## Soutions to Extra Problems for PHY293 Modern Physics (December 2017)

4.73 
$$p = \frac{h}{\lambda} = \frac{h}{4R_{\text{mx}}} = \frac{6.63 \times 10^{-34} \,\text{J} \cdot \text{s}}{16 \times 10^{-15} \,\text{m}} = 4.14 \times 10^{-20} \text{kg·m/s}. \text{ Classically the velocity would be } v = \frac{4.14 \times 10^{-20} \text{kg·m/s}}{9.11 \times 10^{-31} \text{kg}}$$
$$= 4.5 \times 10^{10} \text{ m/s}. \text{ Impossible! An electron confined to such small dimensions will be moving relativistically.}$$
Relativistically,  $p = 4.14 \times 10^{-20} \text{kg·m/s} = \gamma_u mu \rightarrow 4.14 \times 10^{-20} \text{kg·m/s} = (9.11 \times 10^{-31} \text{kg}) \frac{u}{\sqrt{1-u^2/c^2}}.$ Squaring:  $1.72 \times 10^{-39} \text{kg}^2 \cdot \text{m}^2/\text{s}^2 = 8.3 \times 10^{-61} \text{kg}^2 \frac{u^2}{1-u^2/c^2} \text{ or } \frac{1.72 \times 10^{-39} \text{kg}^2 \cdot \text{m}^2/\text{s}^2}{8.3 \times 10^{-61} \text{kg}^2} = \frac{1}{1/u^2 - 1/c^2}$ 
$$\Rightarrow u = 0.9999783 \text{c}. \text{KE} = (\gamma_u - 1)mc^2 = \left(\frac{1}{\sqrt{1-(0.9999783)^2}} - 1\right) (9.11 \times 10^{-31} \text{kg}) (9 \times 10^{16} \text{m}^2/\text{s}^2) = 1.2 \times 10^{-11} \text{J}.$$
PE =  $\frac{(9 \times 10^9 \,\text{N} \cdot \text{m}^2/\text{C}^2)(20 \times 1.6 \times 10^{-19} \,\text{C})(-1.6 \times 10^{-19} \,\text{C})}{4 \times 10^{-15} \,\text{m}}$ 

negative and is a maximum of zero when the electron is infinitely far from the nucleus. If at any point the electron has positive *total* (kinetic plus potential) energy, it may indeed make it out to infinitely far away. We have shown that the confined electron's KE would be *ten times* the potential energy—so the electron cannot be confined by this potential energy.

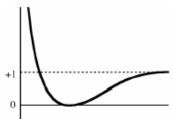
5.15 Its wavelength could not be constant. Since E is constant, the kinetic energy increases on the left, meaning a larger momentum and a smaller wavelength there.

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- 5.21 As  $x \to 0$  this goes to  $+\infty$ ; the  $1/x^2$  diverges faster in the positive direction than 1/x in the negative. As  $x \to \infty$  it goes to  $1 \frac{d}{dx} U(x) = -2 \frac{1}{x^3} + \frac{2}{x^2}$ . Setting this to 0 gives  $x = \infty$  and x = 1. There must be a minimum at x = 1. U(1) = 0.
  - (b) When is KE zero? When E = U.  $0.5 = \frac{1}{x^2} \frac{2}{x} + 1 \rightarrow 0.5 \frac{2}{x} + \frac{1}{x^2} = 0 \rightarrow 0.5x^2 2x + 1 \Rightarrow$

$$x = \frac{2 \pm \sqrt{2^2 - 4(0.5)1}}{2(0.5)} = 2 \pm \sqrt{2}$$
. With turning points on either side, yes, it would be bound.

(c)  $2 = \frac{1}{x^2} - \frac{2}{x} + 1 \rightarrow -1 - \frac{2}{x} + \frac{1}{x^2} = 0 \rightarrow x^2 + 2x - 1 \Rightarrow x = \frac{-2 \pm \sqrt{2^2 - 4(1)(-1)}}{2(1)} = \frac{-2 \pm \sqrt{8}}{2}$ . The only positive root is  $\sqrt{2} - 1$ . No.



5.26 
$$E = \frac{\pi^2 \hbar^2 n^2}{2mL^2} = \frac{\pi^2 (1.055 \times 10^{-34} \,\mathrm{J} \cdot \mathrm{s})^2 n^2}{2(1.67 \times 10^{-27} \,\mathrm{kg})(15 \times 10^{-15} \,\mathrm{m})^2} = 1.5 \times 10^{-13} \,\mathrm{J} \times n^2 \cong 1 \,\mathrm{MeV} \times n^2.$$
 Transitions between various  $n$ 

values should indeed generate photons whose energies are in the MeV range.

5.90 Set the probability integral over all space to unity.  $1 = \frac{2}{\pi} a^3 \int_{-\infty}^{\infty} \frac{1}{(x^2 + a^2)^2} dx$ . The integral can be looked up in table, or done by trig substitution  $x = a \tan \theta$ . In any case, its value is  $\frac{\pi}{2a^3}$ , so the normalization constant is

correct.

5.91  $\overline{x} = \frac{2}{\pi} a^3 \int_{-\infty}^{+\infty} x \frac{1}{(x^2 + a^2)^2} dx$ . This is the integral of an odd function of x over a symmetric interval, so it is zero, as

we would expect by symmetry.  $\overline{x^2} = \frac{2}{\pi} a^3 \int_{-\infty}^{+\infty} x^2 \frac{1}{\left(x^2 + a^2\right)^2} dx$ . The integral can be looked up in table, or done by

trig substitution  $x = a \tan \theta$ . Its value is  $\pi/2a$ , so  $\overline{x^2} = \frac{2a^2}{\pi} \frac{\pi}{2a} = a^2$ . Thus,  $\Delta x = \sqrt{\overline{x^2} - \overline{x}^2} = a$ .

5.92 It must solve the Schrödinger equation.  $-\frac{\hbar^2}{2m}\frac{d^2}{dx^2}\psi(x) + U(x)\psi(x) = 0.$ 

Thus 
$$U(x) = \frac{1}{\psi(x)} \frac{\hbar^2}{2m} \frac{d^2}{dx^2} \psi(x)$$
. But  $\frac{d^2}{dx^2} \psi(x) = \sqrt{\frac{2}{\pi}} a^{3/2} \frac{d}{dx} \frac{-2x}{(x^2 + a^2)^2}$   

$$= \sqrt{\frac{2}{\pi}} a^{3/2} \frac{-2(x^2 + a^2)^2 - (-2x)(4x)(x^2 + a^2)}{(x^2 + a^2)^4} = \sqrt{\frac{2}{\pi}} a^{3/2} \frac{-2(x^2 + a^2) - (-2x)(4x)}{(x^2 + a^2)^3} = \sqrt{\frac{2}{\pi}} a^{3/2} \frac{6x^2 - 2a^2}{(x^2 + a^2)^3}$$
Therefore,  $U(x) = \frac{1}{\sqrt{2/\pi}a^{3/2}} (x^2 + a^2) \frac{\hbar^2}{2m} \sqrt{\frac{2}{\pi}} a^{3/2} \frac{6x^2 - 2a^2}{(x^2 + a^2)^3} = \frac{\hbar^2}{2m} \frac{6x^2 - 2a^2}{(x^2 + a^2)^2}$ 

(c) We find the classical turning points by setting the potential energy equal to the total:

$$U(x) = E \to \frac{\hbar^2}{2m} \frac{6x^2 - 2a^2}{(x^2 + a^2)^2} = 0 \Rightarrow x = \pm a/\sqrt{3} \text{ . Thus: } -a/\sqrt{3} < \mathbf{x} < +a/\sqrt{3}$$

