

# INTERACTIONS OF CHARGED PARTICLES

## WITH MATTER

- FUNDAMENTALLY RESPONSIBLE FOR "WORKING" OF ALL PARTICLE DETECTORS
    - TRACKERS - MINIMAL INTERACTION
    - CALORIMETERS - MAXIMAL INTERACTION
- (EVEN NEUTRAL PARTICLES MUST BE CONVERTED TO CHARGED PARTICLES BEFORE BEING DETECTED)

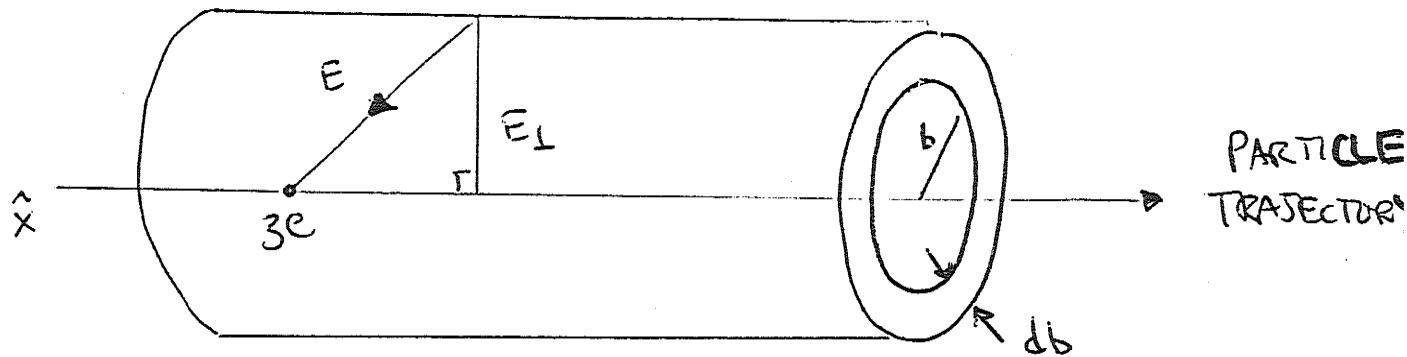
### CONSIDER

#### 1) IONIZATION

- AVERAGE ENERGY LOSS
- LANDAU TAIL

#### 2) MULTIPLE SCATTERING

# INFLUENCE OF PARTICLE (CHARGE $z$ ) ON NEARBY ELECTRONS



MOMENTUM TRANSFERRED  $\vec{P} = P_{\parallel} \hat{u}_{\parallel} + P_{\perp} \hat{u}_{\perp}$

$$P_{\parallel} = 0$$

$$P_{\perp} = \int_{-\infty}^{\infty} e E_{\perp} dt$$

$$= \frac{e}{v} \int E_{\perp} dx$$

$$= \frac{e}{v} \frac{1}{2\pi b} \int_{\text{surface}} E_{\perp} (2\pi b) dx$$

$$= \frac{e}{v} \frac{1}{2\pi b} \left( \frac{3e}{G} \right)$$

(2)

Energy transferred  $E = P^2/2m$

$$E = \left(\frac{3e^2}{\epsilon_0}\right)^2 \frac{1}{2mc^2} \frac{1}{4\pi^2} \frac{1}{b^2} \frac{1}{\beta^2}$$

N.B.  $\frac{dE}{E} = -\frac{2db}{b}$

Now define:  $\phi(E, E') dE' dx$  to be the probability that a particle loses energy between  $E' \rightarrow E + dE$  in path length  $dx$

$$\begin{aligned} \Rightarrow \phi(E, E') dE' dx &= -n_0 (2\pi b) db dx \\ &\quad \hookrightarrow \# \text{ electrons / unit volume} \\ &= n_0 \pi b^2 dx \left(-\frac{2db}{b}\right) \\ &= n_0 \pi \frac{3^2 e^2 (mc^2)}{\beta^2} \frac{dE'}{E'^2} dx \\ &\equiv \frac{A (\rho dx)}{\beta^2} \frac{dE'}{E'^2} \end{aligned}$$

(3)

$$\tilde{A} = 0.154 \frac{Z^2}{A} \frac{\text{MeV cm}^2}{\text{gm}}$$

$Z$  = Charge of Media  
 $A$  = Atomic # of Media

## PRIMARY IONIZATION

- MEAN NUMBER OF "ENCOUNTERS" BETWEEN  $E'_{\min}$  and  $E'_{\max}$  (CAUSE "SINGLE IONIZATION") PER UNIT LENGTH

$$\begin{aligned}\frac{n_{\text{prim}}}{\Delta x} &= \left\langle \frac{dN}{dx} \right\rangle = \int_{E_{\min}}^{E'_{\max}} \phi(E', E) dE' \\ &= \frac{\tilde{A} \beta}{B^2} \int_{E_{\min}}^{E_{\max}} \frac{dE'}{E'^2} \\ &= \frac{\tilde{A} \beta}{B^2} \left( \frac{1}{E_{\min}} - \frac{1}{E_{\max}} \right)\end{aligned}$$

FOR MOST GASES

$$\frac{\tilde{A} \beta}{E_{\max}} \ll 1$$

SO AS  $\beta \rightarrow 1$   
 THIS CAN BE  
 IGNORED

$$\Rightarrow \frac{n_{\text{prim}}}{\Delta x} \approx \frac{\tilde{A} \beta}{B^2} \frac{1}{E_{\min}}$$

$E_{\min} \geq I_0$  (THE THRESHOLD TO IONISE THE GAS)

FOR MOST GASES QUANTUM EFFECTS ARE IMPORTANT (SCREENING ETC.). FOR HYDROGEN BETHE PERFORMED A "FULL CALCULATION". CONCLUDED:

$$E_{\min} = \frac{I_0}{c \left[ \ln \frac{2mc^2}{I_0} \frac{\beta^2}{1-\beta^2} + S - \beta^2 \right]}$$

$$I_0 = 15.4 \text{ eV} \quad c = 0.29 \quad S = 3.04$$

$$\Rightarrow E_{\min} = 3.5 \text{ eV} \quad n_{\text{prim}} = 4.6 / \text{cm}^3$$

FOR OTHER GASES

$$E_{\min} = \frac{\tilde{A}_1 \rho}{A_1 \left[ A_2 + \ln \left( \frac{\beta^2}{1-\beta^2} \right) - \beta^2 \right]}$$

See table for  $A_1$ ,  $A_2$  and plots of experimental determinations of  $n_{\text{prim}}$ .

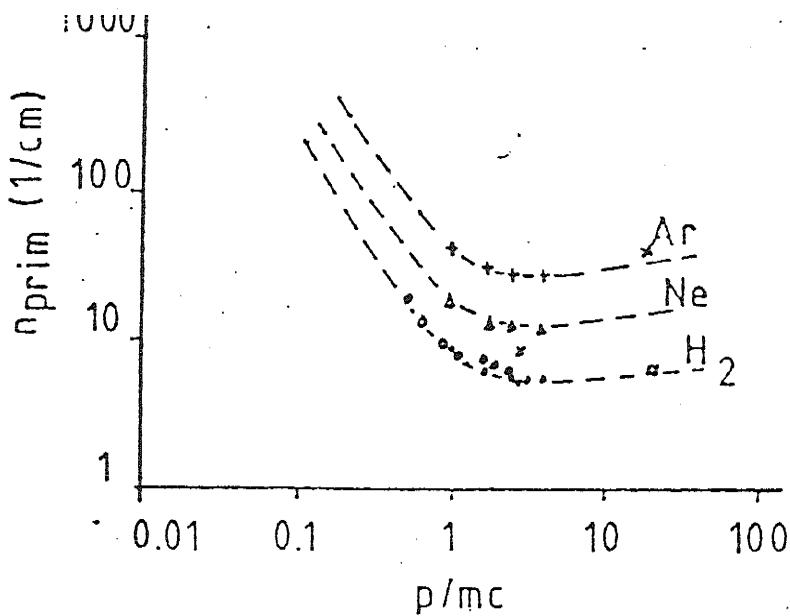


Fig. 1 Primary ionisation. Points:  
measured, --- Bethe's formula  
adjusted

Table I

gas	density (mg/cm <sup>3</sup> ) (0°C, 1atm)	I <sub>0</sub> (eV)	A <sub>1</sub>	A <sub>2</sub>	n <sub>prim</sub> (cm <sup>-1</sup> )
H <sub>2</sub>	0.0899	15.4	Bethe's formula		4.6
H <sub>e</sub>	0.1785	24.6	0.244	11.64	3.5
N <sub>e</sub>	0.9004	21.7	0.844	10.89	11.4
A <sub>r</sub>	1.7837	15.8	1.828	11.45	25.8
X <sub>e</sub>	5.8510	12.1	3.554	11.31	49.6
N <sub>2</sub>	1.2506	15.5	1.941	11.43	27.1
O <sub>2</sub>	1.4290	12.2	2.079	11.28	28.9

## ENERGY LOSS

$$\begin{aligned} \langle \frac{dE}{dx} \rangle &= \int_{E_{\min}}^{E_{\max}} E' \phi(E', E) dE' \\ &= \frac{\bar{A}P}{\beta^2} \int_{E_{\min}}^{E_{\max}} \frac{dE'}{E'} \\ &= \frac{\bar{A}P}{\beta^2} \ln \left( \frac{E_{\max}}{E_{\min}} \right) \end{aligned}$$

START OF DOING FULL Q.E.D. CALCULATION  
 NEED SOME PHYSICAL ARGUMENTS FOR  
 THE ORDER OF MAGNITUDE OF  $E_{\min}$   
 AND  $E_{\max}$ .

$E_{\min}$  CORRESPONDS TO  $b_{\max}$

DURATION OF INTERACTION SHOULD BE SHORT COMPARED  
 TO ORBITAL PERIOD OF THE ATOM.

$$\frac{b}{v} \leq \frac{1}{v} \quad (h\nu = I_0)$$

$$\Rightarrow b_{\max} = \frac{v}{\nu} f_{\max}$$

(6)

AS PARTICLE BECOMES RELATIVISTIC  $\vec{E}$  EXTENDS BY A FACTOR  $\gamma$  ( $= \frac{1}{\sqrt{1-\beta^2}}$ ).

$$\Rightarrow b_{\max} = \frac{\sqrt{\gamma}}{\gamma} f_{\max}$$

$b_{\min}$ : CAN'T BE SHORTER THAN THE deBroglie WAVELENGTH OF THE ELECTRON IN THE CENTRE OF MASS SYSTEM OF THE PROJECTILE (As  $v_{\text{cm}}(\text{electron}) \leq v_{\text{incident}}$  COLLISION CANNOT CAUSE IONIZATION).

$$\lambda_{\text{deBroglie}} = \frac{h}{mc\beta\gamma}$$

$$\Rightarrow b_{\min} = \frac{h}{mc\beta\gamma} f_{\min}$$

$$\begin{aligned} \frac{b_{\max}}{b_{\min}} &= \frac{\sqrt{\gamma}}{\gamma} \frac{mc\beta\gamma}{h} f' \\ &= \frac{mc^2}{hv} \beta^2 \gamma^2 f' \end{aligned}$$

BETHE'S CALCULATION SAYS  $f' = 2$   $\oplus$  TERM INDEPENDENT OF  $\beta$ .

# PUTTING THIS ALL BACK TOGETHER

$$\frac{dE}{dx} = 4\pi r_e^2 \left( 3 \frac{Z N}{A} \right) mc^2 \frac{1}{\beta^2} \left[ \ln \frac{2mc^2}{I} \left( \frac{\beta^2}{1-\beta^2} \right) - \beta^2 - \frac{d}{Z} \right]$$

$$\frac{dE}{dx} \approx 0.31 \frac{Z}{A} \left\{ \frac{\text{MeV cm}^2}{g} \right\} \frac{3^2}{\beta^2} \left[ \ln \left( \frac{2mc^2 \gamma^2 \beta^2}{I} \right) - \beta^2 - \frac{d}{Z} \right] \quad (1)$$

## INTERPRETATION:

a)  $\frac{dE}{dx} \propto \frac{1}{\beta^2}$  AT LOW  $\beta$

Collision time  $\tau \approx \frac{1}{\beta}$

$\Rightarrow$  momentum transfer  $\propto \frac{1}{\beta}$

$\Rightarrow$  Energy  $\propto \frac{1}{\beta^2}$ .

b) As  $\beta \rightarrow 1$

$\ln(\beta^2 \gamma^2) = \ln(\rho)$  dominates

Known as relativistic rise.

STARTS AT  $\approx \beta \gamma \approx 3.5 - 4$ ,

(8)

c) Recall that we argued  $b_{\max} \approx \gamma b_0$   
 since at very high energies  $E_{\perp}$  "spreads out":  $\approx \gamma$ .

This cannot continue indefinitely.

Material begins to polarise - depending on its density (hence symbol  $d$ )

$$\Rightarrow \frac{d}{2} = \ln \frac{\gamma^2 \gamma^2}{\hbar^2 \omega_p^2} - \ln \frac{I^2}{\hbar^2 \omega_p^2} - 1$$

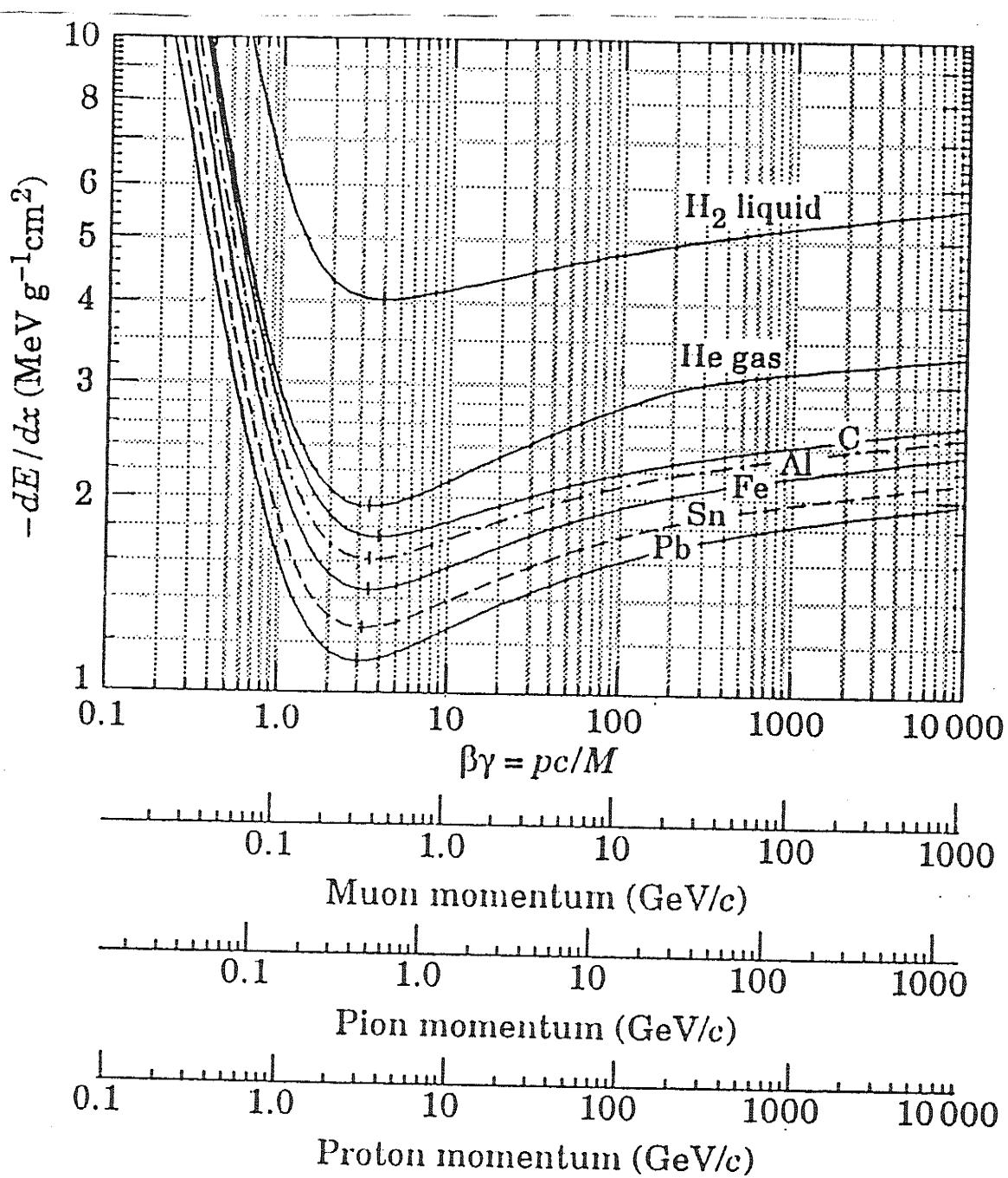
$\omega_p$  is plasma frequency of the material through which the particle is passing

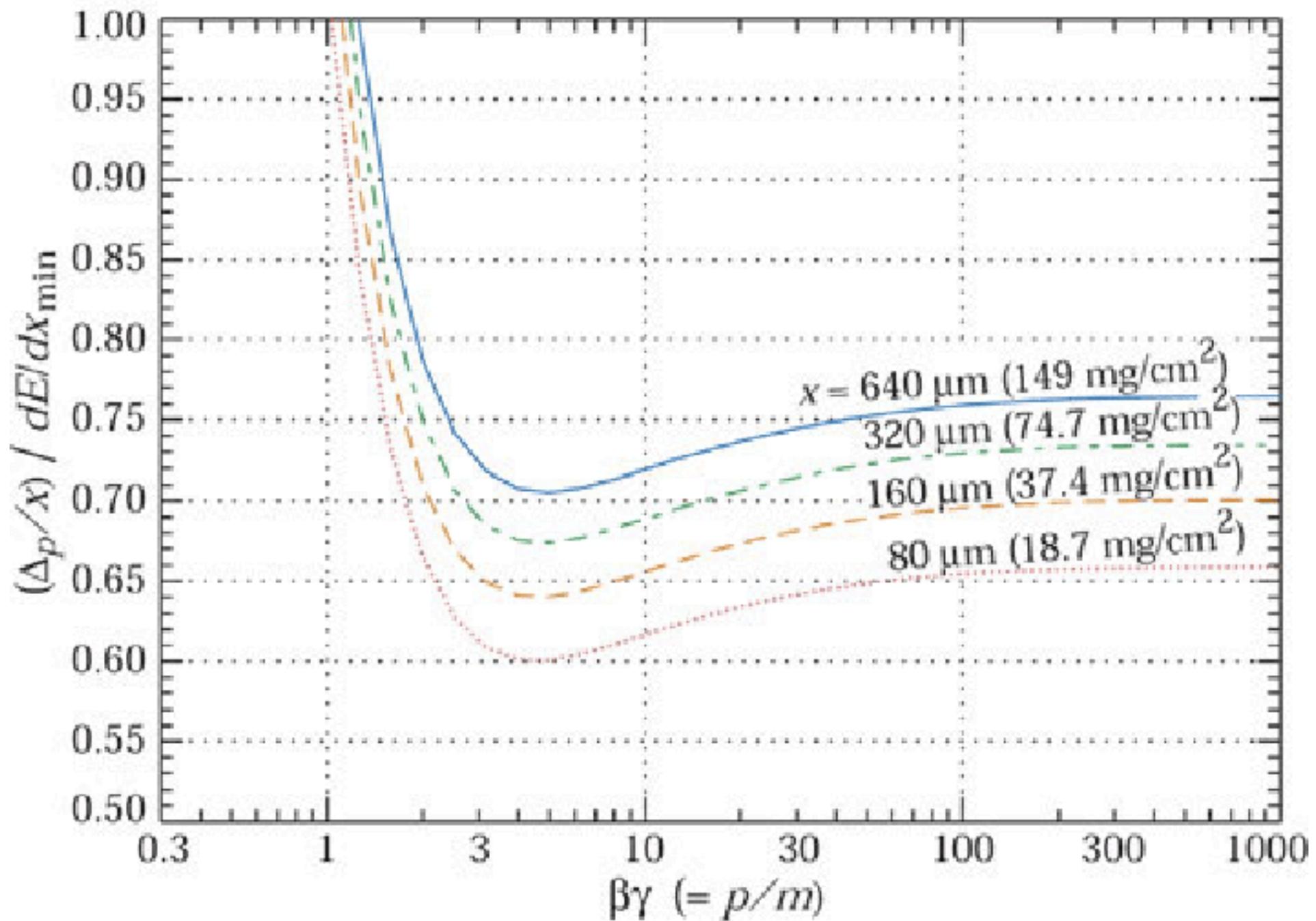
$$\omega_p = \left( \frac{n_0 e^2}{\epsilon_0 m} \right)^{1/2}$$

The density effect cancels the relativistic rise for:

$$\gamma_{\text{plateau}} \approx \frac{I}{\hbar \omega_p}$$

⑨  $\gamma_{\text{plateau}} \approx 1000$  in gases  $\approx 10$  in silicon.





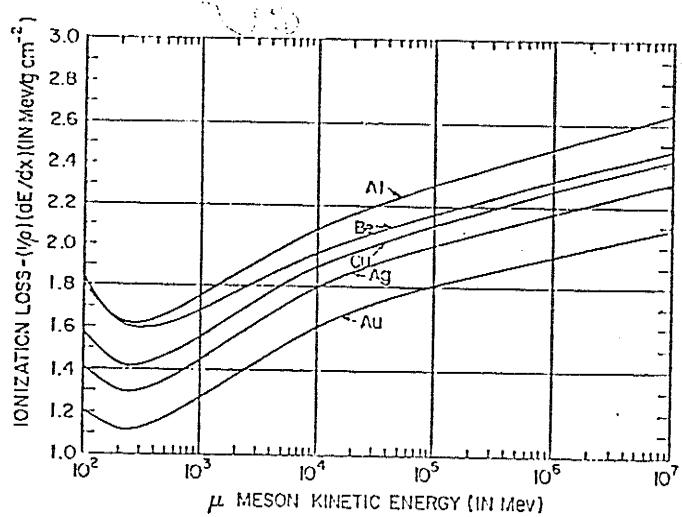


FIG. 4. The average ionization loss of  $\mu$  mesons in Be, Al, Cu, Ag, and Au, as a function of the  $\mu$ -meson kinetic energy [Eq. (1.1.19)].

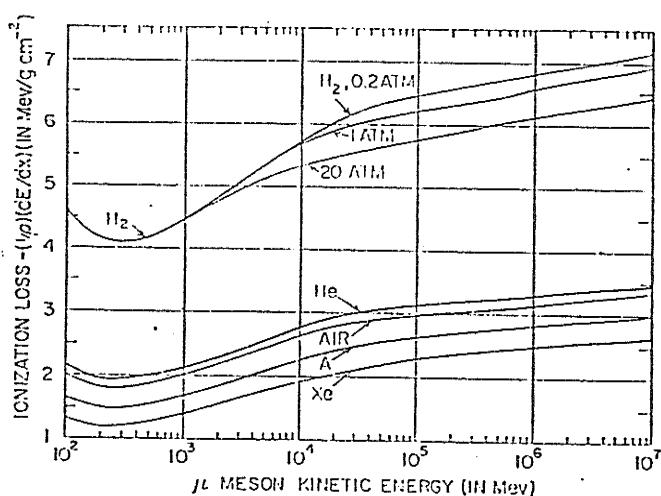


FIG. 5. The average ionization loss of  $\mu$  mesons in  $H_2$ , He, air, Ar, and Xe, as a function of the  $\mu$ -meson kinetic energy [Eq. (1.1.19)]. The curves for He, air, Ar, and Xe pertain to normal pressure.

IF YOU HAVE A MIXTURE OF MATERIALS  
(e.g. A SAMPLING CALORIMETER - SEE LATER)

THEN YOU CAN SUM THE ENERGY LOSS  
FROM THE DIFFERENT COMPONENTS :

$$\frac{dE}{dx} = \sum_j w_j \left. \frac{dE}{dx} \right|_j$$

WHERE  $w_j$  IS THE FRACTION (BY MASS) OF  
THE  $j^{\text{th}}$  COMPONENT.

$\left. \frac{dE}{dx} \right|_j$  IS  $\textcircled{I}$  FOR  $j^{\text{th}}$  COMPONENT.

# FLUCTUATIONS ABOUT MOST PROBABLE

## ENERGY LOSS (LANDAU DISTRIBUTION)

FOR VERY SMALL IMPACT PARAMETER COLLISIONS THERE IS A (LOW PROBABILITY) CHANCE OF PRODUCING A VERY HIGH ENERGY SECONDARY ELECTRON OR  $\gamma$ -RAY.

WHILE IT HAS HIGH-ENERGY COMPARED TO "NORMAL" IONISATION IT HAS LOW ENERGY COMPARED TO RELATIVISTIC PARTICLE IN THE FIRST PLACE.

$\Rightarrow$  ELECTRON IS VERY HEAVILY IONISING  $\left(\frac{1}{\beta^2}\right)$

LANDAU WORKED OUT FIRST APPROXIMATION TO THIS DESCRIBE SUCH A DISTRIBUTION:

$$\phi(\lambda) = \frac{1}{\sqrt{2\pi}} e^{-\frac{1}{2}(\lambda + \lambda^2)}$$

NORMALISED DEVIATION FROM MOST PROBABLE ENERGY LOSS.

$$\lambda = (E - E_{\text{mp}}) / \xi$$

$$\xi = \frac{1}{2} (4\pi r_e^2) mc^2 (Z^2 Z) \left( \frac{N_A \rho}{A} \right) \frac{\Delta x}{\beta^2}$$

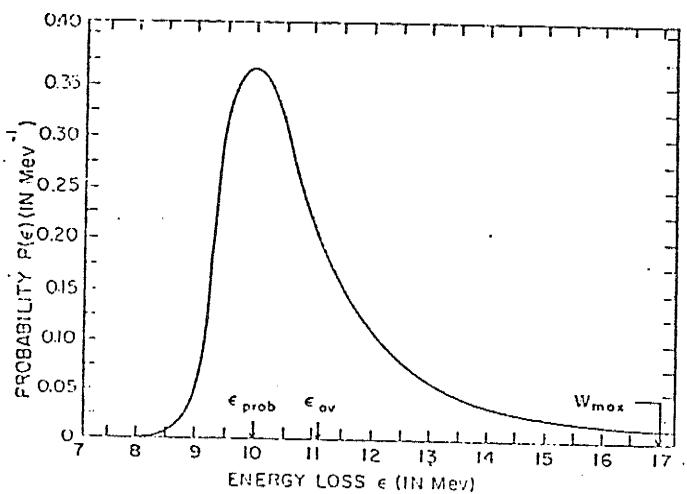


FIG. 3. The Landau distribution of energy losses  $\epsilon$  for 3-Bev protons traversing a thickness 6.97 gm/cm<sup>2</sup> of Be, for which  $\epsilon_{\text{prob}} = 10$  Mev,  $\epsilon_{\text{ov}} = 11.10$  Mev, and  $W_{\text{max}} = 17$  Mev.

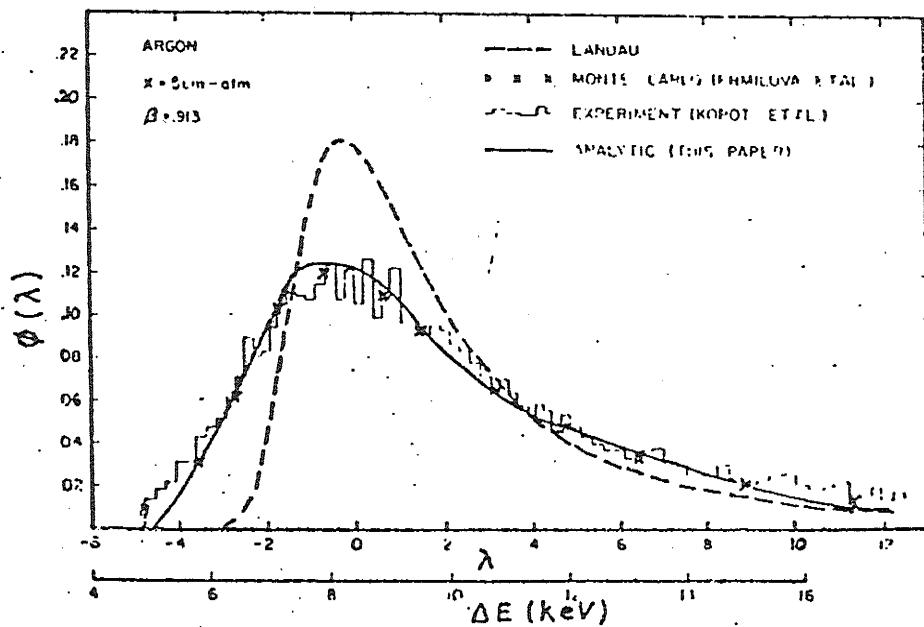


Fig. 17 Energy loss distribution. Comparison of calculations and measurements.

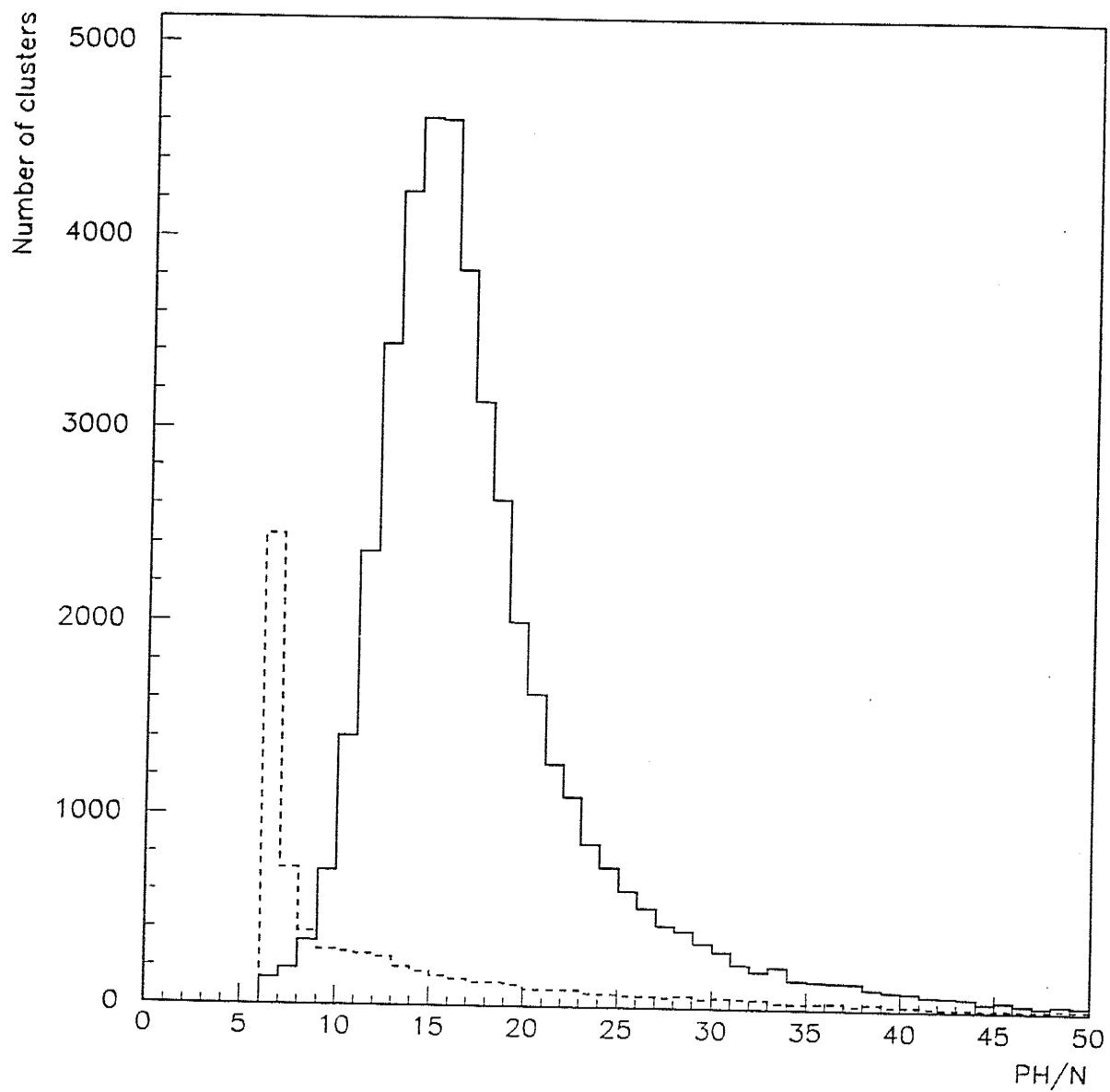
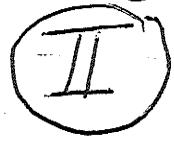


Figure 20:  $\text{PH}/\text{N}$  for clusters with (full-line) and without (dashed-line) a track associated to them for the Inner layer modules. The  $\text{PH}/\text{N}$  has been normalized to the minimum track length in traversing the silicon.

DELTA ELECTRONS HAVE AT LEAST TWO IMPLICATIONS FOR PARTICLE DETECTOR PERFORMANCE.

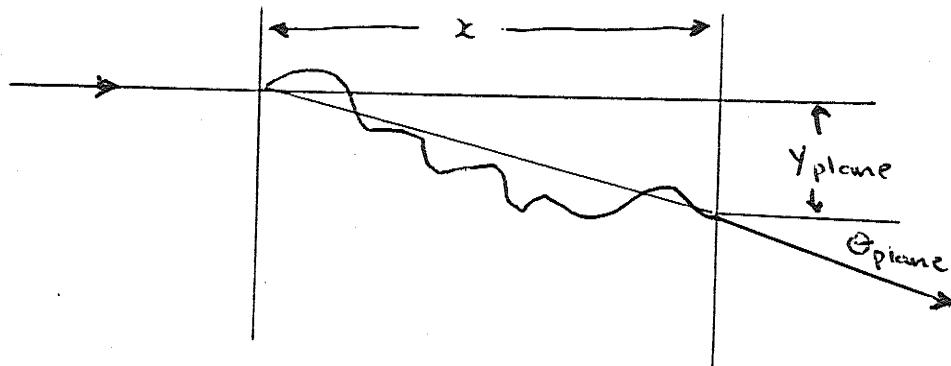
- 1) SOME TRACKERS USE  $dE/dx$  FOR PARTICLE ID (SEE LATER). LANDAU TAIL CAN MASK 20-40% EFFECTS OF DIFFERING PARTICLE MASSES IF NOT PROPERLY ACCOUNTED FOR.
- 2) S-RAYS CAN BE "ORTHOONAL" TO ORIGINAL PARTICLE TRAJECTORY. IN VERY PRECISE TRACKERS (SILICON) THIS CAN DETERIORATE POSITION RESOLUTION.

FAR FROM AN EXACT SCIENCE, BEWARE OF LIMITATIONS OF PARAMETRISATION 

## MULTIPLE SCATTERING

AS PARTICLES TRAVERSE A MEDIUM THEY ARE DEFLECTED BY MANY SMALL ANGLES - COLLISIONS MAINLY WITH NUCLEI IN THE MATERIAL.

THIS IS COULOMB SCATTERING.



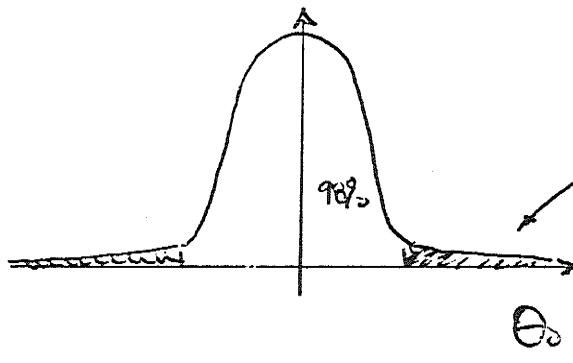
IF WE DEFINE  $\theta_0 = \theta_{\text{plane}}^{\text{rms}} = \frac{1}{\sqrt{2}} \theta_{\text{space}}^{\text{rms}}$

THEN WE CAN APPROXIMATE 98% OF CASES BY A SIMPLE GAUSSIAN WITH:

III  $\theta_0 = \frac{13.6 \text{ MeV}}{\beta c p} \sqrt{\frac{x}{x_0}} \left\{ 1 + 0.038 \ln\left(\frac{x}{x_0}\right) \right\}$

$x_0$  IS THE RADIATION LENGTH OF MEDIUM

NB. DISTRIBUTION OF  $\theta_0$  LOOKS LIKE:



LONG TAILS DUE TO  
HARD SCATTERS

cf LYNCH, DAHL  
NIMBSB (1991)

FOR  $\beta=1$   $\beta \rightarrow 1$  FORMULA III VARIES

FROM MEASUREMENTS BY LESS THAN 10%

FOR :  $10^{-3} < \frac{x}{X_0} < 10^2$

OFTEN:

$$\boxed{\theta_0 \approx \frac{13.6}{\beta c p} \sqrt{\frac{x}{X_0}}}$$

is ENOUGH.

IN PRACTICAL APPLICATIONS (PARTICLE TRAVERSING, MORE THAN 1 MATERIAL) IT IS BETTER TO COMPUTE COMBINED  $x \leq X_0$  AND APPLY FORMULA III ONCE. SCATTERING ANGLES ARE NOT INDEPENDENT  $\Rightarrow$  CANNOT JUST SUM IN QUADRATURE.

## RADIATION LENGTH

- DEFINED AS THE THICKNESS OF MATERIAL IN UNITS OF  $\left(\frac{g}{cm^2}\right)^{-1}$  OVER WHICH A HIGH ENERGY ELECTRON LOSES ALL BUT  $e^{-1}$  OF ITS ENERGY

$$\frac{1}{X_0} = 4\pi \alpha e^2 N_A \left\{ Z^2 [L_{RAD} \cdot f(Z)] + Z L'_{RAD} \right\}$$

$L_{RAD}$	$L'_{RAD}$	Z	Element
5.31	6.14	1	H
4.79	5.62	2	He
4.74	5.81	3	Li
4.71	5.92	4	Be
$\ln(184 Z^{1/3})$	$\ln(1194 Z^{-2/3})$	>4	Others

$$f(Z) = a^2 \left[ \frac{1}{1+a^2} + 0.202 = 0.037 a^2 + \dots \right]$$

$$a = \alpha Z$$

Dahl provides a more succinct form:

$$X_0 = \frac{716.4 \left( \frac{g}{cm^2} \right) A}{Z(Z+1) \ln(287/\sqrt{Z})}$$

$Z^2$  is dominant term.

ENERGY LOSS MECHANISM HERE IS BREMSSTRAHLUNG  
NOT IONISATION (SEE LATER WITH CALORIMETERS)

