

# Review

Degenerate Fermi gas:

- Conduction electrons in metal treat as free particles
   Non Interacting
- Defined Fermi energy,  $\varepsilon_F$ , equal to  $\mu(T=0)$ 
  - $\mathcal{V}_{\alpha} = \left(\frac{\lambda}{\sqrt{2\pi}mkT}\right)^{3}$

- Degenerate gas
  - Almost all states below  $\varepsilon_F$  are occupied and all above are unoccupied

# **Degenerate Fermi Gas** Wavefunctions for free electron in metal block - Represented by standing wave wavefunctions $\varepsilon = \frac{\left|\vec{p}\right|^2}{2m} = \frac{h^2}{8mI^2} \left( n_x^2 + n_y^2 + n_z^2 \right) \qquad \vec{n} \rightarrow n_x \, n_y \, n_z$ - Each "lattice point" $(n_x, n_y, n_z)$ is pair of electron states • One for each spin orientation $n_z$ - Surface of sphere $\Rightarrow$ radius $n_{max}$ $n_{\max}$ $n_y$ $n_x$ **Number of Occupied States** • Energy of state on surface of sphere $\varepsilon_F = \frac{h^2 n_{\text{max}}^2}{2 m I^2}$ number of occupied states $N = 2 \times \frac{1}{8} \times \frac{4}{3} \pi h_{max}^{3}$ = Thmax -> nmax = 23/3N 3 $\Sigma_{f} = \frac{L^{2}}{R_{n-1}^{2}} \left(\frac{3N}{T}\right)^{2/3}$ $=\frac{l^2}{2m}\left(\frac{3N}{TV}\right)^{2/3}$ -> does not depend on size of metal

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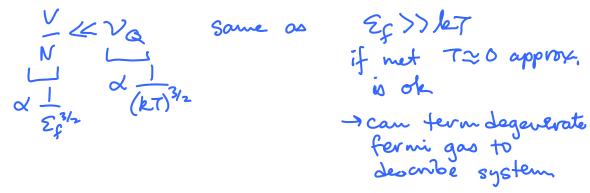
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#### **Comparing Energies**

- Fermi energy for conduction electrons in typical metal – A few eV
- Thermal energy of a particle at room temperature ٠  $- kT \approx 1/40 \text{ eV}$
- Same as comparing quantum volume and volume per particle



k~ 8×10-5 eV/k Ef= ET for Equilev Tf ~ 11,600K fermi temperature **Degeneracy Pressure** • Using  $P = -\left(\frac{\partial U}{\partial V}\right)_{ch}$   $\rightarrow$  thermodynamic identity  $P = -\frac{\partial}{\partial v} \left[ \frac{3}{5} N \cdot \frac{h^2}{2} \left( \frac{3N}{T} \right)^{2/3} V^{-2/3} \right]$  $= -\frac{3}{5}N\frac{l^{2}}{l^{2}}\left(\frac{3N}{l^{2}}\right)^{2/3}\cdot -\frac{2}{3}V^{-5/3}$  $=\frac{2}{5}\frac{N}{V} \epsilon_{f}$  $P = \frac{2}{3} \frac{U}{V} \quad \text{at } T = 0$ 

# Temperatures above Zero K At T=0, cannot determine heat capacity Need small non-zero temperatures to do this Increase temperature from zero, all particles typically gain thermal energy of kT Only electrons within ~kT of Fermi energy can acquire Move to unoccupied states above Fermi energy

# **Temperatures above Zero K**

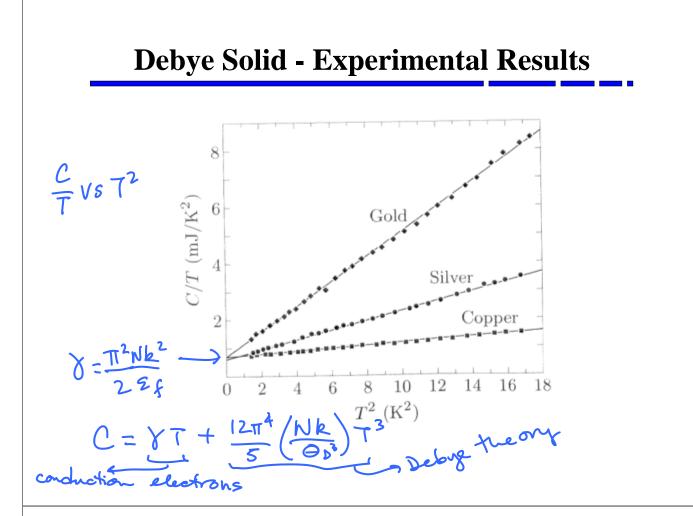
• At T=0, cannot determine heat capacity

- Need small non-zero temperatures to do this

- Increase temperature, all particles gain thermal energy of kT
  - Only electrons within  $\sim kT$  of Fermi energy can acquire
  - Move to unoccupied states above Fermi energy
- Additional energy gained related to temperature
  - Number of affected electrons proportional to kT and N

# **Degenerate Fermi Gas Above 0 K**

see text for further Dimensional analysis approach: Additional energy = (number of affected electrons) x (energy acquired by each) a (NET) (ET)  $\alpha N(kT)^{2}$ - divide by Ef - constant of system - proportionality constant TT2/4 Li from Sonnerfeld Expansion  $U = \frac{3}{5}N\Sigma_{f} + \frac{11^{2}}{4}\frac{N(leT)^{2}}{\Sigma_{f}}$ from additional peol to be in low 7 lowit T << Ef/k room T meets this requirement  $C_{V} = \left(\frac{\partial U}{\partial T}\right)_{V} = \frac{\Pi^{2}}{4} \frac{N}{\Sigma_{C}} 2\left(k^{2}T\right)$  $= \frac{TT^2 N \mu^2}{2 \cdot 2c} T$ linear term in Cu



### **Distribution Functions Used**

Distribution functions describe the average number of particles in  $\bar{h} = \sum_{n} n \bar{P}(n)$ a state Quantum statistics for cases when  $Z_1 >> N$  condition is not met:

- Fermi-Dirac Distribution (fermions)  $\overline{n}_{FD} = \frac{1}{e^{(\varepsilon - \mu)/kT} + 1}$
- Bose-Einstein Distribution (bosons)  $\overline{n}_{BE} = \frac{1}{e^{(\varepsilon \mu)/kT} 1}$

# **Relation to Planck Distribution**

• Bose-Einstein Distribution for photons and phonons becomes the Planck Distribution

$$\overline{n}_{BE} = \frac{1}{e^{(\varepsilon - \mu)/kT} - 1} \longrightarrow \overline{n}_{Pl} = \frac{1}{e^{hf/kT} - 1} = \frac{1}{e^{\varepsilon/kT} - 1}$$

• Because  $\mu=0$  for photons and phonons

$$M = \begin{pmatrix} \partial f \\ \partial N \end{pmatrix}_{T,V}$$
 at eq. = 0 f has to be  
minimum at  
equilibrium  
photons not conserved  
NOTE: Not responsible for density of states  
or sommerfeld Expansion in Sec 7.3