq m$ and solve for $x^{m}$ as a linear combination of these polynomials. This means that, if there exists a function $f$ such that $\left(f, Q_{n}\right)=0 \forall n$, then $\left(x^{m}, f\right)=0$ $\forall m$. Note that since $n$ is not bounded, neither is $m$, and we find that $\left(\mathrm{e}^{\mathrm{i} k x}, f\right)=0$.. This means that the Fourier transform of $f(x) w(x)$ vanishes, which can only be true if $f(x) w(x)=0$. Now, by definition, $w(x)=0$ only outside the interval of interest, so $f(x)=0$ in that interval-except perhaps at a countable number of points. This is precisely the criterion for completeness. In regions where $w(x)=0, f(x)$ is arbitrary because it does not contribute to the inner product; the Hilbert space is restricted to functions on the interval in which $w(x) \neq 0$.

Restricting to the polynomial solutions of a Sturm-Liouville problem has many consequences. Firstly, we can not rely on the vanishing of the Wronskian in condition 5.35 to get a self-adjoint differential operator $L$. There exists no finite interval at both boundaries of which polynomials of any degree all vanish; and polynomials do not vanish at infinity. We must therefore demand that $w \alpha$ vanish at the boundaries of any finite interval we choose. If the interval is infinite, $w \alpha$ must go to zero faster than $1 / x^{m}$ for any $m>0$.

This in turns means that we must exclude operators with $\beta(x)=0$, otherwise condition 5.35 would force $w \alpha$ to be constant, and therefore zero, everywhere! Put another way, Sturm-Liouville operators of the form $\mathrm{d}_{x}^{2}+\gamma(x)$ do not have polynomial eigenfunctions.

Now, demanding that the eigenfunctions of any second-order operator $L$ be polynomials $Q_{n}(x)$ constrains its coefficients to be themselves polynomials, whose degree is easily determined. Since it cannot vary, we can nail it down just by inserting in turn $Q_{0}, Q_{1}$ and $Q_{2}$ in $L\left[Q_{n}\right]=\lambda_{n} Q_{n}$, transforming it into an algebraic system.

When $n=0$, the derivative terms in $L$ contribute nothing, leaving $\gamma(x)=\lambda_{0}$. The other two equations are:

$$
\begin{array}{rlrl}
\beta(x) \mathrm{d}_{x} Q_{1}+\lambda_{0} Q_{1} & =\lambda_{1} Q_{1} & (n=1) \\
\alpha(x) \mathrm{d}_{x}^{2} Q_{2}+\beta(x) \mathrm{d}_{x} Q_{2}+\lambda_{0} Q_{2} & =\lambda_{2} Q_{2} & & (n=2)
\end{array}
$$

Inserting $\beta(x)$ from the first equation into the second, there comes:

$$
\begin{align*}
& \beta(x)=\left(\lambda_{1}-\lambda_{0}\right) \frac{Q_{1}}{Q_{1}^{\prime}} \equiv \beta_{1} x+\beta_{0} \\
& \alpha(x)=\left(\lambda_{2}-\lambda_{0}\right) \frac{Q_{2}}{Q_{2}^{\prime \prime}}-\left(\lambda_{1}-\lambda_{0}\right) \frac{Q_{1}}{Q_{1}^{\prime}} \frac{Q_{2}^{\prime}}{Q_{2}^{\prime \prime}} \equiv \alpha_{2} x^{2}+\alpha_{1} x+\alpha_{0} \tag{5.38}
\end{align*}
$$

where we have used the prime notation for the derivatives so as to minimise clutter.

## Sturm-Liouville Eigenvalues

We can solve for the eigenvalues of a Sturm-Liouville operator $L=\alpha \mathrm{d}_{x}^{2}+\beta \mathrm{d}_{x}+\gamma(x)$ with polynomial eigenfunctions in terms of a given $\alpha$ and $\beta$ to find that they depend only on the coefficient of the highest power of $x$ in $\alpha$ and $\beta$ (EXERCISE):

$$
\begin{equation*}
\lambda_{n}=n(n-1) \alpha_{2}+n \beta_{1} \tag{5.39}
\end{equation*}
$$

where $\gamma=\lambda_{0}$ has been absorbed into $\lambda_{n}$. This shows that the eigenvalues have a multiplicity of 1 , ie., they are non-degenerate.

The form of eq. (5.39) suggests that we shift all the eigenvalues by $\lambda_{0}$, which will not change the eigenfunctions. This is equivalent to taking $\gamma(x)=\gamma_{0}=0$. Now the five parameters $\alpha_{0}, \alpha_{1}, \alpha_{2}, \beta_{0}$ and $\beta_{1}$ will determine the eigenfunctions, but the eigenvalues depend only on $\alpha_{2}$ and $\beta_{1}$.

But some of these parameters can be set arbitrarily. This arbitrariness manifests itself in the fact that we can make the change of variable $x \rightarrow a x+b$, where $a$ and $b$ ar real constants, without changing the eigenvalues, only the polynomials which now are functions of $a x+b$. Also, any change of the coefficients that results in scaling the eigenvalues (multiplying all of them by the same constant), will not be considered as fundamentally distinct.

## Parity of the Sturm-Liouville Polynomials

In one dimension, the parity operator $P$ reverses the sign of $x$ in an expression. Therefore, it is idempotent: $P^{2}=\mathbf{I}$, the identity operator. By operating twice on an arbitrary 1-dim eigenfunction of $P$, its eigenvalues are easily found to be $\pm 1$.

Now suppose that $L$ is invariant under $P$, ie. $\alpha_{1}=\beta_{0}=0$ in eq. (5.38). Then $L P\left[Q_{n}(x)\right]=P\left[L Q_{n}(x)\right]=$ $\lambda_{n}\left(P Q_{n}(x)\right)$, and $P Q_{n}(x)$ is also an eigenfunction of $L$ with the same eigenvalue $\lambda_{n}$. Because there is only one eigenfunction for a given $\lambda_{n}$, we must have $P Q_{n}(x)=c_{n} Q_{n}(x) . Q_{n}$ is thus an eigenfunction of $P$ as well, with well-defined parity $\pm 1$. We conclude that when $\alpha_{1}=\beta_{0}=0$, the Sturm-Liouville polynomials contain only even or odd powers of $x$, not both.

### 5.6.3 The classical Sturm-Liouville operators

We are now ready to find all the polynomials that solve a Sturm-Liouville problem. Since, as we know, $\beta(x) \neq 0$, we must calculate a weight function $w(x)$ that can turn $L$ into its standard Sturm-Liouville form.

- $\alpha=\alpha_{0}$ constant

The eigenvalues are then $n \beta_{1}$. We choose $\alpha_{0}=1$, noting that if it isn't, we can simply multiply $L$ by an appropriate constant, leaving the eigenfunctions the same and re-scaling the eigenvalues. Then, choosing $C=\mathrm{e}^{\beta_{0}^{2} / 2 \beta_{1}}$ in eq. (5.36):

$$
w(x)=C \exp \left[\int^{x} \frac{\beta\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right)} \mathrm{d} x^{\prime}\right]=\exp \left[\frac{\beta_{1}}{2}\left(x+\frac{\beta_{0}}{\beta_{1}}\right)^{2}\right]
$$

$w(x)$ is non-zero over the whole $x$-axis. We are at liberty to choose the origin as we wish without loss of generality, so we consider only $\beta_{0}=0$. We also remember that $w \alpha=w$ must go to zero fast enough at infinity, because the Wronskian in eq. (5.35), being made of polynomials, will not. For this it is sufficient that $\beta_{1}$ be negative. In fact, we can make it any negative number we wish with the knowledge that different choices merely correspond to a rescaling of $x$. Choosing $\beta_{1}=-2$ leads to $w(x)=\mathrm{e}^{-x^{2}}$ and $\lambda_{n}=-2 n$, yielding the differential equation defined over $(-\infty, \infty)$ :

$$
\begin{equation*}
f^{\prime \prime}-2 x f^{\prime}+2 n f=0 \quad \Longleftrightarrow \quad\left[\mathrm{e}^{-x^{2}} f^{\prime}(x)\right]^{\prime}+2 n \mathrm{e}^{-x^{2}} f(x)=0 \tag{5.40}
\end{equation*}
$$

- $\alpha=\alpha_{1} x+\alpha_{0}$

Again, the eigenvalues are $n \beta_{1}$. We can let $\alpha_{1}=1$ since, if it isn't, we can divide $L$ by $\alpha_{1}$, the only effect being a rescaling of all eigenvalues. And we can shift the origin so as to absorb $\alpha_{0}$. Then:

$$
w(x) \alpha(x)=\exp \left[\int^{x} \frac{\beta\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right)} \mathrm{d} x^{\prime}\right]=\exp \left[\int^{x} \frac{\beta_{1} x^{\prime}+\beta_{0}}{x^{\prime}} \mathrm{d} x^{\prime}\right]=C x^{\beta_{0}} \mathrm{e}^{\beta_{1} x}
$$

Provided that $\beta_{0} \geq 0$ so as to prevent a divergence at $x=0$, we can ensure that $w \alpha$ goes to zero at $+\infty$ faster than any finite power of $x$ by taking $\beta_{1}=-1$ (any other negative number is equivalent via a rescaling of $x$ ). Then we must exclude the interval $(-\infty, 0)$ by demanding that $w(x)=0$ for $x<0$. For $x \geq 0$, $w(x)=x^{s} \mathrm{e}^{-x}$, with $s=\beta_{0}-1$. We could choose $\beta_{1}>0$ if the interval is restricted to $(-\infty, 0]$. The eigenvalues are $\lambda_{n}=-n$, and we arrive at a differential equation defined this time only over $[0, \infty)$ :

$$
\begin{equation*}
x f^{\prime \prime}-[x-s-1] f^{\prime}+n f=0 \quad \Longleftrightarrow \quad\left[x^{s+1} \mathrm{e}^{-x} f^{\prime}(x)\right]^{\prime}+n x^{s} \mathrm{e}^{-x} f(x)=0 \tag{5.41}
\end{equation*}
$$

In the two cases we have so far considered, the intervals over which the polynomials are defined extend all the way to infinity because of exponential dependence of $w$. This puts them outside the reach of the Theorem of Weierstrass which asserts the existence of polynomial approximations only over finite intervals. Our third case, though, involves a finite interval.

- $\alpha=\alpha_{2} x^{2}+\alpha_{1} x+\alpha_{0}$

First, let $x \rightarrow a x+b$ (rescaling of $x$ together with a shift of origin). Then we have: $\alpha=\alpha_{2} a^{2} x^{2}+$ $a\left(2 \alpha_{2} b+\alpha_{1}\right) x+\alpha_{2} b^{2}+\alpha_{1} b+\alpha_{0}$. The linear term is eliminated by choosing $b=-\alpha_{1} / 2 \alpha_{2}$, so that $\alpha=\alpha_{2} a^{2} x^{2}+(3 / 4) \alpha_{1}^{2} / \alpha_{2}+\alpha_{0}$. Now we divide $L$ by $-\alpha_{2} a^{2}$, rescaling all the eigenvalues by the same factor, and we adjust $a^{2}$ so that $\alpha=-x^{2} \pm 1$.
Our manipulations have changed $\beta$, but it is still a linear polynomial. From eq. (5.36) we have, using Maple:

$$
\begin{aligned}
w \alpha & =C \exp \left[\int^{x} \frac{\beta\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right)} \mathrm{d} x^{\prime}\right]=C \exp \left[\int^{x} \frac{\beta_{1} x^{\prime}+\beta_{0}}{ \pm 1-x^{\prime 2}} \mathrm{~d} x^{\prime}\right] \\
& =(\sqrt{\alpha})^{-\beta_{1}}\left\{\begin{array}{lc}
\exp \left[\beta_{0} \operatorname{arctanh} x\right] & \alpha=1-x^{2} \\
\exp \left[-\beta_{0} \arctan x\right] & \alpha=-1-x^{2}
\end{array}\right.
\end{aligned}
$$

Since this time the exponential factor is of no help, and $(\sqrt{\alpha})^{-\beta_{1}}$ cannot vanish faster than any power of $x$ so as to counteract the bad asymptotic behaviour of the Wronskian in eq. (5.35), we must demand that the interval be finite and that $(\sqrt{\alpha})^{-\beta_{1}}=0$ at the end-points, and thus that $\alpha$ have real roots. Therefore, choose $\alpha=1-x^{2}$. But then $\beta_{1}<0$ so as to prevent $w \alpha$ from diverging at $x= \pm 1$. Then $w(x) \neq 0$ on $(-1,1)$, and since $w( \pm 1)=0$, we can match it to another solution $w=0$ valid outside $[-1,1]$.
With $\alpha=1-x^{2}$, we can ask Maple again for $w$, taking care to specify that $x \in[-1,1]$ :

$$
w(x)=(1-x)^{-\left(\beta_{0}+\beta_{1}\right) / 2-1}(1+x)^{\left(\beta_{0}-\beta_{1}\right) / 2-1}
$$

It is usual to redefine the exponents so that:

$$
w(x)=(1-x)^{p}(1+x)^{q}
$$

where $\beta_{0}=q-p$ and $\beta_{1}=-(p+q+2)$. The real principal root is implied if $p$ or $q$ are not integers. Since $\beta_{1}<0$ so as to guarantee proper behaviour of $w \alpha$ as $x \rightarrow \pm 1$, then we have: $q>-1, p>-1$.
The eigenvalues are $\lambda_{n}=n(n-1) \alpha_{2}+n \beta_{1}=-n(n+p+q+1)$.
Instead of setting $\alpha_{1}=0$ in $L$, we could transform $x$ to $1-2 x$, in which case we would have $\alpha=x(1-x)$ and $w=x^{p}(1-x)^{q}$ over $[0,1]$ and $w=0$ elsewhere.

### 5.6.4 Generating formula for Sturm-Liouville polynomials

A general Rodrigues formula uniquely determines the $Q_{n}$ for a given weight function:

$$
\begin{equation*}
Q_{n}(x)=K_{n} \frac{1}{w(x)} \frac{\mathrm{d}^{n}}{\mathrm{~d} x^{n}}\left[\alpha^{n}(x) w(x)\right] \tag{5.42}
\end{equation*}
$$

where the $K_{n}$ are normalisation constants which are not determined by the Sturm-Liouville eigenvalue equation since it is linear. They are chosen according to which application we are interested in.

To see how this can be true, we first show that the functions produced by the Rodrigues formula are indeed polynomials of degree $n$, and then that these are orthogonal with respect to the weight function $w(x)$.

First, given a function $\alpha^{k} w p_{l}(x)$, with $p_{l}(x)$ a polynomial of degree $l$, we establish that because:

$$
\mathrm{d}_{x}\left(\alpha^{k} w p_{l}(x)\right)=\alpha^{k-1} w\left[\beta p_{l}(x)+(k-1) \mathrm{d}_{x} \alpha p_{l}(x)+\alpha \mathrm{d}_{x} p_{l}(x)\right]
$$

then, for $k \geq 1$, and since $\beta_{1}<0$ in all cases and $\alpha_{2}<0\left(\alpha_{2} \neq 0\right.$ needed for the last two terms to be of degree $l+1$ ), there can be no cancellations of powers $l+1$, and the term in square brackets must be a polynomial $r_{l}(x)$ of degree $l+1$.

This means that if we start with $\alpha^{n} w p_{0}(x)$ and differentiate it $k \leq n$ times, at each step we lose a power of $\alpha$ and raise the degree of $r_{l}$ by one, so that after the $k^{\text {th }}$ differentiation we are left with:

$$
\mathrm{d}_{x}^{k}\left(\alpha^{n} w\right)=\alpha^{n-k} w r_{k}(x)
$$

After differentiating $n$ times and inserting in eq. (5.42), we get (up to a constant) $r_{n}(x)$. Then the functions $Q_{n}$ coming out of the Rodrigues formula have the form $r_{n}(x)$ and are thus polynomials of degree $n$.

Also, being aware that $w(x)$ can vanish outside a finite interval, we have, up to a constant:

$$
\begin{aligned}
\int_{-\infty}^{\infty} Q_{m}(x) Q_{n}(x) w(x) \mathrm{d} x & =\int_{-\infty}^{\infty} Q_{m}(x) \mathrm{d}_{x}^{n}\left(\alpha^{n} w\right) \mathrm{d} x \\
& =\left.\sum_{j=0}^{n-1}\left[(-1)^{j}\left(\mathrm{~d}_{x}^{j} Q_{m}\right) \mathrm{d}_{x}^{n-1-j}\left(\alpha^{n} w\right)\right]\right|_{-\infty} ^{\infty}+(-1)^{n} \int_{-\infty}^{\infty}\left(\mathrm{d}_{x}^{n} Q_{m}\right) \alpha^{n} w \mathrm{~d} x
\end{aligned}
$$

If $n>m, \mathrm{~d}_{x}^{n} Q_{m}=0$ and the integral on the second line vanishes; if it isn't, we start with the Rodrigues formula for $Q_{m}$ instead of $Q_{n}$. As for the boundary terms, we know that $\mathrm{d}^{n-1-j}\left(\alpha^{n} w\right)=\alpha^{j+1} w r_{n-1-j}(x)$. But $w \alpha$ vanishes at the boundaries faster than any power of $x$, and so, therefore, will this term. This proves that the Rodrigues-generated polynomials are orthogonal. Also, since we have assumed $\alpha(x)$ and $w(x)$ to be real, it is not hard to see from the Rodrigues formula that the $Q_{n}$ themselves will be real polynomials.

Now those polynomials that satisfy the orthogonality relation for a given weight function and over a given interval can also be shown (see last paragraph on p . BF233) to be complete, and therefore unique, We conclude that the Rodrigues formula provides a (nice and compact!) representation of the polynomial eigenfunctions of any Sturm-Liouville operator. All that is needed is $w$ and $\alpha$.

Here is a list of all the Sturm-Liouville polynomials:

## - Hermite polynomials:

$$
\begin{equation*}
H_{n}(x)=K_{n} \frac{1}{w(x)} \frac{\mathrm{d}^{n}}{\mathrm{~d} x^{n}} w(x)=(-1)^{n} \mathrm{e}^{x^{2}} \mathrm{~d}_{x}^{n} \mathrm{e}^{-x^{2}} \tag{5.43}
\end{equation*}
$$

are the (not normalised to 1) polynomial eigenfunctions corresponding to the Sturm-Liouville equation over $(-\infty, \infty)$ :

$$
H^{\prime \prime}-2 x H^{\prime}+2 n H=0
$$

Since $\alpha_{1}=\beta_{0}=0$, the $H_{n}$ have parity $(-1)^{n}$. They satisfy the relation:

$$
\int_{-\infty}^{\infty} H_{m}(x) H_{n}(x) \mathrm{e}^{-x^{2}} \mathrm{~d} x=2^{n} n!\sqrt{\pi} \delta_{m n}
$$

The functions $H_{n}(x) \mathrm{e}^{-x^{2} / 2} /\left[2^{n} n!\sqrt{\pi}\right]^{1 / 2}$ are the basis wave-functions for the 1 -dim quantum harmonic oscillator.

- The associated Laguerre polynomials (or just Laguerre polynomials when $s=0$ ):

$$
\begin{equation*}
L_{m}^{s}(x)=K_{m} \frac{1}{w(x)} \mathrm{d}_{x}^{m}\left(x^{m} w(x)\right)=x^{-s} \mathrm{e}^{x} \mathrm{~d}_{x}^{m}\left(x^{m+s} \mathrm{e}^{-x}\right) \quad x \in[0, \infty) \tag{5.44}
\end{equation*}
$$

are the (not normalised to 1) polynomial eigenfunctions of degree $m$ corresponding to the Sturm-Liouville equation over $[0, \infty)$ :

$$
x L_{m}^{s}{ }^{\prime \prime}-[x-s-1] L_{m}^{s}{ }^{\prime}+m L_{m}^{s}=0
$$

Since $\alpha_{1} \neq 0$, the $L_{m}^{s}(x)$ do not have definite parity. The Laguerre polynomials $(s=0)$ satisfy:

$$
\int_{0}^{\infty} L_{m}^{0}(x) L_{n}^{0}(x) \mathrm{e}^{-x} \mathrm{~d} x=m!\delta_{m n}
$$

The functions $x^{2 l} L_{n+l}^{2 l+1}(x) \mathrm{e}^{-x / 2}$, where $n$ is the energy quantum number and $l$ the orbital quantum number, occur in the radial wave-functions which solve the Schrödinger equation for the Coulomb potential. Showing the equivalence of the associated radial Schrödinger equation and of the above forms does require some nontrivial manipulation (see your favourite quantum mechanics textbook for details).

## - Jacobi polynomials:

$$
\begin{equation*}
J_{n}^{p, q}(x)=K_{n}^{p, q} \frac{1}{w(x)} \mathrm{d}_{x}^{n}\left(\alpha^{n} w\right)=(1-x)^{-p}(1+x)^{-q} \mathrm{~d}_{x}^{n}\left[\left(1-x^{2}\right)^{n}(1-x)^{p}(1+x)^{q}\right] \tag{5.45}
\end{equation*}
$$

are the most general (not normalised to 1) polynomial eigenfunctions corresponding to the Sturm-Liouville eigenvalue equation over $[-1,1]$ :

$$
\left(1-x^{2}\right) J_{n}^{\prime \prime}+[(q-p)-(2+p+q) x] J_{n}^{\prime}+n(1+p+q+n) J_{n}=0
$$

The most interesting subclass of the Jacobi polynomials, $J_{n}^{p, q}$, are the Gegenbauer polynomials, sometimes called ultraspherical, for which $p=q=m>-1$. They all have well-defined parity. Their Rodrigues generating formula is the more tractable:

$$
\begin{equation*}
\left.G_{n}^{m}(x)=K_{n}^{m} \frac{1}{w(x)} \mathrm{d}_{x}^{n}\left[\left(1-x^{2}\right)^{n} w\right)\right]=\left(1-x^{2}\right)^{-m} \mathrm{~d}_{x}^{n}\left[\left(1-x^{2}\right)^{n+m}\right] \tag{5.46}
\end{equation*}
$$

Two subclasses of the Gegenbauer polynomials are more famous and useful:

- The Chebyshev (Tchebycheff, Tschebyscheff) polynomials which have $m= \pm 1 / 2$ and satisfy either:

$$
\left(1-x^{2}\right) f^{\prime \prime}-x f^{\prime}+n^{2} f=0 \quad(m=-1 / 2)
$$

or

$$
\left(1-x^{2}\right) f^{\prime \prime}-3 x f^{\prime}+n(n+2) f=0 \quad(m=1 / 2)
$$

With $m=-1 / 2$ the Chebyshev polynomials of the first kind are generated by:

$$
\begin{equation*}
T_{n}(x)=K_{n} \sqrt{\left(1-x^{2}\right)} \mathrm{d}_{x}^{n}\left[\left(1-x^{2}\right)^{n-1 / 2}\right] \tag{5.47}
\end{equation*}
$$

With the change of variable $x=\cos \theta$, the equation for the $T_{n}$ can be rewritten as $\mathrm{d}_{\theta}^{2} T_{n}+n^{2} T_{n}=0$. The solutions in complex form are:

$$
\mathrm{e}^{\mathrm{i} n \theta}=\cos n \theta+\mathrm{i} \sin n \theta=(\cos \theta+\mathrm{i} \sin \theta)^{n}
$$

Expressing the right-hand side as its binomial expansion and taking the real part, there comes:

$$
\cos n \theta=\sum_{\substack{k=0 \\ k \text { even }}}^{n}(-1)^{-k / 2}\binom{n}{k} \cos ^{n-k} \theta\left(1-\cos ^{2} \theta\right)^{k / 2}
$$

By inspection, this is a polynomial in $\cos \theta$. The imaginary part $\sin n \theta$, however, is a polynomial multiplied by $\sin \theta$, which is not a polynomial in $\cos \theta$, and therefore cannot be a Chebyshev polynomial of the first kind. Its polynomial part, $(\sin n \theta) / \sin \theta$, is known as a Chebyshev polynomial of the second kind.
So we can now assert that $T_{n}(\cos \theta)=\cos n \theta$. This can be used, eg., to express $\cos n \theta$ in terms of powers of $\cos \theta$. More important, Chebyshev expansions are related to Fourier series with which they share most properties, except that they can converge faster.

- The restriction $p=q=0$, or $\beta_{1}=-2, \beta_{0}=0$, leads to $w(x)=1$, and the eigenvalues are $\lambda_{l}=-l(l+1)$. We obtain the differential equation:

$$
\begin{equation*}
\left(1-x^{2}\right) P_{l}^{\prime \prime}-2 x P_{l}^{\prime}+l(l+1) P_{l}=0 \quad \Longleftrightarrow \quad\left[\left(1-x^{2}\right) P_{l}^{\prime}(x)\right]^{\prime}+l(l+1) P_{l}(x)=0 \tag{5.48}
\end{equation*}
$$

with the Legendre polynomials as (not normalised to 1) polynomial solutions:

$$
\begin{equation*}
P_{l}(x)=K_{l} \frac{1}{w(x)} \frac{\mathrm{d}^{l}}{\mathrm{~d} x^{l}}\left[\left(1-x^{2}\right)^{l} w(x)\right]=\frac{(-1)^{l}}{2^{l} l!} \mathrm{d}_{x}^{l}\left(1-x^{2}\right)^{l} \quad x \in[-1,1] \tag{5.49}
\end{equation*}
$$

The $P_{l}(x)$ have parity $(-1)^{l}$ and satisfy:

$$
\int_{-1}^{1} P_{l^{\prime}}(x) P_{l}(x) \mathrm{d} x=\frac{2}{2 l+1} \delta_{l^{\prime} l}
$$

Quite often, the parametrisation $x=\cos \theta$ is used. Legendre polynomials provide a very useful expansion of the function $1 /\left|\mathbf{x}-\mathbf{x}^{\prime}\right|$; therefore they allow potentials with a $1 / r$ dependence to be expressed in terms of a so-called multipole expansion. They also arise as the angular part of the solution of the Laplace equation, $\nabla^{2} f=0$, in spherical coordinates, in cases of azimuthal symmetry, ie. when the solution is independent of the azimuthal angle that rotates around a suitably chosen $z$-axis.

All the Sturm-Liouville polynomials satisfy recurrence (or recursion) relations found in many references (see course webpage). There are general ones that cover all the polynomials, but they are complicated. Often, it is easiest to derive them directly.

Example 5.4. Using the trigonometric representation of the Chebyshev polynomials, $T_{n}(x=\cos \theta)=$ $\cos n \theta$, one almost immediately obtains (EXERCISE):

$$
T_{n+1}(x)=2 x T_{n}(x)-T_{n-1}(x)
$$

Example 5.5. We also derive (EXERCISE) the relations:

$$
H_{n}^{\prime}=2 n H_{n-1}, \quad H_{n+1}=2 x H_{n}-2 n H_{n-1}
$$

From these examples one can see that recurrence relations can provide a quick way of generating classical orthogonal polynomials. They also come in very handy in the evaluation of some integrals.

Although this discussion exhausts all possibilities for complete orthonormal sets of polynomials as solution of a Sturm-Liouville problem, they are not necessarily the only complete orthonormal functions on the same interval.

### 5.7 Associated Legendre Functions and Spherical Harmonics

To discover a complete set of orthonormal functions on the unit sphere in three dimensions, we appeal again to the Theorem of Weierstrass. This asserts the existence of partial sums:

$$
g_{M}(\theta, \phi)=\sum_{\alpha, \beta, \gamma=0}^{M} A_{\alpha \beta \gamma}^{(M)} u^{\alpha} v^{\beta} w^{\gamma}=\sum_{l=0}^{3 M} \sum_{\substack{\alpha, \beta, \gamma=0 \\ \alpha+\beta+\gamma=l}}^{M} A_{\alpha \beta \gamma}^{(M)} \mathrm{e}^{\mathrm{i} m \phi} \sin ^{\alpha+\beta-|m|} \theta \sin ^{|m|} \theta \cos ^{\gamma} \theta \quad l>0
$$

such that the $g_{M}(\theta, \phi)$ converge uniformly to the continuous function $g(\theta, \phi)$ in the limit $M \rightarrow \infty$. We have written $u=x+\mathrm{i} y=\sin \theta \mathrm{e}^{\mathrm{i} \phi}, v=x-\mathrm{i} y=\sin \theta \mathrm{e}^{-\mathrm{i} \phi}, w=z=\cos \theta$, and $m=\alpha-\beta$. Note that the last
term on the right is merely a rearrangement of the summations. Note also that the second sum is restricted to $\alpha+\beta+\gamma=l$.

The powers in the summand can now be written in terms of $l$ and $m$, so that $g_{M}(\theta, \phi)$ becomes:

$$
g_{M}(\theta, \phi)=\sum_{l=0}^{3 M} \sum_{m=-l}^{l} B_{l m}^{(M)} Y_{l m}(\theta, \phi)
$$

where: $Y_{l m}(\theta, \phi)=\mathrm{e}^{\mathrm{i} m \phi} \sin ^{|m|} \theta f_{l m}(\cos \theta)$, and: $f_{l m}(\cos \theta) \equiv\left(1-\cos ^{2} \theta\right)^{(l-|m|-\gamma) / 2} \cos ^{\gamma} \theta$ is a polynomial in powers of $\cos \theta$, with maximum degree $l-|m|$, because $\alpha+\beta-|m| \geq 0$ is even.

Then it is enough to find $f_{l m}(\cos \theta)$ such that:

$$
\int_{\phi=0}^{2 \pi} \int_{\theta=0}^{\pi} Y_{l^{\prime} m^{\prime}}^{*}(\theta, \phi) Y_{l m}(\theta, \phi) \sin \theta \mathrm{d} \theta \mathrm{~d} \phi=\delta_{l^{\prime} l} \delta_{m^{\prime} m}
$$

Following the same kind of reasoning as with the Fourier expansion, this will guarantee that the $Y_{l m}$ form a complete set.

The $\phi$ dependence of $Y_{l m}$ already guarantees part of the orthogonality. The following functions do satisfy (not proved) the full orthonormality relation:

$$
\begin{equation*}
Y_{l m}(\theta, \phi)=(-1)^{m} \sqrt{\frac{2 l+1}{4 \pi} \frac{(l-m)!}{(l+m)!}} P_{l}^{m}(x) \mathrm{e}^{\mathrm{i} m \phi} \quad(x=\cos \theta, 0 \leq m \leq l) \tag{5.50}
\end{equation*}
$$

where the unnormalised associated Legendre functions of order $m, P_{l}^{m}(x)$, are given in terms of the Legendre polynomials $P_{l}(x)=P_{l}^{0}(x)$ by:

$$
\begin{align*}
P_{l}^{m}(x) & =\frac{(-1)^{l}}{2^{l} l!}\left(1-x^{2}\right)^{m / 2} \mathrm{~d}_{x}^{l+m}\left(1-x^{2}\right)^{l} \\
& =\left(1-x^{2}\right)^{m / 2} \mathrm{~d}_{x}^{m} P_{l}(x) \quad(0 \leq m \leq l,-1 \leq x \leq 1) \tag{5.51}
\end{align*}
$$

The spherical harmonics also exist for $m<0$ and are then defined as:

$$
Y_{l,-m}=(-1)^{m} Y_{l m}^{*}
$$

The $\sqrt{(l+1 / 2)(l-m)!/(l+m)!} P_{l}^{m}(x)$ are themselves orthonormal functions on $[-1,1]$, just as the $\sqrt{l+1 / 2} P_{l}(x)$, but they are not polynomials and thus do not contradict the uniqueness of the Legendre polynomials. They too satisfy a Sturm-Liouville equation. Differentiating the Legendre equation (5.48) $m$ times and redefining $\mathrm{d}_{x}^{m} P_{l}(x)$ with eq. (5.51), we obtain:

$$
\begin{equation*}
\mathrm{d}_{x}\left[\left(1-x^{2}\right) \mathrm{d}_{x} P_{l}^{m}(x)\right]+\left[l(l+1)-\frac{m^{2}}{1-x^{2}}\right] P_{l}^{m}(x)=0 \tag{5.52}
\end{equation*}
$$

The $P_{l}^{m}(x)$ are the only solutions of this equation which are regular at $x= \pm 1$.
The completeness relation satisfied by the spherical harmonics reads:

$$
\begin{equation*}
\sum_{l=0}^{\infty} \sum_{m=-l}^{l} Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi)=\delta\left(x-x^{\prime}\right) \delta\left(\phi-\phi^{\prime}\right) \quad(x=\cos \theta) \tag{5.53}
\end{equation*}
$$

The spherical harmonics satisfy other useful relations which can be found in several references (Arfken's Mathematical Methods for Physicists and Handbook of Mathematical Functions by Abramowitz and Stegun are two popular ones). They occur in the solution to many problems in physics: in quantum mechanics for instance, when the potential in the Schrödinger equation is spherically-symmetric, the angular dependence of the wave functions is always given by spherical harmonics. They are also extremely useful in electrostatics.

## 6 MODULE VI - Solution of Differential Equations with Green Functions

Physical quantities are generally represented by functions of up to four (three spatial and one time) variables and therefore satisfy partial differential equations (PDE). More precisely, let $y\left(x^{1}, \ldots, x^{n}\right)$ be a variable dependent on the independent variables $x^{1}, \ldots, x^{n}$, then $y$ may have to satisfy equations of the form:

$$
\begin{equation*}
f\left(y, \frac{\partial y}{\partial x^{i}}, \ldots, \frac{\partial^{m} y}{\partial x^{i} \partial x^{j} \ldots}, x^{i}\right)=0 \tag{6.1}
\end{equation*}
$$

where $0 \leq i, j, \ldots \leq m$, with the constraint: $i+j+\ldots=m$.
If this equation can be split into:

$$
g\left(y, \frac{\partial y}{\partial x^{i}}, \ldots, \frac{\partial^{m} y}{\partial x^{i} \partial x^{j} \ldots}, x^{i}\right)=F\left(x^{i}\right)
$$

it is said to be inhomogeneous. If $F\left(x^{i}\right)=0, g=0$ is said to be a homogeneous equation.
You may be relieved to know that in physics we almost never have to go beyond $m=2$. Still, PDEs can be extremely challenging, and most have to be solved numerically. Very thick books have been written on techniques for numerically solving PDEs, and we will not even attempt to broach the topic. In some important cases, PDEs in $n$ independent variables can be converted into $n$ ordinary differential equations (ODE).via the technique of separation of variables. To test whether a PDE has completely separable solutions, insert $y\left(x^{1}, \ldots, x^{n}\right)=X_{1}\left(x^{1}\right) X_{2}\left(x^{2}\right) \ldots X_{n}\left(x^{n}\right)$ into it, and see if it can be written as a sum of terms, each of which depends on one $x^{i}$ only. If that happens, the PDE can be satisfied only if each term is equal to a constant, called the separation constant, with all the constants summing to zero. Then we are left with $n$ ODEs, one for each $X_{i}\left(x^{i}\right)$. If the solution to each of these ODEs is unique, this solution to the PDE will also be unique.

In the next few sections, we shall discuss ODEs of first and second order, returning to PDEs later.

### 6.1 Ordinary Differential Equations

An ODE of order $n$ for $y(x)$ is said to be linear if $g\left(y, y^{\prime}, y^{\prime \prime}, \ldots\right)$ is linear in $y$ and all its derivatives occurring in the equation. Then, with the notation $L[y] \equiv g\left(y, y^{\prime}, y^{\prime \prime}, \ldots\right)$, linearity means that $L\left[a f_{1}+b f_{2}\right]=a L\left[f_{1}\right]+b L\left[f_{2}\right]$, where $f_{1}$ and $f_{2}$ solve the ODE; $a$ and $b$ are constants. We see immediately that adding the sum of the general solution of $L[y]=0$ and a solution to $L[y]=F$ also solves the latter and is its general solution.

Many important natural phenomena are described by non-linear differential equations, but there exists no general method for solving them, only tricks that may work in a limited number of cases. One such trick is to try a clever redefinition. Take for instance the first-order Bernoulli equation:

$$
\mathrm{d}_{x} y+\beta(x) y=y^{n} F(x)
$$

which is readily rewritten in the inhomogeneous form:

$$
y^{-n} \mathrm{~d}_{x} y+\beta(x) y^{1-n}=F(x)
$$

If we notice that $(1-n) y^{-n} \mathrm{~d}_{x} y=\mathrm{d}_{x}\left(y^{1-n}\right)$ then we should redefine $u=y^{1-n}$ to convert the ODE into the linear equation

$$
\mathrm{d}_{x} u+[(1-n) \beta(x)] u=(1-n) F(x)
$$

A whole lore also exists on finding so-called integrating factors, but again this is very much a hit-and-miss affair, and we shall (with some regret) leave the study of non-linear ODEs for another time.

### 6.1.1 First-order ODEs

The most general form for a first-order ODE is:

$$
\begin{equation*}
y^{\prime}=g(x, y(x)) \tag{6.2}
\end{equation*}
$$

When supplemented by an initial condition, such as $y\left(x_{0}\right)=y_{0}$, this becomes an initial-value problem (IVP). Provided that $g$ is continuous in $x$, a theorem by Peano guarantees the existence of a solution, at least over some interval around $x_{0}$. But such a solution may not be unique! Indeed, let $g(x, y(x))=y^{n}$ for $0<n<1$ when $x<0$ and zero otherwise, with $y(0)=0$. Then $y(x)=[(1-n) x]^{1 /(1-n)}(x<0)$ and zero otherwise solves the IVP (EXERCISE), but so does $y(x)=0 \forall x$. There is even a third solution, $y(x)=-[(1-n) x]^{1 /(1-n)}$ when $n$ is rational and the numerator of $1 /(1-n)$ is an odd integer!. All of these with the same initial condition, so that the solution is not unique.

If, however, one also requires continuity of $\mathrm{d}_{y} g$ in $y$ (which was not true at $x=0$ in our example), the Picard-Lindelöf theorem guarantees unicity. Briefly, to derive this solution one first rewrites eq. (6.2) in the form:

$$
y(x)=y_{0}+\int_{x_{0}}^{x} g\left(x, y\left(x^{\prime}\right)\right) \mathrm{d} x^{\prime}
$$

Then, starting with $y_{0}=0$, one constructs the sequence of functions:

$$
y_{n+1}=y_{0}+\int_{x_{0}}^{x} g\left(x, y_{n}\left(x^{\prime}\right)\right) \mathrm{d} x^{\prime}
$$

This iterative process is called Picard's iteration method. Then one shows that the sequence is Cauchy (see module V ) to prove that it converges toward $y(x)$.

On the other hand, it is not difficult to obtain an explicit general solution to a first-order linear ODE. First, it is convenient to recast it into its normal form:

$$
\mathrm{d}_{x} f+\beta(x) f=F(x)
$$

Then one shows (EXERCISE) that if $\beta(x)$ is continuous over $(a, b)$, the general solution of this ODE is:

$$
\begin{equation*}
f(x)=\frac{i\left(x_{0}\right)}{i(x)} f\left(x_{0}\right)+\int_{x_{0}}^{x} F\left(x^{\prime}\right) \frac{i\left(x^{\prime}\right)}{i(x)} \mathrm{d} x^{\prime} \quad i(x)=\mathrm{e}^{\int^{x} \beta\left(x^{\prime}\right) \mathrm{d} x^{\prime}} \tag{6.3}
\end{equation*}
$$

$\forall x_{0} \in(a, b)$, and where $f\left(x_{0}\right)$ is arbitrary.

### 6.1.2 Second-order linear ODEs

Consider first the most general form for a homogeneous second-order equation:

$$
\begin{equation*}
L[f]=\alpha(x) \mathrm{d}_{x}^{2} f+\beta(x) \mathrm{d}_{x} f+\gamma(x) f=0 \tag{6.4}
\end{equation*}
$$

Note that since $\gamma$ is known, this is not an eigenvalue equation for $L$.
It can be shown (the proof is technical and not very illuminating) that in a finite interval $[a, b]$, the only solution for which $f$ and $\mathrm{d}_{x} f$ both vanish at the initial point $x=a$ is the trivial solution $f=0$. Consequently, if there exist two solutions $f$ and $g$ that satisfy the initial conditions $f(a)-g(a)=0$ and $f^{\prime}(a)-g^{\prime}(a)=0$, then $f=g$ over the whole interval. In other words, the two initial conditions uniquely specify the solution to a homogeneous second-order differential equation.

Let $f_{1}(x)$ and $f_{2}(x)$ be two solutions of eq. (6.4). If there exists no constant $C$ such that $f_{2}=C f_{1} \forall x, f_{1}$ and $f_{2}$ are linearly independent and any linear combination of them is also a solution of the homogeneous equation (principle of linear superposition). But one cannot have both $f_{1}=f_{2}$ and $f_{1}^{\prime}=f_{2}^{\prime}$ at any $x \in[a, b]$.

The Wronskian of two functions $f_{1}(x)$ and $f_{2}(x)$ is defined as: $W(x):=f_{1} f_{2}^{\prime}-f_{2} f_{1}^{\prime}$. It is easy to obtain a first-order differential equation for $W(x)$, whose solution (EXERCISE) is called Abel's formula:

$$
\begin{equation*}
W(x)=W\left(x_{0}\right) \mathrm{e}^{-\int_{x_{0}}^{x}\left[\beta\left(x^{\prime}\right) / \alpha\left(x^{\prime}\right)\right] \mathrm{d} x^{\prime}} \tag{6.5}
\end{equation*}
$$

where $x_{0}$ is any point in the interval $[a, b]$. This form of the Wronskian makes no direct reference to homogeneous solutions of eq. (6.4). We see that eq. (6.4) with $\beta(x)=0$ leads to constant Wronskians. We also see, because the exponential cannot vanish in a finite interval, that if the Wronskian vanishes anywhere, it vanishes everywhere.

And now comes a surprising fact: given one solution of the homogeneous equation, a general solution of the inhomogeneous equation valid over $[a, b]$ :

$$
\alpha(x) \mathrm{d}_{x}^{2} f+\beta(x) \mathrm{d}_{x} f+\gamma(x) f=F(x)
$$

can be generated. We give a sketch of the procedure, also known as the method of variation of parameters:

1. Given $f_{1}$, we seek a second linearly independent solution of the homogeneous equation. Noticing that:

$$
\frac{W(x)}{f_{1}^{2}(x)}=\left(\frac{f_{2}}{f_{1}}\right)^{\prime}
$$

and integrating, we find with eq. (6.5) that:

$$
\begin{equation*}
f_{2}(x)=f_{1}(x) \int_{a}^{x} \frac{W\left(x_{0}\right)}{f_{1}^{2}\left(x^{\prime}\right)} \mathrm{e}^{-\int^{x^{\prime}}(\beta / \alpha) \mathrm{d} x^{\prime \prime}} \mathrm{d} x^{\prime} \tag{6.6}
\end{equation*}
$$

is also a solution. If the result still contains a part that is proportional to $f_{1}$, discarding it leaves a solution that is linearly independent from $f_{1}$. Note that here $W\left(x_{0}\right)$ is an arbitrary non-zero constant.
2. Having generated a second solution to the homogeneous equation, let us obtain a particular solution $h(x)$ to the inhomogeneous equation. The key step is to insert $h(x)=f_{1}(x) g(x)$ to obtain (EXERCISE) a first-order equation for $g^{\prime}$. Using the general first-order equation solution eq. (6.3) and (6.5) again, there comes:

$$
g^{\prime}(x)=\left(\frac{f_{2}}{f_{1}}\right)^{\prime} \int_{a}^{x} \frac{f_{1}\left(x^{\prime}\right) F\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right) W\left(x^{\prime}\right)} \mathrm{d} x^{\prime}
$$

A straightforward integration then leads to:

$$
\begin{equation*}
h(x)=f_{1} g=f_{2}(x) \int_{a}^{x} \frac{f_{1}\left(x^{\prime}\right) F\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right) W\left(x^{\prime}\right)} \mathrm{d} x^{\prime}-f_{1}(x) \int_{a}^{x} \frac{f_{2}\left(x^{\prime}\right) F\left(x^{\prime}\right)}{\alpha\left(x^{\prime}\right) W\left(x^{\prime}\right)} \mathrm{d} x^{\prime} \tag{6.7}
\end{equation*}
$$

The general solution of the inhomogeneous equation is then a linear combination of $f_{1}$ and $f_{2}$ (homogeneous solution) plus this particular solution. The coefficients of the linear combination are fixed by boundary (2-point) or initial (1-point) conditions which we shall soon discuss.

Let us note also that provided $\int \beta(x) / \alpha(x) \mathrm{d} x$ exists within the interval of interest, it is always possible to eliminate the first-order derivative term in any linear second-order ODE, by a redefinition of the form $f(x)=$ $g(x) \mathrm{e}^{\mu(x)}$ (the substitution $f(x)=\mu(x) g(x)$ also works), to arrive at (EXERCISE):

$$
\begin{equation*}
\alpha(x) g^{\prime \prime}(x)+\left(\gamma(x)-\frac{\beta^{\prime}}{2}+\frac{\beta}{2} \frac{\alpha^{\prime}}{\alpha}-\frac{1}{4} \frac{\beta^{2}}{\alpha}\right) g(x)=F(x) \exp \left[\int^{x} \frac{\beta}{2 \alpha} \mathrm{~d} x^{\prime}\right] \tag{6.8}
\end{equation*}
$$

as determined by the requirement that the transformed ODE have no first-order derivative in $g$.
One would do well to remember, however, that making an ODE look simpler does not always make it easier to solve!

### 6.2 Solving One-dimensional Equations with Green's Functions (BF 7.3)

### 6.2.1 Solutions in terms of Green's Functions and Boundary Conditions

We focus on inhomogeneous second-order linear equations of the type:

$$
L[f(t)]=\alpha(t) \mathrm{d}_{t}^{2} f(t)+\beta(t) \mathrm{d}_{t} f(t)+\gamma(t) f(t)=F(t)
$$

with $\alpha, \beta$ and $\gamma$ continuous all over the interval of interest. Such ordinary (one-dimensional) differential equations often have time as the independent variable, and in physics $F(t)$ is then called a driving term. Of course, its solutions will be twice (and most often more than twice) differentiable.

Obtaining a very useful expression for a solution to this equation is not so difficult if $L$ admits a complete set of orthonormal eigenfunctions $\phi_{j}$ on the interval (which is the case if $L$ is self-adjoint), with real eigenvalues $\lambda_{j} \neq 0$. Then $f$ can be expanded over the set (with unknown coefficients $a_{j}$ ), and so can $F$, with known coefficients $b_{j}$. Both sets of coefficients are, as usual, projections of $f$ and $F$ on the eigenfunctions. The eigenvalue equation then yields a relation between them, and there comes (EXERCISE), assuming that integral and summation signs can be interchanged) the inhomogeneous solution:

$$
f(t)=\int\left[\sum_{j} \frac{\phi_{j}(t) \phi_{j}^{*}\left(t^{\prime}\right)}{\lambda_{j}}\right] F\left(t^{\prime}\right) \mathrm{d} t^{\prime}=\int G\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}
$$

where we have identified a two-point function:

$$
\begin{equation*}
G\left(t, t^{\prime}\right)=\sum_{j} \frac{\phi_{j}(t) \phi_{j}^{*}\left(t^{\prime}\right)}{\lambda_{j}} \quad \lambda_{j} \neq 0 \tag{6.9}
\end{equation*}
$$

Acting on $G\left(t, t^{\prime}\right)$ with $L$ yields an associated differential equation:

$$
L\left[G\left(t, t^{\prime}\right)\right]=\sum_{j} \frac{L\left[\phi_{j}(t)\right] \phi_{j}^{*}\left(t^{\prime}\right)}{\lambda_{j}}=\sum_{j} \phi_{j}(t) \phi_{j}^{*}\left(t^{\prime}\right)=\delta\left(t-t^{\prime}\right)
$$

where the last equality comes from the completeness relation for the eigenfunctions. The defining equation for a Green function $G\left(t, t^{\prime}\right)$ is:

$$
\begin{equation*}
L\left[G\left(t, t^{\prime}\right)\right]=\delta\left(t-t^{\prime}\right) \tag{6.10}
\end{equation*}
$$

Suppose that we know two linearly independent solutions, $f_{1}(t)$ and $f_{2}(t)$, to the associated homogeneous equation $L[f(t)]=0$. Then, if boundary conditions allow the homogeneous solution to exist, the general solution is the sum of the general solution to the homogeneous equation plus a solution to the inhomogeneous equation.

In 1-dim problems, boundary conditions are generally of two types:
(1) One-Point (Initial) conditions, aka IVP, or Initial-Value Problem: $f$ and its first-order derivative are known at some time, usually called the initial time $t_{0}$. Then the general solution of our inhomogeneous equation can be written:

$$
\begin{equation*}
f(t)=A f_{1}(t)+B f_{2}(t)+\int_{t_{0}}^{t} G_{\mathrm{ivp}}\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime} \tag{6.11}
\end{equation*}
$$

where $A$ and $B$ are two constants to be determined from the initial conditions, and the Green function $G_{\mathrm{ivp}}\left(t, t^{\prime}\right)$ associated with a given differential operator $L$ corresponds to this specific type of boundary conditions. We see that: $f\left(t_{0}\right)=A f_{1}\left(t_{0}\right)+B f_{2}\left(t_{0}\right)$, and $\dot{f}\left(t_{0}\right)=\left[A \dot{f}_{1}(t)+B \dot{f}_{2}(t)\right]_{t_{0}}$, so that $A$ and $B$ do not depend on the inhomogeneous term $F(t)$. With the (always non-zero) Wronskian of $f_{1}$ and $f_{2}$ : $W \equiv f_{1} \dot{f}_{2}-\dot{f}_{1} f_{2}$, they are found to be:

$$
A=\frac{\dot{f}_{2}\left(t_{0}\right) f\left(t_{0}\right)-f_{2}\left(t_{0}\right) \dot{f}\left(t_{0}\right)}{W\left(t_{0}\right)}, \quad B=-\frac{\dot{f}_{1}\left(t_{0}\right) f\left(t_{0}\right)-f_{1}\left(t_{0}\right) \dot{f}\left(t_{0}\right)}{W\left(t_{0}\right)}
$$

In time-evolution problems with initial conditions, we most often demand that $G_{\mathrm{ivp}}\left(t, t^{\prime}\right)=0$ for $t^{\prime}>t$ so as to preserve causality. This allows us, if we wish, to extend the upper bound of the integral all the way to infinity.
(2) Two-point boundary conditions, or Boundary-Value Problem: In one example of this case, $f\left(t_{0}\right)$ and $f\left(t_{1}\right)$ $\left(t_{1}>t_{0}\right)$ are known (the case for which the derivatives of $f$ are known will be addressed in section 6.2.3). Then the general solution would be written:

$$
\begin{equation*}
f(t)=A f_{1}(t)+B f_{2}(t)+\int_{t_{0}}^{t_{1}} G_{\mathrm{bvp}}\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime} \tag{6.12}
\end{equation*}
$$

Note the different upper limit of integration, with the integral still a function of $t$. This time, $A$ and $B$ do not depend on the inhomogeneous term if the appropriate Green function always obeys the homogeneous boundary conditions: $G_{\mathrm{bvp}}\left(t_{0}, t^{\prime}\right)=G_{\mathrm{bvp}}\left(t_{1}, t^{\prime}\right)=0$; we shall prove that this is indeed the case below in section 6.2.3. If such a $G_{\mathrm{bvp}}$ is to exist, we see from eq. (6.9) that $L$ cannot have a zero eigenvalue; in other words, there can be no non-trivial solution of the homogenous equation $L[f]=0$ that satisfies homogeneous B.C. $f\left(t_{0}\right)=f\left(t_{1}\right)=0$. And indeed, in that case one easily computes $A=B=0$ in eq. (6.12).

Moreover, symbolically writing the inhomogeneous solution as $f(t)=L_{t}^{-1} F(t)=\int G\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}$, we see that knowledge of the Green function amounts to inverting the inhomogeneous equation, and gives us an inhomogeneous solution in the form of an integral. Our interest in Green functions entirely resides in this property.

### 6.2.2 A General Expression for 1-d Green Functions for $L[f(t)]=F(t)$

What restrictions does eq. (6.10) impose on $G\left(t, t^{\prime}\right)$ ? Three, in fact:
(a) $G\left(t, t^{\prime}\right)$ is a continuous function of $t$ everywhere, including at $t=t^{\prime}$, otherwise its second derivative at $t=t^{\prime}$ would be the derivative of a $\delta$-function, and the differential equation would not be satisfied. Note, however, that the Green function for a first-order operator can be discontinuous, eg., $L=-\mathrm{id}_{t}$ has as Green function the step-function $\mathrm{i} \theta\left(t-t^{\prime}\right)$.
(b) $\dot{G}$ must have a discontinuity at $t=t^{\prime}$. To see this, integrate eq. (6.10) from $t=t^{\prime}-\epsilon$ to $t=t^{\prime}+\epsilon$. Since the coefficients in $L$ are continuous, they hardly vary when the interval is taken to be arbitrarily small $(\epsilon \rightarrow 0)$. In that limit, the integrals of $G$ and $\dot{G}$ both vanish because $G$ is continuous, and we are left with the contribution of the integral of $\ddot{G}$ :

$$
\left.\lim _{\epsilon \rightarrow 0} \dot{G}\left(t, t^{\prime}\right)\right|_{t=t^{\prime}-\epsilon} ^{t=t^{\prime}+\epsilon}=\frac{1}{\alpha\left(t^{\prime}\right)}
$$

So, when the coefficient of the second-order derivative in $L$ is 1 , the derivative of $G$ must jump by 1 at $t=t^{\prime}$.
(c) When $t \neq t^{\prime}, G$ must satisfy the homogeneous equation $L\left[G\left(t, t^{\prime}\right)\right]=0$, and can be written in terms of $f_{1}$ and $f_{2}$ :

$$
G\left(t, t^{\prime}\right)= \begin{cases}a_{1}\left(t^{\prime}\right) f_{1}(t)+a_{2}\left(t^{\prime}\right) f_{2}(t) & t>t^{\prime} \\ b_{1}\left(t^{\prime}\right) f_{1}(t)+b_{2}\left(t^{\prime}\right) f_{2}(t) & t<t^{\prime}\end{cases}
$$

The continuity of $G$ and the expression for the discontinuity in $\dot{G}$ at $t=t^{\prime}$ then yield:

$$
\begin{aligned}
& a_{1} f_{1}\left(t^{\prime}\right)+a_{2} f_{2}\left(t^{\prime}\right)-b_{1} f_{1}\left(t^{\prime}\right)-b_{2} f_{2}\left(t^{\prime}\right)=0 \\
& a_{1} \dot{f}_{1}\left(t^{\prime}\right)+a_{2} \dot{f}_{2}\left(t^{\prime}\right)-b_{1} \dot{f}_{1}\left(t^{\prime}\right)-b_{2} \dot{f}_{2}\left(t^{\prime}\right)=\frac{1}{\alpha\left(t^{\prime}\right)}
\end{aligned}
$$

These can be solved for $a_{1}$ and $a_{2}$ in terms of $b_{1}$ and $b_{2}$ :

$$
\begin{aligned}
& a_{1}\left(t^{\prime}\right)=b_{1}\left(t^{\prime}\right)-\frac{f_{2}\left(t^{\prime}\right)}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)} \\
& a_{2}\left(t^{\prime}\right)=b_{2}\left(t^{\prime}\right)+\frac{f_{1}\left(t^{\prime}\right)}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)}
\end{aligned}
$$

where the Wronskian of $f_{1}$ and $f_{2}$ is $W \equiv f_{1} \dot{f}_{2}-\dot{f}_{1} f_{2} \neq 0$ since $f_{1}$ and $f_{2}$ are linearly independent. Then the Green function for $L$ must take the form:

$$
G\left(t, t^{\prime}\right)= \begin{cases}b_{1}\left(t^{\prime}\right) f_{1}(t)+b_{2}\left(t^{\prime}\right) f_{2}(t)-\frac{f_{1}(t) f_{2}\left(t^{\prime}\right)-f_{2}(t) f_{1}\left(t^{\prime}\right)}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)} & t>t^{\prime}  \tag{6.13}\\ b_{1}\left(t^{\prime}\right) f_{1}(t)+b_{2}\left(t^{\prime}\right) f_{2}(t) & t<t^{\prime}\end{cases}
$$

The term with the Wronskian vanishes at $t=t^{\prime}$, guaranteeing the continuity of $G$ as required.
The adjustable parameters $b_{1}$ and $b_{2}$ can now be chosen so that $G$ satisfies suitable boundary conditions. In the case (IVP) where initial conditions are specified on $f(t)$ and $\dot{f}(t)$, requiring $G\left(t, t^{\prime}\right)=0$ for $t<t^{\prime}$ immediately leads to $b_{1}=b_{2}=0$, and there comes the simple but very general expression:

$$
\begin{equation*}
G_{\mathrm{ivp}}\left(t, t^{\prime}\right)=\theta\left(t-t^{\prime}\right) \frac{f_{2}(t) f_{1}\left(t^{\prime}\right)-f_{1}(t) f_{2}\left(t^{\prime}\right)}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)} \tag{6.14}
\end{equation*}
$$

where the step-function, or Heaviside function, $\theta(u)$ is defined ${ }^{*}$ by:

$$
\theta(u)= \begin{cases}1 & u>0  \tag{6.15}\\ 0 & u \leq 0\end{cases}
$$

If the initial conditions are homogeneous, ie., if $f(0)=\left.\dot{f}\right|_{0}=0$, there is no non-trivial homogeneous solution for $f$, and $\int G_{\mathrm{ivp}}\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}$ in eq. (6.11), which can always be evaluated, either analytically or numerically, is the general solution to the inhomogeneous equation $L[f]=F$.

On the other hand, for two-point boundary conditions (BVP) at $t_{0}$ and $t_{1}, G\left(t_{0}, t^{\prime}\right)=0\left(t_{0}<t^{\prime}\right)$ immediately leads to: $b_{2}\left(t^{\prime}\right)=-b_{1}\left(t^{\prime}\right) f_{1}\left(t_{0}\right) / f_{2}\left(t_{0}\right)$, whereas $G\left(t_{1}, t^{\prime}\right)=0\left(t_{1}>t^{\prime}\right)$ gives:
$b_{1}\left(t^{\prime}\right)=\frac{f_{2}\left(t_{0}\right)}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)} \frac{f_{1}\left(t_{1}\right) f_{2}\left(t^{\prime}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t^{\prime}\right)}{f_{1}\left(t_{1}\right) f_{2}\left(t_{0}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t_{0}\right)} \quad \Longrightarrow \quad b_{2}\left(t^{\prime}\right)=\frac{f_{1}\left(t_{0}\right)}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)} \frac{f_{2}\left(t_{1}\right) f_{1}\left(t^{\prime}\right)-f_{1}\left(t_{1}\right) f_{2}\left(t^{\prime}\right)}{f_{1}\left(t_{1}\right) f_{2}\left(t_{0}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t_{0}\right)}$
The BVP Green function factorises (EXERCISE) in $t$ and $t^{\prime}$ :

$$
G_{\mathrm{bvp}}\left(t, t^{\prime}\right)=\frac{1}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)} \frac{\left[f_{1}\left(t_{1}\right) f_{2}\left(t_{>}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t_{>}\right)\right]\left[f_{1}\left(t_{<}\right) f_{2}\left(t_{0}\right)-f_{1}\left(t_{0}\right) f_{2}\left(t_{<}\right)\right]}{f_{1}\left(t_{1}\right) f_{2}\left(t_{0}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t_{0}\right)}
$$

where $t_{>}:=\max \left(t, t^{\prime}\right)$ and $t_{<}:=\min \left(t, t^{\prime}\right)$. If $f_{1}\left(t_{0}\right)=0$, then:

$$
G_{\mathrm{bvp}}\left(t, t^{\prime}\right)=\frac{1}{\alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)}\left[f_{2}\left(t_{>}\right) f_{1}\left(t_{<}\right)-\frac{f_{2}\left(t_{1}\right)}{f_{1}\left(t_{1}\right)} f_{1}\left(t_{>}\right) f_{1}\left(t_{<}\right)\right]
$$

The most simple case occurs when $f_{1}\left(t_{0}\right)=f_{2}\left(t_{1}\right)=0$ (homogeneous B.C.); then: $G_{\mathrm{bvp}}\left(t, t^{\prime}\right)=f_{1}\left(t_{<}\right) f_{2}\left(t_{>}\right) / \alpha\left(t^{\prime}\right) W\left(t^{\prime}\right)$.
Note that, unlike for the IVP problem, a Green function for the one-dim BVP for a given $L$ exists only if $b_{1}$ and $b_{2}$ are finite, ie. if $f_{1}\left(t_{1}\right) f_{2}\left(t_{0}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t_{0}\right) \neq 0$. Also, if the boundary conditions allow the existence of a homogeneous solution, ie. an eigenfunction of $L$ with eigenvalue zero satisfying these same conditions, eq. (6.9) forbids a Green function.

## Example 6.1. A Simple Example: the Harmonic Oscillator

Consider the undamped oscillator operator $\mathrm{d}_{t}^{2}+\omega_{0}^{2}$ with initial conditions on $f$ and $\dot{f}$ at a single point (IVP). We choose $f_{1}=\sin \omega_{0} t$ and $f_{2}=\cos \omega_{0} t$. Also, noting that $\alpha=1$ and $W=-\omega_{0}$, eq. (6.13) yields, with $b_{1}=b_{2}=0$, the IVP Green function:

$$
G_{\mathrm{ivp}}\left(t, t^{\prime}\right)=G_{\mathrm{ivp}}\left(t-t^{\prime}\right)=\theta\left(t-t^{\prime}\right) \frac{\sin \left[\omega_{0}\left(t-t^{\prime}\right)\right]}{\omega_{0}}
$$

[^21]Note that the dependence of the IVP Green function is on the difference ${ }^{\dagger} t-t^{\prime}$. Indeed, it can be shown (EXERCISE) that for the second-order linear differential equation: $L[f(t)]=F(t)$ with constant coefficients, Green functions for a one-dimensional IVP must satisfy $G\left(t, t^{\prime}\right)=G\left(t-t^{\prime}\right)$, just by using the general form of the homogeneous solutions: $f_{ \pm}(t)=\mathrm{e}^{\lambda_{ \pm} t}$. This is a manifestation of the invariance of the differential operator with constant coefficients under translations of the variable $t$ (eg. time). $G\left(t-t^{\prime}\right)$ can be viewed as a convolution operator or a response function in the sense of section 4.7.
By contrast, for the same $L$ but with a BVP at $t_{0}=0$ and $t_{1}$, we immediately obtain:

$$
b_{1}\left(t^{\prime}\right)=\frac{\cot \omega_{0} t_{1} \sin \omega_{0} t^{\prime}-\cos \omega_{0} t^{\prime}}{\omega_{0}}=-\frac{\sin \omega_{0}\left(t_{1}-t^{\prime}\right)}{\omega_{0} \sin \omega_{0} t_{1}}
$$

and, with $t_{>}:=\max \left(t, t^{\prime}\right), t_{<}:=\min \left(t, t^{\prime}\right)$, eq. (6.13) gives (EXERCISE):

$$
\begin{equation*}
G_{\mathrm{bvp}}\left(t, t^{\prime}\right)=\frac{1}{\omega_{0} \sin \omega_{0} t_{1}} \sin \left[\omega_{0}\left(t_{>}-t_{1}\right)\right] \sin \omega_{0} t_{<} \tag{6.16}
\end{equation*}
$$

and a particular solution to $\left(\mathrm{d}_{t}^{2}+\omega_{0}^{2}\right) f(t)=F(t)$ is:

$$
f(t)=\frac{\sin \left[\omega_{0}\left(t-t_{1}\right)\right]}{\omega_{0} \sin \omega_{0} t_{1}} \int_{t_{0}}^{t} \sin \omega_{0} t^{\prime} F\left(t^{\prime}\right) \mathrm{d} t^{\prime}+\frac{\sin \omega_{0} t}{\omega_{0} \sin \omega_{0} t_{1}} \int_{t}^{t_{1}} \sin \left[\omega_{0}\left(t^{\prime}-t_{1}\right)\right] F\left(t^{\prime}\right) \mathrm{d} t^{\prime}
$$

If $\omega_{0} t_{1}=n \pi(n \in \mathbb{Z})$, ie. if $t_{1}$ is an integer multiple of the half-period, the condition for the existence of a BVP Green function, $f_{1}\left(t_{1}\right) f_{2}\left(t_{0}\right)-f_{2}\left(t_{1}\right) f_{1}\left(t_{0}\right) \neq 0$, is violated.

Example 6.2. An interesting feature of eq. (6.13) is that the explicit dependence of the Green functions on the form of $L$ is only through the coefficient of the second-order derivative. This can be exploited to minimise the work required to find the Green functions for a damped oscillator operator, $L=\mathrm{d}_{t}^{2}+$ $2 \gamma \mathrm{~d}_{t}+\omega_{0}^{2}$, by using instead the appropriate homogeneous solutions: $f_{1}(t)=\mathrm{e}^{-\gamma t} \sin \left[\sqrt{\omega_{0}^{2}-\gamma^{2}} t\right]$, $f_{2}(t)=\mathrm{e}^{-\gamma t} \cos \left[\sqrt{\omega_{0}^{2}-\gamma^{2}} t\right]$. Now $W=-\sqrt{\omega_{0}^{2}-\gamma^{2}} \mathrm{e}^{-2 \gamma t^{\prime}}$, and a straightforward substitution into eq. (6.13) for an IVP ( $b_{1}=b_{2}=0$ ) gives:

$$
\begin{equation*}
G_{(3)}\left(t, t^{\prime}\right)=\theta\left(t-t^{\prime}\right) \mathrm{e}^{-\gamma\left(t-t^{\prime}\right)} \frac{\sin \left[\sqrt{\omega_{0}^{2}-\gamma^{2}}\left(t-t^{\prime}\right)\right]}{\sqrt{\omega_{0}^{2}-\gamma^{2}}} \tag{6.17}
\end{equation*}
$$

Example 6.3. While we are talking about the damped harmonic oscillator, let us use it to illustrate another way to solve differential equations that combines Fourier and Green techniques. The idea is to write the equation:

$$
\ddot{f}(t)+2 \gamma \dot{f}(t)+\omega_{0}^{2} f(t)=F(t)
$$

in the frequency domain, assuming that the driving force dies at $t \rightarrow \pm \infty$ or, alternatively, is turned on at, say, $t=0$, and then off at some later time. In this case the Fourier transform of $F(t)$ exists, and eq. (5.14) can be used to obtain:

$$
f(\omega)=\frac{F(\omega)}{-\omega^{2}+2 \mathrm{i} \gamma \omega+\omega_{0}^{2}}
$$

The differential equation in the time domain becomes an algebraic equation in the frequency domain! To go back to the time domain, we just write a solution to the inhomogeneous equation:

$$
\begin{aligned}
f(t) & =\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} f(\omega) \mathrm{e}^{\mathrm{i} \omega t} \mathrm{~d} \omega=\int\left[\frac{1}{2 \pi} \int \frac{\mathrm{e}^{\mathrm{i} \omega\left(t-t^{\prime}\right)} \mathrm{d} \omega}{-\omega^{2}+2 \mathrm{i} \gamma \omega+\omega_{0}^{2}}\right] F\left(t^{\prime}\right) \mathrm{d} t^{\prime} \\
& =\int_{-\infty}^{\infty} G\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}
\end{aligned}
$$

[^22]where:
$$
G\left(t, t^{\prime}\right)=G\left(t-t^{\prime}\right)=-\frac{1}{2 \pi} \int_{-\infty}^{\infty} \frac{\mathrm{e}^{\mathrm{i} \omega\left(t-t^{\prime}\right)}}{\left(\omega-\omega_{+}\right)\left(\omega-\omega_{-}\right)} \mathrm{d} \omega
$$
with $\omega_{ \pm}= \pm \sqrt{\omega_{0}^{2}-\gamma^{2}}+\mathrm{i} \gamma$.
To calculate $G$ for $t>t^{\prime}$, we use contour integration in the complex $\omega$ plane, with the contour $C$ chosen to be counterclockwise around the upper infinite half-plane. Both poles $\omega=\omega_{ \pm}$lie in the upper half-plane. Breaking up the contour into the real axis plus the semi-circle at infinity, we have:
$$
-\frac{1}{2 \pi} \int_{-\infty}^{\infty} \frac{\mathrm{e}^{\mathrm{i} \omega\left(t-t^{\prime}\right)}}{\left(\omega-\omega_{+}\right)\left(\omega-\omega_{-}\right)} \mathrm{d} \omega=-\frac{1}{2 \pi} \oint_{C} \frac{\mathrm{e}^{\mathrm{i} \omega\left(t-t^{\prime}\right)}}{\left(\omega-\omega_{+}\right)\left(\omega-\omega_{-}\right)} \mathrm{d} \omega+\frac{1}{2 \pi} \int_{|\omega| \rightarrow \infty} \frac{\mathrm{e}^{\mathrm{i} \omega\left(t-t^{\prime}\right)}}{\left(\omega-\omega_{+}\right)\left(\omega-\omega_{-}\right)} \mathrm{d} \omega
$$

With $t-t^{\prime}>0$, the numerator in the second integral on the right goes to zero as $|\omega| \rightarrow \infty$, and the integral vanishes. The contour integral is evaluated with the Residue theorem (4.34). Then:

$$
\begin{aligned}
G\left(t-t^{\prime}\right) & =-\frac{1}{2 \pi} \int_{-\infty}^{\infty} \frac{\mathrm{e}^{\mathrm{i} \omega\left(t-t^{\prime}\right)}}{\left(\omega-\omega_{+}\right)\left(\omega-\omega_{-}\right)} \mathrm{d} \omega=2 \pi \mathrm{i} \frac{-1}{2 \pi}\left(\frac{\mathrm{e}^{\mathrm{i} \omega_{+}\left(t-t^{\prime}\right)}}{\omega_{+}-\omega_{-}}-\frac{\mathrm{e}^{\mathrm{i} \omega_{-}\left(t-t^{\prime}\right)}}{\omega_{+}-\omega_{-}}\right) \\
& =\mathrm{e}^{-\gamma\left(t-t^{\prime}\right)} \frac{\sin \left[\sqrt{\omega_{0}^{2}-\gamma^{2}}\left(t-t^{\prime}\right)\right]}{\sqrt{\omega_{0}^{2}-\gamma^{2}}}
\end{aligned}
$$

When $t-t^{\prime}<0$, we must use a contour enclosing the lower infinite half-plane. But the integrand in the contour integral is analytic in this region, and the integral vanishes by the Cauchy-Goursat theorem. Thus, $G\left(t, t^{\prime}\right)=0$ for $t<t^{\prime}$, and we have recovered the result obtained in eq. (6.17). Unlike that approach, here no knowledge of the homogeneous solutions was needed to find the Green function!
The general solution follows by adding a linear combination of homogeneous solutions.

### 6.2.3 Green's second 1-dim identity and general solution in terms of Green functions

Consider the expression, quadratic and symmetric in $u(t)$ and $v(t): Q(u, v)=\alpha(t) \dot{u} \dot{v}+\lambda(t)(\dot{u} v+u \dot{v})-$ $\mu(t) u v$. Then, subtracting $Q(v, u)$, one derives after a few manipulations (EXERCISE) Lagrange's identity:

$$
v L[u]-u L[v]=\mathrm{d}_{t}[v(\alpha(\dot{u}-u \dot{v})]
$$

where $L[u]=\mathrm{d}_{t}(\alpha \dot{u})+(\dot{\lambda}+\mu) u$. Integrate this identity over an interval $\left[t_{0}, t_{1}\right]$ to obtain Green's second identity (1-d):

$$
\begin{equation*}
\int_{t_{0}}^{t_{1}}(v L[u]-u L[v]) \mathrm{d} t^{\prime}=\left.\alpha(v \dot{u}-u \dot{v})\right|_{t_{0}} ^{t_{1}} \tag{6.18}
\end{equation*}
$$

Now suppose that $u=G\left(t, t^{\prime}\right)$ and that $v=f\left(t^{\prime}\right)$ satisfies the inhomogeneous equation $L\left[f\left(t^{\prime}\right)\right]=F\left(t^{\prime}\right)$. Then one easily shows from Green's identity that for $t \in\left[t_{0}, t_{1}\right]$,

$$
\begin{equation*}
f(t)=\int_{t_{0}}^{t_{1}} G\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}-\left[\alpha\left(G \partial_{t^{\prime}} f-f \partial_{t^{\prime}} G\right)\right]_{t^{\prime}=t_{0}}^{t^{\prime}=t_{1}} \tag{6.19}
\end{equation*}
$$

where $G\left(t, t^{\prime}\right)$ is a Green function for the self-adjoint differential operator $L$. We are already familiar with the first (inhomogeneous) term, but the second one warrants careful examination. Obviously, it must be related to the homogeneous solutions. But wait-is $f(t)$ actually the general solution? Not yet! It is still just an identity. The second term is evaluated at the end-points of the interval, so it depends on the boundary conditions for $f$. We cannot freely specify $f$ and $\dot{f}$ at both $t_{0}$ and $t_{1}$ as this would be in general inconsistent. If $f$ if specified at the end-points, then we must find the solution for $f$ in order to know what its derivatives are at the end-points.

But we can use the fact that any function $\tilde{G}\left(t, t^{\prime}\right)$ which satisfies $L[\tilde{G}]=0$ can always be added to $G$ to choose different types of boundary conditions for the Green function. For instance, specifying $f$ at $t_{0}$ and $t_{1}$, we can in
principle find a $\tilde{G}$ such that the new $G$ vanishes at $t_{0}$ and $t_{1}$. This kind of boundary-value problem, where the solution (not its derivative) is specified on the boundary, is called a Dirichlet problem, with the corresponding Green function denoted $G_{\mathrm{D}}$. Incidentally, this justifies our earlier setting of $G$ to zero at the end-points. Now we can write the general solution for the two-point (Dirichlet) problem:

$$
\begin{equation*}
f(t)=\int_{t_{0}}^{t_{1}} G_{\mathrm{D}}\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}+\left[\alpha f \partial_{t^{\prime}} G_{\mathrm{D}}\right]_{t^{\prime}=t_{0}}^{t^{\prime}=t_{1}} \quad G_{\mathrm{D}}\left(t, t_{0}\right)=G_{\mathrm{D}}\left(t, t_{1}\right)=0 \tag{6.20}
\end{equation*}
$$

Compare this form of the general solution, which explicitly depends only on $F(t)$ and $G_{\mathrm{D}}$, plus $f\left(t_{0}\right)$ and $f\left(t_{1}\right)$, to the solution of eq. (6.12) in terms of the homogeneous solutions. It is a very instructive EXERCISE to show their equivalence. We also see that if $f$ happens to obey homogeneous B.C., $f\left(t_{0}\right)=f\left(t_{1}\right)=0$, there is no homogeneous part, consistent with our earlier statement near the end of section 6.2.1 that the existence of $G_{\mathrm{bvp}}$ is conditional on this absence.

We could instead want to specify $\dot{f}$ at $t_{0}$ and $t_{1}$. This Neumann problem has a solution if it is possible to set $\left.\dot{G}_{\mathrm{N}}\right|_{t_{0}, t_{1}}=0$ (or it has a solution up to a constant if $\left.\dot{G}_{\mathrm{N}}\right|_{t_{0}, t_{1}}$ are constants). It is not nearly as frequent in physics as Dirichlet problems. Then:

$$
\left.f(t)=\int_{t_{0}}^{t_{1}} G_{\mathrm{N}}\left(t, t^{\prime}\right) F\left(t^{\prime}\right) \mathrm{d} t^{\prime}-\left.\left[\alpha G_{\mathrm{N}} \partial_{t^{\prime}} f\right]_{t^{\prime}=t_{0}}^{t^{\prime}=t_{1}} \quad \quad \dot{G}_{\mathrm{N}}\left(t, t^{\prime}\right)\right|_{t^{\prime}=t_{0}}=\dot{G}_{\mathrm{N}}\left(t, t^{\prime}\right)\right)\left.\right|_{t^{\prime}=t_{1}}=0
$$

For this to be possible, it will often be necessary to add an extra term to the defining equation of the Green function (more about this when we look at 3-dim Green functions). Much more exotic, we might want to specify $f$ at one end-point and $\dot{f}$ at the other (so-called Robin B.C.); the B.C. on Green's function can be worked out by using Green's second identity.

I would argue that eq. (6.20) is a much nicer representation of the general solution to a Dirichlet problem than the more standard eq. (6.12). First, it has contributions from both volume and boundary terms. Second, the Dirichlet boundary conditions on $f$ (which are really on the homogeneous solution since $G_{\mathrm{D}}\left(t, t^{\prime}\right)=0$ at $t=t_{0}$ and $t=t_{1}$ ) appear explicitly in the solution. Third, unlike eq. (6.12), there is no explicit dependence on $f_{1}$ and $f_{2}$, and thus no need to calculate integration constants. It is true that $f_{1}$ and $f_{2}$ are still needed if $G_{\mathrm{D}}\left(t, t^{\prime}\right)$ is found via eq. (6.13), but we have seen in example 6.3 that Green functions can be calculated directly. Also, contrary to eq. (6.12), eq. (6.20) can be generalised to higher dimensions. Note, however, that this particular approach is not suited to an IVP, in which B.C. are specified at one point.

One important property of Dirichlet Green functions may be derived by letting $v=G_{\mathrm{D}}\left(t^{\prime}, t\right)$ and $u=$ $G_{\mathrm{D}}\left(t^{\prime}, t^{\prime \prime}\right)$ in Green's second 1-dim identity (6.18), which holds for differential operators of the form $L=$ $\mathrm{d}_{t}\left(\alpha \mathrm{~d}_{t}\right)+\gamma$. Because $G_{\mathrm{D}}=0$ at the end-points and $L\left[G\left(t, t^{\prime}\right)\right]=\delta\left(t-t^{\prime}\right)$, we immediately find that $G_{\mathrm{D}}$ for such operators is symmetric in its arguments:

$$
\begin{equation*}
G_{\mathrm{D}}\left(t, t^{\prime \prime}\right)=G_{\mathrm{D}}\left(t^{\prime \prime}, t\right) \tag{6.21}
\end{equation*}
$$

The differential operator for the undamped harmonic oscillator is indeed of the right form (with $\alpha=1$ ); we expect that $G_{\mathrm{D}}$ will be symmetric, and $G_{(1)}$ in eq. (6.16) is indeed symmetric. This property can provide a useful check on calculations.

In closing this treatment of 1-dim Green functions, let us remark that whenever $G\left(t, t^{\prime}\right)=G\left(t-t^{\prime}\right)$, the volume integral $\int F\left(t^{\prime}\right) G\left(t-t^{\prime}\right) \mathrm{d} t^{\prime}$ can be seen as an example of the convolution integral introduced in section 5.5.4.

## Problems in More than One Dimension (BF 7.4)

In one dimension, Green's function for a second-order linear differential operator $L$ always exists and is unique for an IVP. If it exists for a BVP (no zero eigenvalue for $L$ ), it is unique. This is closely related to the fact that boundary conditions are specified at one or two points only. In two or more dimensions, the boundaries contain an infinite number of points, and Green functions are no longer guaranteed to exist, even for an IVP, But they do exist in important cases of physical interest.

### 6.3 Differential Equations with Partial Derivatives

Unless you are working on superstrings, you will find that it is sufficient to study PDEs in no more than four dimensions ${ }^{\dagger}$.

In accordance with modern usage, we shall use Greek indices in four-dimensional (three spatial and one time) problems, and roman indices in three spatial dimensions. We also implement the Einstein summation convention according to which repeated indices in factors are to be summed over; in any such pair, we will try to write one index as a superscript and one as a subscript so as to spot them more easily. Then the form of a second-order linear differential operator that we shall use is:

$$
\begin{equation*}
L=\alpha^{\mu}(\mathbf{x}) \partial_{\mu}^{2}+\beta^{\nu}(\mathbf{x}) \partial_{\nu}+\gamma(\mathbf{x}) \tag{6.22}
\end{equation*}
$$

where $\mathbf{x}$ is the generalised position and it should be emphasised that Cartesian coordinates are implied. The coefficients are assumed to be continuous in $\mathbf{x}$.

We follow Hadamard (1923) and classify such equations according to the coefficients of the second-order derivatives:

Definition 6.1. - If at least one of the $\alpha^{\mu}$ vanishes at some point, the operator (and corresponding homogeneous PDE will be said to be parabolic (eg. heat equation, Schrödinger equation, in which there is no second-order time-derivative).

- If the sign of one $\alpha^{\mu}$ coefficient is different from all others, we say that $L$ is hyperbolic (expected in a pseudo-Euclidean spacetime, eg. with the wave equation).
- If all $\alpha^{\mu}$ coefficients have the same sign (expected in a Euclidean space), $L$ is elliptic (eg. Laplace and Helmholtz operators - static 3-dim problems).


### 6.4 Separation of Variables in Elliptic Problems

Since the Laplacian operator occurs in all elliptic problems, it is worth taking a closer look at it. Our first task is to separate it into two convenient parts; at the same time this will get us acquainted with a very powerful technique.

### 6.4.1 An Important and Useful 3-dim Differential Operator

To do this, we introduce the self-adjoint vector operators $-\mathrm{i} \boldsymbol{\nabla}$ and $\mathbf{L}=-\mathrm{i} \mathbf{x} \times \nabla$, or $L_{i}=-\mathrm{i} \mathrm{t}_{i j k} x^{j} \partial^{k}$, where $\epsilon_{i j k}$ is the completely antisymmetric Levi-Civita symbol, and summation over repeated indices is implied. With the identity: $\epsilon_{i j k} \epsilon^{i m n}=\delta_{j}{ }^{m} \delta_{k}{ }^{n}-\delta_{j}{ }^{n} \delta_{k}{ }^{m}$, the scalar product of $\mathbf{L}$ with itself is, in Cartesian coordinates:

$$
\begin{aligned}
\mathbf{L} \cdot \mathbf{L} & =-\epsilon_{i j k} \epsilon^{i m n} x^{j} \partial^{k} x_{m} \partial_{n} \\
& =-x^{j}\left(\partial_{j}+x_{j} \partial^{k} \partial_{k}-3 \partial_{j}-x_{k} \partial^{k} \partial_{j}\right)=-x^{j} x_{j} \partial^{k} \partial_{k}+2 x^{j} \partial_{j}-x^{j} \partial_{j}+x^{j} \partial_{j} x^{k} \partial_{k}
\end{aligned}
$$

Extracting the Laplacian and reverting to coordinate-free notation, there comes:

$$
\begin{equation*}
\nabla^{2}=-\frac{\mathbf{L}^{2}}{r^{2}}+\frac{1}{r}\left[\partial_{r}+\partial_{r}\left(r \partial_{r}\right)\right] \tag{6.23}
\end{equation*}
$$

[^23]The distance $r$ to the origin can be expressed in any coordinates we wish, yet this expression obviously wants to single out the direction along $\mathbf{x}=r \hat{\mathbf{n}}$ from the other two. Also, it would be nice if $\mathbf{L}$ only involved derivatives in directions perpendicular to $\hat{\mathbf{n}}$. This is most easily realised in a spherical coordinate system, since its radial coordinate naturally correpsonds to the direction along x ; the other two coordinates are angular. By transforming the Cartesian components of $\mathbf{L}$ to spherical coordinates ( $r, \theta, \phi$ ), we obtain (the calculation is rather tedious, but Maple/Mathematica will readily do it for us):

$$
\begin{aligned}
L_{x} & =-\mathrm{i}\left(y \partial_{z}-z \partial_{y}\right)=-\mathrm{i}\left(-\sin \phi \partial_{\theta}-\cot \theta \cos \phi \partial_{\phi}\right) \\
L_{y} & =-\mathrm{i}\left(z \partial_{x}-x \partial_{z}\right)=-\mathrm{i}\left(\cos \phi \partial_{\theta}-\cot \theta \sin \phi \partial_{\phi}\right) \\
L_{z} & =-\mathrm{i}\left(x \partial_{y}-y \partial_{x}\right)=-\mathrm{i} \partial_{\phi}
\end{aligned}
$$

The derivatives with respect to $r$ have cancelled out! We also find that:

$$
\begin{equation*}
\mathbf{L}^{2}=-\left[\frac{1}{\sin \theta} \partial_{\theta}\left(\sin \theta \partial_{\theta}\right)+\frac{1}{\sin ^{2} \theta} \partial_{\phi}^{2}\right] \tag{6.24}
\end{equation*}
$$

So $\mathbf{L}^{2}$ depends only on the angular coordinates. Also, eq. (6.23) makes it obvious that the commutator $\left[\nabla^{2}, \mathbf{L}^{2}\right]:=$ $\nabla^{2} \mathbf{L}^{2}-\mathbf{L}^{2} \nabla^{2}=0$, so that $\left[\nabla^{2}, \mathbf{L}\right]=0$.

Now one readily shows that the following important relations hold:

$$
\begin{equation*}
\left[L_{x}, L_{y}\right]=\mathrm{i} L_{z}, \quad\left[L_{y}, L_{z}\right]=\mathrm{i} L_{x}, \quad\left[L_{z}, L_{x}\right]=\mathrm{i} L_{y} \tag{6.25}
\end{equation*}
$$

Using these commutator rules, we find that:

$$
\begin{aligned}
{\left[\mathbf{L}^{2}, L_{z}\right] } & =\left[L_{x}^{2}, L_{z}\right]+\left[L_{y}^{2}, L_{z}\right]+\left[L_{z}^{2}, L_{z}\right] \\
& =L_{x}\left[L_{x}, L_{z}\right]+\left[L_{x}, L_{z}\right] L_{x}+L_{y}\left[L_{y}, L_{z}\right]+\left[L_{y}, L_{z}\right] L_{y} \\
& =-\mathrm{i} L_{x} L_{y}-\mathrm{i} \Sigma_{y} L_{x}+\mathrm{i} L_{y} L_{x}+\mathrm{i} L_{x} L_{y} \\
& =0
\end{aligned}
$$

By symmetry, we have immediately that $\left[\mathbf{L}^{2}, \mathbf{L}\right]=0$.

### 6.4.2 Eigenvalues of $\mathbf{J}^{2}$ and $J_{z}$

The importance of eq.(6.25) cannot be overstated. It says that $\mathbf{L}$ belongs to the class of self-adjoint operators $\mathbf{J}$ whose Cartesian components satisfy the canonical commutation relations:

$$
\begin{equation*}
\left[J_{i}, J_{j}\right]=\mathrm{i} \epsilon_{i j k} J_{k}, \tag{6.26}
\end{equation*}
$$

Just from these properties, it is possible to derive the eigenvalues $\lambda$ of $\mathbf{J}^{2}$, ie. such that $\mathbf{J}^{2} g=\lambda g$, and the eigenvalues $m$ of $J_{z}$, such that $J_{z} f=m f$, where $\lambda$ and $m$ are expected to be real numbers since the operators are self-adjoint.

1. Introduce the ladder operators $J_{ \pm}=J_{x} \pm \mathrm{i} J_{y}$. Then $\left[\mathbf{J}^{2}, J_{ \pm}\right]=0$. Since $J_{x}$ and $J_{y}$ are self-adjoint, we also have $J_{+}^{\dagger}=J_{-}$, and $J_{-}^{\dagger}=J_{+}$. Therefore:

$$
\left[J_{z}, J_{ \pm}\right]=\left[J_{z}, J_{x}\right] \pm \mathrm{i}\left[J_{z}, J_{y}\right]=\mathrm{i} J_{y} \pm \mathrm{i}\left(-\mathrm{i} J_{x}\right)= \pm J_{ \pm}
$$

Now, using $\left[\mathbf{J}^{2}, J_{ \pm}\right]=0$ and $\left[J_{z}, J_{ \pm}\right]= \pm J_{ \pm}$, we apply first $J_{ \pm}$and then $J_{z}$ on $f$ :

$$
J_{z}\left(J_{ \pm} f\right) \equiv\left[J_{z}, J_{ \pm}\right] f+J_{ \pm} J_{z} f= \pm J_{ \pm} f+m J_{ \pm} f=(m \pm 1)\left(J_{ \pm} f\right)
$$

Now $f$ is also an eigenfunction of $\mathbf{J}^{2}$. Indeed, $J_{z} \mathbf{J}^{2} f=\mathbf{J}^{2} J_{z} f=m \mathbf{J}^{2} f$, and $\mathbf{J}^{2} f$ is an eigenfunction of $J_{z}$ with the same eigenvalue as $f$. Since the eigenvalues of $J_{z}$ are non-degenerate, $\mathbf{J}^{2} f$ must be a multiple of $f: \mathbf{J}^{2} f=\lambda f$. Also:

$$
\mathbf{J}^{2}\left(J_{ \pm} f\right)=J_{ \pm}\left(\mathbf{J}^{2} f\right)=\lambda\left(J_{ \pm} f\right)
$$

These results tell us what $J_{ \pm}$do for a living: they raise $\left(J_{+}\right)$or lower $\left(J_{-}\right)$the eigenvalues of $J_{z}$ by 1 , whence their name. In other words, if $f$ is an eigenfunction of $J_{z}$ with eigenvalue $m$, so is $J_{ \pm} f$, but with eigenvalue $(m \pm 1)$. We also have found that all the eigenfunctions of $J_{z}$ reachable with the ladder operators are eigenfunctions of $\mathbf{J}^{2}$ as well, with the same eigenvalue $\lambda$.
It is also reassuring to find that applying $J_{ \pm}$to eigenfunctions of $J_{z}$ and $J^{2}$ gives a result that belongs to a Hilbert space; because $J_{ \pm}$are not self-adjoint, this was not guaranteed.
2. Another expression will be needed. Use the definition of $J_{ \pm}$to write:

$$
J_{ \pm} J_{\mp}=J_{x}^{2}+J_{y}^{2} \mp \mathrm{i}\left[J_{x}, J_{y}\right]=J_{x}^{2}+J_{y}^{2} \pm J_{z}=\mathbf{J}^{2}-J_{z}^{2} \pm J_{z}
$$

so that:

$$
\begin{equation*}
\mathbf{J}^{2}=J_{ \pm} J_{\mp}+J_{z}^{2} \mp J_{z} \tag{6.27}
\end{equation*}
$$

3. For a given value of $\lambda$, we expect that $m$ should have a maximum value, $m_{\max } \equiv j$, as well as a minimum value, $m_{\text {min }} \equiv j^{\prime}$.
Now act with $\mathbf{J}^{2}$ on the eigenfunction of $J_{z}$ with the maximum value of $m$, which we call $f_{j}$. Then $\mathbf{J}_{+} f_{j}=0$ and, from the identity (6.27), we find:

$$
\mathbf{J}^{2} f_{j}=J_{z}^{2} f_{j}+J_{z} f_{j}=j(j+1) f_{j}=\lambda f_{j}
$$

Similarly, act with $\mathbf{J}^{2}$ on the eigenfunction of $\mathbf{J}_{z}$ with the minimum value of $m, f_{j^{\prime}}$, keeping in mind that $f_{j^{\prime}}$ is also an eigenfunction of $\mathbf{J}^{2}$ with the same eigenvalue, $\lambda$, as $f_{j}$ :

$$
\mathbf{J}^{2} f_{j^{\prime}}=\left(j^{\prime}\right)^{2} f_{j^{\prime}}-j^{\prime} f_{j^{\prime}}=j^{\prime}\left(j^{\prime}-1\right) f_{j^{\prime}}=\lambda f_{j^{\prime}}
$$

Comparing yields $\lambda=j(j+1)=j^{\prime}\left(j^{\prime}-1\right)$, and thus $j^{\prime}=-j$. It follows that $m$ goes from $-j$ to $j$ in $N$ integer steps, ie, $j=-j+N$, so $j=N / 2$.

We conclude that:

- The eigenvalues of $\mathbf{J}^{2}$ are $j(j+1)$, where $j$ is a positive integer or a half-integer.
- For a given value of $j, m$ can take $2 j+1$ values, from $-j$ to $j$.

It is worth stressing that these results were found without any knowledge of an explicit form for the eigenfunctions of $\mathbf{J}^{2}$, or indeed $\mathbf{J}$. And they apply to all self-adjoint operators which satisfy the canonical commutation relations (6.26).

With the help of eq. (6.27), we can now exhibit the full action of $J_{-}$on a normalised eigenfunction $f_{j m}$ of $\mathbf{J}^{2}$ and $J_{z}$. Let $J_{-} f_{j m}=c_{-} f_{j, m-1}$. Then, using the rules for taking adjoints, and with $(f, g)$ the inner product of $f$ and $g$ :

$$
\left(f_{j m}, J_{+} J_{-} f_{j m}\right)=\left(J_{-} f_{j m}, J_{-} f_{j m}\right)=\left(c_{-} f_{j, m-1}, c_{-} f_{j, m-1}\right)=\left|c_{-}\right|^{2}\left(f_{j, m-1}, f_{j, m-1}\right)=\left|c_{-}\right|^{2}
$$

But since $J_{ \pm} J_{\mp}=\mathbf{J}^{2}-J_{z}^{2} \pm J_{z}$, we also have that:

$$
\left(f_{l m}, J_{+} J_{-} f_{j m}\right)=\left(f_{j m},\left(\mathbf{J}^{2}-J_{z}^{2}+J_{z}\right) f_{j m}\right)=j(j+1)-m^{2}+m
$$

Comparing yields $c_{-}$up to an unimportant exponential phase factor which we put equal to 1 . We find the coefficient in $J_{+} f_{j m}=c_{+} f_{j, m+1}$ in a strictly analogous way. The results for both ladder operators are:

$$
\begin{equation*}
J_{ \pm} f_{j m}=\sqrt{j(j+1)-m(m \pm 1)} f_{j, m \pm 1} \tag{6.28}
\end{equation*}
$$

### 6.4.3 Eigenfunctions of $L^{2}$ and $L_{z}$

To find the common eigenfunctions for $\mathbf{J}$ and $\mathbf{J}^{2}$ operators, we must know their form. Here, we will be interested in the $\mathbf{L}$ operator whose form we do know and which makes up the angular part of the Laplacian in spherical coordinates.

The eigenfunctions of $L_{z}$ are readily obtained by solving the differential equation:

$$
L_{z} f(\theta, \phi)=-\mathrm{i} \partial_{\phi} f(\theta, \phi)=m f(\theta, \phi)
$$

With a separation ansatz: $f(\theta, \phi)=F(\theta) G(\phi)$, the solution for $G$ is:

$$
\begin{equation*}
G(\phi)=\mathrm{e}^{\mathrm{i} m \phi} \tag{6.29}
\end{equation*}
$$

Now we require that $G$ (and $f$ ) be single-valued, that is, $G(\phi+2 \pi)=G(\phi)$. Thus:

$$
\mathrm{e}^{\mathrm{i} m(\phi+2 \pi)}=\mathrm{e}^{\mathrm{i} m \phi} \Longrightarrow \mathrm{e}^{2 \mathrm{i} m \pi}=\cos 2 m \pi+\mathrm{i} \sin 2 m \pi=1
$$

which constrains $m$ to be any integer. Therefore, $l:=m_{\max }$ must also be an integer. Thus, we find that the particular form $\mathbf{L}=-\mathrm{ix} \times \nabla$ rules out the possibility of half-integer values of $j$ allowed for a self-adjoint $\mathbf{J}$ that satisfies the canonical commutation relations (6.26).

The $\theta$ dependence of the eigenfunctions must be derived from the eigenvalue equation for $\mathbf{L}^{2}$. Call $f(\theta, \phi)=$ $Y_{l}^{m}(\theta, \phi)=F(\theta) G(\phi)$; these must satisfy:

$$
-\left[\frac{1}{\sin \theta} \partial_{\theta}\left(\sin \theta \partial_{\theta}\right)+\frac{1}{\sin ^{2} \theta} \partial_{\phi}^{2}\right] Y_{l}^{m}(\theta, \phi)=l(l+1) Y_{l}^{m}(\theta, \phi)
$$

Inserting $Y_{l}^{m}(\theta, \phi)=F(\theta) \mathrm{e}^{\mathrm{i} m \phi}$ into this equation leaves:

$$
-\left[\frac{1}{\sin \theta} \mathrm{~d}_{\theta}\left(\sin \theta \mathrm{d}_{\theta}\right)-\frac{m^{2}}{\sin ^{2} \theta}\right] F(\theta)=l(l+1) F(\theta)
$$

Instead of solving this equation by brute force, we use a clever technique involving the ladder operators $L_{ \pm}$:

$$
L_{ \pm}= \pm \mathrm{e}^{\mathrm{i} \phi}\left(\partial_{\theta} \pm \mathrm{i} \cot \theta \partial_{\phi}\right)
$$

Now, when $m=l$, we have:

$$
L_{+} Y_{l}^{l}=\mathrm{e}^{\mathrm{i} \phi}\left(\partial_{\theta}+\mathrm{i} \cot \theta \partial_{\phi}\right) Y_{l}^{l}(\theta, \phi)=0
$$

Inserting $Y_{l}^{l}=F(\theta) \mathrm{e}^{\mathrm{i} l \phi}$, this reduces to the much simpler

$$
\mathrm{d}_{\theta} F(\theta)-l \cot \theta F(\theta)=0
$$

whose solution is $F(\theta)=(\sin \theta)^{l}$. Therefore, $Y_{l}^{l}=(\sin \theta)^{l} \mathrm{e}^{\mathrm{i} l \phi}$. Applying $L_{-}$the requisite number of times generates the other $Y_{l}^{m}(0<m<l): Y_{l}^{m} \propto L_{-}^{l-m} Y_{l}^{l}$. When normalised, these are the spherical harmonics already found in eq. (5.50):

$$
\begin{equation*}
Y_{l}^{m}(\theta, \phi)=\frac{(-1)^{m}}{2^{l} l!} \sqrt{\frac{2 l+1}{4 \pi} \frac{(l-m)!}{(l+m)!}}\left(1-x^{2}\right)^{m / 2}\left[\mathrm{~d}_{x}^{l+m}\left(x^{2}-1\right)^{l}\right] \mathrm{e}^{\mathrm{i} m \phi} \quad x=\cos \theta \tag{6.30}
\end{equation*}
$$

### 6.4.4 General Solution of a Spherically-Symmetric, 2nd-order, Homogeneous, Linear Equation

Suppose we are presented with the equation $\left[\nabla^{2}+\gamma(\mathbf{x})\right] \Psi(\mathbf{x})=0$. Work in spherical coordinates, and make the ansatz: $\Psi(\mathbf{x})=R(r) F(\theta, \phi)$. Using the form for $\nabla^{2}$ derived earlier, eq. (6.23), we have:

$$
\begin{aligned}
\nabla^{2} \Psi+\gamma(\mathbf{x}) \Psi & =-\frac{\mathbf{L}^{2} \Psi}{r^{2}}+\frac{1}{r}\left[\partial_{r} \Psi+\partial_{r}\left(r \partial_{r} \Psi\right)\right]+\gamma(\mathbf{x}) \Psi \\
& =-R(r) \frac{\mathbf{L}^{2} F(\theta, \phi)}{r^{2}}+\frac{F(\theta, \phi)}{r}\left[\mathrm{~d}_{r} R(r)+\mathrm{d}_{r}\left(r \mathrm{~d}_{r} R(r)\right)\right]+\gamma(\mathbf{x}) R(r) F(\theta, \phi)
\end{aligned}
$$

Dividing the second line by $R(r) F(\theta, \phi)$ and multiplying by $r^{2}$, we see that the equation is separable provided $\gamma(\mathbf{x})=\gamma(r)$ :

$$
\mathbf{L}^{2} F(\theta, \phi)=\lambda F(\theta, \phi) \quad \mathrm{d}_{r} R(r)+\mathrm{d}_{r}\left(r \mathrm{~d}_{r} R(r)\right)+r \gamma(r) R(r)=\lambda \frac{R(r)}{r}
$$

The first equation is the eigenvalue equation for $\mathbf{L}^{2}$, whose eigenvalues are $\lambda=l(l+1)(l \geq 0 \in \mathbb{Z})$ with the spherical harmonics $Y_{l}^{m}(\theta, \phi)$ as eigenfunctions.

The radial equation can thus be written: $\frac{1}{r^{2}} \mathrm{~d}_{r}\left(r^{2} \mathrm{~d}_{r} R_{l}(r)\right)+\left(\gamma(r)-\frac{l(l+1)}{r^{2}}\right) R_{l}(r)=0$
When $\gamma(r)=0$, this is the radial part of the Laplace equation which becomes, after the change of variable $r=\mathrm{e}^{x}, \mathrm{~d}_{x}^{2} R+\mathrm{d}_{x} R-l(l+1) R=0$. Inserting a solution of the form $\mathrm{e}^{p x}$ turns the equation into $p^{2}+p-l(l+1)=0$, that is, $p=l$ or $p=-(l+1)$, which leads to $R=A \mathrm{e}^{l x}+B \mathrm{e}^{-(l+1) x}=A r^{l}+B r^{-(l+1)}$. Therefore, the general solution to the Laplace equation in spherical coordinates is:

$$
\begin{equation*}
\Psi(r, \theta, \phi)=\sum_{l=0}^{\infty} \sum_{m=-l}^{l}\left(A_{l m} r^{l}+\frac{B_{l m}}{r^{l+1}}\right) Y_{l}^{m}(\theta, \phi) \tag{6.31}
\end{equation*}
$$

The coefficients $A_{l m}$ and $B_{l m}$ are determined from boundary or matching conditions. In regions either containing the origin, or extending all the way to infinity, $B_{l m}=0$ or $A_{l m}=0$, respectively. Clearly, if this solution is to be regular, and if it holds everywhere, it must vanish. In other words, if the Laplace equation is valid everywhere, it has no non-vanishing regular solution. For a non-trivial solution, there must be a region of space where there exists an inhomogeneous term acting as a source.

Note, however, that the general solution holds at any point where there is no source. The effect of sources is encoded in the coefficients $A_{l m}$ and $B_{l m}$.

When $\gamma(r)=k^{2}>0$, we get the radial part of the Helmholtz equation in spherical coordinates:

$$
\mathrm{d}_{r}^{2} R_{l}(r)+\frac{2}{r} \mathrm{~d}_{r} R_{l}(r)+\left(k^{2}-\frac{l(l+1)}{r^{2}}\right) R_{l}(r)=0
$$

which the substitutions $R_{l}=u_{l} / \sqrt{r}$ and $x=k r$ readily transform into:

$$
\mathrm{d}_{x}^{2} u_{l}(x)+\frac{1}{x} \mathrm{~d}_{x} u_{l}(x)+\left(1-\frac{(l+1 / 2)^{2}}{x^{2}}\right) u_{l}(x)=0
$$

that is, the Bessel equation (4.33) with $n=l+1 / 2$. The solutions are the spherical Bessel functions of the first and second (Neumann) kind, usually written as (see also Jackson's Classical Electrodynamics, section 9.6):

$$
\begin{align*}
& j_{l}(x)=\sqrt{\frac{\pi}{2 x}} J_{l+1 / 2}(x)=(-x)^{l}\left(\frac{1}{x} \frac{\mathrm{~d}}{\mathrm{~d} x}\right)^{l}\left(\frac{\sin x}{x}\right) \quad \sim \begin{cases}x^{l} & x \ll(1, l) \\
\frac{1}{x} \sin (x-l \pi / 2) & x \gg l\end{cases}  \tag{6.32}\\
& n_{l}(x)=\sqrt{\frac{\pi}{2 x}} N_{l+1 / 2}(x)=-(-x)^{l}\left(\frac{1}{x} \frac{\mathrm{~d}}{\mathrm{~d} x}\right)^{l}\left(\frac{\cos x}{x}\right) \sim \begin{cases}-\frac{1}{x^{l+1}} & x \ll(1, l) \\
-\frac{1}{x} \cos (x-l \pi / 2) & x \gg l\end{cases} \tag{6.33}
\end{align*}
$$

Here are a few spherical Bessel and Neumann functions as plotted on Maple, with $\rho=x$ :


The $n_{l}$ diverge at the origin and thus are excluded from any solution regular at the origin.
(Spherical) Bessel functions of the third kind, aka Hankel functions of the first and second kind, sometimes come in handy: $h_{l}^{(1,2)}(x)=j_{l}(x) \pm \mathrm{i} n_{l}(x)$. One can express the general solution of the Helmholtz equation in terms of the $j_{l}$ and $n_{l}$, or in terms of the $h_{l}^{(1,2)}$.

### 6.5 Second 3-dim Green Identity, or Green's Theorem

Before discussing the all-important subject of boundary conditions, we derive a result that will prove very useful in the study of 3-dim elliptic problems. We assume that $L[f]=\partial^{i}\left(\alpha(\mathbf{x}) \partial_{i} f\right)+\gamma(\mathbf{x}) f$ (Sturm-Liouville form). If it isn't, it is possible to bring it into that form, provided that the solutions of $L[g]=0$ go to zero at infinity or else that $\alpha \rightarrow 0$ sufficiently fast.

Write the divergence theorem for $\boldsymbol{\nabla} \cdot(\alpha f \nabla g)$ defined over a connected volume, and expand the divergence to get:

$$
\begin{equation*}
\int_{V}[f \boldsymbol{\nabla} \cdot(\alpha \boldsymbol{\nabla} g)+\alpha \boldsymbol{\nabla} f \cdot \boldsymbol{\nabla} g] \mathrm{d}^{3} x=\oint_{\partial V} \alpha f \boldsymbol{\nabla} g \cdot \mathrm{~d} \mathbf{S} \tag{6.34}
\end{equation*}
$$

where $\partial V$ is the boundary of the volume $V$ of integration. This is Green's first identity in three dimensions; when $\alpha$ is a constant, it reduces to the more familiar form:

$$
\begin{equation*}
\int_{V}\left[f \nabla^{2} g+\boldsymbol{\nabla} f \cdot \boldsymbol{\nabla} g\right] \mathrm{d}^{3} x=\oint_{\partial V} f \boldsymbol{\nabla} g \cdot \mathrm{~d} \mathbf{S} \tag{6.35}
\end{equation*}
$$

Interchanging $f$ and $g$ in the first identity (6.34) and subtracting, we easily find, after adding and subtracting $\gamma f g$ in the volume itnegral, the second Green identity in three dimensions:

$$
\begin{equation*}
\int_{V}(f L[g]-g L[f]) \mathrm{d}^{3} x=\oint_{\partial V} \alpha(f \boldsymbol{\nabla} g-g \boldsymbol{\nabla} f) \cdot \mathrm{d} \mathbf{S} \tag{6.36}
\end{equation*}
$$

With $\alpha$ a constant, this becomes the well-known Green theorem:

$$
\begin{equation*}
\int_{V}\left(f \nabla^{2} g-g \nabla^{2} f\right) \mathrm{d}^{3} x=\oint_{\partial V}(f \boldsymbol{\nabla} g-g \boldsymbol{\nabla} f) \cdot \mathrm{d} \mathbf{S} \tag{6.37}
\end{equation*}
$$

Note that in the case of conpact regions without boundaries (sphere, torus), the right-hand side vanishes.

### 6.5.1 Uniqueness and existence of solutions for the inhomogeneous Laplace equation with B.C.

The inhomogeneous Laplace (aka Poisson) equation is of the form $\nabla^{2} \Psi(\mathbf{x})=F(\mathbf{x})$, where the right-hand side is called a source term. As justified below, we also specify B.C. for either $\Psi$ or $\hat{\mathbf{n}} \cdot \nabla \Psi$ on a closed boundary.

Now, with $f=g=\Psi_{3}$ and $\alpha$ constant, Green's first identity-eq. (6.35)— becomes:

$$
\int_{V}\left[\Psi_{3} \nabla^{2} \Psi_{3}+\left(\nabla \Psi_{3}\right)^{2}\right] \mathrm{d}^{3} x=\oint_{\partial V} \Psi_{3} \partial_{n} \Psi_{3} \mathrm{~d} S
$$

where we have introduced the normal derivative $\partial_{n} \Psi_{3}$, ie. the component of $\nabla \Psi_{3}$ normal outward to $\partial V$.
Suppose there exist two solutions, $\Psi_{1}$ and $\Psi_{2}$, of $\nabla^{2} \Psi(\mathbf{x})=F(\mathbf{x})$ that satisfy the same conditions on the surface. Define $\Psi_{3}:=\Psi_{2}-\Psi_{1}$. Then $\nabla^{2} \Psi_{3}=0$ inside the volume. The surface integral is zero because either $\Psi_{3}=0$ or $\partial \Psi_{3} / \partial n=0$ on the surface, and $\int\left(\nabla \Psi_{3}\right)^{2} \mathrm{~d}^{3} x=0$ everywhere. Also, $\Psi_{3}$ being twice differentiable at all points in the volume, $\nabla \Psi_{3}$ is continuous and therefore zero everywhere inside the volume, so that $\Psi_{3}$ is a constant. It follows immediately that if $\Psi_{3}=0$ on the boundary, $\Psi_{3}=0$ everywhere; on the other hand, when $\partial \Psi_{3} / \partial n=0$ on the boundary, $\Psi_{3}$ can be a non-zero constant inside.

We conclude that $\Psi_{1}=\Psi_{2}$ inside the volume (up to a possible additive constant), and that the solution, if it exists, is uniquely determined. The importance of this result cannot be overemphasised: any function that satisfies the inhomogeneous Laplace (aka Poisson) equation and the B.C. is the solution, no matter how it was found! Moreover, we see that we cannot arbitrarily specify both $\Psi$ and $\partial \Psi / \partial n$ on the boundary since one suffices to determine the unique solution.

The BC determine the solution, but only if it exists. Further conditions must be met for this to happen. Indeed, with $\phi(\mathbf{x}):=\left.\partial_{n} \Psi\right|_{\mathbf{x} \in \partial V}$, integrate $\nabla^{2} \Psi(\mathbf{x})=F(\mathbf{x})$ over (connected!) $V$; the divergence theorem immediately yields a condition linking $\phi$ to the source $F$ :

$$
\begin{equation*}
\int_{V} F(\mathbf{x}) \mathrm{d}^{3} x=\int_{\partial V} \phi(\mathbf{x}) \mathrm{d} S \tag{6.38}
\end{equation*}
$$

Another condition for the existence of a solution is that the enclosing boundary be "reasonably" smooth (eg. no pointy parts), otherwise it may prove impossible even to impose B.C.

### 6.6 3-dim Boundary Value (Elliptic) Problems with Green's Functions

Introduce Green functions that satisfy $L\left[G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\right]=\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ in compact regions with closed boundaries or in non-compact regions ${ }^{\dagger}$ (again, some authors multiply the right-hand side by $\pm 4 \pi$ ). If we are a little careful, we will find that for some important cases this kind of problem can admit unique Green functions.

### 6.6.1 Dirichlet and Neumann Boundary Conditions

Suppose that $\Psi(\mathbf{x})$ satisfies $^{\dagger} L \Psi(\mathbf{x})=F(\mathbf{x})$; then take $f=\Psi$ and $g=G$ in Green's second identity (eq. (6.36)):

$$
\int_{V}(\Psi L[G]-G L[\Psi]) \mathrm{d}^{3} x^{\prime}=\oint_{\partial V} \alpha\left(\Psi \partial_{n^{\prime}} G-G \partial_{n^{\prime}} \Psi\right) \mathrm{d} S^{\prime}
$$

where, as in example 6.5.1, $\boldsymbol{\nabla} f \cdot \mathrm{~d} \mathbf{S}=\partial_{n} f \mathrm{~d} S$. We obtain:

$$
\int_{V}\left[\Psi\left(\mathbf{x}^{\prime}\right) \delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right)-F\left(\mathbf{x}^{\prime}\right) G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\right] \mathrm{d}^{3} x^{\prime}=\oint_{\partial V} \alpha\left(\Psi \partial_{n^{\prime}} G-G \partial_{n^{\prime}} \Psi\right) \mathrm{d} S^{\prime}
$$

With x inside the volume, re-arranging then yields:

$$
\begin{equation*}
\Psi(\mathbf{x})=\int_{V} F\left(\mathbf{x}^{\prime}\right) G\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime}+\oint_{\partial V} \alpha\left(\Psi \partial_{n^{\prime}} G-G \partial_{n^{\prime}} \Psi\right) \mathrm{d} S^{\prime} \tag{6.39}
\end{equation*}
$$

where the normal derivatives in the integrand on the right-hand side are to be evaluated on $\partial V$, the boundary of the arbitrary volume. This expression for $\Psi$ cannot be considered a solution yet; it is still "just" an identity.

[^24]Again, note that $\Psi$ and $\partial \Psi / \partial n$ are in general not independent on the boundary. We are not free to specify them both arbitrarily at the same point on $\partial V$ as such values will in general be inconsistent.

Specifying $\Psi$ on the boundary gives Dirichlet B.C., whereas specifying $\partial \Psi / \partial n$ gives Neumann B.C.
How do we get a solution for $\Psi$ then? In principle, this is simple. We use the fact that the Green functions we find by solving $L\left[G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\right]=\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ are not unique; indeed, we can add to them any function $G_{1}$ that satisfies $L G_{1}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=0$. "All" we have to do then is find a $G_{1}$ that eliminates one of the two surface integrals.

Suppose we wish to specify $\Psi$ freely on the boundary (Dirichlet problem). Then we should ensure that $G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=0 \forall \mathbf{x}^{\prime} \in \partial V$. The solution for $\Psi$ would then be:

$$
\begin{equation*}
\Psi(\mathbf{x})=\int_{V} F\left(\mathbf{x}^{\prime}\right) G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime}+\oint_{\partial V} \alpha \Psi\left(\mathbf{x}^{\prime}\right) \partial_{n^{\prime}} G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \mathrm{d} S^{\prime} \tag{6.40}
\end{equation*}
$$

The solution is now uniquely determined by the B.C. on $\Psi$ via $G_{\mathrm{D}}$. Note that the total surface $\partial V$ enclosing the volume may be disjoint, as occurs for instance with the volume between two concentric spheres.

If we have managed to find $G_{\mathrm{D}}$ for a particular type of boundary, the source-free solution $\left(F\left(\mathbf{x}^{\prime}\right)=0\right)$ is just the surface integral; on the other hand, if it happens that $\Psi=0$ on $\partial V$, only the volume integral contributes. Many boundary-value problems in electrostatics, for which the boundary conditions are reasonably simple, can be solved this way.

With Green's second identity, it is also straightforward to prove (exercise) that $G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)$ is symmetric in its arguments.

Similar considerations apply to Neumann boundary conditions, ie. when $\partial \Psi / \partial n$ rather than $\Psi$ is known on the boundary. But we must be a little careful about the boundary conditions on $\partial_{n} G_{\mathrm{N}}$ : we cannot simply put this equal to 0 in eq. (6.39). Indeed, take for instance $L=\nabla^{2}$; then, from the divergence theorem and the defining equation $L\left[G_{\mathrm{N}}\right]=\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$ :

$$
\int \nabla \cdot \nabla G_{\mathrm{N}} \mathrm{~d}^{3} x=\oint_{\partial V} \partial_{n} G_{\mathrm{N}} \mathrm{~d} S=1
$$

A consistent boundary condition is $\left.\partial_{n} G_{\mathrm{N}}\right|_{\partial V}=1 / S$, and we obtain:

$$
\begin{equation*}
\Psi(\mathbf{x})=<\Psi>_{\partial V}+\int_{V} F\left(\mathbf{x}^{\prime}\right) G_{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime}-\oint_{\partial V} G_{\mathrm{N}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \partial_{n^{\prime}} \Psi\left(\mathbf{x}^{\prime}\right) \mathrm{d} S^{\prime} \tag{6.41}
\end{equation*}
$$

Up to the a priori unknown average of $\Psi$ over the surface, $<\Psi>_{\partial V}$, a constant, this is the solution to that Neumann problem. Often (but not always!) the volume is bounded by two surfaces, one closed and finite and the other at infinity, in which case the normal derivative of $G_{\mathrm{N}}$ can be set to zero on the entire boundary, and the average of $\Psi$ over $\partial V$ (the first term) vanishes.

### 6.6.2 Green's function for the 3-d Elliptic Helmholtz operator without boundary conditions

We proceed to find a Green function for the operator $\nabla^{2}+\lambda$, with $\lambda$ a constant. Using eq. (5.13), the Fourier transform of $\left(\nabla^{2}+\lambda\right) \Psi(\mathbf{x})=F(\mathbf{x})$ is $\left(-k^{2}+\lambda\right) \psi(\mathbf{k})=F(\mathbf{k})$. We must distinguish between two possibilities:

1. $\lambda=-\kappa^{2} \leq 0, \kappa \geq 0$

Then, similarly to what happens in one dimension (example 6.3), an "inhomogeneous" solution is:

$$
\Psi(\mathbf{x})=-\frac{1}{(2 \pi)^{3 / 2}} \int \frac{F(\mathbf{k})}{k^{2}+\kappa^{2}} \mathrm{e}^{\mathrm{i} \mathbf{k} \cdot \mathbf{x}} \mathrm{~d}^{3} k=-\frac{1}{(2 \pi)^{3}} \iint \mathrm{~d}^{3} x^{\prime} \mathrm{e}^{-\mathrm{i} \mathbf{k} \cdot \mathbf{x}^{\prime}} \frac{F\left(\mathrm{x}^{\prime}\right)}{k^{2}+\kappa^{2}} \mathrm{e}^{\mathrm{i} \mathbf{k} \cdot \mathbf{x}} \mathrm{~d}^{3} k
$$

Compare with the Green-function form of the inhomogeneous solution, $\int_{V} F\left(\mathbf{x}^{\prime}\right) G\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime}$ (EXERCISE):

$$
G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=-\frac{1}{(2 \pi)^{3}} \int \frac{\mathrm{e}^{\mathrm{i} \mathbf{k} \cdot\left(\mathbf{x}-\mathrm{x}^{\prime}\right)}}{k^{2}+\kappa^{2}} \mathrm{~d}^{3} k=\frac{\mathrm{i}}{(2 \pi)^{2}\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \int_{-\infty}^{\infty} \frac{k \mathrm{e}^{\mathrm{i} k\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{k^{2}+\kappa^{2}} \mathrm{~d} k
$$

This integral is easily evaluated as part of a contour integral around a semi-circle at infinity in the upper complex $k$ half-plane. As before, the contribution at infinity vanishes, and the residue due to the pole at
$k=\mathrm{i} \kappa$ is $\mathrm{e}^{-\kappa\left|\mathrm{x}-\mathrm{x}^{\prime}\right|} / 2$. The Residue theorem then yields the (sometimes called fundamental, or singular) solution:

$$
\begin{equation*}
G\left(\mathbf{x}, \mathrm{x}^{\prime}\right)=-\frac{1}{4 \pi} \frac{\mathrm{e}^{-\kappa\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathrm{x}-\mathrm{x}^{\prime}\right|} \tag{6.42}
\end{equation*}
$$

This is very well-behaved at infinity. For $\lambda=0(\kappa=0)$, we obtain a Green function for the Laplacian operator.
With $\kappa=0$ and $F(\mathbf{x})=-4 \pi \rho(\mathbf{x})$ (Gaussian units!), for instance, an inhomogeneous solution is the generalised Coulomb Law for the electrostatic potential of a localised charge density $\rho(\mathbf{x})$, or one that vanishes at infinity faster than $1 /\left|x-x^{\prime}\right|$.
2. $\lambda=\kappa^{2} \geq 0$

In order to invert the algebraic equation for $\psi(\mathbf{k})$, we write $\lambda=(q \pm \mathrm{i} \epsilon)^{2}(\epsilon \geq 0)$. Then we arrive at:

$$
\begin{equation*}
G_{q}^{( \pm)}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=-\frac{1}{(2 \pi)^{3}} \lim _{\epsilon \rightarrow 0} \int \frac{\mathrm{e}^{\mathrm{i} \mathbf{k} \cdot\left(\mathbf{x}-\mathbf{x}^{\prime}\right)}}{k^{2}-(q \pm \mathrm{i} \epsilon)^{2}} \mathrm{~d}^{3} k=-\frac{1}{4 \pi} \frac{\mathrm{e}^{ \pm \mathrm{i} q\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \tag{6.43}
\end{equation*}
$$

For details of the calculation, see pp. BF415-416.
Do check that these Green functions satisfy $\left(\nabla^{2}+\lambda\right) G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right)$. But note that they are not the general solution of this equation, since any function that satisfies the homogeneous equation can be added to them!

If the volume integral extends over all space, the surface integral in the Dirichlet solution for the case $\lambda<0$ certainly vanishes at infinity for fairly weak conditions on $\Psi(\mathbf{x})$, because of the exponential factor in Green's function. When $\lambda \geq 0$, the surface integral also vanishes provided $\Psi(x) \rightarrow 0$ faster than $1 /\left|x-x^{\prime}\right|^{2}$, (since $\mathrm{d} S \sim\left|\mathrm{x}-\mathrm{x}^{\prime}\right|^{2}$ ), and we are left with just the inhomogeneous integral:

$$
\begin{equation*}
\Psi_{q}^{( \pm)}(\mathbf{x})=-\frac{1}{4 \pi} \int_{V} \frac{F\left(\mathbf{x}^{\prime}\right) \mathrm{e}^{ \pm i q\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \mathrm{d}^{3} x^{\prime} \tag{6.44}
\end{equation*}
$$

If, however, $\Psi(\mathbf{x})$ does not vanish fast enough at infinity, it is more convenient to write it in terms of the solution of the homogeneous equation $\left(\nabla^{2}+q^{2}\right) \Psi(\mathbf{x})=0$, plus the volume integral:

$$
\begin{equation*}
\Psi_{q}^{( \pm)}(\mathbf{x})=A \mathrm{e}^{\mathrm{i} \mathbf{q} \cdot \mathbf{x}}-\frac{1}{4 \pi} \int_{V} \frac{F\left(\mathbf{x}^{\prime}\right) \mathrm{e}^{ \pm \mathrm{i} q\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \mathrm{d}^{3} x^{\prime} \tag{6.45}
\end{equation*}
$$

Note that these expressions for Green's functions assume no boundary surfaces (except at infinity)!

### 6.6.3 Dirichlet Green function for the Laplacian

When there are no boundary conditions for $\Psi$ on finite surfaces, the volume integral $\int F\left(\mathrm{x}^{\prime}\right) G\left(\mathrm{x}, \mathrm{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime}$ can be taken as the solution to $L[\Psi]=F$. For instance, in the case of a point-source located at $\mathbf{y}: F\left(\mathbf{x}^{\prime}\right)=-4 \pi q \delta\left(\mathbf{y}-\mathbf{x}^{\prime}\right)$, with $q$ some constant, we see that $\Psi(\mathbf{x})=-4 \pi q G(\mathbf{x}, \mathbf{y})=q /|\mathbf{x}-\mathbf{y}|$ in the case of $L=\nabla^{2}$.

When there are finite boundaries, however, as in a Dirichlet problem, we know that we have to ensure that $G_{\mathrm{D}}\left(\mathrm{x}, \mathrm{x}^{\prime}\right)=0$ when either x or $\mathrm{x}^{\prime}$ is a point on the surface that encloses the volume in which our solution is valid. Obviously, with the Green function given in eq. (6.42), which vanishes only on a boundary at infinity, this is impossible. It is time to exercise our freedom to add to $G$ a function that satisfies the homogeneous equation $L[G]=0$ and contains free parameters that can be set so as force the combined Green function to vanish on the boundary. In the case of the Laplacian, we take:

$$
G_{\mathrm{D}}\left(\mathrm{x}, \mathrm{x}^{\prime}\right)=-\frac{1}{4 \pi}\left(\frac{1}{\left|\mathrm{x}-\mathrm{x}^{\prime}\right|}+\frac{g}{\left|\mathbf{x}-\mathrm{x}^{\prime \prime}\right|}\right)
$$

where $g$ and $\mathbf{x}^{\prime \prime}$ will ensure that the second term satisfies the Laplace equation $\forall \mathrm{x}$ inside the volume where we are looking for a solution, as well as vanishing on the boundary.

## Example 6.4. Solution of the Dirichlet problem on a sphere for the Laplacian

Consider a sphere of radius $a$ centered on the origin. We want: $G_{\mathrm{D}}\left(a \hat{\mathbf{n}}, \mathbf{x}^{\prime}\right)=G_{\mathrm{D}}\left(\mathbf{x}, a \hat{\mathbf{n}}^{\prime}\right)=0$ Symmetry dictates that $\mathbf{x}^{\prime \prime}$ and $\mathbf{x}^{\prime}$ be collinear, which means that, at $|\mathbf{x}|=r=a$, we can write:

$$
G_{\mathrm{D}}\left(a \hat{\mathbf{n}}, \mathbf{x}^{\prime}\right)=-\frac{1}{4 \pi}\left(\frac{1}{a\left|\hat{\mathbf{n}}-\frac{r^{\prime}}{a} \hat{\mathbf{n}}^{\prime}\right|}+\frac{g}{r^{\prime \prime}\left|\frac{a}{r^{\prime \prime}} \hat{\mathbf{n}}-\hat{\mathbf{n}}^{\prime \prime}\right|}\right)
$$

where $r \hat{\mathbf{n}}=\mathbf{x}$, etc. By inspection, we see that if $G_{\mathrm{D}}\left(a \hat{\mathbf{n}}, \mathbf{x}^{\prime}\right)$ is to vanish for $\hat{\mathbf{n}}$ in an arbitrary direction, we must have: $1 / a=-g / r^{\prime \prime}$ and $r^{\prime} / a=a / r^{\prime \prime}$, as well as: $\hat{\mathbf{n}} \cdot \hat{\mathbf{n}}^{\prime}=\hat{\mathbf{n}} \cdot \hat{\mathbf{n}}^{\prime \prime}$ ( $\hat{\mathbf{n}}^{\prime}$ and $\hat{\mathbf{n}}^{\prime \prime}$ collinear). Then:

$$
\begin{equation*}
g=-a / r^{\prime}, \quad \quad r^{\prime} r^{\prime \prime}=a^{2} \tag{6.46}
\end{equation*}
$$

Thus, $\mathrm{x}^{\prime \prime}$ lies outside the sphere if $\mathrm{x}^{\prime}$ is inside, and vice-versa. We replace $a \hat{\mathbf{n}}$ by $r \hat{\mathbf{n}}=\mathrm{x}$ to obtain:

$$
G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=-\frac{1}{4 \pi}\left[\frac{1}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}-\frac{1}{\left|\left(r^{\prime} / a\right) \mathbf{x}-\left(a / r^{\prime}\right) \mathbf{x}^{\prime}\right|}\right]
$$

If this is evaluated in spherical coordinates centered on the sphere, the angle $\gamma$ between $\mathbf{x}$ amd $\mathbf{x}^{\prime}$ is, from spherical trigonometry: $\cos \gamma=\cos \theta \cos \theta^{\prime}+\sin \theta \sin \theta^{\prime} \cos \left(\phi-\phi^{\prime}\right)$, and there comes:

In this form it is most easy to see that $G_{\mathrm{D}}\left(a \hat{\mathbf{n}}, \mathrm{x}^{\prime}\right)=G_{\mathrm{D}}\left(\mathbf{x}, a \hat{\mathbf{n}}^{\prime}\right)=0$, as desired. The Dirichlet Green function we have found is valid for any ball since it does not care about which particular B.C. is specified for $\Psi(\mathbf{x})$ on its spherical boundary. When $\Psi\left(r^{\prime}=a\right)=0$, the surface integral in eq. (6.40) vanishes; the volume integral remains the same since it is independent of the B.C. for $\Psi$. If $\Psi\left(r^{\prime}=a\right) \neq 0$, we must evaluate $\partial_{n^{\prime}} G_{\mathrm{D}}$. In spherical coordinates, this is:

$$
\frac{\partial G_{\mathrm{D}}}{\partial n^{\prime}}= \pm\left.\frac{\partial G_{\mathrm{D}}}{\partial r^{\prime}}\right|_{r^{\prime}=a}= \pm \frac{1}{4 \pi a^{2}} \frac{a\left(a^{2}-r^{2}\right)}{\left(r^{2}+a^{2}-2 a r \cos \gamma\right)^{3 / 2}}
$$

depending on whether $\mathrm{d} \mathbf{S}^{\prime}$, the normal to the surface which always points out of the volume, is in the direction of $\mathrm{x}^{\prime}$ or in the opposite direction. For instance, the general solution of the inhomogeneous Laplace equation with B.C. specified on the surface $r=a$ for $\Psi$ is:

$$
\begin{gather*}
\Psi(\mathbf{x})=\frac{1}{4 \pi} \int F\left(\mathbf{x}^{\prime}\right)\left[\frac{1}{\sqrt{r^{2}{r^{\prime}}^{2} / a^{2}+a^{2}-2 r r^{\prime} \cos \gamma}}-\frac{1}{\sqrt{r^{2}+{r^{\prime 2}}^{2}-2 r r^{\prime} \cos \gamma}}\right] \mathrm{d}^{3} x^{\prime} \\
\pm \frac{1}{4 \pi} \oint \Psi\left(r^{\prime}=a\right) \frac{a^{2}-r^{2}}{a\left(r^{2}+a^{2}-2 a r \cos \gamma\right)^{3 / 2}} \mathrm{~d} S^{\prime} \tag{6.48}
\end{gather*}
$$

where the $(+)$ sign refers to the solution for $r<a$ and the $(-)$ sign applies to $r>a$. In the latter case, there is an implicit assumption that the integrand, $\Psi \partial_{n^{\prime}} G_{\mathrm{D}}$, of the surface integral vanishes at infinity faster than $1 / r^{\prime 2}$. When $F(\mathbf{x})=0$ everywhere inside the volume where the solution is valid, we are left with the Laplace equation $\nabla^{2} \Psi=0$, with solution:

$$
\begin{equation*}
\Psi(\mathbf{x})= \pm \oint \Psi\left(a, \theta^{\prime}, \phi^{\prime}\right)\left[\frac{1}{4 \pi a^{2}} \frac{a\left(a^{2}-r^{2}\right)}{\left(r^{2}+a^{2}-2 a r \cos \gamma\right)^{3 / 2}}\right] \mathrm{d} S^{\prime} \tag{6.49}
\end{equation*}
$$

This has the form of a Poisson integral for the solution of $\nabla^{2} \Psi=0$ in three dimensions, with the term in square brackets the Poisson kernel that we first met in 2-dim form for a disc in the complex plane in section 4.3.2. In fact, Poisson kernels $P_{r}(\cos \gamma)$ are normal derivatives of Dirichlet Green functions for the Laplacian! Indeed, one shows that for a ball of radius $a$ and surface $\Omega_{n-1}$ in $\mathbb{R}^{n}$ :

$$
G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\left\{\begin{array}{lr}
\frac{1}{2 \pi}\left(\ln \left|\mathbf{x}-\mathbf{x}^{\prime}\right|-\ln \left|\frac{r^{\prime}}{a} \mathbf{x}-\frac{a}{r^{\prime}} \mathbf{x}^{\prime}\right|\right) & (n=2)  \tag{6.50}\\
-\frac{1}{(n-2) \Omega_{n-1}}\left(\frac{1}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right| n-2}-\frac{1}{\left|\left(r^{\prime} / a\right) \mathbf{x}-\left(a / r^{\prime}\right) \mathbf{x}^{\prime}\right|^{n-2}}\right) & (n>2)
\end{array}\right.
$$

leads (EXERCISE) to a unified exprssion valid for $n \geq 2$ :

$$
\begin{equation*}
P_{r}(\cos \gamma)=\left.\partial_{n^{\prime}} G_{\mathrm{D}}\right|_{r^{\prime}=a}=\left.\frac{1}{\Omega_{n-1}} a^{n-2} \frac{a^{2}-r^{2}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|^{n}}\right|_{r^{\prime}=a} \tag{6.51}
\end{equation*}
$$

Clearly, if $\Psi\left(a, \theta^{\prime}, \phi^{\prime}\right) \neq 0$ and $r>a, F(\mathbf{x}) \neq 0$ somewhere in the region $r<a$, and vice-versa.

### 6.6.4 An important expansion for Green's Functions in Spherical Coordinates

The angular dependence in the Green functions such as derived above is quite complicated and may well not yield a solution in closed form when integrated, so it is often sensible to use a expansion appropriate to the coordinate system selected for the problem. Indeed, let us do this for the Laplacian in spherical coordinates.

In spherical coordinates, Green functions for the Laplacian operator all satisfy:

$$
\begin{align*}
\nabla_{\mathbf{x}}^{2} G\left(\mathbf{x}, \mathbf{x}^{\prime}\right) & =\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \\
& =\frac{1}{r^{2}} \delta\left(r-r^{\prime}\right) \sum_{l=0}^{\infty} \sum_{m=-l}^{l} Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi) \tag{6.52}
\end{align*}
$$

where the completeness relation (5.53) for spherical harmonics has been invoked.
We shall look for an expansion over separable terms of the form:

$$
G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\sum_{l=0}^{\infty} \sum_{m=-l}^{l} g_{l}\left(r, r^{\prime}\right) Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi)
$$

Inserting into eq. (6.52), we immediately find that $g_{l}\left(r, r^{\prime}\right)$ must satisfy the radial equation:

$$
r^{2} \nabla_{r}^{2} g_{l}\left(r, r^{\prime}\right)=\mathrm{d}_{r}\left[r^{2} \mathrm{~d}_{r} g_{l}\left(r, r^{\prime}\right)\right]-l(l+1) g_{l}\left(r, r^{\prime}\right)=\delta\left(r-r^{\prime}\right)
$$

We now find ourselves in the familiar territory of 1-dim Green-function problems and Sturm-Liouville operators. For instance, we can connect with eq. (6.13) for a 1 -dim Dirichlet Green function. We take $t_{0}=a$ and $t_{1}=b$, corresponding to two concentric spheres of radius $a$ and $b$,

We have $\alpha\left(r^{\prime}\right)=r^{\prime 2}$ and, with $f_{1}=r^{l}$ and $f_{2}=r^{-(l+1)}, W\left(r^{\prime}\right)=-(2 l+1) / r^{\prime 2}$. Also, let $r_{<} \equiv \min \left(r, r^{\prime}\right)$ and $r_{>} \equiv \max \left(r, r^{\prime}\right)$. It takes only a straightforward computation to arrive at (EXERCISE):

$$
\begin{equation*}
G_{\mathrm{D}}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\sum_{l=0}^{\infty} \sum_{m=-l}^{l} \frac{Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi)}{(2 l+1)\left[1-(a / b)^{2 l+1}\right]}\left(r_{<}^{l}-\frac{a^{2 l+1}}{r_{<}^{l+1}}\right)\left(\frac{r_{>}^{l}}{b^{2 l+1}}-\frac{1}{r_{>}^{l+1}}\right) \tag{6.53}
\end{equation*}
$$

Inspection of the last two factors shows that this expression vanishes at $r=a$ and $r=b$ (and when $r^{\prime}=a$ or $r^{\prime}=b$ ), as it should. We did not have to require this since it is built in the derivation of the 1 -dim Dirichlet Green function. Two important cases:

$$
\begin{array}{ll}
G_{\mathrm{D}}\left(\mathrm{x}, \mathbf{x}^{\prime}\right)=\sum_{l=0}^{\infty} \sum_{m=-l}^{l} \frac{Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi)}{(2 l+1)} r_{<}^{l}\left(\frac{r_{>}^{l}}{b^{2 l+1}}-\frac{1}{r_{>}^{l+1}}\right) & (a=0) \\
G_{\mathrm{D}}\left(\mathrm{x}, \mathbf{x}^{\prime}\right)=\sum_{l=0}^{\infty} \sum_{m=-l}^{l} \frac{Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi)}{(2 l+1)} \frac{1}{r_{>}^{l+1}}\left(\frac{a^{2 l+1}}{r_{<}^{l+1}}-r_{<}^{l}\right) & (b \rightarrow \infty) \tag{6.55}
\end{array}
$$

The first expression gives the Green function inside a sphere of radius $b$; the second one, outside a sphere of radius $a$ and all the way to infinity. When there are no boundary surfaces, we obtain over all space:

$$
\begin{equation*}
G\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=-\sum_{l=0}^{\infty} \sum_{m=-l}^{l} \frac{1}{2 l+1} \frac{r_{<}^{l}}{r_{>}^{l+1}} Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) Y_{l m}(\theta, \phi) \tag{6.56}
\end{equation*}
$$

Comparing with eq. (6.42) with $\kappa=0$ also yields a useful expansion of the ubiquitous distance factor $1 /\left|\mathbf{x}-\mathbf{x}^{\prime}\right|$.
When $0 \leq r \leq b$ (interior case) we can immediately rewrite (EXERCISE) the surface integral in eq. (6.40) as:

$$
\sum_{l=0}^{\infty} \sum_{m=-l}^{l}\left[\int \Psi\left(b, \theta^{\prime}, \phi^{\prime}\right) Y_{l m}^{*}\left(\theta^{\prime}, \phi^{\prime}\right) \mathrm{d} \Omega^{\prime}\right]\left(\frac{r}{b}\right)^{l} Y_{l m}(\theta, \phi)
$$

where $\Psi\left(b, \theta^{\prime}, \phi^{\prime}\right)$ is specified on the surface $r=b$. The normal derivative of the Green function on the surface, $\partial G / \partial n^{\prime}=\partial G /\left.\partial r^{\prime}\right|_{r^{\prime}=b}$, has been evaluated for $r_{<}=r$ and $r_{>}=r^{\prime}$ since $r<r^{\prime}=b$. Also, the surface element on a sphere of radius $b$ is $d S^{\prime}=b^{2} \mathrm{~d} \Omega^{\prime}$. This expression is still rather complicated, but it simplifies considerably if $\Psi\left(b, \theta^{\prime}, \phi^{\prime}\right)$ exhibits a symmetry (eg. azimuthal). Also, if one can write $\Psi\left(b, \theta^{\prime}, \phi^{\prime}\right)$ as a linear combination of spherical harmonics, the angular integration becomes trivial due to the orthonormality of the harmonics, and only a few terms in the sums might contribute.

### 6.6.5 An Elliptic Problem with a Twist: the Time-independent Schrödinger Equation

The time-independent Schrödinger equation (TISE) for a potential $V(\mathbf{x})$ takes the following suggestive form:

$$
\begin{equation*}
\left(\nabla^{2}+k^{2}\right) \psi(\mathbf{x})=\frac{2 m}{\hbar^{2}} V(\mathbf{x}) \psi(\mathbf{x}) \tag{6.57}
\end{equation*}
$$

where $k^{2}=2 m E / \hbar^{2}$. Although the right-hand side is not inhomogeneous, this in no way invalidates our previous results.

For bound states $(E<0)$ of an attractive potential, $k^{2}=\lambda<0$, and we have the integral equation:

$$
\psi(\mathbf{x})=-\frac{m}{2 \pi \hbar^{2}} \int \frac{\mathrm{e}^{-\kappa\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} V\left(\mathbf{x}^{\prime}\right) \psi\left(\mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime}
$$

with $\kappa^{2}=-2 m E / \hbar^{2}$. A somewhat simpler integral expression may be derived by writing the left-hand side of eq. (6.57) as its Fourier representation, and viewing $V(x) \psi(x)$ as the Fourier transform of the convolution $[V * \psi](\mathbf{k})$ (see section 5.5.4):

$$
-\int\left(k^{2}+\kappa^{2}\right) \psi(\mathbf{k}) \mathrm{e}^{\mathrm{i} \mathbf{k} \cdot \mathbf{x}} \mathrm{~d}^{3} k=\frac{2 m}{\hbar^{2}} \int\left[\frac{1}{(2 \pi)^{3 / 2}} \int V\left(\mathbf{k}-\mathbf{k}^{\prime}\right) \psi\left(\mathbf{k}^{\prime}\right) \mathrm{d}^{3} k^{\prime}\right] \mathrm{e}^{\mathrm{i} \mathbf{k} \cdot \mathbf{x}} \mathrm{~d}^{3} k
$$

Extracting $\psi(\mathbf{k})$, there comes:

$$
\psi(\mathbf{k})=-\frac{2 m}{(2 \pi)^{3 / 2} \hbar^{2}} \int \frac{V\left(\mathbf{k}-\mathbf{k}^{\prime}\right) \psi\left(\mathbf{k}^{\prime}\right)}{k^{2}+\kappa^{2}} \mathrm{~d}^{3} k^{\prime}
$$

See p. BF414 for more details and an application to the Yukawa potential.
For unbound states $(E>0), k^{2}=\lambda>0$, and we can immediately write the Lippmann-Schwinger equation:

$$
\begin{equation*}
\psi_{q}^{( \pm)}(\mathbf{x})=\frac{A}{(2 \pi)^{3 / 2}} \mathrm{e}^{\mathrm{i} \mathbf{q} \cdot \mathbf{x}}-\frac{m}{2 \pi \hbar^{2}} \int \frac{\mathrm{e}^{ \pm \mathrm{i} q\left|\mathbf{x}-\mathbf{x}^{\prime}\right|}}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} V\left(\mathbf{x}^{\prime}\right) \psi_{q}^{( \pm)}\left(\mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime} \tag{6.58}
\end{equation*}
$$

with $q=\sqrt{2 m E / \hbar^{2}}$.

The asymptotic form of the Lippmann-Schwinger equation is of particular interest. When $r \gg r^{\prime}$, we can expand $\left|\mathbf{x}-\mathbf{x}^{\prime}\right|=\sqrt{r^{2}-2 \mathbf{x} \cdot \mathbf{x}^{\prime}+r^{\prime 2}} \approx r-\hat{\mathbf{n}} \cdot \mathbf{x}^{\prime}$, with $\hat{\mathbf{n}}=\mathbf{x} / r$. Inserting into the integral equation yields:

$$
\begin{aligned}
\psi_{\mathbf{q}}^{( \pm)}(\mathbf{x}) & \underset{r \rightarrow \infty}{=} \frac{A}{(2 \pi)^{3 / 2}} \mathrm{e}^{\mathrm{i} \mathbf{q} \cdot \mathbf{x}}-\frac{m}{2 \pi \hbar^{2}} \frac{\mathrm{e}^{ \pm \mathrm{i} q r}}{r} \int \mathrm{e}^{\mp \mathrm{i} \mathrm{q} \hat{\mathbf{n}} \cdot \mathbf{x}^{\prime}} V\left(\mathbf{x}^{\prime}\right) \psi^{( \pm)}\left(\mathbf{x}^{\prime}\right) \mathrm{d}^{3} x^{\prime} \\
& =\frac{A}{(2 \pi)^{3 / 2}}\left[\mathrm{e}^{\mathrm{i} \mathbf{q} \cdot \mathbf{x}}+f_{ \pm}(\mathbf{q}) \frac{\mathrm{e}^{ \pm \mathrm{i} q r}}{r}\right]
\end{aligned}
$$

This expression represents the spatial dependence of a superposition of a plane wave and a scattered spherical wave propagagating inward or outward from the origin. The function $f_{ \pm}(\mathbf{q})$ is called the scattering amplitude; it also obeys an integral equation, eq. BF7.75, and its square modulus is directly related to experimental data.

### 6.7 A Hyperbolic Problem: the d'Alembertian Operator

With the Fourier integral representation (note the different normalisation and sign in the exponentials!):

$$
\begin{align*}
\Psi(\mathbf{x}, t) & =\frac{1}{2 \pi} \int_{-\infty}^{\infty} \Psi(\mathbf{x}, \omega) \mathrm{e}^{-\mathrm{i} \omega t} \mathrm{~d} \omega \\
\Psi(\mathbf{x}, \omega) & =\int_{-\infty}^{\infty} \Psi(\mathbf{x}, t) \mathrm{e}^{\mathrm{i} \omega t} \mathrm{~d} t \tag{6.59}
\end{align*}
$$

we can transform a typical inhomogeneous wave equation:

$$
\square \Psi(\mathbf{x}, t)=\nabla^{2} \Psi(\mathbf{x}, t)-\frac{1}{c^{2}} \partial_{t}^{2} \Psi(\mathbf{x}, t)=F(\mathbf{x}, t)
$$

where $F(\mathbf{x}, t)$ is a known source, to its Helmholtz form:

$$
\begin{equation*}
\left(\nabla^{2}+k^{2}\right) \Psi(\mathbf{x}, \omega)=F(\mathbf{x}, \omega) \tag{6.60}
\end{equation*}
$$

where $k^{2}>0$ can be taken as a short form for $(\omega / c)^{2}$.
Just as for the Laplacian operator, there exist Green functions for $\nabla^{2}+k^{2}$; we have found them a little earlier in eq. (6.43):

$$
\begin{equation*}
G^{( \pm)}(R)=-\frac{1}{4 \pi} \frac{\mathrm{e}^{ \pm \mathrm{i} k R}}{R} \tag{6.61}
\end{equation*}
$$

where $R=\left|\mathbf{x}-\mathbf{x}^{\prime}\right|$.
Now we are ready to derive the full Green functions for the d'Alembertian operator, which satisfy:

$$
\begin{equation*}
\square_{\mathbf{x}} G\left(\mathbf{x}, t ; \mathbf{x}^{\prime}, t^{\prime}\right)=\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \delta\left(t-t^{\prime}\right) \tag{6.62}
\end{equation*}
$$

or, in the frequency domain:

$$
\left(\nabla_{x}^{2}+k^{2}\right) G\left(\mathbf{x}, \mathbf{x}^{\prime}, \omega, t^{\prime}\right)=\delta\left(\mathbf{x}-\mathbf{x}^{\prime}\right) \mathrm{e}^{\mathrm{i} \omega t^{\prime}}
$$

Assume separable solutions of the form $G\left(\mathbf{x}, \mathbf{x}^{\prime}\right) \mathrm{e}^{\mathrm{i} \omega t^{\prime}}$; inserting into this equation, we get from (6.60) the solutions $G^{ \pm}\left(\mathbf{x}, \mathbf{x}^{\prime}, \omega, t^{\prime}\right)=-\mathrm{e}^{\mathrm{i}\left( \pm k R+\omega t^{\prime}\right)} / 4 \pi R$. Then, transforming back to the time domain and using the representation (5.16) for the $\delta$-function yields the Green functions:

$$
\begin{equation*}
G^{( \pm)}\left(\mathbf{x}, t ; \mathbf{x}^{\prime}, t^{\prime}\right)=-\frac{1}{8 \pi^{2} R} \int_{-\infty}^{\infty} \mathrm{e}^{\mathrm{i} \omega\left[ \pm R / c+\left(t^{\prime}-t\right)\right]} \mathrm{d} \omega=-\frac{1}{4 \pi} \frac{\delta\left(t^{\prime}-[t \mp R / c]\right)}{R} \tag{6.63}
\end{equation*}
$$

In higher dimensions, Green functions can contain $\delta$-functions and so may not be actual functions!
Using eq. (6.62), we also recognise that:
$\square_{\mathbf{x}} \int_{\text {allspace }} \mathrm{d}^{3} x^{\prime} \int_{-\infty}^{\infty} G^{( \pm)}\left(\mathbf{x}, t ; \mathbf{x}^{\prime}, t^{\prime}\right) F\left(\mathbf{x}^{\prime}, t^{\prime}\right) \mathrm{d} t^{\prime}=\int \mathrm{d}^{3} x^{\prime} \int_{-\infty}^{\infty} F\left(\mathbf{x}^{\prime}, t^{\prime}\right) \square_{\mathbf{x}} G^{( \pm)}\left(\mathbf{x}, t ; \mathbf{x}^{\prime}, t^{\prime}\right) \mathrm{d} t^{\prime}=F(\mathbf{x}, t)$
has the generic form $\square \Psi(\mathbf{x}, t)=F(\mathbf{x}, t)$, which shows that the general solution of a wave equation with sources can be written either as the retarded solution:

$$
\begin{align*}
\Psi(\mathbf{x}, t) & =\Psi_{\text {in }}(\mathbf{x}, t)+\iint_{-\infty}^{\infty} G^{(+)}\left(\mathbf{x}, t ; \mathbf{x}^{\prime}, t^{\prime}\right) F\left(\mathbf{x}^{\prime}, t^{\prime}\right) \mathrm{d}^{3} x^{\prime} \mathrm{d} t^{\prime} \\
& =\Psi_{\text {in }}(\mathbf{x}, t)-\frac{1}{4 \pi} \int \frac{F\left(\mathbf{x}^{\prime}, t_{\mathrm{ret}}\right)}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|_{\text {ret }}} \mathrm{d}^{3} x^{\prime} \tag{6.64}
\end{align*}
$$

or, equivalently, as the advanced solution:

$$
\begin{align*}
\Psi(\mathbf{x}, t) & =\Psi_{\text {out }}(\mathbf{x}, t)+\iint_{-\infty}^{\infty} G^{(-)}\left(\mathbf{x}, t ; \mathbf{x}^{\prime}, t^{\prime}\right) F\left(\mathbf{x}^{\prime}, t^{\prime}\right) \mathrm{d}^{3} x^{\prime} \mathrm{d} t^{\prime} \\
& =\Psi_{\text {out }}(\mathbf{x}, t)-\frac{1}{4 \pi} \int \frac{F\left(\mathbf{x}^{\prime}, t_{\text {adv }}\right)}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|_{\text {adv }}} \mathrm{d}^{3} x^{\prime} \tag{6.65}
\end{align*}
$$

where the suffixes ret and $a d v$ stand for the fact that $t^{\prime}$ must be evaluated at the retarded time $t_{\text {ret }}=t-R / c$, or the advanced time $t_{\mathrm{adv}}=t+R / c$. This ensures the proper causal behaviour of the solutions, in the sense that, eg., the solution at time $t$ is only influenced by the behaviour of the source point $\mathrm{x}^{\prime}$ at time $t-R / c . \Psi_{\text {in }}$ and $\Psi_{\text {out }}$ are possible plane-wave solutions of the homogeneous wave equation for $\Psi$. Most often they can be taken to be zero.

### 6.8 Initial Value Problem with Constraints

The Initial Value Problem (IVP) consists in finding which data must be specified at a given time for the time evolution of variables to be uniquely determined by their equations of "motion".

By initial data, one means the state of the system of variables and their first-order derivatives on a threedimensional spacelike hypersurface; usually, this means at some time $t_{0}$ everywhere in space. The IVP together with the evolution equations constitute the Cauchy Problem of the theory. If the IVP can be solved, the dynamical behaviour of the system can be uniquely predicted from its initial data.

Most often, the equations of "motion" take the form of a set of wave equations, each of the form $\square f=F$. If they always told the whole story, the Cauchy problem would be solved by specifying the value of $f$ and its firstorder time derivatives at $t=t_{0}$. Things are not so simple, however, when there are inherent, built-in constraints on the initial data. Those constraint equations must be discovered and solved. Also, we must find which initial data we are allowed to specify freely.

We study in some depth a very important example: Maxwell's theory. In linear, unpolarised and unmagnetised media, Maxwell's equations are:

$$
\begin{array}{rlr}
\boldsymbol{\nabla} \cdot \mathbf{E}=4 \pi k_{\mathrm{e}} \rho & \boldsymbol{\nabla} \times \mathbf{B}-\frac{k_{\mathrm{m}}}{k_{\mathrm{e}}} \partial_{t} \mathbf{E}=4 \pi k_{\mathrm{m}} \mathbf{J} \\
\boldsymbol{\nabla} \cdot \mathbf{B}=0 & \boldsymbol{\nabla} \times \mathbf{E}+\partial_{t} \mathbf{B}=0 \tag{6.66}
\end{array}
$$

where $k_{\mathrm{e}}$ and $k_{\mathrm{m}}$ are constants that depend on the system of units, and $k_{\mathrm{e}} / k_{\mathrm{m}}=c^{2}$, with $c$ the speed of light. The source terms $\rho$ and $\mathbf{J}$ are not independent; indeed, one derives from eq. (6.66) a continuity equation:

$$
\begin{equation*}
\partial_{t} \rho=\frac{1}{4 \pi k_{\mathrm{e}}} \boldsymbol{\nabla} \cdot \partial_{t} \mathbf{E}=-\boldsymbol{\nabla} \cdot \mathbf{J} \tag{6.67}
\end{equation*}
$$

The two homogeneous equations are equivalent to:

$$
\begin{equation*}
\mathbf{E}=-\partial_{t} \mathbf{A}-\boldsymbol{\nabla} \Phi \quad \mathbf{B}=\boldsymbol{\nabla} \times \mathbf{A} \tag{6.68}
\end{equation*}
$$

Observe that $\mathbf{A}$ is determined only up to a term $\boldsymbol{\nabla} f$, where $f(\mathbf{x}, t)$ is an arbitrary differentiable function. In fact, if we perform the gauge transformations $\Phi \rightarrow \Phi-\partial_{t} f$ and $\mathbf{A} \rightarrow \mathbf{A}+\boldsymbol{\nabla} f$, neither $\mathbf{E}$ nor $\mathbf{B}$ change! We say that Maxwell's theory is gauge-invariant.

The inhomogeneous Maxwell equations (6.66) become second-order equations for $\Phi$ and $\mathbf{A}$ :

$$
\begin{align*}
\nabla^{2} \Phi+\partial_{t}(\nabla \cdot \mathbf{A}) & =-4 \pi k_{\mathrm{e}} \rho  \tag{6.69}\\
\square \mathbf{A}-\nabla\left(\nabla \cdot \mathbf{A}+\frac{1}{c^{2}} \partial_{t} \Phi\right) & =-4 \pi k_{\mathrm{m}} \mathbf{J}
\end{align*}
$$

### 6.8.1 Second-order Cauchy problem using transverse/longitudinal projections

While eq. (6.69) are gauge-invariant, $\mathbf{A}$ and $\Phi$ themselves are not, at least at first sight. What this means is that the time-evolution of at least some of the four quantities $\Phi$ and $\mathbf{A}$ cannot be uniquely determined from their initial conditions and eq. (6.69) since we can always perform an arbitrary gauge transformation on them at some arbitrary later time $t$, as often as we wish. This is a serious issue which must be understood and addressed if $\Phi$ and $\mathbf{A}$ are to be of any use at all.

One instructive approach is to note that according to the Helmholtz theorem any differentiable 3-dim vector field that goes to zero at infinity faster than $1 / r$ may be written as the sum of two vectors:

$$
\mathbf{A}=\underbrace{\boldsymbol{\nabla} u}_{\mathbf{A}_{\mathrm{L}}}+\underbrace{\boldsymbol{\nabla} \times \mathbf{w}}_{\mathbf{A}_{\mathrm{T}}}
$$

The first term, $\mathbf{A}_{\mathrm{L}}=\boldsymbol{\nabla} u$, whose curl vanishes identically, is called the longitudinal part (or projection) of $\mathbf{A}$; the second, $\mathbf{A}_{\mathrm{T}}=\boldsymbol{\nabla} \times \mathbf{w}$, whose divergence vanishes identically, is called the transverse part (or projection) of $\mathbf{A}$. Since the longitudinal and transverse projections are perpendicular to each other, we can decompose (project out) Maxwell's equations for the fields $\mathbf{E}$ and $\mathbf{B}$ and for the potential $\mathbf{A}$ into longitudinal and tranverse parts.

Before we do this, however, we note that since by definition $\nabla \cdot \mathbf{J}_{\mathrm{T}}=0$, the continuity equation for the sources does not involve the transverse part of $\mathbf{J}$. Also, $\boldsymbol{\nabla} \cdot \mathbf{A}$ is really $\boldsymbol{\nabla} \cdot \mathbf{A}_{\mathrm{L}}$ and contains no information about $\mathbf{A}_{\mathrm{T}}$

Project the second equation (6.69). The transverse projection immediately gives:

$$
\begin{equation*}
\square \mathbf{A}_{\mathrm{T}}=-4 \pi k_{\mathrm{m}} \mathbf{J}_{\mathrm{T}} \tag{6.70}
\end{equation*}
$$

where we have used the fact a gradient is a longitudinal object. The two transverse components $\mathbf{A}_{\mathrm{T}}$ satisfy a proper wave equation and correspond to physically observable quantities, in the sense that being transverse, they are unaffected by $\mathbf{A} \rightarrow \mathbf{A}+\boldsymbol{\nabla} f$, which can change only the longitudinal component $\mathbf{A}_{\mathrm{L}}$. Therefore, the time evolution of the two transverse $\mathbf{A}_{\mathrm{T}}$ is not arbitrary and they have a well-posed Cauchy problem.

Now, remembering that $\square=\nabla^{2}-\left(\partial_{t}^{2}\right) / c^{2}$, take the divergence of the longitudinal projection of (6.69):

$$
\begin{aligned}
\boldsymbol{\nabla} \cdot\left[\square \mathbf{A}_{\mathrm{L}}-\boldsymbol{\nabla}\left(\boldsymbol{\nabla} \cdot \mathbf{A}_{\mathrm{L}}+\frac{1}{c^{2}} \partial_{t} \Phi\right)+4 \pi k_{\mathrm{m}} \mathbf{J}_{\mathrm{L}}\right] & =\square\left(\boldsymbol{\nabla} \cdot \mathbf{A}_{\mathrm{L}}\right)-\nabla^{2}\left(\boldsymbol{\nabla} \cdot \mathbf{A}_{\mathrm{L}}\right)-\frac{\partial_{t} \nabla^{2} \Phi}{c^{2}}+4 \pi k_{\mathrm{m}} \boldsymbol{\nabla} \cdot \mathbf{J}_{\mathrm{L}} \\
& =-\frac{1}{c^{2}} \partial_{t}\left[\partial_{t}\left(\boldsymbol{\nabla} \cdot \mathbf{A}_{\mathrm{L}}\right)+\nabla^{2} \Phi+4 \pi k_{\mathrm{e}} \rho\right]
\end{aligned}
$$

where eq. (6.67) has been invoked in the second line. But the terms in the square bracket on that line are just the first of equations (6.69). Therefore, the longitudinal projection of the second Maxwell equation for the 3-vector potential contains no information about $\boldsymbol{\nabla} \cdot \mathbf{A}$ that is not in the first equation. But that is really an equation for $\Phi$ with $\boldsymbol{\nabla} \cdot \dot{\mathbf{A}}$ (more precisely, $\boldsymbol{\nabla} \cdot \dot{\mathbf{A}}_{\mathrm{L}}$ ) as a source together with $\rho$. Therefore, Maxwell's theory cannot uniquely determine the time evolution of the divergence of the 3 -vector potential. Nor can it uniquely determine the time evolution of $\Phi$, since $\Phi$ is gauge-variant. Systems whose time-evolution involves arbitrary functions are often called singular.

### 6.8.2 Choices for the divergence of $A$

Since the theory does not know $\boldsymbol{\nabla} \cdot \mathbf{A}$ (and its first-order time derivative for that matter), we have to tell it what it is by making an arbitrary choice. If we choose $\boldsymbol{\nabla} \cdot \mathbf{A}$ to vanish (Coulomb condition), the vector potential becomes pure transverse (hence the name "transverse gauge also given to this choice), and the equation for $\Phi$ becomes a Poisson-type equation with solution:

$$
\begin{equation*}
\Phi(\mathbf{x}, t)=k_{\mathrm{e}} \int \frac{\rho\left(\mathbf{x}^{\prime}, t\right)}{\left|\mathbf{x}-\mathbf{x}^{\prime}\right|} \mathrm{d}^{3} x^{\prime} \tag{6.71}
\end{equation*}
$$

This looks innocuous enough until we realise that any change in the source is instantaneously reflected in the scalar potential. The Coulomb condition leads to acausal behaviour, which is also a reflection of the fact that the condition is not relativistically covariant, in the sense that it is not necessarily the same in all inertial frames. But the equation for $\Phi$ is not a classical wave equation, and $\Phi$ does not really propagate as a wave, so one should not expect proper causal behaviour from it.

The problem is seemingly resolved just by choosing instead the Lorenz condition: $\boldsymbol{\nabla} \cdot \mathbf{A}=-\partial_{t} \Phi / c^{2}$, which turns eq. (6.69) into standard wave equations of the type $\square$ (potential) $=$ source with causal solution eq. (6.64). Then one can calculate the energy radiated to infinity following standard treatments (eg. chapter 10 in Griffiths or chapter 14 in Jackson) and find that the scalar potential does make a mathematical contribution to the energy radiated to infinity.. This, however, would not have happened if we had chosen the Coulomb condition. Therefore, we should not attach any physical significance to that contribution: it arises simply out of consistency with this particular choice of condition on $\boldsymbol{\nabla} \cdot \mathbf{A}$.

Conditions that do not continue to hold automatically in the future must be enforced by hand at all times. This could be the case for conditions such as the ones we have imposed. Fortunately, one can show that the Coulomb and Lorenz conditions propagate forward in time once imposed at initial time. Take for instance the wave equation for $\mathbf{A}$ in eq. (6.69). Impose $\boldsymbol{\nabla} \cdot \mathbf{A}=0$ and $\partial_{t}(\boldsymbol{\nabla} \cdot \mathbf{A})=0$ at some time in both the equations for $\Phi$ and $\mathbf{A}$. Then take the divergence of the resulting wave equation for $\mathbf{A}$ and the time derivative of the resulting equation for $\Phi$, and use the continuity equation to obtain:

$$
\square(\nabla \cdot \mathbf{A})=0
$$

showing that if we choose $\boldsymbol{\nabla} \cdot \mathbf{A}=0$ and $\partial_{t}(\boldsymbol{\nabla} \cdot \mathbf{A})=0$ at $t=t_{0}$, it remains the same for all time. Similarly, it is trivial to show that imposing the Lorenz condition everywhere in space at $t=t_{0}$ also leads to a wave equation for $\boldsymbol{\nabla} \cdot \mathbf{A}$.

Note also that the energy radiated to infinity can be calculated solely in terms of $\mathbf{B}$ and, therefore, of $\mathbf{A}$, without $\Phi$ being involved. In fact, since $\mathbf{B}=\boldsymbol{\nabla} \times\left(\mathbf{A}_{L}+\mathbf{A}_{T}\right)=\boldsymbol{\nabla} \times \mathbf{A}_{\mathrm{T}}$, only the two transverse components of A contribute! These are independent of any choice, and thus entirely physical (contrary to assertions sometimes made that the electromagnetic potential is not as physical as the fields because it is not gauge-invariant-now we know that this only applies to $\mathbf{A}_{\mathrm{L}}$ and $\Phi$ ).

### 6.8.3 First-order Cauchy problem

Now consider this same Cauchy Problem from the point of view of the fields $\mathbf{E}$ and $\mathbf{B}$. Taking the curl of the first-order curl equations (6.66), we arrive at:

$$
\begin{align*}
& \square \mathbf{E}=4 \pi k_{\mathrm{e}} \boldsymbol{\nabla} \rho+4 \pi k_{\mathrm{m}} \partial_{t} \mathbf{J}  \tag{6.72}\\
& \square \mathbf{B}=-4 \pi k_{\mathrm{m}} \boldsymbol{\nabla} \times \mathbf{J}
\end{align*}
$$

These look like wave equations for six quantities. But only those of their solutions which also satisfy the first-order field equations (6.66), including at $t=t_{0}$, are acceptable.

The two first-order divergence equations contain no time derivatives and are thus constraints on $\mathbf{E}$ and $\mathbf{B}$ at $t=t_{0}$. The constraint equation on $\mathbf{E}$ can be rewritten $\nabla^{2} u=\rho$, a Poisson-type equation which can be solved for $u$ at initial time so long as $\rho$ falls off faster than $1 / r^{2}$ at infinity). In the case of $\mathbf{B}$, the scalar field $u$ satisfies
a Laplace equation everywhere and is therefore zero. So $\mathbf{B}$ has no longitudinal component, only a transverse one just as we had found at the end of the previous section. In both cases, the longitudinal component is either zero or can be solved for at $t_{0}$, so cannot be freely specified.

Now look at the two first-order Maxwell field equations (6.66) which contain time derivatives. Suppose we specify $\mathbf{E}$ and $\partial_{t} \mathbf{E}$ at $t=t_{0}$, which are needed to solve the $2^{\text {nd }}$-order equations, eq. (6.72). Then the two transverse components of $\mathbf{B}$ are determined by $\nabla \times \mathbf{B}=4 \pi k_{\mathrm{m}} \mathbf{J}+\partial_{t} \mathbf{E} / c^{2} ; \partial_{t} \mathbf{B}$ is determined, also at $t=t_{0}$, by the curl equation for $\mathbf{E}$. Therefore, once we have specified the two transverse components of $\mathbf{E}$ and their time derivatives, Maxwell's first-order equations take over and determine the others at $t=t_{0}$. Alternatively, we could have started with the two transverse components of $\mathbf{B}$; specifying them and their time derivatives at $t=t_{0}$ constrains all the other field components and time derivatives.

You can also use (exercise) the transverse/longitudinal projections of the first-order equations (6.66) to show that in source-free space, only the transverse components of $\mathbf{E}$ and $\mathbf{B}$ obey a classical wave equation.

Thus, the results of the first-order Cauchy-data analysis are fully consistent with the second-order analysis on A: only two transverse components correspond to independent, physical dynamical degrees of freedom, Also, one of the advantages of this Cauchy analysis is that it does not rely on some particular solution, but is valid for any electromagnetic field and potential.
Addendum: The Lorenz condition $\boldsymbol{\nabla} \cdot \mathbf{A}=-\partial_{t} \Phi / c^{2}$, which is imposed in almost all treatments of electromagnetic radiation, could lead you to believe that $\Phi$ and the three components of A propagate to infinity, whereas I hope to have convinced you that only the transverse components of $\mathbf{A}$ do.

The Lorenz condition relates the longitudinal component of $\mathbf{A}$ to $\Phi$. Now I will show that $\mathbf{A}_{L}$ can be made to disappear without affecting Maxwell's equations for the fields and the potentials.

The key observation is that one can change both $\mathbf{A}$ and $\Phi$ to new functions that still obey the Lorenz condition. Indeed, let $f$ be some scalar function that satisfies the homogeneous wave equation $\square f=\nabla^{2} f-\frac{1}{c^{2}} \partial_{t}^{2} f=0$. Then add $\nabla^{2} f$ to $\nabla \cdot \mathbf{A}$ and $\frac{1}{c^{2}} \partial_{t}^{2} f$ to $-\partial_{t} \Phi / c^{2}$ to obtain:

$$
\begin{equation*}
\nabla \cdot(\mathbf{A}+\nabla f)=-\frac{1}{c^{2}} \partial_{t}\left(\Phi-\partial_{t} f\right) \tag{6.73}
\end{equation*}
$$

This shows that gauge-transformed potentials still satisfy the Lorenz condition! As noted before, it is important to keep in mind that since the transformation shifts $\mathbf{A}$ by a gradient, which is a longitudinal object, it does not affect the transverse components of $\mathbf{A}$.

Now, for the first time, we shall have to look at actual solutions of the wave equations for $\mathbf{A}$ and $\Phi$. To make things as simple as possible, take plane-wave solutions $\mathbf{A}=\mathbf{A}_{0} \mathrm{e}^{\mathrm{i}(k x-\omega t)}$, where the $x$-axis has been aligned along the direction of propagation, and $\Phi=\Phi_{0} \mathrm{e}^{\mathrm{i}(k x-\omega t)}$. Then:

$$
\boldsymbol{\nabla} \cdot \mathbf{A}=\partial_{x} A_{x}=\mathrm{i} k A_{0 x} \mathrm{e}^{\mathrm{i}(k x-\omega t)}, \quad \quad \partial_{t} \Phi=-\mathrm{i} \omega \Phi_{0} \mathrm{e}^{\mathrm{i}(k x-\omega t)}
$$

Inserting into the Lorenz condition with $\omega / k=c$ yields, as expected, a relation between the longitudinal component $A_{x}$ and $\Phi: A_{0 x}=\Phi_{0} / c$.

Now fold in $f=f_{0} \mathrm{e}^{\mathrm{i}(k x-\omega t)}$ into eq. (6.73) for the gauge-transformed potentials, to get:

$$
\mathrm{i} k\left(A_{0 x}+\mathrm{i} k f_{0}\right) \mathrm{e}^{\mathrm{i}(k x-\omega t)}=\mathrm{i} \frac{\omega}{c^{2}}\left(\Phi_{0}+\mathrm{i} \omega f_{0}\right) \mathrm{e}^{\mathrm{i}(k x-\omega t)}
$$

Since $f_{0}$ is arbitrary, we can choose it to cancel $A_{0 x}$, which at the same time gets rid of $\Phi_{0}$, leaving us with the transverse components of A only!

Although the analysis under the Lorenz condition is quite a bit more involved than with the Coulomb condition, the conclusions are the same: only the two transverse components of $\mathbf{A}$ propagate, in the sense that they carry energy to infinity.


[^0]:    ${ }^{\dagger}$ These notes generally follow the conventions set by the ISO (International Standards Organisation) for mathematical typography, with one important exception: as in BF, vectors and tensors are in bold upright (u) instead of bold italic font ( $\boldsymbol{u}$ ). Sans-serif fonts denote matrices, eg. M.

[^1]:    ${ }^{\dagger}$ Infinite linear combinations (series) require extra topological structure on $\mathcal{V}$ so as to allow the notion of convergence.

[^2]:    ${ }^{\dagger}$ This assumes that $\mathcal{V}$ 's dimension is finite!

[^3]:    ${ }^{\dagger}$ So far as I know, this metaphor was first proposed by Misner, Thorne and Wheeler (MTW) in their monumental textbook, Gravitation.

[^4]:    ${ }^{\dagger}$ Here, the meaning of "dual" has no relation to its other use in "dual" space or basis.

[^5]:    ${ }^{\dagger}$ Some authors prefer the notation $\nabla \wedge$ for the exterior derivative.

[^6]:    ${ }^{\dagger}$ Again, with a mostly negative metric, such as in Jackson's Classical Electrodynamics, there would be no minus sign on the right-hand side. This is because $\mathbf{F}$ has opposite sign between the two conventions so as to obtain the same relations between the electric and magnetic fields and the vector and scalar potentials.

[^7]:    ${ }^{\dagger}$ Since a binary composition is not necessarily a multiplication, we avoid the potentially misleading $f \cdot g$ notation of Byron and Fuller (and many others).

[^8]:    ${ }^{\dagger}$ Actually, conjugacy is only a particular type of equivalence.

[^9]:    ${ }^{\dagger}$ Since there is little scope for confusion in the context of $S_{n}$, we replace "group composition" with "group product".

[^10]:    ${ }^{\dagger}$ We are getting a little bit ahead of ourselves. If we acted from the left, vectors would be row-vectors, and we would have $x^{i} \pi=$ $x^{j} T_{j}{ }^{i}(\pi)$, with $\pi_{5}$ and $\pi_{6}$ interchanged. The difference between the right and left actions of a group will be discussed in section 2.4.2.
    ${ }^{\ddagger}$ Note that this subgroup being Abelian is not sufficient to establish that it is invariant; it must be self-conjugate with respect to all elements in $S_{3}$.

[^11]:    ${ }^{\dagger}$ This is because $\operatorname{Tr} A B=A_{i}{ }^{j} B_{j}{ }^{i}=B_{j}{ }^{i} A_{i}{ }^{j}=\operatorname{Tr} B A$.

[^12]:    ${ }^{1}$ See, eg: http://demonstrations.wolfram.com/RotatingCubesAboutAxesOfSymmetry3DRotationIsNonAbelian/.

[^13]:    ${ }^{\dagger}$ As a mnemonic device, in the commutator on the left take the first and third indices for the $\delta$, and the second and fourth for $J$, in that order, for the first term on the right; then cycle through to generate the last three terms, using first $J_{i j}=-J_{j i}$ so that after cycling one index on the delta belongs to the $(m n)$ pair and the other to the $(p q)$ pair.

[^14]:    ${ }^{\dagger}$ The fact that for a given $i, j Y_{i}$ and $Z_{j}$ in the decoupled basis of eq. (3.18) form an Abelian subalgebra does not invalidate our argument, because this subalgebra is not invariant.
    ${ }^{\ddagger}$ Another inner product which is sometimes introduced is the Hilbert-Schmidt product: $(\mathbf{A}, \mathbf{B})=\operatorname{Tr}\left(\mathbf{A}^{\dagger} \mathbf{B}\right)$.

[^15]:    ${ }^{\dagger}$ For an accessible introduction, see http://dx.doi.org/10.1063/1. 530244 (free access from a UofT computer) and references cited therein.

[^16]:    ${ }^{\dagger}$ The best proof of the Cauchy-Goursat theorem relies only on $f(z)$ being holomorphic on $S$ and its boundary $C$, but it is a lot fussier than the one we present, which strictly speaking applies to the Cauchy theorem (continuity of derivatives assumed).

[^17]:    ${ }^{\dagger}$ It could be argued that this is not sufficient because there could exist singularities whose contributions to the contour integral cancel out. But it is easy to show that the $n^{\text {th }}$-order derivative of $g(\omega)$ always exists, simply by differentiating eq. (4.25) $n$ times and noting that the resulting integrand remains bounded when $\tau>0$ and $\Im(\omega)>0$. On the real $\omega$ axis, $g(\omega)$ is bounded, and any branch point can be bypassed without changing anything.

[^18]:    ${ }^{\dagger}$ As noted above, the sign of the argument of the exponentials is also a matter of convention, but it must change from one transform to the other.

[^19]:    ${ }^{\dagger}$ Sometimes, $\operatorname{sinc}(x)$ is defined as $(\sin x) / x$.

[^20]:    ${ }^{\dagger}$ Here we shall not bother with the distinction between regular and singular Sturm-Liouville systems.

[^21]:    *See pp. BF399-401 for useful properties of the step-function.

[^22]:    ${ }^{\dagger}$ The response function in eq. (4.24)is now seen to be a Green function.

[^23]:    ${ }^{\dagger}$ Anyway, it is straightforward to generalise our discussion to any number of spatial dimensions plus one time dimension.

[^24]:    ${ }^{\dagger}$ In compact domains without boundaries (see remark in last section), this defining equation is inconsistent and must be modified (EXERCISE: Can you see why, and how?)
    ${ }^{\dagger}$ Although we call it "inhomogeneous", nothing in what we will do here prevents $F(\mathbf{x})$ from depending on $\Psi(\mathbf{x})$

